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Detectors for Absolute Luminosity Measurement for Crabbed Waist Collisions at DAFNE Phi-factory



Three different processes are used to measure the DAFNE luminosity: -The Bhabha elastic scattering e+ e- to e+ e-; polar angle range covered by calorimeters 18-27 degrees, expected rate (~440 Hz at a luminosity of 10^{32} cm-2 s-1)

-The very high rate e+ e- to e+ e- g (radiative Bhabha process); 95% of the signal in contained in a cone of 1.7 mrad aperture, but suffers heavily from beam losses due to interactions with residual gas and from Touschek effect.

- e+e- to Φ to K+K-, at a rate of about 25 Hz at 10^{32} is expected in the SIDDHARTA experiment monitor at ~90 degrees.



The SIDDHARTA preliminary setup installed at DAFNE.

Bhabha calorimeters construction

habha calorimeters

Trigger

42908

triggering

not triggering

Entries

abs(adc20)-peds(21)+abs(adc23)-peds(24)

Total energy deposited in module



1 cm scintillator tiles [Protvino], each read by 3 wavelength shifting fibers [Bicron, 1mm] 36 fibers/sector, 1 photomultiplier/sector

All tiles wrapped in a Tyvek foil







trackers

Triple-GEM





In front of each calorimeter, at a distance of 18.5cm from the IR, a ring of triple-GEM detectors is installed around the beam pipe. The two GEM trackers are divided in two units, with an half-moon shape; the top (bottom) half covers azimuthal angles [14,166] degrees ([194,346] degrees).

Each of the four GEM units is segmented into *32 pads:*

eight cells in azimuth (covering 19 degrees) are arranged in four rings of equal radial extension.

When a charged particle crosses the 3 mm drift gap, it generates electrons that will be multiplied by the three GEM foils separated *by 2/1/2 mm*.

Each of the GEM planes is made of a thin (50mm) kapton foil sandwiched between two copper clads and perforated by a dense set of holes (70mm diameter, 140 mm pitch).

As a high potential difference (about 400 kV) is applied between the copper sides, the holes act as multiplicating channels and the gain of each layer is about 20 (and hence roughly 8,000 in total). The GEM trackers, are included into the main DAQ system.

Bhabha elastic scattering e+ e- to e+ ehas a very clean signature: two energy deposits in two back-to-back s ectors.

At the trigger level, we require two energy *deposits in the sum of the 5 sectors of a couple* of back-to-back modules, with an equivalent threshold of ~200 MeV

Trigger: back-to-back coincidences of 2 modules [1 module=5 sectors] (**0** AND **3**) OR (**1** AND **2**)

In order to correct the Bhabha event rate measured using the calorimeters and the GEM trackers for detectors' acceptance and selection efficiency, we have developed a full Monte Carlo simulation of the whole experimental set-up, based on the GEANT3 package. This includes all the materials and fields

SIDDHARTA

present in the interaction region, as well as a simulation of the detectors response.



A crucial point in the Monte Carlo simulation is the evaluation of the calorimeters acceptance as a *function of:*

the minimum impact (polar) angle - the effective threshold in the trigger

BKG Accide

These are also the main contributions to the systematic error



8 lead plates [5 mm thick] + 3 lead plates [1 cm]



The four half-calorimeters installed around the DAFNE interaction region





on luminosty. We estimate an overall systematic uncertainty of 11% without using the GEM information, which improves to 7% using the GEM tracks.

Data vs. Monte Carlo

rate vs. azimuthal angle



σ _{syst}

2 %

6 %

0%

1 %

0%

0%

~7 %

Determine acceptance from Monte Carlo, after having measured the *effective threshold* using data, taking the ratio of triggering/all events Energy scale nicely set by the Bhabha peak (510 MeV).

Sources	σ _{syst}		Sources
Calo	4%		Calo
Soyuz	8%		Soyuz
Siddharta	2%		Siddharta
Threshold	5%		Threshold
SKG Accidentals	3%		BKG Accidentals
BKG	0.1%		BKG
Total (Σ_{quad})	11%		Total (Σ_{quad})

Systematic error

Acceptance & threshold

Adding GEM tracking

Radiative Bhabha monitors







A filtering process has been implemented in the DAQ system, for on-line correction of the rate by subtracting the percentage of background events.

The correction is estimated analyzing blocks of events (typically $1000 \sim 1s$ at 1 kHz), and by performing a cut on the distribution of the time of the two triggering modules. In the difference of the arrival time of Bhabha *candidates (for a couple of triggering modules)* we can see two contributions: - a Gaussian distribution peaked at $\Delta t=0$ (good Bhabha's); - a flat distribution due to **background**.







The diagnostics installed on the new DAFNE IR in order to measure luminosity for the test of the new crab waist scheme, started to operate at the beginning of February 2008 and has collectied encouraging results from the machine.



by DAFNE slow control system with a sampling time of 15 s, and are available for offline analysis and on the word wide web for online performance presentation.





*PbWO*₄ *crystals*

Two gamma monitor detectors are located 170 cm away from the IR, collecting photons radiated by electron or positron beam. Each counter is made by **4 PbWO crystals** (squared section, 30mm×30mm, and 110mm height) assembled together along z: the 30 mm face towards the photon beam, with a total depth of 120 mm corresponding to about $13X_{0}$. Each crystal is readout by a Hamamatsu R7600 compact photomultiplier. Prompt estimate of the luminosity is provided on-line by the DAFNE control system in order to perform machine optimization and characterization.

An on-line vertical beam-beam scan performed by acquiring the relative luminosity estimate of the crystal gamma monitor.



