

To be submitted to  
Nuovo Cimento

ISTITUTO NAZIONALE DI FISICA NUCLEARE  
Laboratori Nazionali di Frascati

LNF -78 / 1(P)  
2 Gennaio 1978

S. Bartalucci, G. Basini, S. Bertolucci, J. K. Bienlein, M. Fiori,  
P. Giromini, R. Laudan, E. Metz, C. Rippich and A. Sermoneta:  
MEASUREMENT OF THE PHOTOPRODUCTION PHASES  
OF THE  $\varrho$ ,  $\omega$ , AND  $\Phi$  MESONS.

INFN - Laboratori Nazionali di Frascati  
Servizio Documentazione

LNF-78/1(P)  
2 Gennaio 1978

MEASUREMENT OF THE PHOTOPRODUCTION PHASES OF THE  $\varrho$ ,  $\omega$ , AND  $\Phi$  MESONS.

S. Bartalucci, G. Basini, S. Bertolucci, J.K. Bienlein, M. Fiori, P. Giromini, R. Laudan,  
E. Metz, C. Rippich and A. Sermoneta

Deutsches Elektronen-Synchrotron DESY, Hamburg  
and  
INFN - Laboratori Nazionali di Frascati.

ABSTRACT. -

We measured 30000 wide-angle electron-positron pairs from the reaction  $\gamma + p \rightarrow p + e^+ + e^-$  in the invariant mass region  $500 \leq m \leq 1060$  MeV. The photoproduction amplitudes of the  $\varrho$ ,  $\omega$ , and  $\Phi$  mesons were measured to deviate from being pure imaginary by  $37.5^\circ {}^{+2.8^\circ}_{-3.1^\circ}$ ,  $29.6^\circ {}^{+15.5^\circ}_{-12.9^\circ}$ ,  $3.4^\circ {}^{+5.3^\circ}_{-4.2^\circ}$  respectively.

-----

We determined, for the process  $\gamma p \rightarrow pV$  ( $V \rightarrow e^+e^-$ ), the ratio of the real to imaginary part of the forward photoproduction cross section

$$\left. \frac{d\sigma}{dt} \right|_{t=0} \left[ \begin{array}{l} \gamma p \rightarrow pV \quad (V \rightarrow e^+e^-) \end{array} \right]$$

for the three vector mesons  $V = \varrho$ ,  $\omega$ , and  $\Phi$ , by studying with the 2.7 GeV bremsstrahlung beam of DESY the  $e^+e^-$  yields of the reaction



in the energy region 3.0-7.2 GeV and in the  $e^+e^-$  invariant mass range  $500 < m < 1060$  MeV.

Four diagrams contribute to the reaction (1) (Fig. 1a). The first two are the Bethe-Heitler diagrams which can be calculated<sup>(1)</sup>. The third one is the diffraction production diagram where the  $e^+e^-$  pairs are coming from leptonic decays of  $\varrho$ ,  $\omega$ , and  $\Phi$ . The last diagram represents the  $e^+e^-$  contribution when the vector mesons are produced via one-pion exchange or inelastic channels.

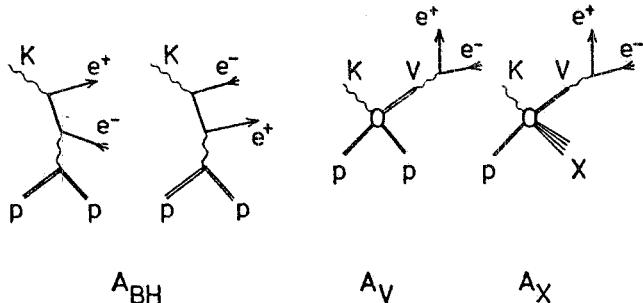


FIG. 1a - Feynman diagrams  
for reaction (1).

The second order, the  $e^+e^-$  pair production rate is expressed by

$$R \propto |A_{BH}|^2 + \left| \sum_{V(V=\varrho, \omega, \Phi)} A_V \right|^2 + |A_X|^2 + 2 A_{BH} \operatorname{Re} \left\{ \sum_V A_V \right\}^{(*)} + \text{QED terms connected to more than two } \gamma \text{ rays.} \quad (2)$$

In the framework of the VDM,  $A_V$  is parametrized as

$$A_V \propto \frac{m_V^2}{m^2} \frac{i e^{i\varphi_V}}{m_V^2 - m^2 - i m_V \Gamma_V} \sqrt{\frac{d\sigma_V}{dt} \frac{\Gamma_V}{m_V}}, \quad (3)$$

using the notation

$$\frac{d\sigma_V}{dt} = \frac{d\sigma}{dt} (\gamma p \rightarrow pV (V \rightarrow e^+e^-)) = \frac{d\sigma}{dt} (\gamma p \rightarrow Vp) \Big|_{t=0} e^{b_V t} BR_{e^+e^-}^{(2)}.$$

Here  $m_V$  is the mass of the vector meson,  $\Gamma_V$  its width,  $\varphi_V + \pi/2$  is the relative phase between the vector meson photoproduction amplitude and the BH amplitude (which is real),  $b_V$  is the slope of the differential cross section for vector meson photoproduction on protons, and  $m$  and  $t$  are the invariant mass and the momentum transfer of the  $e^+e^-$  pair, respectively.

It follows from charge-conjugation invariance that the interference term in (2) between the Bethe-Heitler and the Compton amplitudes is antisymmetric under exchange of  $e^+$  and  $e^-$ . Hence, the effect of the interference term consists in producing an asymmetric distribution of the experimental events as a function of any kinematical variable which is antisymmetric in the four-momenta  $p_+$  and  $p_-$ . Such an asymmetry is a measurement of the interference term and consequently of the photoproduction phase  $\varphi_V$ . Any contributions to the interference term due to second order QED terms have been estimated<sup>(3)</sup> to be about 0.1% of the BH rate and have therefore been neglected.

The apparatus (Fig. 1b) and the experimental procedure are the same as described elsewhere<sup>(4)</sup> in detail. In particular let us recall that our mass resolution is 20 MeV (FWHM) at  $m = 1$  GeV.

(\*) OPE (spin-flip amplitude) and inelastic amplitudes do not interfere with the Bethe-Heitler term.

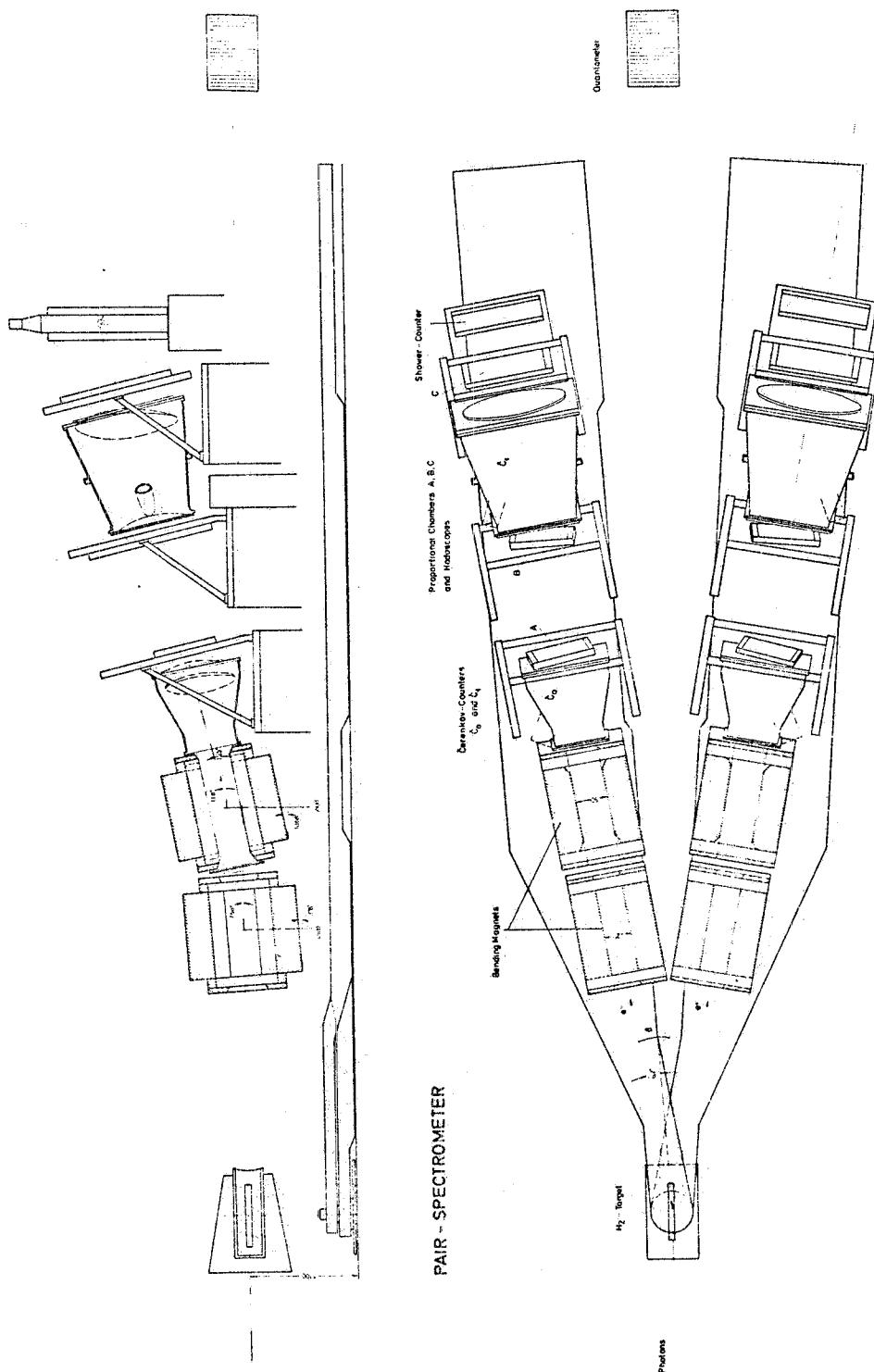


FIG. 1b - Plan and side view of the spectrometer

We have taken data with a beam top-energy  $K_{\max} = 7.2$  GeV, a central spectrometer momentum  $p_0 = 2700$  MeV at three spectrometer opening angles  $\theta = 13^\circ, 15^\circ$  and  $16^\circ$ , and a fourth setting was defined by  $K_{\max} = 6.0$  GeV,  $p_0 = 1700$  MeV, and  $\theta = 13^\circ$ . For each setting, the same amount of data was collected for each of the two spectrometer polarity combinations in order to cancel out minor asymmetries of the apparatus.

The total number of  $e^+e^-$  events, measured in the mass range  $500 \leq m \leq 1060$  MeV, is 29217 which is to be compared with Monte-Carlo predictions of  $22576 \pm 280$  events attributed to Bethe-Heitler and  $7106 \pm 160$  to  $\rho + \omega + \Phi$  production. The Bethe-Heitler rate is corrected for second order<sup>(5)</sup> and inelastic<sup>(1), (6)</sup> terms and is normalized to the collected number of equivalent quanta.

The invariant mass spectrum of the data collected at  $13^\circ$  is shown in Fig. 2. The interference spectra for the four different setting are shown in Fig. 3.

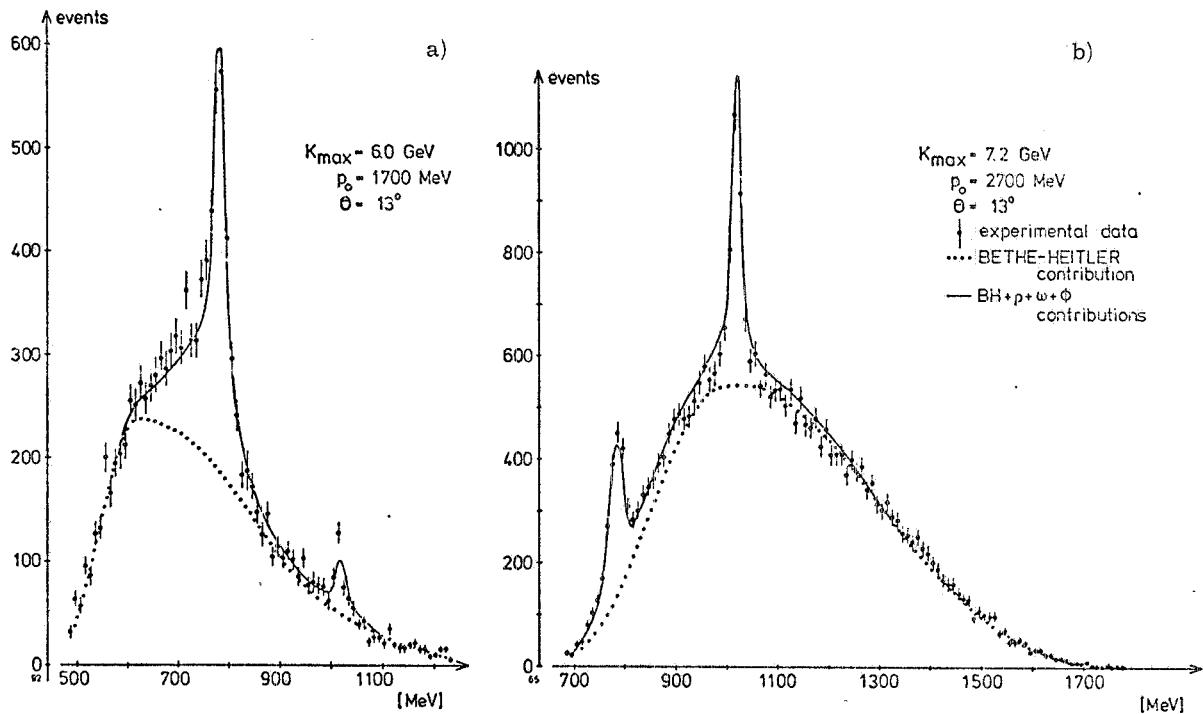


FIG. 2 -  $e^+e^-$  effective distribution. The open circles are the calculated contributions of the BH processes, normalized to the collected number of equivalent quanta. The solid line shows the calculated contribution of the BH processes together with the diffractive  $\rho$ ,  $\omega$ , and  $\Phi$  photoproduction, the  $\rho + \omega$  inelastic effects, and  $\omega$ -OPE production.  
a)  $K_{\max} = 6.0$  GeV,  $p_0 = 1700$  MeV,  $\theta = 13^\circ$ ; b)  $K_{\max} = 7.2$  GeV,  $p_0 = 2700$  MeV,  $\theta = 13^\circ$ .

Due to the large uncertainties in the calculation of the  $\rho$ ,  $\omega$ , and  $\Phi$  inelastic and of the  $\omega$  OPE contributions to the invariant mass yield, only the measured interference pattern has been used to fit the photoproduction phases.

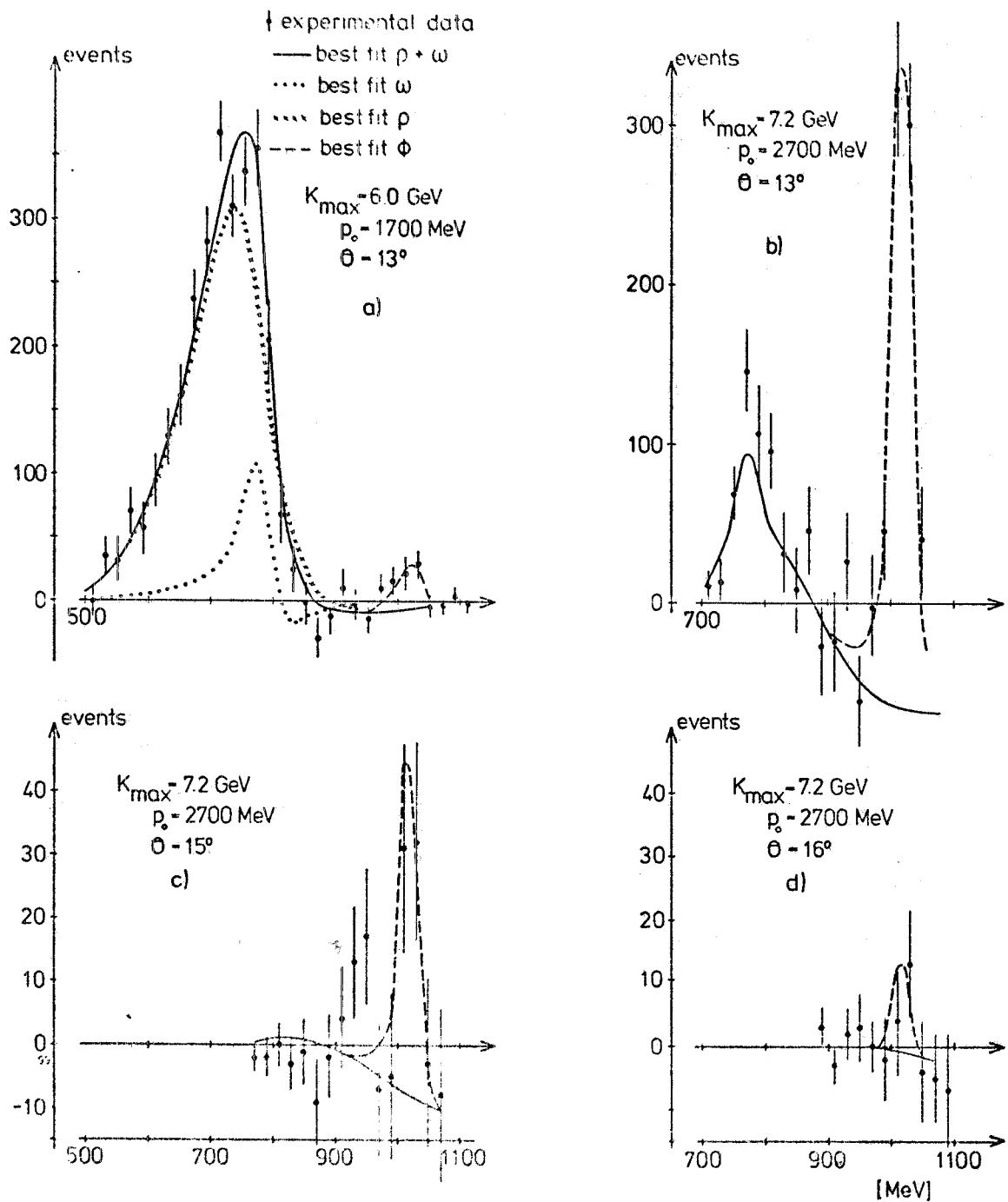


FIG. 3 - Spectrum of interference events. The solid line shows the best fit to the data with the  $\rho + \omega$  interference term of eq. (2), using  $m_\rho = 767.6 \text{ MeV}$ ,  $\Gamma_\rho = 150.9 \text{ MeV}$ ,  $C_\rho = 4.55 \text{ nb GeV}^{-2}$ ,  $\varphi_\rho = 37.5^\circ$ ,  $C_\omega = 0.735 \text{ nb GeV}^{-2}$ ,  $\varphi_\omega = 29.6^\circ$ . In Fig. a) the open circles and the crosses show the  $\rho$  and  $\omega$  contributions also separately. The dashed line is the best fit to the data with the  $\phi$  interference term of eq. (2), using  $m_\phi = 1023.6 \text{ MeV}$ ,  $C_\phi = 0.98 \text{ nb GeV}^{-2}$ , and  $\varphi_\phi = 3.44^\circ$ .  
 a)  $K_{\max} = 6.0 \text{ GeV}$ ,  $p_0 = 1700 \text{ MeV}$ ,  $\theta = 13^\circ$ ; b)  $K_{\max} = 7.2 \text{ GeV}$ ,  $p_0 = 2700 \text{ MeV}$ ,  $\theta = 13^\circ$ ;  
 c)  $K_{\max} = 7.2 \text{ GeV}$ ,  $p_0 = 2700 \text{ MeV}$ ,  $\theta = 15^\circ$ ; d)  $K_{\max} = 7.2 \text{ GeV}$ ,  $p_0 = 2700 \text{ MeV}$ ,  $\theta = 16^\circ$ .

Results for the  $\varrho$  and the  $\omega$ .

We fitted the  $\varrho$  and  $\omega$  interference contribution of eq. (2) to the data between 500 and 960 MeV, using  $m_\varrho$ ,  $\Gamma_\varrho$ ,  $\varphi_\varrho$ ,  $d\sigma_\varrho/dt$ ,  $m_\omega$ ,  $d\sigma_\omega/dt$ , and  $\varphi_\omega$  as free parameters.

We have used  $\Gamma_\varrho = 10 \text{ MeV}^{(7)}$ ,  $b_\varrho = 6.3 \text{ GeV}^{-2} {}^{(9)}$  and  $b_\omega = 6.7 \text{ GeV}^{-2} {}^{(9)}$ .  $\Gamma_\omega$  is assumed to be mass independent, while  $\Gamma_\varrho$  has been parametrized<sup>(8)</sup> as

$$\Gamma_\varrho(m) = \Gamma_\varrho \frac{m_\varrho}{m} \left( \frac{m^2 - 4m_\pi^2}{m_\varrho^2 - 4m_\pi^2} \right)^{3/2} \frac{1 + \frac{R^2}{4} (m_\varrho^2 - 4m_\pi^2)}{1 + \frac{R^2}{4} (m^2 - 4m_\pi^2)}$$

The skewness parameter  $R$  was also left free in the fit.

The energy dependence of the photoproduction cross section has been introduced using the measured values of<sup>(9, 7)</sup> as

$$\frac{d\sigma}{dt} \Big|_{t=0} (\gamma p \rightarrow p\varrho \rightarrow (\varrho \rightarrow e^+e^-)) = C_\varrho (1 + \frac{6 \text{ GeV}^2}{K^2}) \text{ nb GeV}^{-2},$$

$$\frac{d\sigma}{dt} \Big|_{t=0} (\gamma p \rightarrow p\omega \rightarrow (\omega \rightarrow e^+e^-)) = C_\omega (1 + \frac{1.4 \text{ GeV}}{K}) \text{ nb GeV}^{-2}.$$

The published values of  $C_\varrho = 3.65 \pm 0.73 {}^{(7, 9)}$ ,  $C_\omega = 0.7 \pm 0.14 {}^{(7, 9)}$ ,  $m_\varrho = 770 \pm 3 \text{ MeV} {}^{(7)}$ ,  $\Gamma_\varrho = 150 \pm 3 \text{ MeV} {}^{(7)}$  and  $m_\omega = 782.7 \pm 3 \text{ MeV} {}^{(7)}$  have been further used as additional degrees of freedom of the fit<sup>(11)</sup>.

In fitting the interference term of eq. (2), we used the program MINUIT<sup>(12)</sup>. The errors quoted are true (MINOS)  $1\sigma$  errors. The best fit is plotted in Fig. 3 (solid line), and the corresponding parameters are listed in Table I. The resulting contributions from the  $\varrho$  and the  $\omega$  are shown separately in Fig. 3a.

TABLE I - Fitted parameters<sup>(x)</sup>

$m_\varrho$	767.6	$\pm$	2.7	MeV	$\chi^2_{\text{NDF}} = \frac{58}{47}$
$\Gamma_\varrho$	150.9	$\pm$	3	MeV	
$R$	0.05	$\pm$	0.52	$\text{GeV}^{-1}$	
$C_\varrho$	4.55	$\pm$	0.32	$\text{nb GeV}^{-2}$	
$m_\omega$	783.3	$\pm$	2.9	MeV	
$C_\omega$	0.735	$\pm$	0.13	$\text{nb GeV}^{-2}$	
$\varphi_\varrho$	$37.5^\circ$	$\pm$	$2.9^\circ$		
$\varphi_\omega$	$29.6^\circ$	$\pm$	$15.5^\circ$		
				$\pm 12.9^\circ$	

(x) - The quoted errors are  $1\sigma$  MINOS errors.

To be sure that the phase  $\varphi_\Phi$  is not affected by the  $\omega$  parameters, we repeated the fit leaving out the mass bins from 760 to 820 MeV, where the  $\omega$  contribution depends strongly on the  $\varphi_\omega$  value. The best fit results did not change.

### Results for the $\Phi$ .

Fitting for  $\varphi_\Phi$  was done in the mass region  $950 < m < 1050$  MeV. The cross section

$$\left. \frac{d\sigma}{dt} \right|_{t=0} (\gamma p \rightarrow p\Phi(\Phi \rightarrow e^+e^-))$$

was parametrized<sup>(7, 10, 11)</sup> as

$$C_\Phi \left( 1 - e^{-\frac{K}{2.73(\text{GeV})}} \right) \text{nb GeV}^{-2}$$

In order to study the effect of  $C_\Phi$  and  $m_\Phi$  on the phase value, these parameters were allowed in the fit to vary around the published values  $C_\Phi = 0.9 \pm 0.13$  nb GeV $^{-2}$ <sup>(7, 10)</sup> and  $m_\Phi = 1020 \pm 3$  MeV<sup>(7, 10)</sup>. The slope  $b_\Phi$  was taken as  $5.5$  GeV $^{-2}$ <sup>(10)</sup> and the width  $\Gamma_\Phi$  was set = 4 MeV<sup>(7)</sup>. The result of the fit is shown in Figs. 3 and 4 and the best fit parameters are given in Table II.

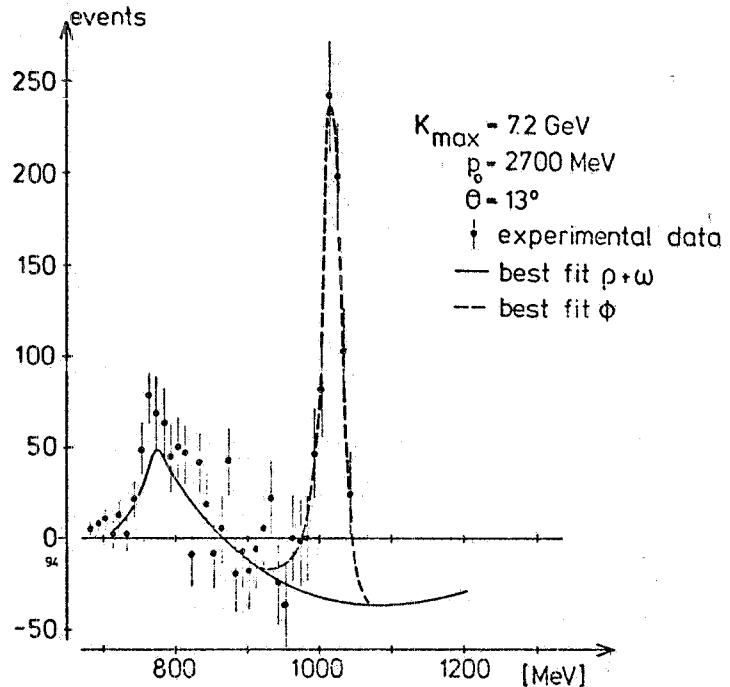


FIG. 4 - Interference spectrum as in Fig. 3b, plotted in 10 MeV bins.

TABLE II - Fitted parameters<sup>(x)</sup>

$\varphi_\Phi$	$3.44^\circ \pm 5.3^\circ$ $\quad \quad \quad \pm 4.2^\circ$	$\chi^2_{\text{NDF}} = \frac{7.6}{9}$
$C_\Phi$	$0.98 \pm 0.12 \text{ nb GeV}^{-2}$	
$m_\Phi$	$1023.6 \pm 2 \text{ MeV}$	

(x) - The quoted errors are 1σ MINOS errors.

The sensitivity of  $\varphi_\Phi$  to the  $\varrho$  tail under the  $\Phi$  peak was found negligible (a change of  $\pm 20\%$  of the  $\varrho$  tail values affects  $\varphi_\Phi$  by  $\pm 1.5^\circ$ ).

In Figs. 2 and 5 we show the various  $\varrho$ ,  $\omega$ , and  $\Phi$  contributions to the  $e^+e^-$  invariant mass yield. The  $\varrho$  contributions are calculated according to our best fit parameter. The  $\omega$  and  $\Phi$  contributions are calculated using photoproduction cross sections into  $e^+e^-$  pairs as from refs. (9, 10) and  $\varphi_V$  from our fit. The inelastic  $\varrho$  and  $\omega$  production was estimated from ref. (13) and the  $\omega$  OPE contribution from ref. (14).

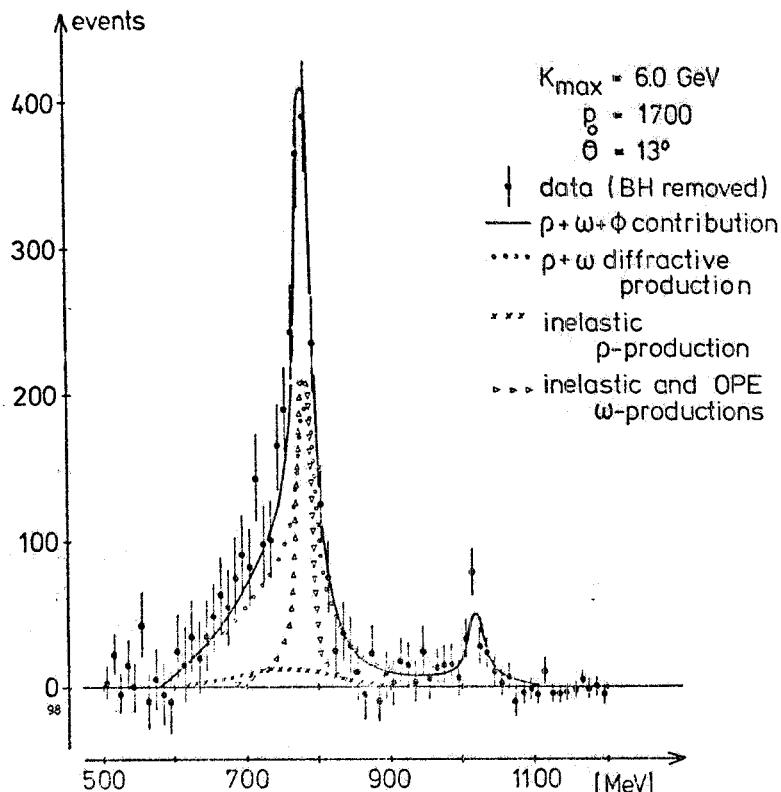


FIG. 5 -  $e^+e^-$  mass yield for  $K_{\max} = 6.0$  GeV,  $p_0 = 1700$  MeV, and  $\theta = 13^\circ$  after removing the calculate BH contributions (see also Fig. 2a). Shown are the separate effects of  $\varrho + \omega + \Phi$  diffractive production, inelastic  $\varrho$  production,  $\omega$  production via OPE, and inelastic  $\omega$  production. The combined sum of these effects is indicated by the solid line.

### Summary.

By studying the interference term between the Compton and the Bethe-Heitler amplitudes, we measured again the  $\varrho$  photoproduction cross section into lepton pairs, quite well in agreement with the published values<sup>(9)</sup>, but our measurement does not prove more sensitive than previous experiments<sup>(7, 9, 10)</sup> with regard to the parameters  $C_\omega$ ,  $C_\Phi$ , as well as to the  $\varrho$ ,  $\omega$ , and  $\Phi$  mass values.

The major result of the experiment concerns the measurement of the  $\varrho$ ,  $\omega$ , and  $\Phi$  photoproduction phases.  $\varphi_\Phi$  is measured to be  $3.4^\circ {}^{+5.3^\circ}_{-4.2^\circ}$ ; this is consistent with the  $\Phi N$  scattering being purely imaginary. The  $\varrho$  and  $\omega$  nucleon scattering are found to deviate from a pure absorptive mechanism by  $37.5^\circ {}^{+2.8^\circ}_{-3.1^\circ}$  and  $29.6^\circ {}^{+15.5^\circ}_{-12.9^\circ}$ , respectively.

A breakdown of the simple " $\varrho$ -photon analogy" and an anomalous large real part in the  $\varrho$  photoproduction amplitude is expected as a dynamical consequence of meson instability inside the nucleous. This effect has been qualitatively estimated to be of the order of

$$\frac{2\Gamma_V}{m_V} \frac{\sigma_T(\pi N)}{\sigma_T(VN)} \quad (13)$$

. Numerically, this amounts to  $\sim 22^\circ$  for the  $\varrho$ , independent of energy, while

it is negligible for the  $\omega$ . On the other hand, because of the  $\varrho - \omega$  mass mixing effect,  $\varphi_\omega$  has to be reduced by  $\sim 11^\circ$  according to ref. (16).

Taking into account such corrections, our result is consistent, inside the errors, with a small refractive part of the Compton amplitude also for the  $\varrho$  and  $\omega$  nucleon forward scattering, as is expected if the reaction proceeds via Pomeron exchange.

Our measured phase values may be compared with the values  $\varphi_\omega - \varphi_\varrho = 41^\circ \pm 22^\circ$ ,  $\varphi_\varrho = 11.8^\circ \pm 4.4^\circ$ , and  $\varphi_\Phi = 25^\circ \pm 15^\circ$  obtained by the DESY-MIT group at  $K = 5.6$  GeV (17, 18). The  $\varrho$  and  $\omega$  phase values have also been measured at DNPL<sup>(19)</sup> to be  $\varphi_\omega - \varphi_\varrho = 118^\circ \pm 13^\circ$  and  $\varphi_\varrho = 16.5^\circ \pm 6.2^\circ$  at  $K = 3.6$  GeV.

#### Acknowledgements.

We are grateful for the support of the DESY Direktorium, who made this collaboration possible. The encouragement of Profs. P. Waloschek, G. Bellettini, and I. Mannelli, was a valuable help. We thank F. Gutbrod, P. Söding, M. Greco for many interesting comments. We thank V. Chiarella for participation in parts of the runs, and D. Fong, C. Bradaschia, and T. Mc-Corriston who, in the past, largely contributed to the development of the data analysis programs we used in this experiment. We are grateful to Fr. U. Rehder, D. Habercorn, H. Lenzen, K. Löffler, and M. Schneider for their skillful technical assistance. The warm cooperation of the Accelerator Group, of the Technical Support Groups, and of the Computer Center of DESY has been essential for the success of the experiment. A. S. thanks the Blanceflor Stieftelsen for financial support.

#### REFERENCES AND NOTES.

- (1) - S. D. Drell and J. D. Walecka, Ann. Phys. 28, 18 (1964); M. L. Perl, T. Braunstein, F. Cox, F. Martin, W. T. Toner, B. D. Dieterle, T. F. Zipf, W. L. Lakin and H. C. Bryant, Phys. Rev. Letters 23, 1191 (1969).
- (2) - Br<sub>e+e-</sub> is defined as  $\frac{a^2}{12} \frac{4\pi}{\gamma_V^2} \frac{m_V}{\Gamma_V}$ , where  $\frac{\gamma_V^2}{4\pi}$  is the vector meson-photon coupling constant.
- (3) - J. Brodsky and J. G. Gillespie, Phys. Rev. 173, 1011 (1968).
- (4) - S. Bartalucci, S. Bertolucci, M. Fiori, D. Fong, T. Mc-Corriston, P. Giromini, S. Guiducci, C. Rippich, M. Rohde, A. Sermoneta and L. Trasatti, Nuovo Cimento 39A, 374 (1977).
- (5) - B. Huld, Phys. Rev. 168, 1782 (1968).
- (6) - M. Damashek and F. Gilman, Phys. Rev. D1, 1319 (1970).
- (7) - Review of Particle Properties, Rev. Mod. Phys. 48, No. 2, Part II (1976).
- (8) - J. Pisut and M. Roos, Nuclear Phys. B6, 352 (1974).
- (9) - J. Ballam, G. B. Chadwick, R. Gearhart, Z. G. T. Guiragossian, F. F. Murray, P. Seyboth, A. H. Rosenfeld, C. K. Sinclair, I. O. Skillicorn, H. Spitzer, G. Wolf, H. H. Bingham, W. B. Fretter, K. C. Moffeit, W. F. Podolsky, M. S. Rabin, A. H. Rosenfeld and R. Windmolders, Phys. Rev. D5, 545 (1972); J. Ballam, G. B. Chadwick, Y. Eisenberg, E. Kogan, K. C. Moffeit, P. Seyboth, I. O. Skillicorn, H. Spitzer, G. Wolf, H. H. Bingham, W. B. Fretter, W. F. Podolsky, M. S. Rabin, A. H. Rosenfeld and G. Smadja, Phys. Rev. D7, 3150 (1973).

- (10) - H. J. Behrend, J. Bodenkamp, W. P. Hesse, D. C. Fries, P. Heine, H. Hirschmann, W. A. Mc Neely, A. Marcou and E. Seitz, Phys. Letters 56B, 403 (1975).
- (11) - We have taken from ref. (7) the branching ratio values  $(4.3 \pm 0.5) \times 10^{-5}$  for the  $\varrho$ ,  $(7.6 \pm 0.17) \times 10^{-5}$  for the  $\omega$ , and  $(3.2 \pm 0.2) \times 10^{-4}$  for the  $\Phi$  mesons. The energy dependence and the value of the  $\omega$  photoproduction cross section are taken from ref. (9) to be
- $$\frac{d\sigma}{dt} \Big|_{t=0} (\gamma p \rightarrow \omega p) = D \left(1 + \frac{1.4 \text{ (GeV)}}{K}\right) \quad \text{with} \quad D = 9.3 \pm 1.7 \mu b \text{ GeV}^{-2}.$$
- The energy dependence of the  $\varrho$  photoproduction cross section has been read off from the data in ref. (9). The cross section values are taken, according to the Yennie approach, from ref. (9) as
- $$\frac{d\sigma}{dt} \Big|_{t=0} (\gamma p \rightarrow p\varrho) = \frac{d^2\sigma}{dt dM} \Big|_{t=0} \frac{1}{2} \pi \Gamma_\varrho,$$
- where  $d^2\sigma/dt dM$  comes from the phenomenological Söding fit to their data using  $m_\varrho = 770$  MeV and  $\Gamma_\varrho = 150$  MeV. The energy dependence of the  $\Phi$  photoproduction cross section has been read off from the data in ref. (10). The errors attached in the text to the above quantities, added as further degrees of freedom to our fit, include errors quoted by the authors (refs. (7, 9, 10)) and normalization uncertainties in our experiment. The errors attached to  $m_\varrho$ ,  $m_\omega$ , and  $m_\Phi$  are a pessimistic estimate of systematic incertitudes in the mass reconstruction.
- (12) - F. James and M. Roos, CERN Computer Program Library (0.506).
- (13) - E. Kogan, J. Ballam, G. B. Chadwick, K. C. Moffeit, P. Seyboth, I. O. Skillicorn, H. Spitzer, G. Wolf, H. H. Bingham, W. B. Fretter, W. J. Podolsky, M. S. Rabin, A. H. Rosenfeld, G. Smadja and Y. Eisenberg, SLAC-Pub-1857, LBL 5590 (1976), Submitted to Nuclear Physics B.
- (14) - C. A. Nelson, E. N. May, J. Abramson, D. E. Andrews, J. Harvey, F. Lobkowicz, M. Singer, E. H. Thorndike, UR-599 (1975), Submitted to Phys. Rev. D.
- (15) - T. Bauer, Phys. Rev. Letters 25, 485 (1970).
- (16) - R. Marshall, Meson Resonances and Related Electromagnetic Phenomena, 101 (1971).
- (17) - H. Alvensleben, U. Becker, W. K. Bertram, M. Chen, K. J. Kohen, R. T. Edwards, T. M. Kna sel, R. Marshall, D. J. Quinn, M. Rohde, G. H. Sanders, H. Schubel and S. C. C. Ting, Nuclear Phys. 25B, 342 (1971); Nuclear Phys. 25B, 333 (1971).
- (18) - H. Alvensleben, U. Becker, W. Busza, M. Chen, K. J. Kohen, R. T. Edwards, P. M. Mantsch, R. Marshall, T. Nash, M. Rohde, H. F. Sadrozinski, G. H. Sanders, H. Schubel, S. C. C. Ting and Sau Lan Wu, Phys. Rev. Letters 27, 444 (1971).
- (19) - P. J. Biggs, D. W. Braben, R. W. Cliff, E. Gabathuler and R. E. Randt, Phys. Rev. Letters 27, 1157 (1971).