

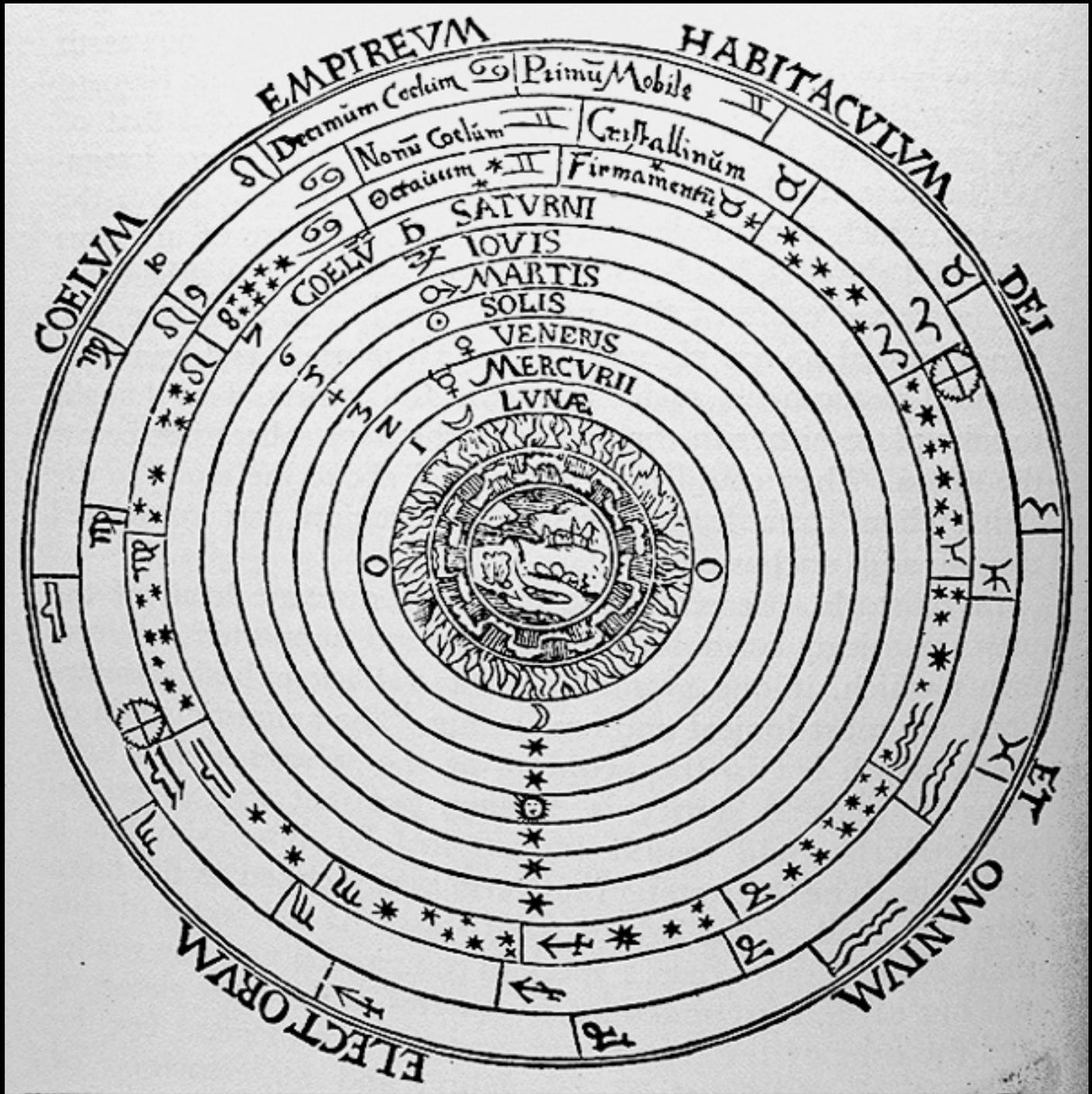


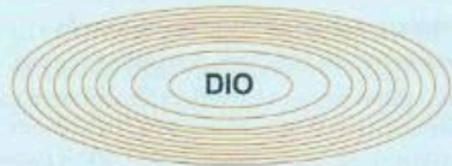
Gravit 

Eugenio Coccia

Gran Sasso Science Institute and INFN

Tolomeo - Aristotele
Modello geocentrico



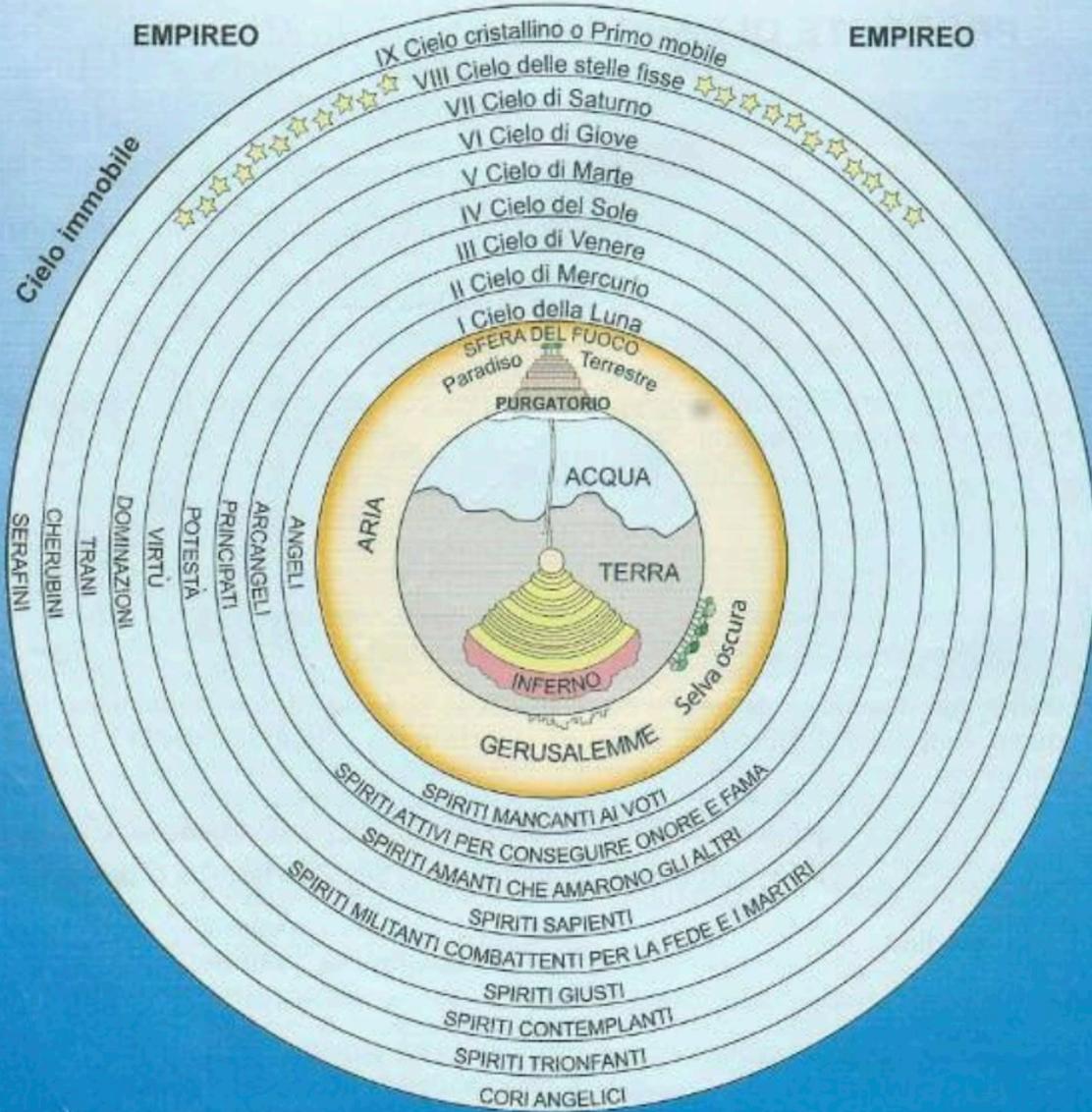


DIO

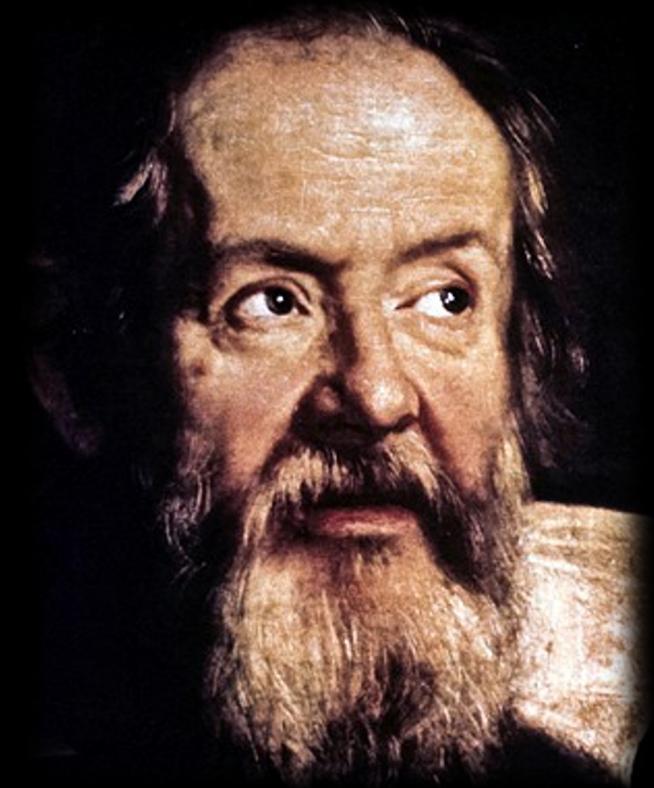
9 Cerchi angelici



Candida Rosa



Galileo Galilei



Die 25. Julij. Summo mane Efecto tempore
in Jacobi die Dominica Potavi primam
obsequiam 7^o orientale matutina cui
substant Greg. Plausq. Merici orientales
ab ipso in huc ordinu

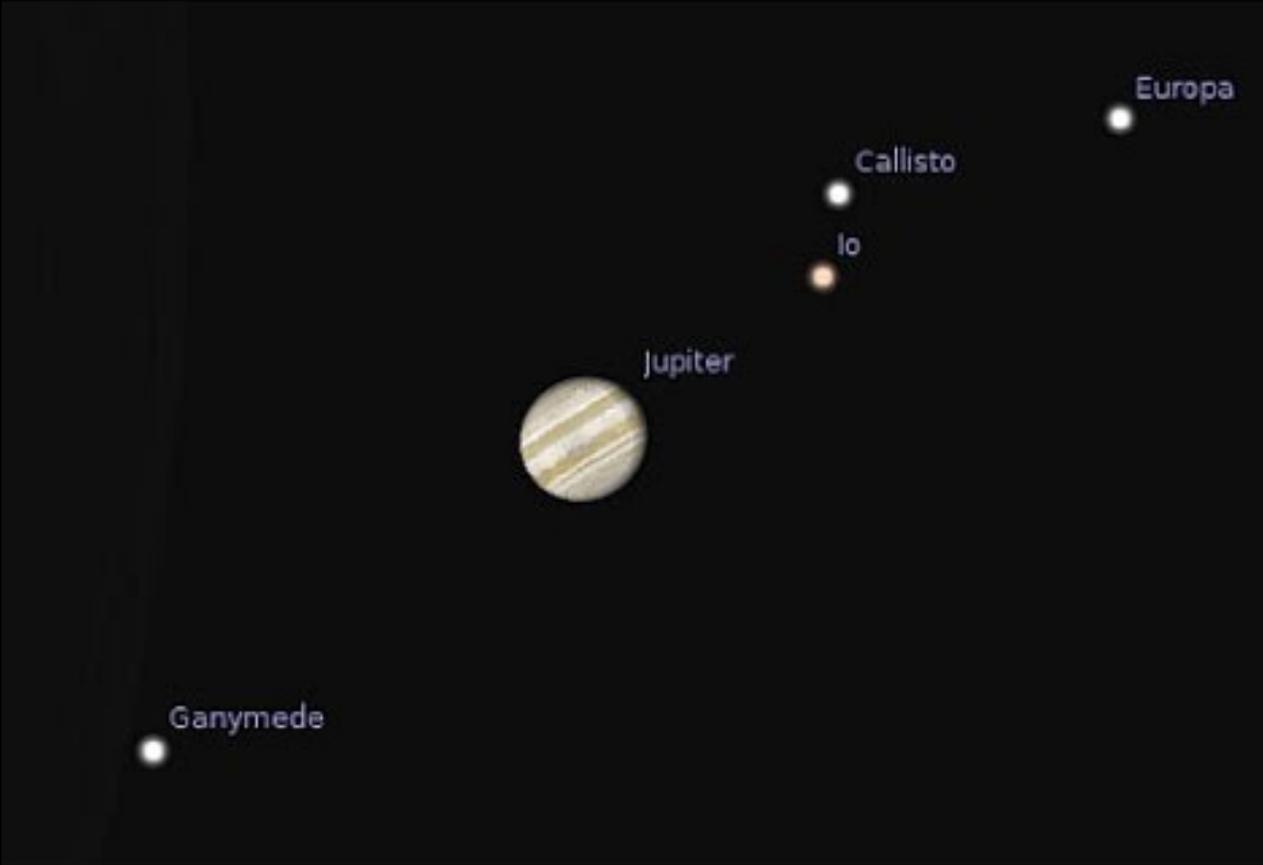
- * * * O
- Die .29. sic * * * O * *
- D. 5. Aug. * * * O * * *
occidentalis; paululu offerbat
- D. 8. * * * O * *
- Die .15. * * * O prox. 4^o adde
but' e' Notion
- D. 17. * * * O * *
- D. 20. O * *
- D. 21. * * * O * * *
- D. 22. * * * O * *
- D. 24. * * * O * * *
3^o 110 12 10 * media occid
Ber: attollat.
- D. 25. O * * *
- D. 31. * * * O * *
- D. 7. Septemb: * * * O * *
- D. 25. Octob * * * O * *
- D. 7. Novemb. * * * O * *
- D. 5. * * * O * *
- D. 14. H. noct. 7. * * * O * *
- D. 15. H. 5. * * * O * *
ca 7^o sint caliginosus, ex pro occidentalis. paulu
conspiciat.
- D. 18. H. 5. * * * O * *
- D. 19. * * * O * *
- D. 20. H. 5. * * * O * *
Beran: et H. 7. 2. autuq. sese. Jere. capabot
- D. 21. * * * O * *

- D. 30. * * * O * * orientalis
2^o hronu paululu 3^o Bor: attollat; et ca 11^o
prop. post Hor. 1/3 conucluy fuit.
- D. 2. Decemb. H. 7. * * * O * *
Hor. 5. 4^o hronus ca eo iunty est; claru. aer.
- D. 3. H. 5. * * * O * * *
- D. 4. H. 5. * * * O * *
- D. 6. H. 5. * * * O * *
- D. 7. H. 5. * * * O * *
H. 7. * * * O. ex ortuq. orion:
calu paululu 3^o Bor: offerbat
- D. 9. H. 5. * * * O * * *
- D. 10. H. 4. * * * O * * *
- D. 12. H. 8. * * * O * * *
- D. 13. H. 3. 30. * * * O. Secuda in 4. 3^o Bor:
attollatur.
Hu. 7. uia in ortu 4. conucluy fuit.
Hu. 5. * * * O. medius uidebat in
austuq. dechnare
- D. 14. H. 5. 30. * * * O
- D. 19. H. 3. 30. O * * *
140 130 8^o * Legu 3:
carte dro po

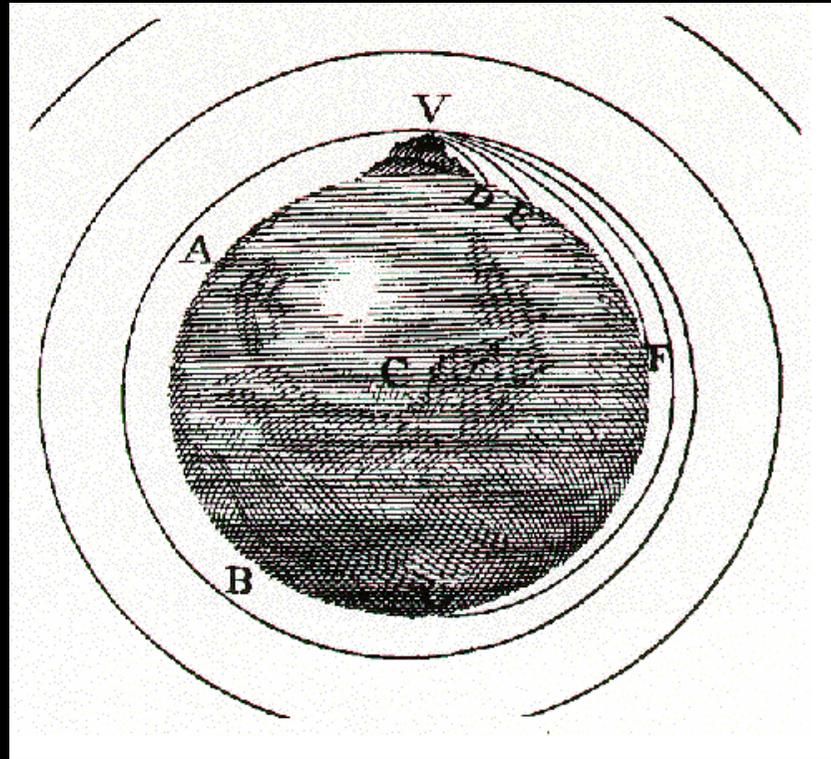
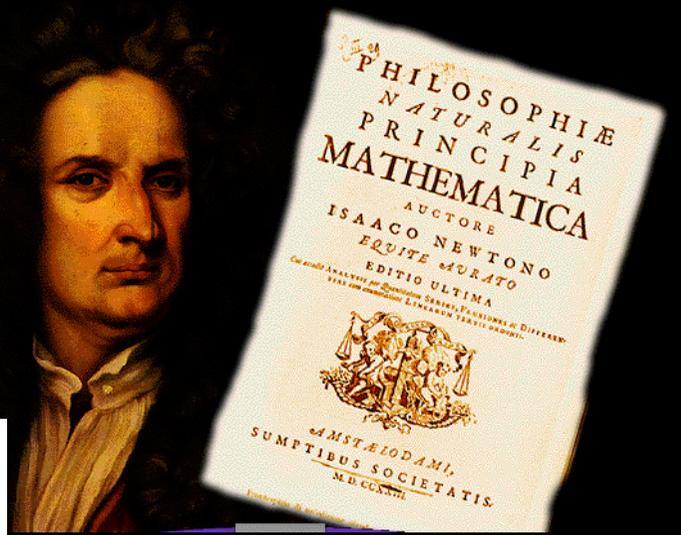


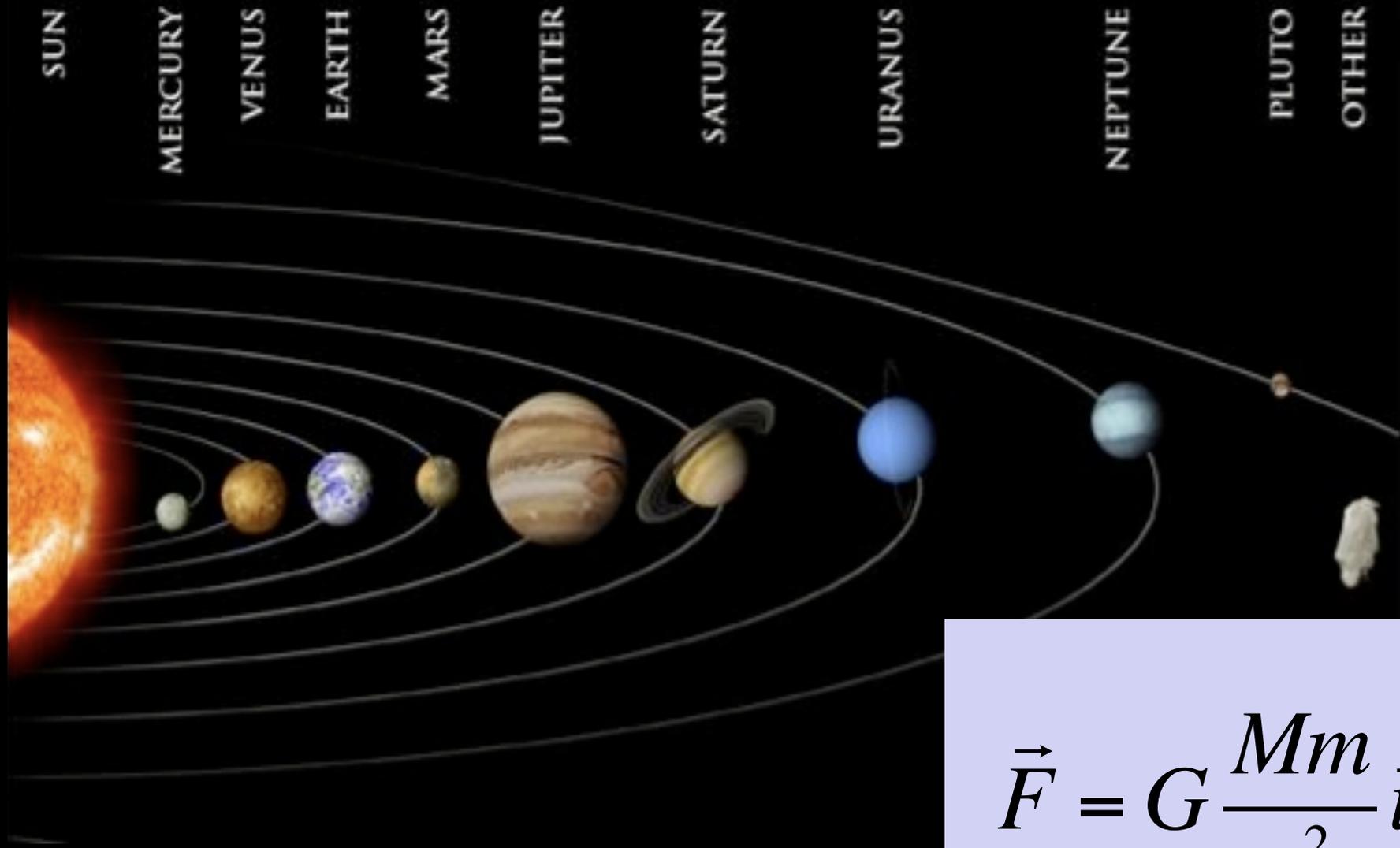
30

Adi 7. di Gennaio 1610 Giove si vedeva col cannone ³ ^{* uen.} ³
 3. stelle fisse così *  delle quali se la il cannone
 miuno si vedeva. ^{ori: *} a di 8. appariva così  * * * era dug
 diretto et no retrogrado come sogono i calculatori.
 Adi 9. si rugolo. a di 10. si vedeva così * *  cio' è d-
 giato in la piu' occidentale si che in occultava p quanto si puo credere.



Newton





$$\vec{F} = G \frac{Mm}{r^2} \vec{u}_r$$

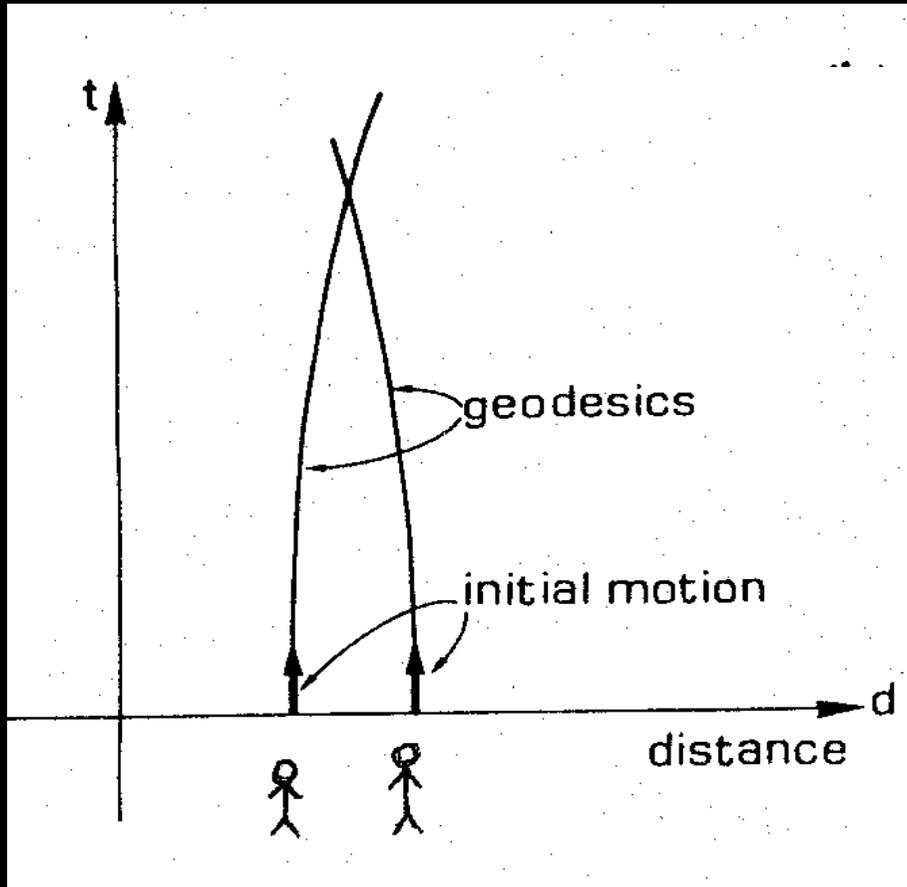
$$^1 \quad F = m_i a \quad F = G \frac{M m_g}{r^2}$$

Experimental result: $\frac{m_i}{m_g} = \text{const} = 1$

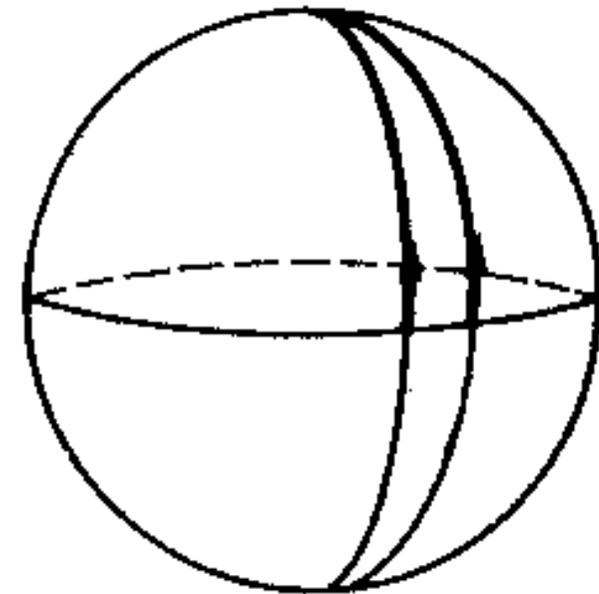
a coincidence?

2

Newtonian gravity propagates with infinite speed!



Einstein



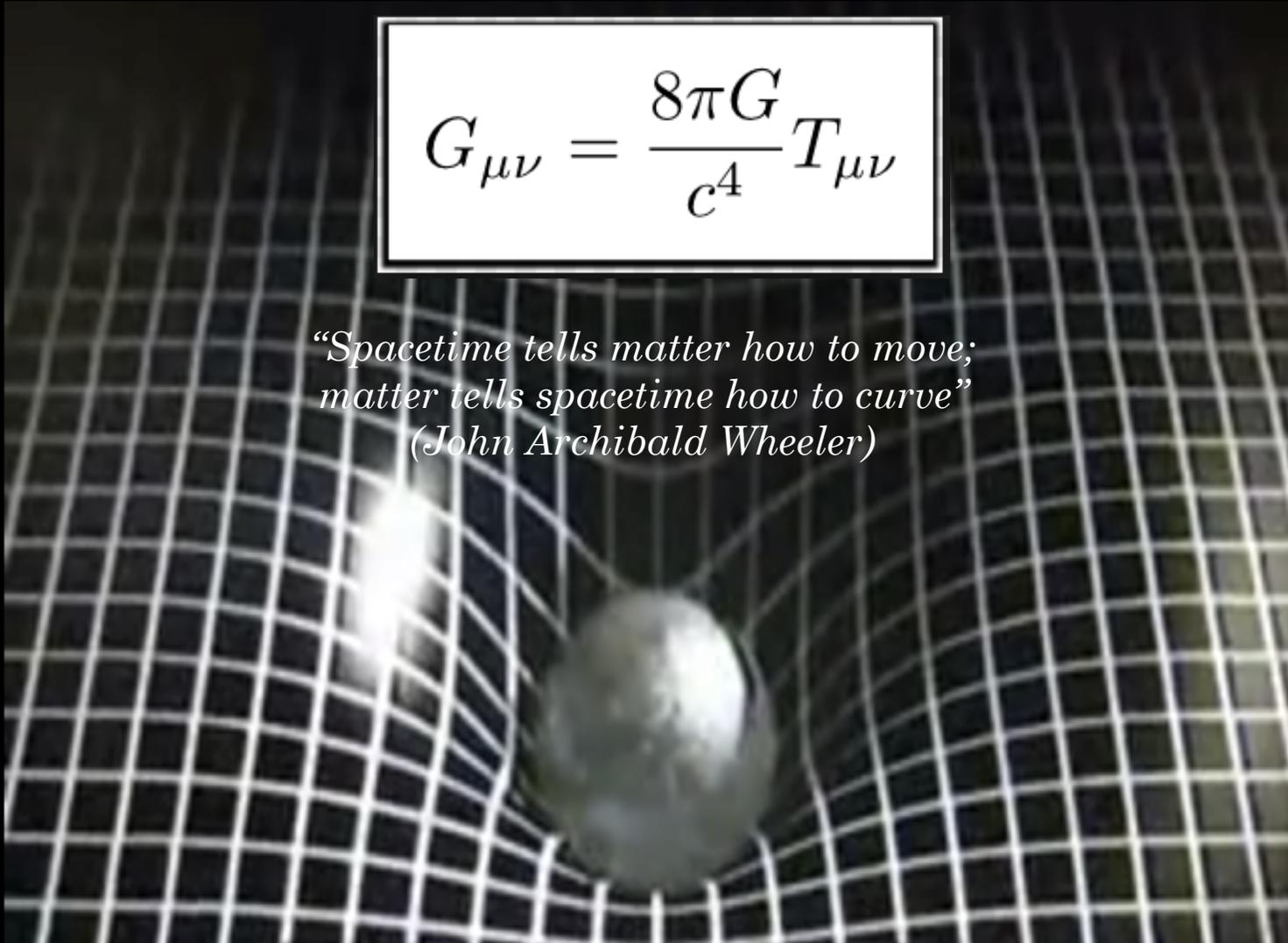
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Einstein's Theory of Gravitation

Gravity is a manifestation of spacetime curvature induced by mass-energy

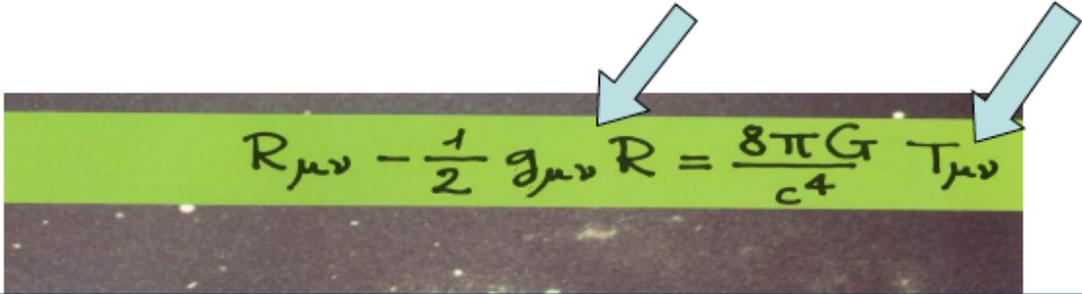
$$G_{\mu\nu} = \frac{8\pi G}{c^4} T_{\mu\nu}$$

*“Spacetime tells matter how to move;
matter tells spacetime how to curve”
(John Archibald Wheeler)*



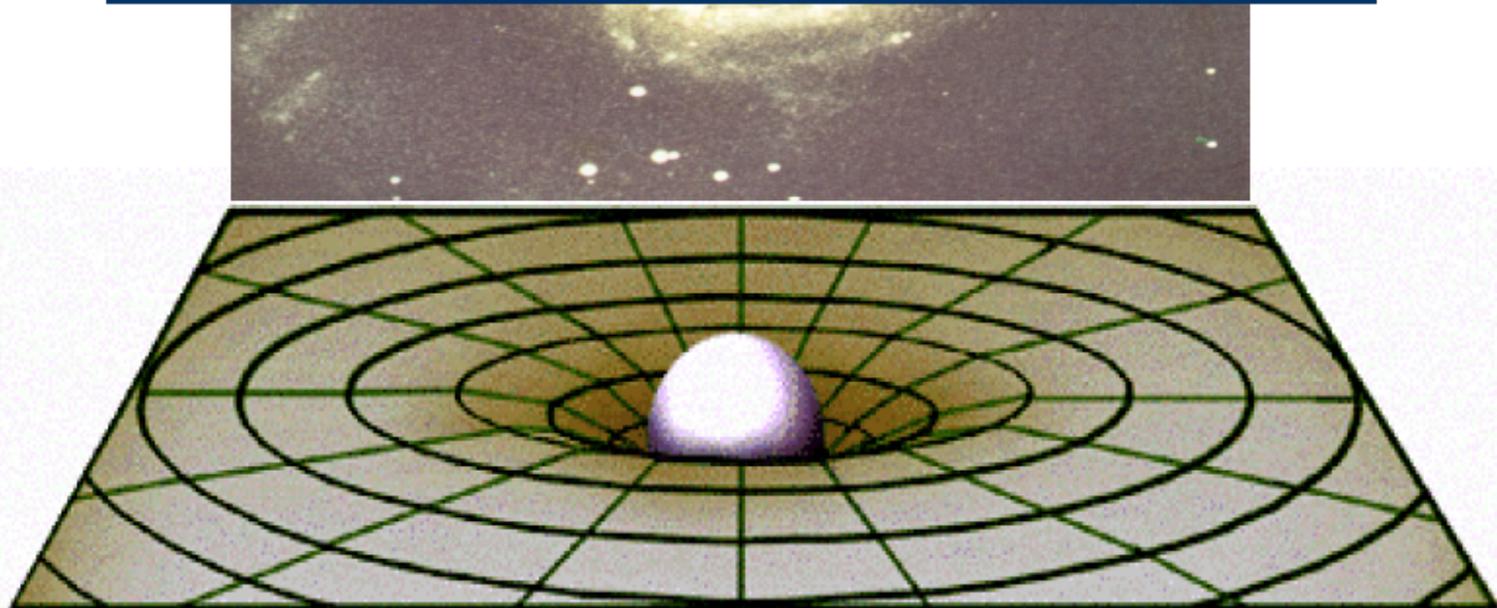
General Relativity (1915)

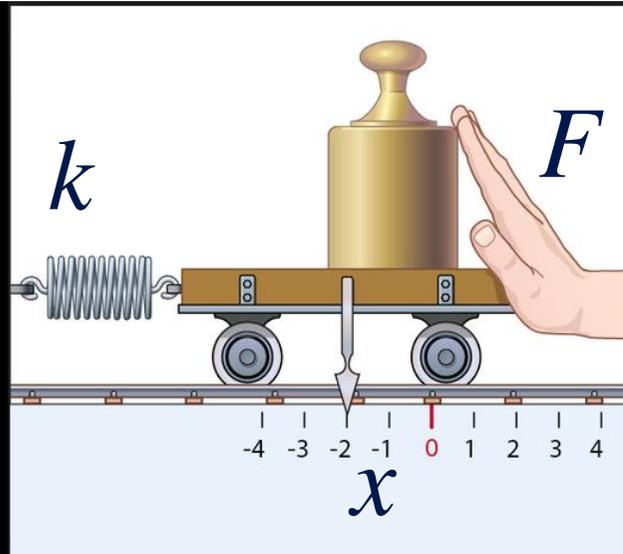
Gravity is a manifestation of spacetime curvature induced by mass-energy


$$R_{\mu\nu} - \frac{1}{2} g_{\mu\nu} R = \frac{8\pi G}{c^4} T_{\mu\nu}$$

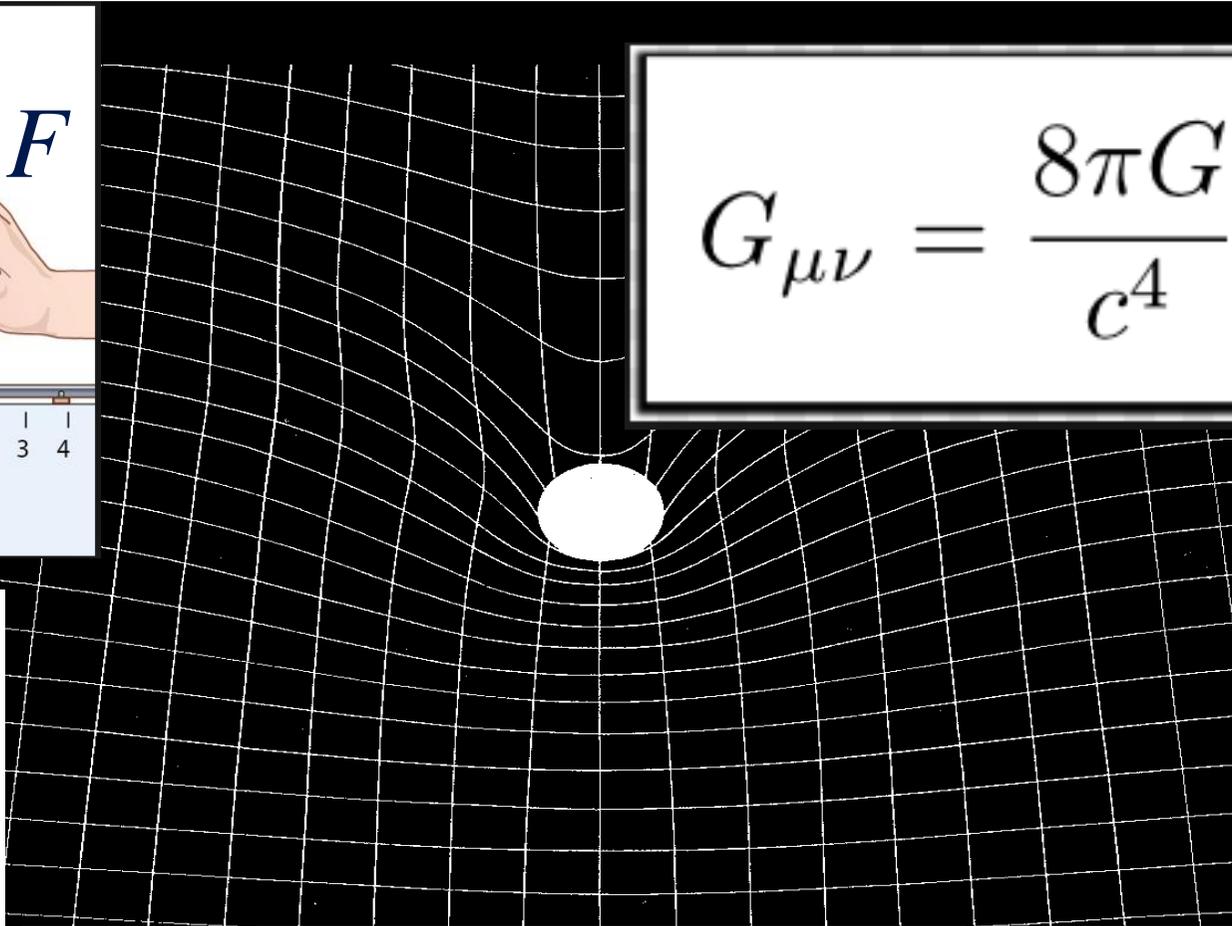
10 non linear equations in the unknown $g_{\mu\nu}$

$$ds^2 = g_{\mu\nu} dx^\mu dx^\nu$$





$$G_{\mu\nu} = \frac{8\pi G}{c^4} T_{\mu\nu}$$



$$F = -kx$$

$$F \Leftrightarrow T_{\mu\nu}$$

$$x \Leftrightarrow G_{\mu\nu}$$

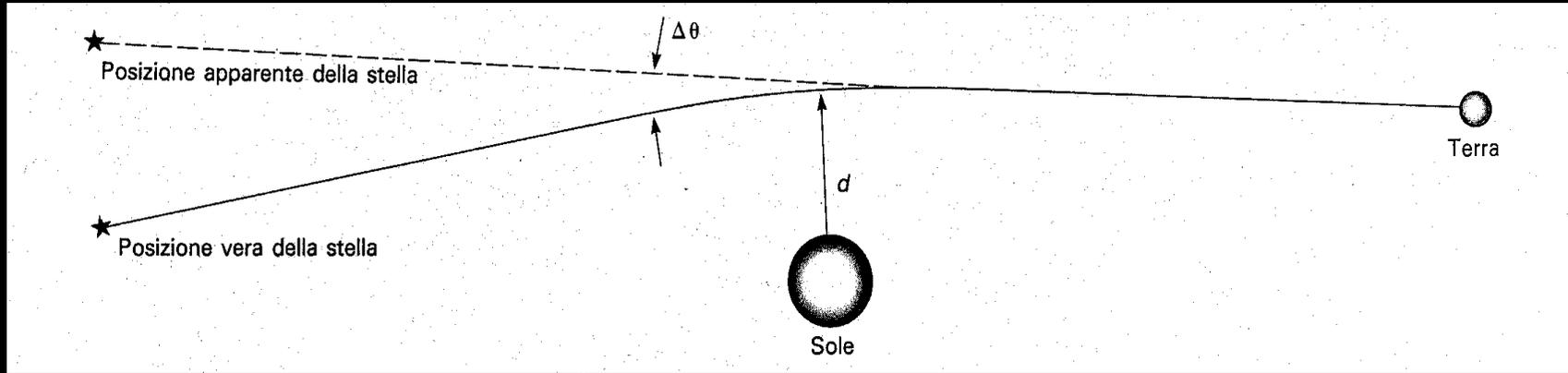
$$k \Leftrightarrow \frac{c^4}{8\pi G}$$

$$c = 299\,792\,458 \text{ m/s} = 3 \times 10^8 \text{ m/s}$$

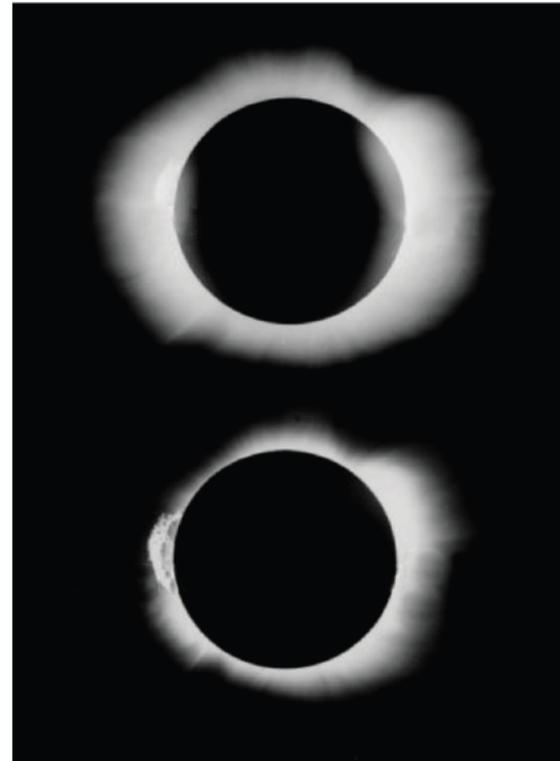
$$G = 0,000\,000\,000\,066\,7 \frac{\text{m}^3}{\text{kg s}^2} = 6,67 \times 10^{-11} \frac{\text{m}^3}{\text{kg s}^2}$$

$$k \approx 10^{45} \frac{\text{kg}}{\text{s}^2} \quad \text{STIFF!}$$

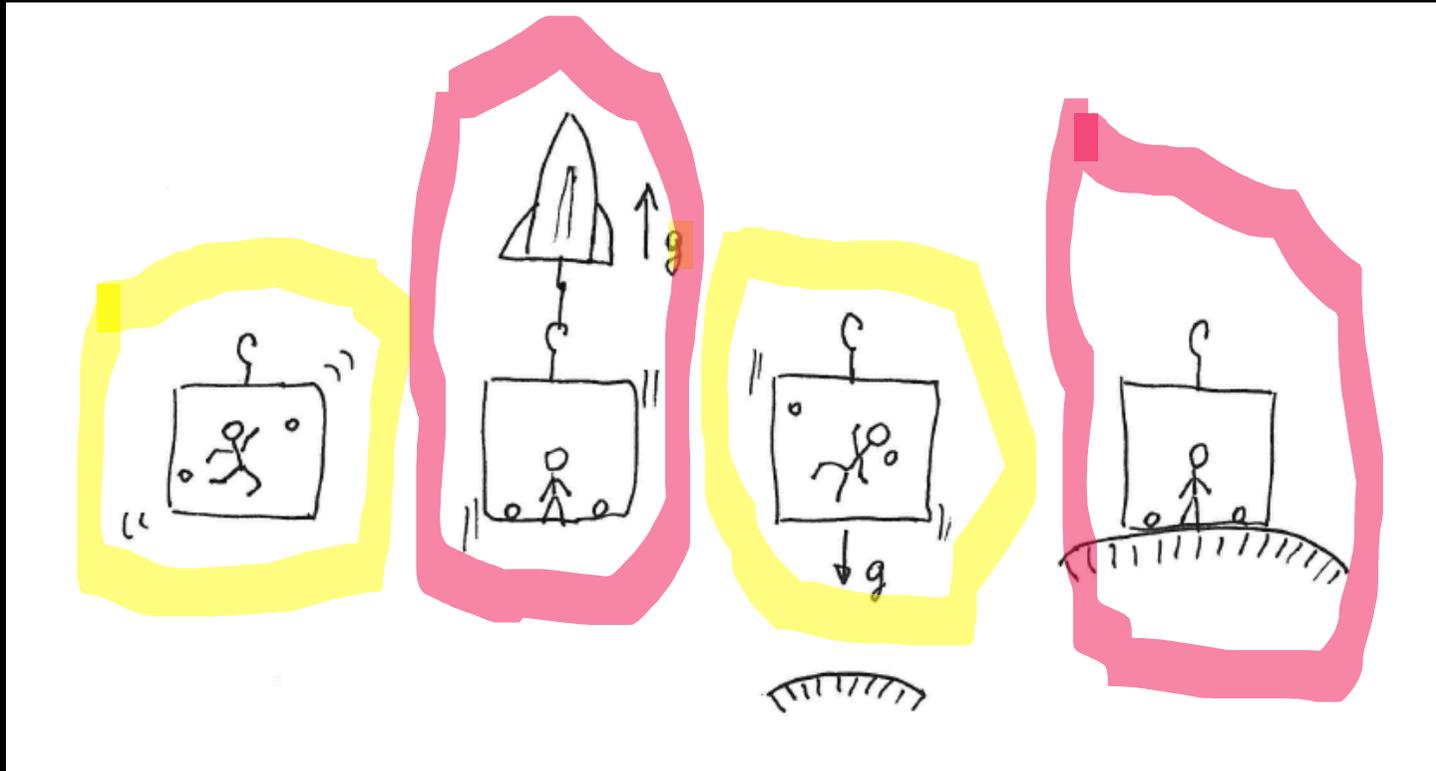
Light deflection



Eddington



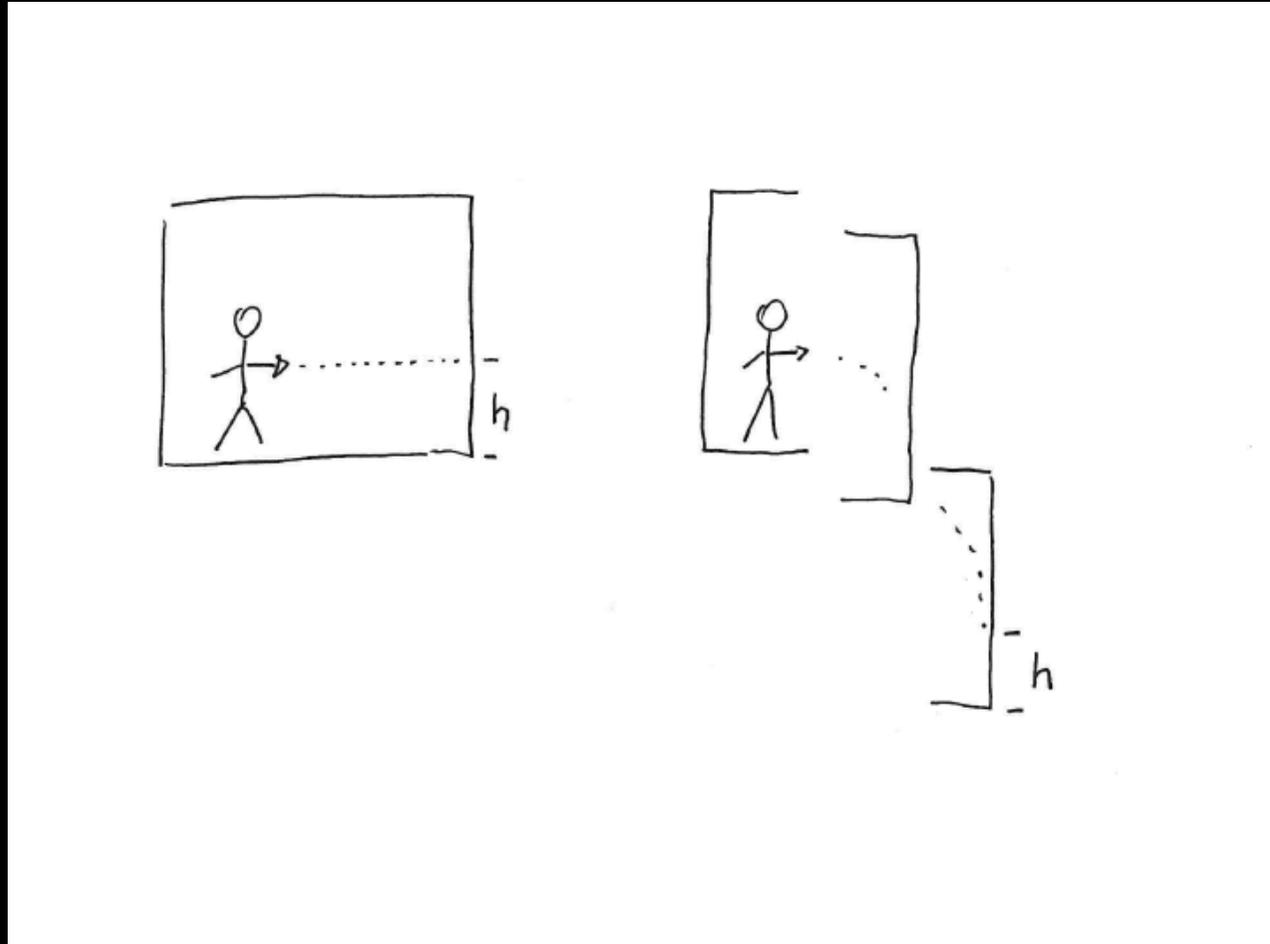
«L'idea più felice della mia vita»



Un sistema fermo in un campo gravitazionale è equivalente a un sistema accelerato.

Un sistema in caduta libera in un campo gravitazionale è equivalente a un sistema inerziale.

Questo è la nuova formulazione del principio di equivalenza, che Einstein trasformò in un principio fondamentale della fisica.



In un campo gravitazionale la traiettoria della luce è una linea curva.

Lo spazio-tempo non è più euclideo e le linee geodetiche non sono più delle rette.

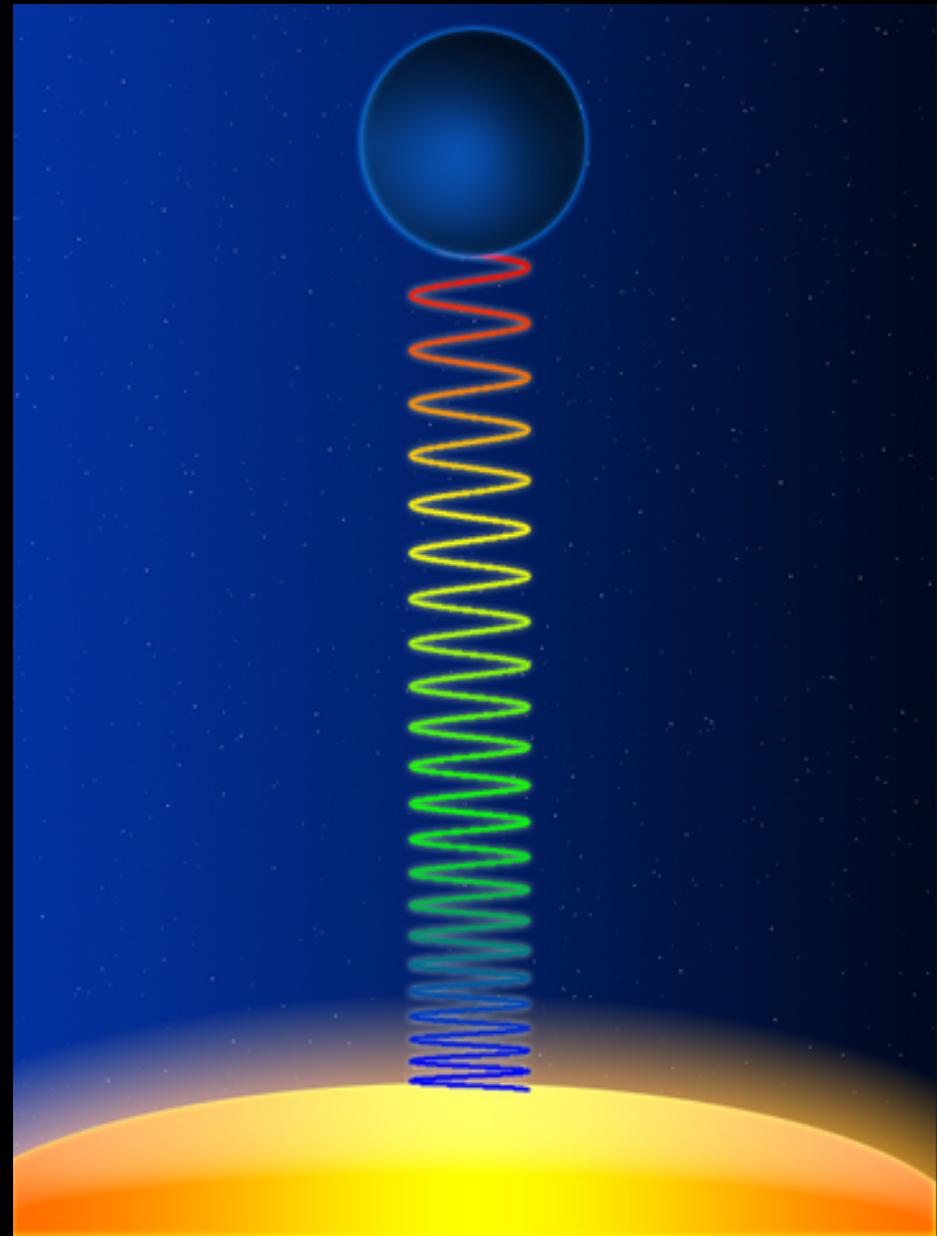


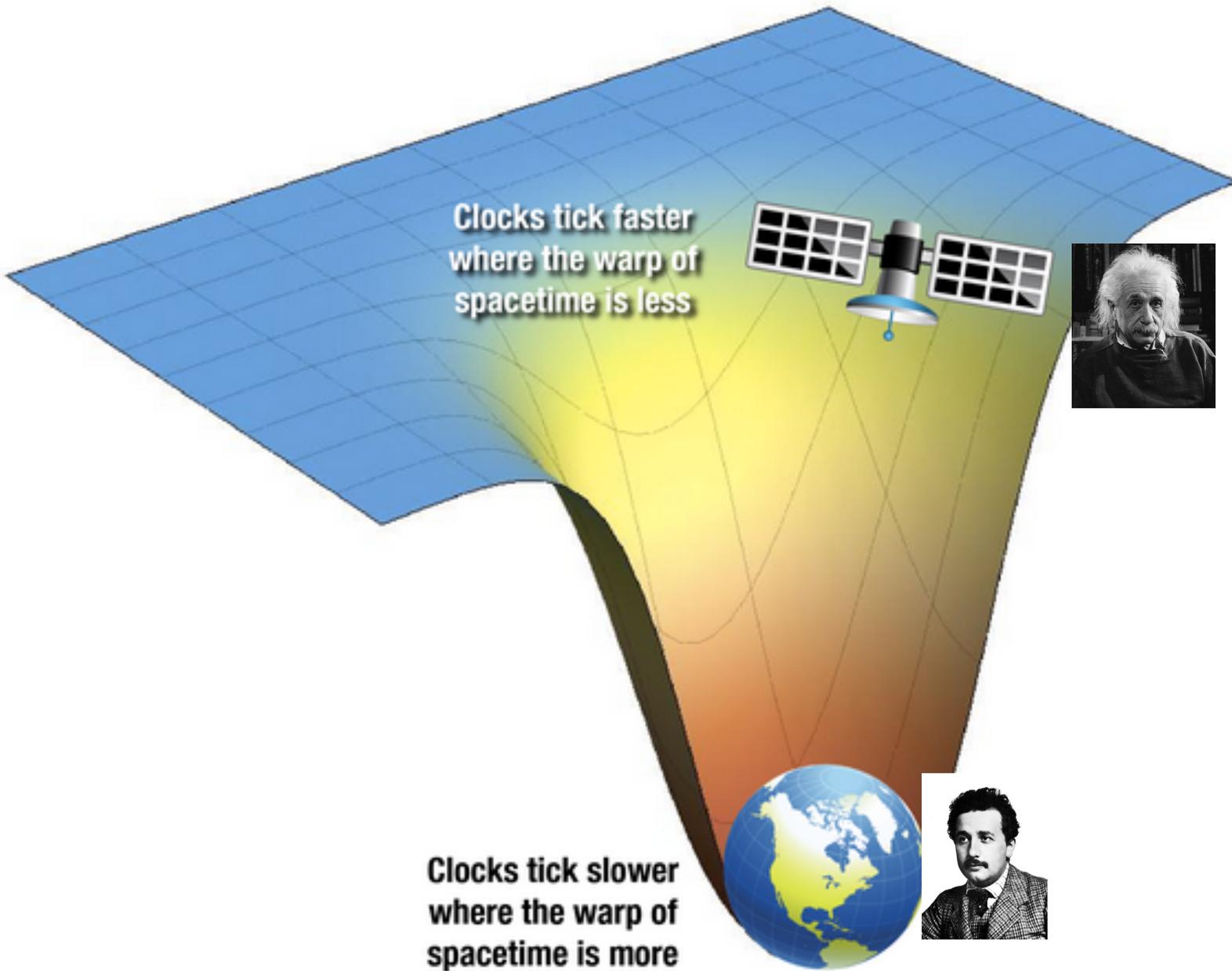
Harvard

Redshift

Pound and Rebka

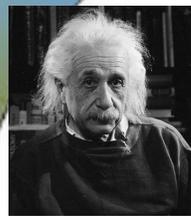
1959





Clocks tick faster
where the warp of
spacetime is less

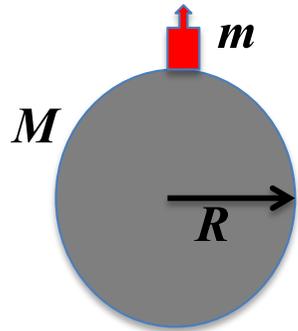
Clocks tick slower
where the warp of
spacetime is more



L'equazione di Einstein si rivela una miniera d'oro. Da questa miniera sono state estratte nel tempo altre pepite preziose come i buchi neri, le onde gravitazionali, l'universo in espansione.

John Michell

1783 → 'dark star'



But not even the light can escape if

i.e. if matter is squeezed inside a sphere of radius

KINETIC Energy

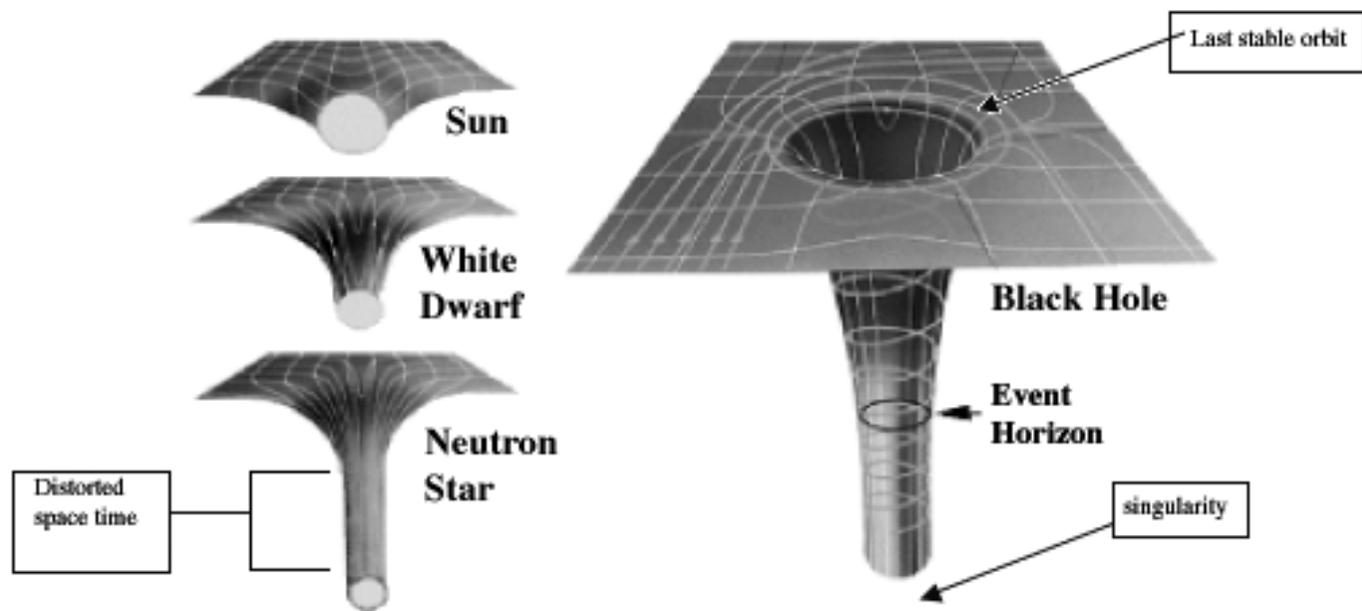
POTENTIAL Energy

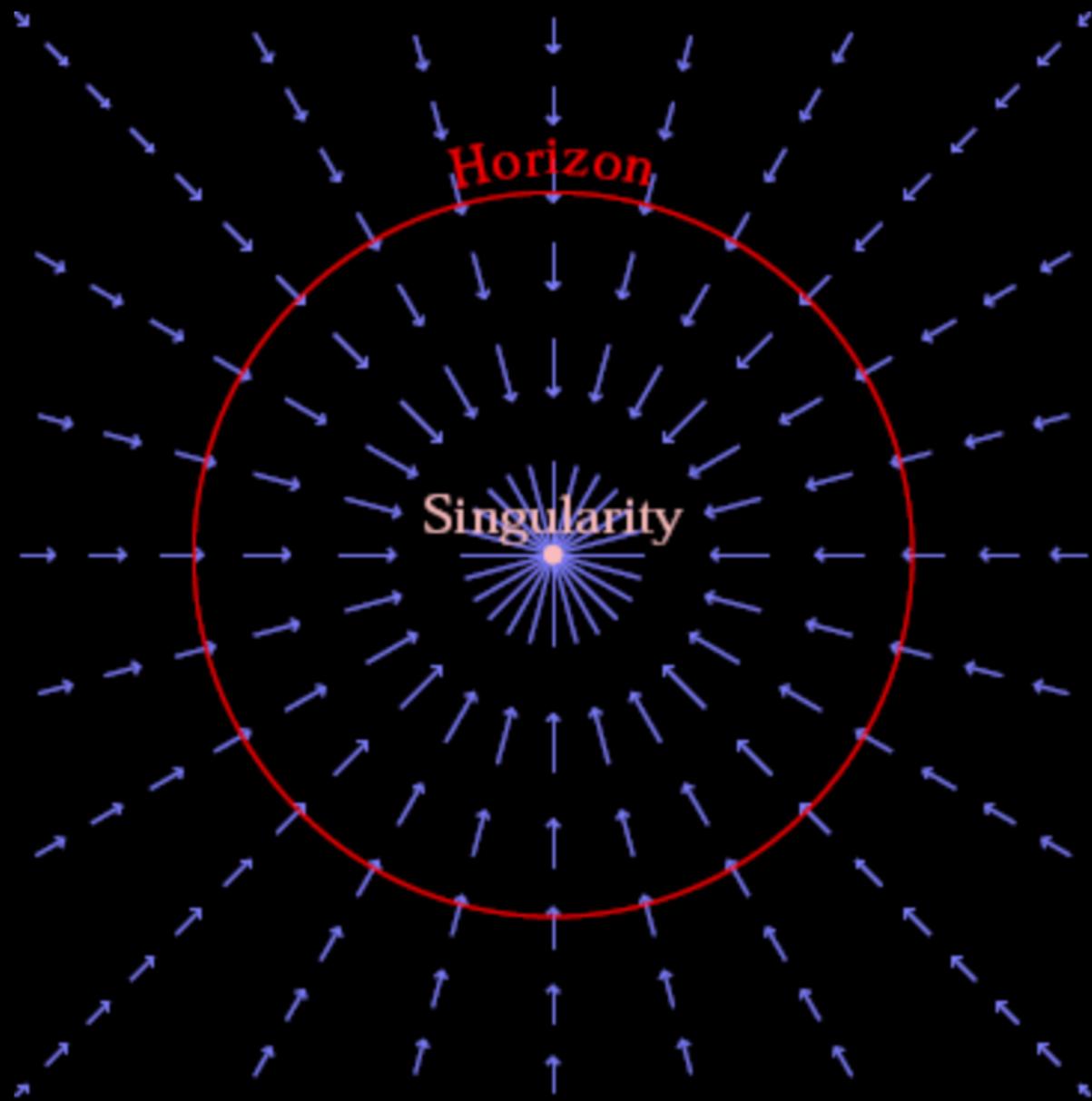
if $\frac{1}{2}mv^2 \geq G \frac{mM}{R}$ → m escapes

$G \frac{M}{R} \geq \frac{1}{2}c^2$ → M is dark

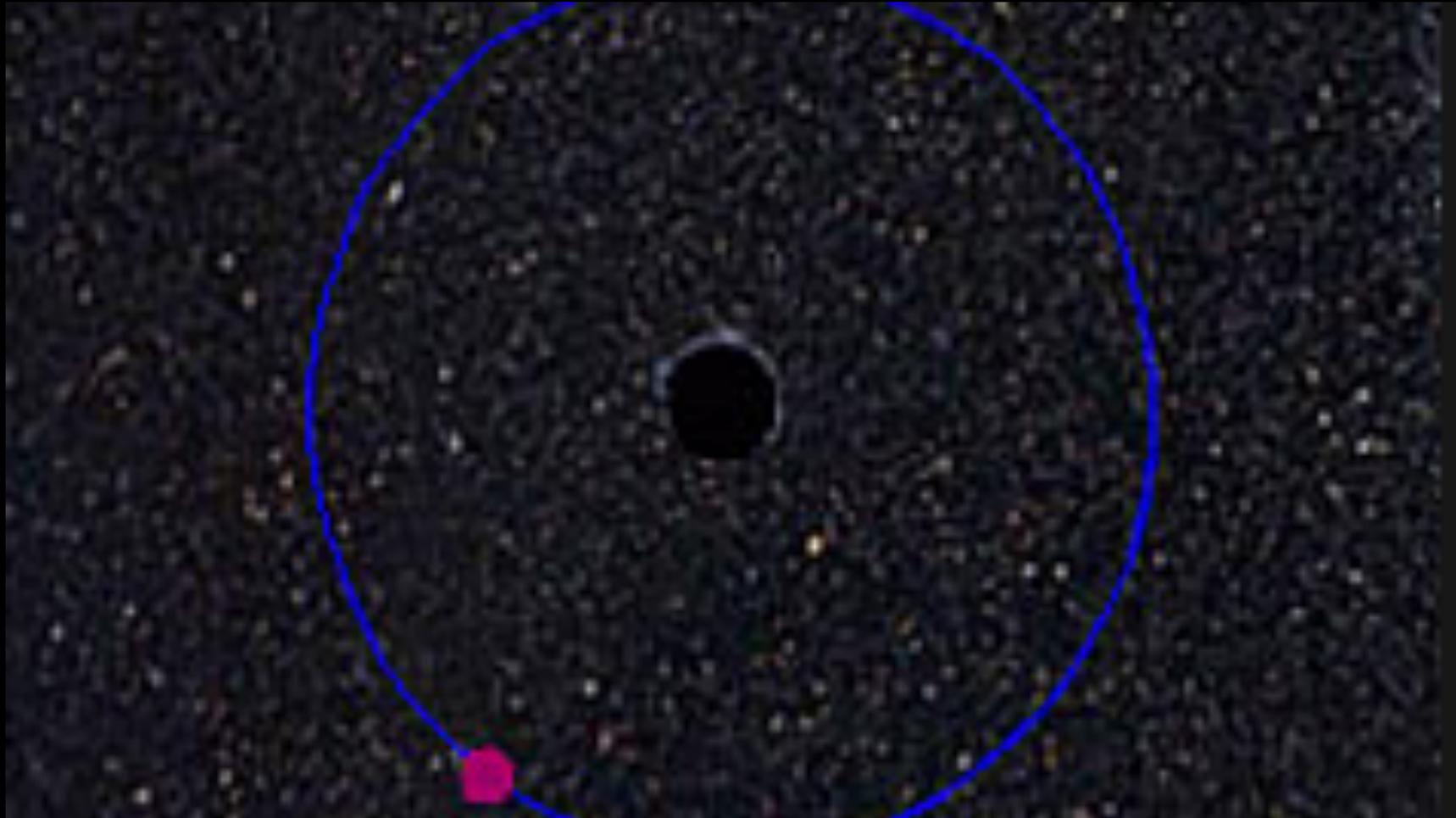
$$R = R_s = \frac{2GM}{c^2}$$

Schwarzschild radius





Time traveling

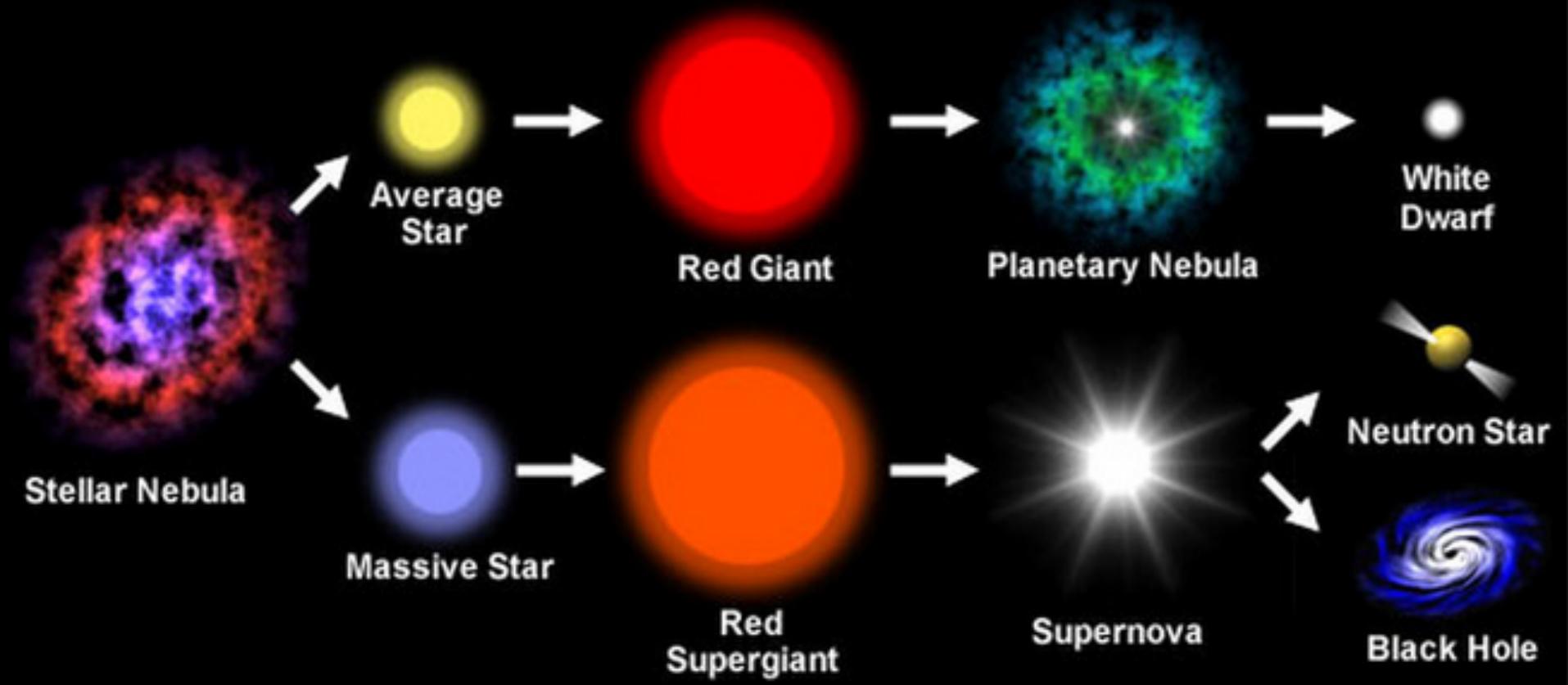


Spaghettification!



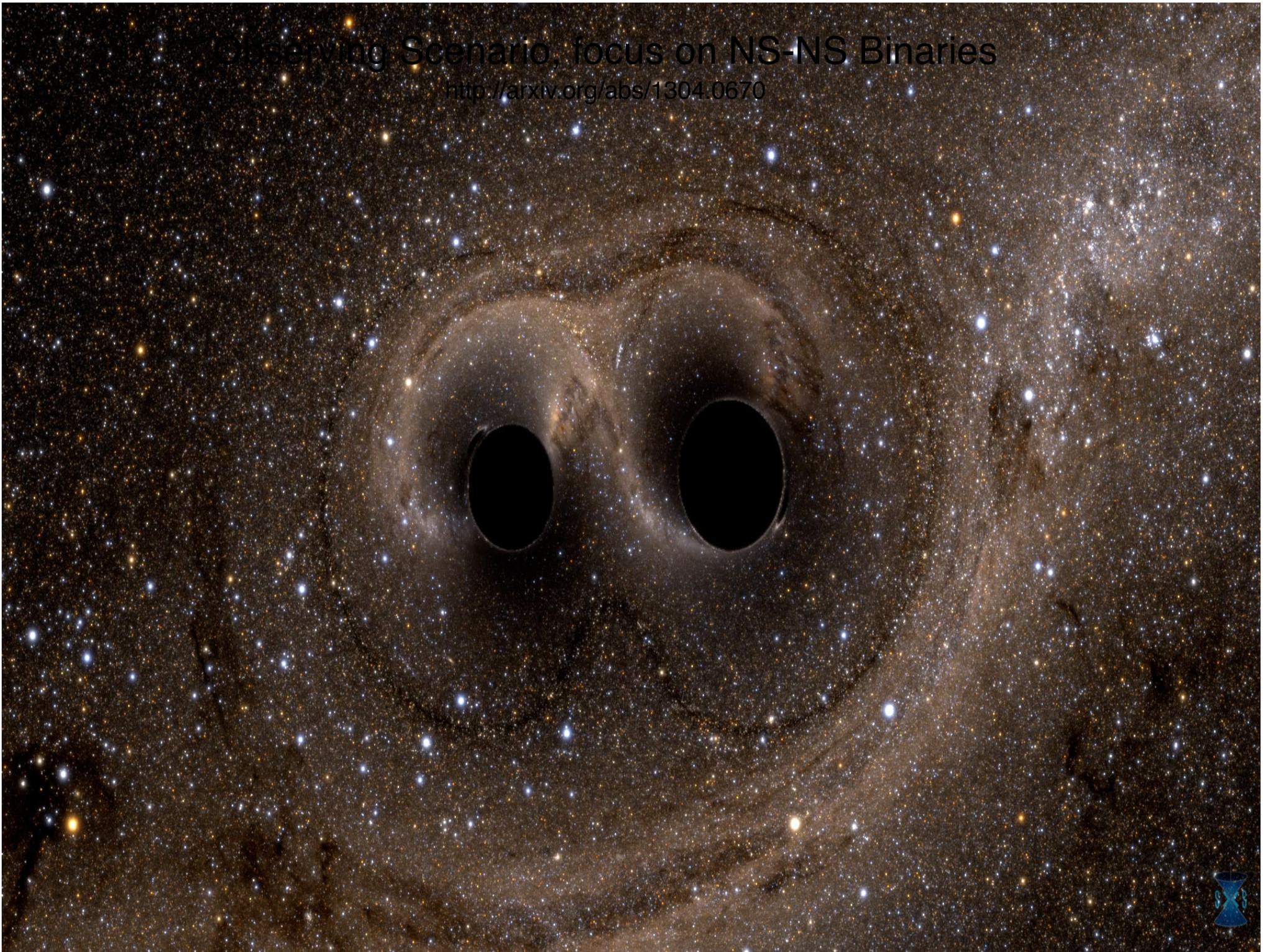
To Black Hole

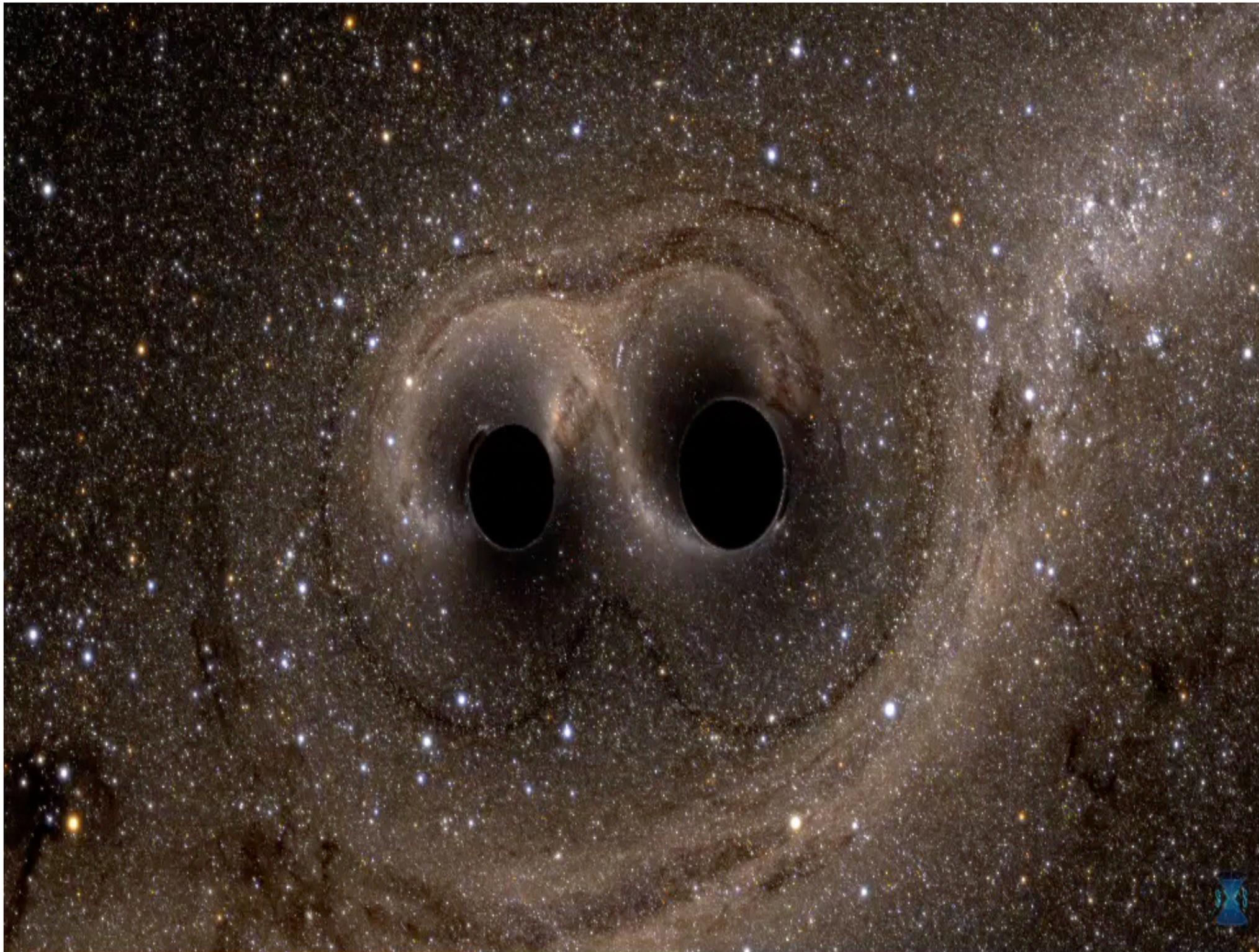
Life Cycle of a Star



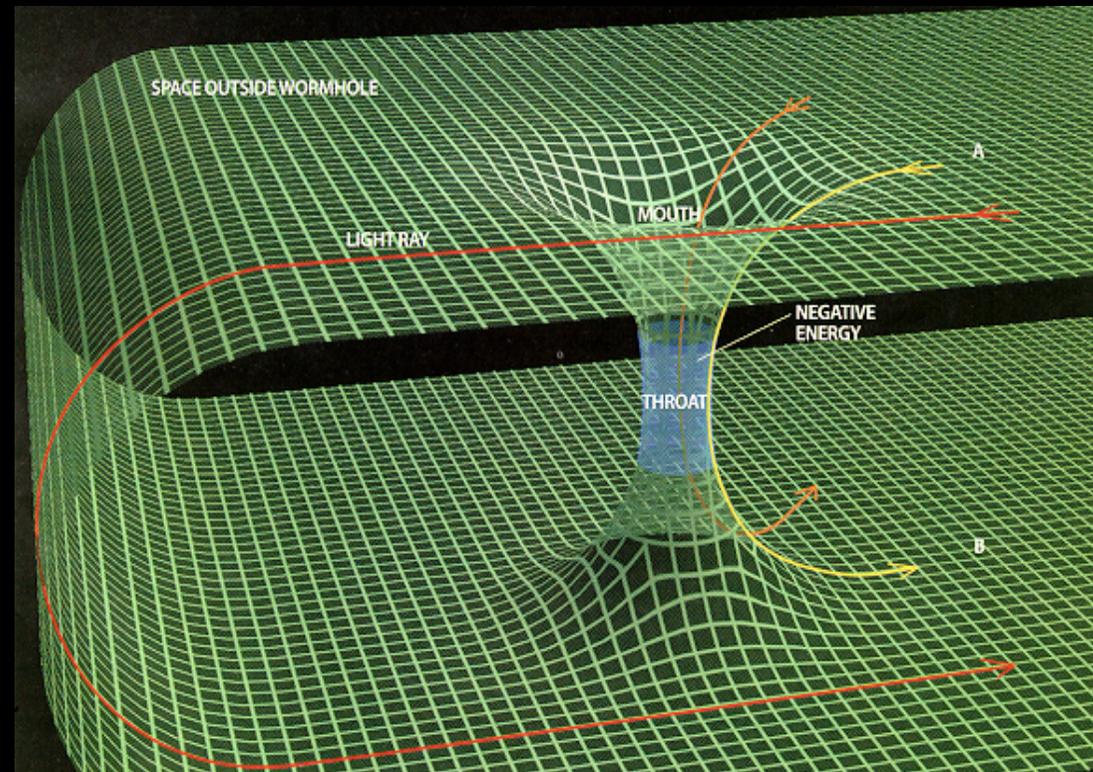
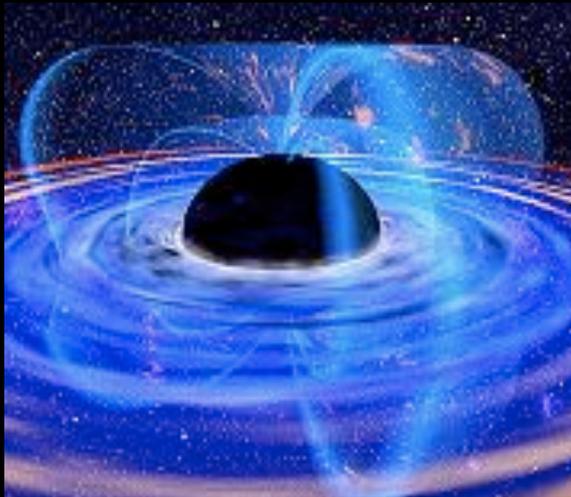
Observing Scenario, focus on NS-NS Binaries

<http://arxiv.org/abs/1304.0670>

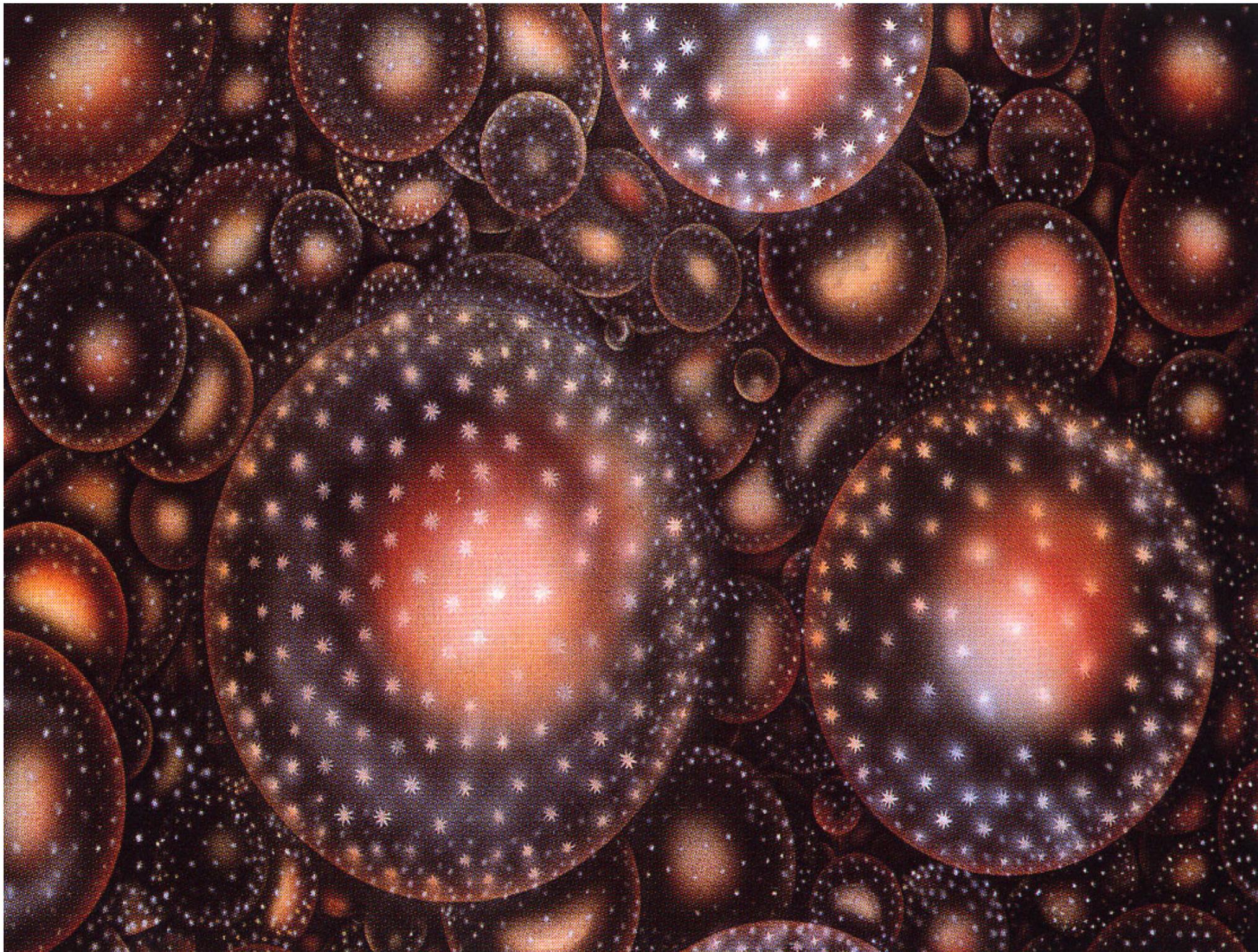


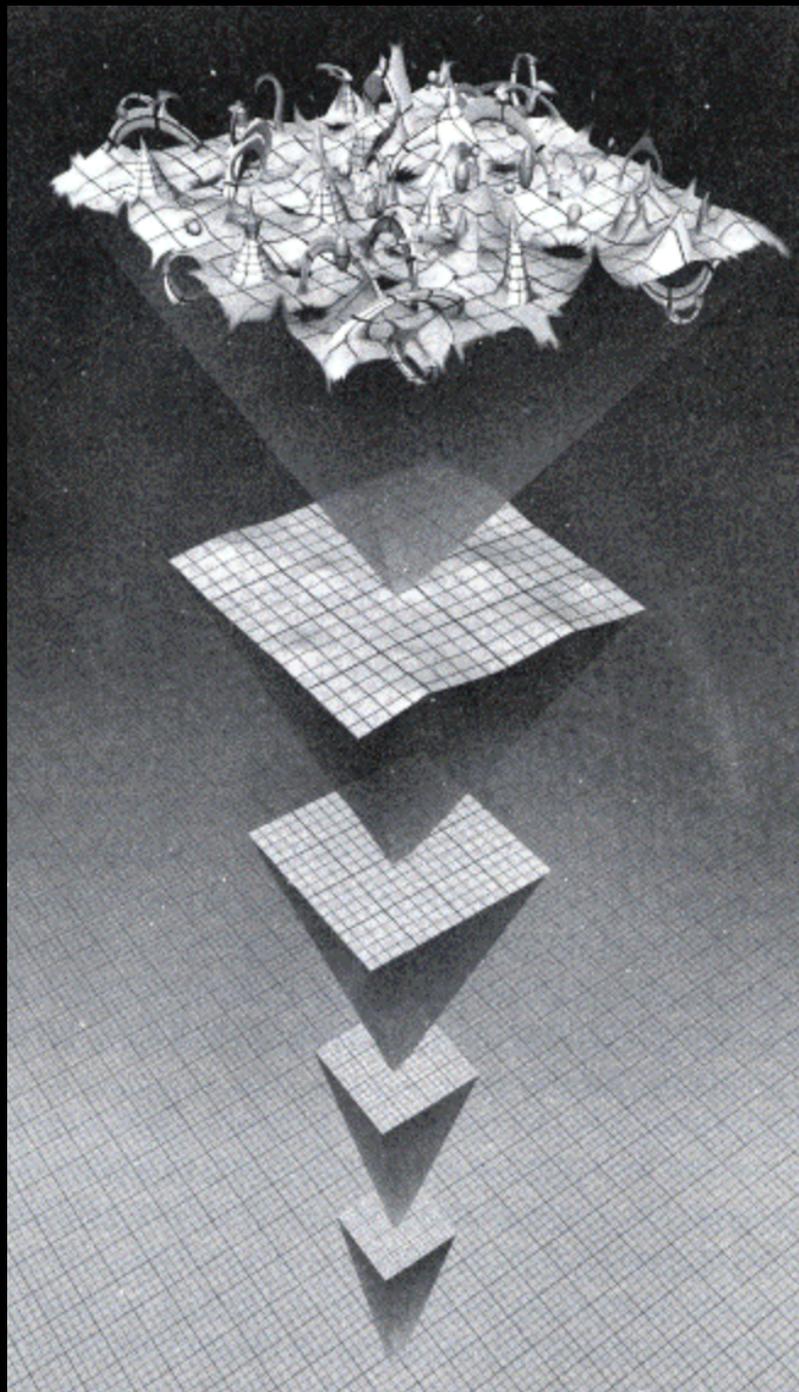


In relatività generale, al centro del buco nero la materia è compressa a densità infinita, e lo spazio tempo ha una curvatura infinita.



Questo significa che la massa del buco nero viene compressa in una regione con zero volume: la **singolarità gravitazionale**. Attorno alla singolarità c'è un confine sferico chiamato **l'orizzonte degli eventi**. Questo segna il **punto di non ritorno** oltre il quale tutto precipita inevitabilmente verso la singolarità.





$$T_{\text{H}} = \frac{\hbar c^3}{8\pi G M k_{\text{B}}}$$

