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ON THE EXTRACTION OF THE PARTICLE NEUTRON AMPLITUDE FROM PARTICLE DEUTERON SCATTERING WITH BREAK UP (I)

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ABSTRACT

In this paper we put the basis for the treatment of the hadron-deuteron reactions, characterized by the break-up of the deuteron.

To begin with we consider spinless hadrons and processes without production. We study the impulse approximation for a general form of the elementary T matrix; the phase space integration is accomplished keeping fixed the angle of the final hadron in the C.M. system of the hadron and the nucleon, which is chosen as participating to the reaction.

The impulse model is eventually checked on the momentum and angular distribution of the so called "spectator" nucleon.

1. - INTRODUCTION

In the series of papers beginning with this one, we will be concerned with the deuteron break-up reaction induced by a particle X

$$X + D \rightarrow Y + N + N$$
.

In the present paper we consider only the case in which Y belongs to the same isospin multiplet, but in the following ones we will consider the case of production, i.e. when Y is different from X, as in pion photoproduction $(\gamma D \rightarrow \pi NN)$, or when Y represents 2 or more particles, as in pion production $(KD\rightarrow K\pi NN)$. When Y is the same particle as X, with the same charge as well, the process is called "elastic incoherent"; in all the other cases, as in charge exchange or production we will refer to the process as "inelastic incoherent". The word "incoherent" refers to the contribution of this process to the differential cross section because this is obtained by summing incoherently all possible states of the 2-nucleon pair (*).

Both processes, elastic and inelastic, were explored experimentally on a large scale in the last few years (1). The main reason for doing such experiments is that they give an information on the particle-neutron system which otherwise would not be possible to obtain; the information is on the behaviour of the differential cross-section over the whole angular range, of the particle neutron elastic and inelastic scattering. It is very well known that if the energy is high enough $(P_{lab} \ge 1 \text{ GeV/c})$, the coherent elastic and inelastic process is characterized by the dominance of the double scattering for momentum transfers -t larger than .3 (GeV/c)²; therefore the information on particle neutron scattering, which can be extracted from the single scattering is limited to the range of small momentum transfers (2,3). Here, on the contrary, the shadow correction is very small even at large angles and can in principle be isolated kinematically. The problem of the shadow correction in deuteron break up reactions is very interesting and was already studied by many authors (4). However we will leave this subject to another separate report and we will concern ourselves here only with the impulse approximation on the phase space. That particular problem was considered already by Stenger in his PH.D. Thesis: his method is reported in the paper by Goldhaber et al. (5), which gives the first analysis of the process $K^+D \rightarrow K^{\circ}pp$ around 1 GeV/c. The approximation, which Stenger uses to reduce the number of integrations is not always valid, as it will be shown in the following, but the paper (5) is still a starting point for the analysis of that process. A more recent work was done by Jew in his Ph.D. Thesis: he uses a Lorentz invariant formalism to calculate the phase space volume. His computer code was used to analyze the process $\overline{K} D \rightarrow \overline{K}$ np and the result of this analysis is reported in the paper by Jew and Kalmus (6).

^(*) When the symbol Y represents more than 1 particle, the scattering angle refers to the C.M. momentum of the particles contained in Y.

Beyond this, we will consider the distribution on the angle and the modulus of the momentum of the "spectator". In connection with this, we quote the work by Butenschoen(), who considered the process $\gamma D \rightarrow \pi^- pp$ in the energy range between .2 and 2 GeV: he looked at the problem of the asymmetry in the "spectator" angle, which is always present in this type of reaction and he explained it as a flux factor effect. The asymmetry was explained in the same way by Hirata (8), who calls this phenomenon Doppler effect: he gives numbers for the excess of high momentum "spectators", but he doesn't give a satisfactory explanation for it. The most recent work on the subject is due to Dean (9), who considered the symmetrization effect on the distribution of the tranverse part of the "spectator" monumentum.

In all these papers, as in this one, the main purpose is to establish a connection between the experimental data and the two body parameters for the hadron-neutron system.

In this paper the invariant formalism (0,1) is introduced from the very beginning (Sec. 2), giving the set of Feynman rules, known from quantum electrodynamics (1^2) and two additional rules, which give the expression of the deuteron vertex and the scattering boxes. This formalism is elegant and gives the possibility of evaluating relativistic effects. It is besides shown in the literature $(1^{\circ}, 1^{\circ})$ that this formalism contains Glauber theory as its limit at small momentum transfers and high energy. The case of a spinless particle impinging on a deuteron is considered afterwards (Sec. 3) and a formula is given for the differential cross-section for the deuteron break-up. In the same section we give the definition of the weight factors and we compare their expression with the results of the previous literature.

In the next section (Sec. 4) we discuss about the general behaviour of the crosssection in terms of the momentum of one nucleon: then we define the "spectator" and we examine the distribution on the modulus and the angle of its momentum.

Finally (Sec. 5) we consider the behaviour of the weight factors as function of the scattering angle, with special reference to the comparison between the exact calculation and the Stenger (5) low energy approximation.

In the appendix A we give a practical method to determine the Feynman rule for the deuteron vertex.

In the appendix B we write down in full detail the method to determine the kinematics of the 3- Body final state if the scattering angle in the particle nucleon C.M. system, and the tri-momentum of one of the two nucleons in the laboratory system are fixed.

2. - INVARIANT FORMALISM

Before starting the actual treatment we give here the rules of the game, which are very well known, but we find worthwhile to state at the beginning to make this report as

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much as possible self consistent. We have to tell that the particles or lines, we consider are all spinless: this makes life easier, concerning the definition of the flux factor and the normalization factors $(^{3})$. However, this assumption doesn't forbid to include spin inside the scattering boxes and in the deuteron vertex. We will not give here the correspondence for the external lines, since in the above conventions it amounts only to a normalization factor, which appears in the definition of the S matrix, but disappears in the T matrix.

Choosing the metric (1, -1, -1, -1), the rules are



The definition of the deuteron vertex in terms of the non-relativistic wave function of the deuteron ψ is very simple, as shown above, if the momentum of the deuteron is small enough that we can neglect relativistic effects. This definition can be obtained from the general form of the deuteron vertex in the non relativistic limit. To the interested reader, we give the reference for the general invariant expression of the deuteron vertex, in terms of the non relativistic wave functions: this is a paper by Gross (⁴⁴); in this paper the non-relativistic limit is actually done and the result is identical to ours apart from trivial normalization errors. The application of these rules to any diagram defines the corresponding S matrix, which is connected by the following standard relation to the T matrix

(2.1)
$$S_{fi} = \delta_{fi}^{+i}(2\pi)^{4}\delta(P_{i}^{-}P_{f}^{-}) \frac{T_{fi}}{\prod_{j=1}^{l} N_{j} \prod_{k=3}^{l} N_{k}}$$

This relation contains the normalization factor for the 2 particles in the initial state and the n-2 particles in the final state: in the above simplifying convention N_i is the normalization factor used for the boson in the literature

(2.2)
$$N_{i} = \sqrt{(2\pi)^{3} 2E_{i}}$$

The calculation of such factors from the Feynman diagram are not necessary, since we shall, eventually take them out to extract the T matrix.

The normalization of the T matrix given by the relation (2.1) is such that the differential cross section, is defined using

(2.3)
$$d\sigma = (2\pi)^4 \delta (P_1 - P_f) \frac{\overline{z} \ \underline{z} \ |T_{fi}|^2}{\Phi} \frac{d^3 p_3 \ \dots \ d^3 p_n}{k=3} N_k^2$$

where Φ is the flux factor, defined by the Lorentz invariant

$$(2.4) \qquad \Phi = 4 \sqrt{(p_1 p_2)^2 - (m_1 m_2)^2}$$

The flux factor is connected with the relative velocity between the two particles in the initial state by the following relation

$$\Phi = 4 E_1 E_2 v$$

The flux factor assumes the simple expressions in the rest system of particle 1 and the total C.M. system

(2.5)
$$\Phi = 4 \text{ m}_1 p_2 = 4 \text{ k} \sqrt{s}$$

where k is the C.M. momentum.

The above definition of the differential cross section (2.3) is consistent with the following normalization of the amplitude in the forward direction, given by the optical theorem

(2.6)
$$\operatorname{Im} \mathfrak{T} (0,s) = \frac{\Phi}{2} \sigma_{\mathfrak{T}}$$

We will use in the foregoing section the following notation for eq. (2.3)

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(2.7)
$$d\sigma = \overline{\Sigma} \Sigma |T_{fi}|^2 \frac{dV^{(n-2)}}{\Phi}$$

where $dV^{(n-2)}$ represents the invariant volume element of the phase space of the n-2 final particles. From the above formula it is clear that $\bar{\xi} \in |T_{fi}|^2$ is an invariant too.

We will consider now the problem of two particles, one with spin 0 and the other with spin 1/2 (meson-nucleon), which scatter elastically: this example, not only clarifies the last few points of the discussion, but provides the background for the foregoing section, where we limit ourselves to this case or to the isospin flip channel. This problem as an application of the invariant formalism is really trivial, but since we know already the answer, we learn something about the method.

The diagram is shown in Fig. 1. We indicate the meson with a dashed line and the nucleon with a continuous one. The T matrix element is

(2.8)
$$\mathbb{T}_{\mathbf{f}_{i}} = 2\mathfrak{m} \, \tilde{u}_{\mathbf{f}} \, (\mathbf{p}') \, (A+B_{\mathbf{f}} Q) \, u_{i}(\mathbf{p})$$

where $Q = \frac{q+q!}{2}$

The Dirac spinors are defined

(2.9)
$$u_{i}(p) = \frac{m + \gamma p}{\sqrt{2m(m + E)}} \qquad \begin{pmatrix} |i\rangle \\ 0 \end{pmatrix}$$

(2.10)
$$\overline{u}_{f}(p) = u_{f}^{+}(p) \gamma_{0}$$

where

(2.11)
$$\gamma p = \gamma_0 p_0 - \overline{\gamma} \cdot \overline{p}$$

$$\vec{\Upsilon} = \begin{pmatrix} 0 & \vec{\sigma} \\ -\vec{\sigma} & 0 \end{pmatrix} \qquad \qquad \Upsilon_0 = \begin{pmatrix} 1 & 0 \\ 0 & -1 \end{pmatrix}$$

The matrix element T_{fi} is written in the four dimensional spin space, but since we don't consider antiparticles, we can easily think it in two dimensional space. The technique is to perform all the products of the two dimensional symbolic matrices and take only the first diagonal element (³):

$$T_{fi} = m < f | (m+E)^{\frac{1}{2}} (m+E')^{\frac{1}{2}} [A + \frac{B}{2} (\omega + \omega')] - \frac{1}{(m+E)^{\frac{1}{2}} (m+E')^{\frac{1}{2}}} \left\{ [A - \frac{B}{2} (\omega + \omega')] (P^2 - \Delta^2/4) + B(2m+E+E') \overline{P} \cdot \overline{Q} + \frac{B}{2} (E - E') \overline{Q} \cdot \overline{\Delta} + \frac{1}{2} (M + \omega')] \overline{\sigma} \cdot (\overline{\Delta} \times \overline{P}) + i B(E' - E) \overline{\sigma} \cdot (\overline{Q} \times \overline{P}) + \frac{1}{2} \frac{B}{2} (2m+E+E') \overline{\sigma} (\overline{\Delta} \times \overline{Q}) \right\} | i > = < f | \hat{T} | i > 1$$

where ω , ω' are the initial and final energies of the meson and E, E' are the initial and final energies of the nucleon

$$\vec{\hat{Q}} = \frac{\vec{\hat{q}} + \vec{\hat{q}'}}{2}$$

$$\vec{\hat{P}} = \frac{\vec{\hat{p}} + \vec{\hat{p}'}}{2}$$

$$\vec{\hat{\Delta}} = \vec{\hat{q}} - \vec{\hat{q}'} = \vec{\hat{p}'} - \vec{\hat{p}}$$

This expression is very clumsy, but it becomes very simple for the laboratory system and the C.M. system, where the number of the spin flip amplitudes reduce to one. In the C.M., we obtain with the constraint of $\vec{p} + \vec{q} = \vec{p}' + \vec{q}' = 0$ that

(2.13)
$$\hat{\mathbb{T}} = \alpha_{f} + i \alpha_{\sigma} \vec{\sigma} \cdot \hat{n}$$

where

$$\hat{n} = \frac{\vec{q} \times \vec{q}'}{|\vec{q} \times \vec{q}'|}$$

and

$$\begin{aligned} \alpha_{f} &= m \left\{ (E+m) [A + (W-m)B] + (E-m) [A + (W+m)B] \cos \vartheta^{*} \right\} \\ \alpha_{g} &= -m(E-m) [-A_{+}(W+m)B] \sin \vartheta^{*} \end{aligned}$$

 $W = \sqrt{s} = E + \omega$, E, ω are the energies of the nucleon and the meson in the C.M. system.

These linear relations can be inverted and A and B can be expressed in terms of the C.M. amplitudes

$$A = \frac{1}{2W} \left[\frac{W \pm m}{E + m} \alpha_{f} + \left(\frac{W \pm m}{E + m} \cos \vartheta^{*} \pm \frac{W - m}{E - m} \right) \frac{\alpha_{g}}{\sin \vartheta^{*}} \right]$$

$$B = \frac{1}{2W} \left[\frac{1}{E \pm m} \alpha_{f} \pm \left(\frac{\cos \vartheta^{*}}{E \pm m} - \frac{1}{E - m} \right) \frac{\alpha_{g}}{\sin \vartheta^{*}} \right]$$

The same reduction is done in the laboratory system and we find

(2.15)
$$\hat{\mathbb{T}} = a_{\hat{\mathbb{f}}} + i a_{g} \vec{\sigma} \cdot \hat{k}$$
 $\hat{k} = \hat{n}$

where

$$a_{f} = 2 m^{2} \sqrt{1-t/4m^{2}} [A + \frac{B}{2} (\epsilon + \epsilon')] - \frac{m}{2} \frac{q^{2}-q'^{2}}{\sqrt{1-t/4m^{2}}} B$$

(2.16)

$$a_g = - \frac{m}{\sqrt{1 - t/4m^2}} B \left| \overline{\Delta} \times \overline{q} \right|$$

where

 ϵ , ϵ' are the initial and final energies of the meson in the laboratory system. If we substitute here to A, B the expressions (2.14) in terms of $\alpha_{\rm f}$, $\alpha_{\rm g}$, we obtain a linear relation between the lab amplitudes and the C.M. amplitudes, which defines the Lorentz transformation

(2.17)
$$a_{f} = C_{ff} \alpha_{f} + C_{fg} \alpha_{g}$$
$$a_{g} = C_{gf} \alpha_{f} + C_{gg} \alpha_{g}$$

where

$$C_{\text{ff}} = \frac{1}{W} \text{ m } \sqrt{1 - t/4m^2} \left(\frac{W + m}{E + m} + \frac{1}{2} - \frac{\varepsilon + \varepsilon}{E + m} \right) - \frac{1}{4W} - \frac{q^2 - q^{\prime 2}}{\sqrt{1 - t/4m^2}} - \frac{1}{E + m}$$

$$C_{fg} = \left\{ \frac{1}{W} \text{ m } \sqrt{1 - t/4m^2} \left[\left(\frac{W+m}{E+m} + \frac{1}{2} \quad \frac{\epsilon + \epsilon'}{E+m} \right) \cos \varphi^* + \left(\frac{W-m}{E-m} - \frac{1}{2} \quad \frac{\epsilon + \epsilon'}{E-m'} \right) \right] - \frac{1}{4W} \frac{q^2 - q'^2}{\sqrt{1 - t/4m^2}} \left(\frac{\cos \varphi^*}{E+m} - \frac{1}{E-m} \right) \right\} \frac{1}{\sin \varphi^*}$$
(2.18)

$$C_{gf} = -\frac{1}{2W} \frac{1}{\sqrt{1-t/4m^2}} \left| \overline{\Delta}_{\times \overline{q}} \right| \frac{1}{E+m}$$

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$$C_{gg} = -\frac{1}{2W} \frac{1}{\sqrt{1-t/4m^2}} \left| \bar{\Delta}_{\times} \bar{q} \right| \left(\frac{\cos \vartheta^*}{E+m} - \frac{1}{E-m} \right) \frac{1}{\sin \vartheta^*}$$

The intuitive reason, why in general the spin amplitudes mix in the transformation from the C.M. to the laboratory system, is that the spin doesn't transform as a vector in a Lorentz transformation, but as a second rank tensor like the electromagnetic field.

The distintive feature of this treatment is the relativistic invariance of $\sum_{fi} |T_{fi}|^{\frac{1}{2}}$: this property can be verifield directly starting from (2.8) and calculating the tr(T T⁺), by means of the projector operator technique (¹²). This invariance fixes the properties of the transformation of the amplitudes: that is

$$|C_{ff}|^{2} + |C_{gf}|^{2} = 1$$
(2.19)
$$|C_{fg}|^{2} + |C_{gg}|^{2} = 1$$

$$C_{ff} C_{fg} + C_{gf} C_{gg} = 0$$

which are the conditions for the invariance property

(2.20)
$$|\alpha_{f}|^{2} + |\alpha_{g}|^{2} = |a_{f}|^{2} + |a_{g}|^{2}$$

This can be written in terms of the conventional spin amplitudes in the C.M. System, as

(2.21)
$$64 \pi^2 s (|f|^2 + |g|^2) = |a_f|^2 + |a_g|^2$$

The factor 64 π^2 s is the simple expression of $\frac{d\Omega^*}{dV^{(2)}} \Phi^{(2)}$, where the index $\binom{2}{m}$ means that we are considering a two body process.

The relations (2.19) are direct consequences of the invariance property (2.20): however they can be checked directly, for instance in the forward direction. If we use the relations

$$E \pm m = \frac{(W \pm m)^2 - \mu^2}{2W}$$

we realize very soon from (2.16) that for $\vartheta^* = 0$, $C_{ff} = 1$ and $C_{fg} = 0$. For small angles $\Delta \sim \sqrt{-t} \sim q \, \vartheta^*$, then $C_{gg} = 1$ and $C_{gf} = 0$.

3. - WEIGHT FACTORS

We pursue here the aim to give a formula, which connects as directly as possible the data of deuteron break up reaction, induced by a spinless particle and the two body amplitudes for the scattering of this particle on nucleon. This connection is established through the "weight factors" for the spin matrix elements or even more directly for the square modulus of the spin amplitudes in the C.M. system, which take in account the complex structure of the deuteron. These factors, which should be constant for the scattering on a pure neutron target, weight differently in different angular regions the various spin matrix elements.

For the moment this calculation is performed in the approximation that in the time lapse, needed to cross the target region, only one of the two nucleon interact with the incoming object, while the other one is unperturbed in the process (impulse approximation). This approximation is valid if the De Broglie wave length of the incoming object is smaller than the average distance between the two nucleons. At low incident momentum we have to consider three-body effects, but for about 600 MeV/c the wave length of the incident particle is about .3 F which is small compared with the m.r. of the deuteron 2.3 F, and therefore the interaction with the nucleons will be completely independent. We will neglect for the moment corrections due to successive scattering of the fast incident particle with the two nucleons and to rescattering of the nucleons in the final state, which are delayed to the already announced report.

The calculation is done in the rest system of the deuteron: this system has the advantage of allowing the straightforward use of the nonrelativistic wave function of the deuteron, but the drawback of requiring a Lorentz transformation of the amplitude. On the other hand, if we work in the C.M. frame of the particle nucleon system, the complication of the Lorentz transformation are shifted from the amplitude to the deuteron vertex.

The impulse approximation is equivalent to assume that the Feynman graphs of Fig. 2 describe the process. The application of the rules of the previous section gives for the T matrix the following expression in the rest system of the deuteron:

(3.1)
$$T = (16 \pi^{3} m_{1})^{\prime 2} [T_{4}(s,t;m^{*2}) \psi(\overrightarrow{-p_{5}}) + T_{5}(s';t;m^{*2}) \psi(\overrightarrow{-p_{4}})]$$

The dependence of the amplitude on the virtual mass of the nucleon is formal: in practice we will neglect this dependence and therefore we will go on thinking it as onmass-shell amplitude. While the definition of the momentum transfer is unambiguous, there is a double possibility for the energy

(3.2.a)

$$s = (p_3+p_4)^2 = (E_2+m_1-E_5)^2 - (\vec{p}_2 - \vec{p}_5)^2$$

 $t = (p_2-p_3)^2$
 $s = (E_2+E_5)^2 - (\vec{p}_2 - \vec{p}_5)^2$

and analogously for the second diagram. For the first definition of energy, the interacting nucleon is thought off-mass-shell, as it should be if we want to conserve the energy in the vertex; in the second case it has been forced to have the real mass: the reason is that, as stated before, in the analysis the amplitude is thought on the mass shell and therefore we might prefer to define the energy using the real mass of the nucleon. As we shall see in the following this ambiguity is not important, if the amplitude is slowly energy dependent, but becomes important in the region of resonances. As a first approximation we will take the second definition and fix $\vec{p}_5 = 0$. This approximation is based on the observation that the wave function is very peaked around small values of its argument and the amplitude is usually slowly varying with energy. The same assumption will be done concerning the spin structure: in this way we will avoid the complication due to the moving target (see formula 2.12). Therefore the invariant amplitude is expressed in the two dimensional spin space

(3.3)
$$T_{i}(s,t;m^{*2}) = a_{p} + b_{p} \bar{\tau} \cdot \bar{\tau}_{i} + i(a_{p} + b_{p} \bar{\tau} \cdot \bar{\tau}_{i}) \bar{\sigma}_{i} \cdot \hat{n}$$

where $\vec{\tau}$ is the isotopic spin operator of the scattered particle

 $\vec{\tau}_i$ is the isotopic spin operator of the i-nucleon

 $\vec{\sigma}_{i}$ is the spin operator of the i-nucleon

n is the versor orthogonal to the scattering plane.

The spin amplitudes a_f , a_g are defined in the previous section in terms of the invariant amplitudes and the C.M. amplitudes. The same can be done for the b_f , b_g , i.e. the i-spin flip amplitudes. We are not performing this transformation now, but we will do it at the level of the differential cross section, were we can use invariance properties.

Before starting to use the above amplitudes, we want to underline that no assumptions at all are done on the isospin of the incident particle, nor on the normalization of the spin and isospin operators. So, although the following treatment is used only for isospin 1/2 particles like K, \overline{K} , we could check the procedure for the very well known isospin 1 objects, the π -mesons.

Let us now calculate the matrix elements of the above operator on the spin and isospin state of the two nucleons and the isospin state of the particle, which are necessary ingredients for the calculation of the cross section. The isospin third component of the particle is determined by the total isospin conservation and it is not written in the following. Since the deuteron is broken in the final state, we have to consider triplet states as well as singlet states for the nucleon pair. We choose the polarization axis along the direction of \hat{n} .

The matrix elements are listed here with the convention that the first two numbers in the ket represent the isospin and the isospin third component of the nucleon pair and the second two, the spin and its third component.

They are:

a) for the singlet state of the nucleon pair

(3.4)
(3.4)
(3.5)
(3.5)
(3.5)
(3.6)
$$|T| = \delta_{\nu,0} i b_{g} < N N | \bar{\tau}.\bar{\tau}_{4}| D > \times K [\psi(-\vec{p}_{5}) + \psi(-\vec{p}_{4})]$$

b) for the triplet

(3.6)

$$<0 \ 0; \ 1 \ \nu' \ | \mathbb{T} | \ 0 \ 0; \ 1 \ \nu > = < \nu' | a_{\mathbf{f}} + \mathbf{i} \ \sigma_{4Z} \ a_{\mathbf{g}} | \nu > \times \\ \times \ K \ [\psi(\overrightarrow{p_5}) + \psi(\overrightarrow{p_4})]$$

$$(3.7) \qquad <1 \ \mathbb{T}_{z}; \ 1 \ \nu' \ |\mathbb{T}| \ 0 \ 0; \ 1 \ \nu > = <\mathbb{NN} |\vec{\tau}.\vec{\tau}_{4}| \mathbb{D} > \times \\ \times <\nu' | \mathbb{b}_{\mathbf{f}^{+}} \ \mathbf{i} \ \sigma_{4z} \ \mathbb{b}_{g} | \nu' > \times \\ \times \ \mathbb{K} [\psi(-\vec{p}_{5}) - \psi(-\vec{p}_{4})] \end{cases}$$

where $K = (16 \pi^2 m_1)^{\frac{1}{2}}$ and the isospin reduced matrix element on the right hand side $\langle NN | \bar{\tau}.\bar{\tau}_4 | D_2 = \langle 1 T_z; \tau \tau'_z | \bar{\tau}.\bar{\tau}_4 | 0,0; \tau \tau_z \rangle$; in the latter τ_z, τ'_z represent the isospin third component in the initial and final state of the meson. Here we have neglected the D-state of the deuteron; but the above formula can easily be generalized to include it.

The differential cross section for the three body process as function of the angle of particle 3 in the center of mass of the particles 3 and 4, or 3 and 5 is (*)

(3.8)
$$\frac{\mathrm{d}\sigma}{\mathrm{d}\Omega^*} = \frac{1}{\Phi} \sum_{\mathrm{f}} \overline{\Sigma} |\mathrm{T}_{\mathrm{f}}|^2 \frac{\mathrm{d}\mathrm{V}^{(3)}}{\mathrm{d}\Omega^*}$$

or if we assume

(3.9)
$$\mathbb{T}_{\text{fi}} = \widetilde{\mathbb{T}}_{\text{fi}} \times \left[\psi(-\vec{p}_5) \pm \psi(-\vec{p}_4)\right]$$

we obtain, integrating on all variables except Ω^*

(3.10) $\frac{\mathrm{d}\sigma}{\mathrm{d}\Omega^*} = \sum_{\mathrm{f}} \sum_{\mathrm{i}} |\widetilde{\mathrm{T}}_{\mathrm{fi}}|^2 \omega^{\pm}$

(*) We call this angle Ω^* , from now onwards

where

$$\omega^{\pm} = \frac{K^2}{\Phi} \int \left| \psi(-\mathbf{p}_5) \pm \psi(-\mathbf{p}_4) \right|^2 \frac{\mathrm{d}V^{(3)}}{\mathrm{d}\Omega^*}$$

The ambiguity in the definition of the system where Ω_3^* is defined is easily found to be only apparent: the weight factor ω^{\pm} can be expressed in the following way, using the symmetry of the invariant phase space volume for the exchange of p_4 with p_5

(3.11)
$$\omega^{\pm} = \frac{2K^2}{\Phi} \left[\int |\psi(\mathbf{p}_5)|^2 \frac{dV^{(3)}}{d\Omega^*} \pm \int \psi(\mathbf{p}_4)\psi(\mathbf{p}_5) \frac{dV^{(3)}}{d\Omega^*} \right]$$

The second term has an integrand function, still symmetric for the interchange of p_4 with p_5 : therefore the ambiguity is completely eliminated. We can choose for instance the C.M.S. of the particles 3 and 4 to calculate Ω^* . To complete our task, to express the differential cross-section (3.10) in terms of the C.M. amplitudes, we perform the Lorentz transformation (2.17) of the amplitudes. For this purpose we calculate separately the charge exchange and the quasi-elastic scattering.

$$(3.12) \qquad \frac{d\sigma}{d\Omega^{*}} \Big|_{CEX} = |\langle NN|\bar{\tau}.\bar{\tau}_{4}|D\rangle|^{2} \left[\frac{|b_{g}|^{2}}{3}\omega^{+} + (|b_{f}|^{2} + \frac{2}{3}|b_{g}|^{2})\omega^{-}\right]$$

$$(3.13) \qquad \frac{d\sigma}{d\Omega^{*}} \Big|_{QE} = \frac{|a_{E}|^{2}}{3}\omega^{-} + (|a_{f}|^{2} + \frac{2}{3}|a_{g}|^{2})\omega^{+} + |\langle NN|\bar{\tau}.\bar{\tau}_{4}|D\rangle|^{2} \times \left[\frac{|b_{E}|^{2}}{3}\omega^{+} + (|b_{f}|^{2} + \frac{2}{3}|b_{g}|^{2})\omega^{-}\right]$$

We will go through all steps, only for the charge exchange cross section, while for the quasi elastic scattering we give later on, the final result, in order to avoid useless repetitions. We substitute now, to the laboratory amplitudes, the expression in terms of the C.M. amplitudes, taking in account the invariance property (2.20)

$$(3.14) \qquad \frac{\mathrm{d}\sigma}{\mathrm{d}\Omega^*}\Big|_{\mathrm{CEX}} = |\langle \mathrm{NN}|\bar{\tau}.\bar{\tau}_4|\mathrm{D}\rangle|^2 \left[\frac{|\beta_{\mathrm{g}}|^2}{3}\omega^+ + (|\beta_{\mathrm{f}}|^2 + \frac{2}{3}|\beta_{\mathrm{g}}|^2)\omega^- + C\right]$$

where

(3.15)
$$C = \frac{|\beta_{f}|^{2} - |b_{f}|^{2}}{3} (\omega^{+} - \omega^{-})$$

C is the relativistic correction to the traditional formula (⁴). This correction seems to be important only near the forward direction, because for high momentum transfer is negligible since $\omega^+ \sim \omega^- \sim 2 \frac{K^2}{\Phi}$. But around the forward direction as from (2.16), $\beta_{\varphi} \sim b_{\varphi}$, therefore the correction is very small and will be neglected for the moment.

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The last step is to change the normalization of the C.M. amplitudes, such as to connect the usual formalism of the phase shift analysis

$$(3.16) \qquad \frac{\mathrm{d}\sigma}{\mathrm{d}\Omega^{*}} \Big|_{\mathrm{CEX}} = \frac{|\langle \mathrm{NN}|\bar{\tau}.\bar{\tau}_{4}|\mathrm{D}\rangle|^{2}}{|\langle \mathrm{N}|\bar{\tau}.\bar{\tau}_{4}|\mathrm{N}\rangle|^{2}} \times \left[\frac{|g^{\mathrm{CEX}}|^{2}}{3} \mathrm{W}^{+} \left(|f^{\mathrm{CEX}}|^{2} + \frac{2}{3}|g^{\mathrm{CEX}}|^{2}\right) \mathrm{W}^{-}\right]$$
$$\frac{\mathrm{d}\sigma}{\mathrm{d}\Omega^{*}} \Big|_{\mathrm{CEX}} = \frac{1}{2} \left[\frac{|g^{\mathrm{CEX}}|^{2}}{3} \mathrm{W}^{+} + \left(|f^{\mathrm{CEX}}|^{2} + \frac{2}{3}|g^{\mathrm{CEX}}|^{2}\right) \mathrm{W}^{-}\right]$$

where

 $f^{CEX} = \frac{1}{2} [f_1 - f_2] \qquad \text{for i-spin } 1/2$ $\frac{\sqrt{2}}{3}$ [f₃-f₁] for i-spin 1

and

$$f^{CEX} = \frac{\sqrt{2}}{3}$$
 [

and analogously for the spin flip amplitude.

$$(3.17) \qquad \qquad \mathbb{W}^{\pm} = \frac{\mathrm{d}\Omega^{\ast}}{\mathrm{d}V}(z) \quad \Phi^{(z)} \omega^{\pm}$$

The factor 1/2 is the squared ratio between the isospin matrix element for the deuterium and for the nucleon; that is

$$\langle PP | \vec{\tau} \cdot \vec{\tau}_4 | D \rangle = \frac{1}{\sqrt{2}} \langle P | \vec{\tau} \cdot \vec{\tau}_4 | N \rangle$$

where

$$\langle \mathbb{P} | \vec{\tau} . \vec{\tau}_4 | \mathbb{N} \rangle = \langle \tau \tau'_{z} ; 1/2 \ 1/2 \ | \vec{\tau} . \vec{\tau}_4 | \tau \tau_{z} ; \ 1/2 \ 1/2 \rangle$$

the above relation in general and holds for K, $\bar{\rm K}$ and π .

The relation between the conventional amplitudes f^{CEX} , g^{CEX} and the amplitudes $\beta_{\rm f}$, $\beta_{\rm g}$

$$\beta_{\rm f} < \mathbb{P} \left| \, \overline{\tau}_1 \, \overline{\tau}_4 \, \right| \, \mathbb{N} > \sqrt{\frac{{\rm dV}^{\left(2\right)}}{{\rm d}\Omega^{\ast}}} \quad \frac{1}{\Phi^{\left(2\right)}} \, = \, f^{\rm CEX}$$

Going through the same steps, we find for the charge preserving scattering

(3.18)
$$\frac{d\sigma}{d\Omega^{*}} = \frac{|g^{+}|^{2}}{3} W^{-}_{+} \left(|f^{+}|^{2}_{+} \frac{2}{3} |g^{+}|^{2} \right) W^{+}_{+} + \frac{|g^{-}|^{2}}{3} W^{+}_{+} \left(|f^{+}|^{2}_{+} \frac{2}{3} |g^{-}|^{2} \right) W^{-}_{-}$$

where for isospin 1/2

$$e^{+} = \frac{1}{4} \left(3 f_{1} + f_{0} \right)$$

$$e^{-} = \frac{1}{4} \left(f_{1} - f_{0} \right)$$

$$e^{+} = \frac{1}{4} \left(2f_{3} + f_{1} \right)$$

and for isospin 1

$$f^{+} = \frac{1}{3} (2f_{3} + f_{1})$$
$$f^{-} = \frac{1}{3} (f_{3} - f_{1})$$

The connection with the α and β amplitudes is

$$\begin{aligned} \mathbf{f}^{+} &= \sqrt{\frac{\mathrm{d} \mathbf{V}^{(2)}}{\mathrm{d} \Omega^{*}}} \quad \frac{1}{\Phi^{(2)}} \quad \alpha_{\mathbf{f}} \\ \mathbf{f}^{-} &= \sqrt{\frac{\mathrm{d} \mathbf{V}^{(2)}}{\mathrm{d} \Omega^{*}}} \quad \frac{1}{\Phi^{(2)}} \quad \beta_{\mathbf{f}} < \mathbf{N} | \, \vec{\tau}_{1} \, \vec{\tau}_{4} \, | \, \mathbf{N} > \end{aligned}$$

In the charge preserving differential cross section, there are no factors coming from the isospin matrix elements because the matrix elements for the bound nucleon are equal to the ones for free nucleon

$$\langle \mathrm{NP} | \overline{\tau}, \overline{\tau}_{4} | \mathrm{D} \rangle = \frac{1}{2} \left(\langle \mathrm{P} | \overline{\tau}, \overline{\tau}_{4} | \mathrm{P} \rangle - \langle \mathrm{N} | \overline{\tau}, \overline{\tau}_{4} | \mathrm{N} \rangle \right) = \langle \mathrm{P} | \overline{\tau}, \overline{\tau}_{4} | \mathrm{P} \rangle$$

The W^{\pm} are the weight factors calculated by Stenger et al. (⁵). He uses certain approximations to calculate the kinematics and the factor $\frac{dV(3)}{d\Omega^*}$, which will be examined in detail later on. The result of Glauber and Franco (⁴) is found by noticing that the three body invariant phase space volume element contains the two body phase space volume element

(3.21)
$$dV^{(3)} = dV^{(2)} \frac{d^3 p_5}{(2\pi)^3 2E_5}$$

This volume element is calculated with the exact kinematics with $p_5 \neq 0$: however it is easy to convince oneself that it is a slowly varying function of p_5 , compared with the wave function, and can be calculated for $p_5 = 0$, taken out of the integration symbol and cancelled (*): then

(3.22)
$$\mathbb{W}^{\pm} = \frac{\Phi^{(2)}}{\Phi} \frac{2K^2}{2(2\pi)^3} \left[\int \psi^2(\mathbf{p}_5) \frac{d^3 \mathbf{p}_5}{\mathbf{E}_5} \pm \int \psi(\mathbf{p}_4) \psi(\mathbf{p}_5) \frac{d^3 \mathbf{p}_5}{\mathbf{E}_5} \right]$$

(*) This implies that the definition of Ω^* is given in the situation $p_5 = 0$: as shown by Stenger (⁵) this makes a very small difference.

As

$$\Phi^{(2)} = 4 p_2 m_5$$

$$\Phi = 4 p_2 m_1$$

$$E_5 \sim m_5$$

 $(3.23) \qquad \mathbb{W}^{\pm} \sim 2 \left[\int \psi^2(\mathbf{p}_5) d^3\mathbf{p}_5 \pm \int \psi(\mathbf{p}_4) \psi(\mathbf{p}_5) d^3\mathbf{p}_5 \right] = 2 \left[1 \pm S(\Delta) \right] \quad (\text{closure})$

if

where S is the deuteron form factor

 $\vec{\Delta} = \vec{p}_2 - \vec{p}_3$

$$s = \int \psi (\vec{p}_5) \psi(\vec{p}_5 - \vec{\Delta}) d^3 p_5 = \int e^{i\vec{\Delta} \cdot \vec{\hat{r}}} \psi(\vec{\hat{r}}) d^3 r$$

The result of Jew $(^6)$ is obtained taking into account the Fermi motion in the factor $\frac{d\Omega^*}{dV(\frac{2}{2})} \,\,\Phi(^2)$, that is considering this kinematical factor for an inelastic scattering of the incident particle on a light nucleon $(m^{*2} \cong m^2 - 2 \,\,p_5^2)$.

In the interference term there are choice problems: we assume that the scattering occurs on the particle 4 like in the first diagram of Fig. 1; this is not a bad approximation, as the overlapping between the two diagrams is important when p_4 and p_5 have similar values. In this way the cancellation occurs inside the integral and we obtain

(3.24)
$$\mathbb{W}^{\pm} = 2 \operatorname{m} \left[\int \psi^{2}(\mathbf{p}_{5}) \frac{\Phi^{(2)}}{\Phi} \frac{d^{3}\mathbf{p}_{5}}{\mathbf{E}_{5}} \pm \int \psi(\mathbf{p}_{4})\psi(\mathbf{p}_{5}) \frac{\Phi^{(2)}}{\Phi} \frac{d^{3}\mathbf{p}_{5}}{\mathbf{E}_{5}} \right]$$

Here the first definition of the energy (3.2a) was implicitly chosen and the flux factor is

(3.25)
$$\Phi^{\binom{2}{2}} = 4 \sqrt{(p_2 p^*)^2 - (m_2 m^*)^2} = 2 \sqrt{p_2^2 (m^{*2} + p_5^2) + p_2^2 p_5^2} + 2E_2 \sqrt{m^{*2} + p_5^2 p_2 p_5 z_5}$$

As seen from the above formula the factor is strongly dependent on the angle of p_5 with the beam direction. This variation is actually seen in the experimental data (⁷), as it will be discussed in the next section. Since only this definition of the weight factors has this asymmetry built in, we believe that this is a good reason to prefer this last one respect to others.

The same procedure can be followed in the C.M. system of the particle 3 and 4: if we neglect the relativistic correction of the deuteron, or in other words we mantain the rule for the deuteron vertex, given in Sec. 2, we arrive to the same formula for the differential cross-section (3.14), with the difference that C is rigorously zero. However the main difference is that we don't need to make the approximation that nucleon 5 is at rest in the laboratory system, to simplify the spin structure of the amplitude: this is very simple and uniquely defined in the C.M. system.

4. - NUCLEON MOMENTUM DISTRIBUTIONS

Before actually doing the integration on the phase space and to discuss the behaviour of the weight factors with the scattering angle, we study here the features of the integrand function

(4.1)
$$E^{\pm} = |\psi(p_4) \pm \psi(p_5)|^2$$

that we call extraction factor. We will think for the moment that particles 4 and 5 are physically distinguishable, as it actually occurs for a charge preserving process, where we have a neutron and a proton in the final state. In this case we can keep the form (4.1) for the extraction factor and study its behaviour, as function of $p_5 \equiv (p_5, \vartheta_5, \varphi_5)$: Fig. 3 and 4 represent the perspective of the function $E^+ \cdot dV^{(3)}/d\Omega_3 dp_5 d\Omega_5$ for $\varphi_5 = \pi$, $\cos \vartheta_3^* = 0$ and $\cos \vartheta_3^* = .9$ in the process $K^+D - K^\circ pp$ at .6 GeV/c. In both drawings we see a moraine and a peak: the first is the region where the particle 5 has small momentum and therefore is called spectator, the second is characterized by large momenta of particle 5, which is recoiling after having interacted with the incident particle. While the first is characterized by a negligible dependence on ϑ_5 , the second is highly dependent on the angle. The top of the peak (socalled "quasi-elastic peak") is moving with the scattering angle and its coordinates p_5 , ϑ_5 represents the recoil momentum of the corresponding two body process: in other words, with a hydrogen target we would obtain a peak, much thinner, exactly in the same position.

It is seen from the drawing at large angles that there is no overlapping between the two bumps: in experimental terms at large angles, we know that the slower nucleon is the spectator. However at small angles there is overlapping or, in other words, the wave character of the phenomenon shows off: therefore it is not possible to make the choice of the spectator, in the same way as in the two slits experiment it is not possible to determine the slit crossed by the particle. The range of angles, where this difficulty occurs, is obviously energy dependent: it tends to become more and more limited with raising energy.

One can ask oneself: why the choice of the spectator is necessary? The answer is that we want to determine the energy and the C.M. angle of the two body process for any event. This can be done, only if we assume a spectator model and know the momentum of the spectator. In the forward region, it is therefore not possible to define the kinematics of the two body process, and the analysis has to be done in terms of the external variables, as for instance the invariant momentum transfer.

However, if we are interested in the whole angular distribution, we can forget about this problem and define the "spectator": the procedure is to give a definition for it and to check afterwards its momentum and angle distribution. We define it as the nucleon which has lower energy. If we consider the distribution of events on the momentum of the nucleon satisfying this condition, as shown of Fig. 5 for the case $K^{+}D--K^{0}pp$ at .98 GeV/c, we realize that up to 250 MeV/c, the experimental spectrum is reproduced by the simple function depending only on the deuteron wave function and the kinematics

(4.2)
$$\frac{\mathrm{d}\sigma}{\mathrm{d}p_5} \sim \mathrm{E}^{-} \frac{\Phi^{(2)}}{\Phi^{(3)}} \frac{\mathrm{p}_5^2}{\mathrm{E}_5} \qquad (*)$$

The region of high momentum (so called "tail") is however showing the still unexplained disagreement.

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In the same way we can study the distribution in the spectator angle with the same model, for various momentum region. We consider to gain statistic, the forward-backward asymmetry for different momentum regions, that is the difference between the number of events with $0 \le \cos \vartheta_5 \le 1$ and $-1 \le \cos \vartheta_5 \le 0$, normalized with the sum of all events in the considered momentum bin. As shown in Table I, for .98 GeV/c the agreement between theory and experiment is satisfactory between 0. and .2 GeV/c, but surprisingly the agreement is broken in the .2 + .3 GeV/c bin. On the "tail" region the agreement is restored, but it is not very significant because of the disagreement in the momentum distribution. If we go to higher energies, the simple function (4.2) tends to be not sufficient to explain the angular behaviour. The reason is that at 1.13 and 1.5 GeV/c the total charge exchange cross section is rapidly decreasing with energy and the energy of the K⁺n system is very sensitive to the spectator angle

$$(4.3) s = m_2^2 + m_4^2 + 2E_2E_5 + 2p_2p_{5z}$$

This fact is an indication that the analysis of these data should take in account of the Fermi motion.

TABLE I

The forward backward asymmetry as function of the momentum (in percentage).

p ₅ GeV/c	$p_2 = .98 \text{ GeV/c}$		p ₂ = 1.13		p ₂ = 1.5	
	Exp	Th	Exp	Th	Exp	Th
0 ÷ .1	10.0+ 4.3	9.1	3.2± 4.8	8.9	0.0±7.3	8.9
.1 + .2	16.7± 9.8	19.1	4.3± 9.5	13.9	8.4±13.0	13.7
.2 + .3	2.1±13.5	24.7	23.4±14.2	22.3	0.0±20.4	12.8
.3 + .6	44.7±10.7	60.0	28.6± 9.7	51.7	42.9±15.3	43.2

^(*) We choose E because we assume that the process occurs through a charge exchange without spin-flip (see (3.4),(3.7)). It is reasonable to assume that the amplitudes contribute in this distribution only as a normalization factor.

5. - COMPARATIVE DISCUSSION AND RESULTS

The present section is devoted to the analysis of the Stenger (⁵) procedure and to the comparison of his and our results. Eventually, we consider the weight factors of Jew (⁶) and compare it with ours.

The procedure of Stenger is very similar to ours: the main difference is in the normalization of the scattering amplitude, which doesn't affect the result for the cross-section:

(5.1)
$$t_{fi}^{s} = \frac{T_{fi}}{\sqrt{8E_{3}E_{4}E_{5}}} \sqrt{E_{1}E_{2}}$$

Therefore the kinematical factor is in our notation

(5.2)
$$\frac{\omega(\mathbf{k'})}{\mathbf{k'}^2} \frac{d\mathbf{E'}_{\mathbf{f}}}{d\mathbf{k'}} \frac{\mathbf{k'}^2 \mathbf{d}_{\mathbf{k}}}{\omega(\mathbf{k})} P^2 \frac{d\mathbf{P}}{d\mathbf{E}_{\mathbf{f}}} d\Omega_p \equiv \frac{d\ \Omega^*}{d\mathbf{V}(^2)} \frac{\Phi(^2)}{\Phi} \frac{d\mathbf{V}(^3)}{d\Omega^*} \mathbf{k}^2$$

where $\vec{k} = \vec{p}$, $E_{\vec{f}} = E_3 + E_4 + E_5$ and k', $E_{\vec{f}}$ are the corresponding quantities for the two body process on the nucleon at rest in the laboratory system, and $\vec{P} = \frac{\vec{p}_4 - \vec{p}_5}{2}$

His approximation is to take out of the integration symbol in k and Ω_p , the phase space volume element and the flux factor for the two body process, but mantaining the dependence of the integration variables in the three body phase space volume element. The value of P is calculated for any value of k and Ω_p using energy conservation

(5.3)
$$\mathbb{E}_{\hat{1}} = m_1 + \mathbb{E}_2 = \mathbb{E}_3 + \mathbb{E}_4 + \mathbb{E}_5 = \mathbb{E}_3 + \sqrt{m^2 + Q^2 + P^2 + 2} \vec{P} \cdot \vec{Q} + \sqrt{m^2 + Q^2 + P^2 - 2} \vec{P} \cdot \vec{Q}$$

where $\vec{Q} = \frac{\vec{p}_4 + \vec{p}_5}{2}$.

Q is thought fixed in the integration and determined with the two body kinematics on the nucleon at rest, from the C.M. scattering angle.

Squaring twice the relation (5.3), one obtains

(5.4)
$$P^{2} = \frac{A^{2} - 4B^{2}}{4(1 - 4C^{2}/A^{2})}$$

where

 $B^{2} = m^{2} + Q^{2}$ $C = Q \cos P^{2}Q$

The main approximation of Stenger consists in neglecting the term $4C^2/A^2$ with respect to 1 and obtaining therefore for P the simple relation

(5.5)
$$P^2 = \frac{A^2 - 4B^2}{4}$$

This approximation allows the immediate integration on the angle Ω_p . He claims to obtain this relation simply expanding the square roots in (5.3) in terms of 2PQ cos P^Q/(m²+ +P²+Q²); but this expansion is possible only if the scattering angle is sufficiently small: in the backward direction $\vec{P} \sim \vec{Q} \sim \vec{p}_4/2 \sim \frac{\omega}{E+\omega} \vec{p}_2$ (where E, ω are energies of the meson and the nucleon in their C.M. system) therefore $\vec{P} \cdot \vec{Q} \simeq P^2$; the above expansion parameter becomes $2P^2/(m^2+2P^2)$, which is very small only if $m^2 >> 2 P^2$.

This condition is obviously violated for high energy and we expect this procedure to fail in this region. To point out the dependence on the angle P^Q of the value of P^2 we have inserted in the Stenger calculation the exact determination of P^2 in terms of various fixed values of cos (PQ) (Fig. 4).

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As shown on the Fig. 6 the Stenger calculation corresponding to $\cos P^Q = 0$ (dotdashed curve) at .98 GeV/c for $K^+D \rightarrow K^opp$ is very low in the backward direction. If we give to $\cos P^Q$ values .5 and 1, (dashed and full curve respectively), the behaviour in the backward direction changes considerably and it tends to the exactly calculated curve (see Fig. 7). This last calculation is done using the same method of Stenger, that is

1) taking out of the integration the term $\frac{d\Omega^*}{dV(^2)} \Phi(^2)$

2) determining the 3 body kinematics in terms of $\cos \vartheta^*$, using the approximation $\vec{p}_5 = 0$.

but no approximation is done to reduce the number of integrations and a numerical method is used to calculate the three-fold integral on the volume element $d^3p_5 = p_5^2 d p_5 d \phi_5 d \cos \theta_5$.

On Fig. 7 we show the results for W^+ at the different momenta of the K^+ .65 (- - -) and 1.51 GeV/c (------). From that figure we see that the overlapping region which is characterized by $W^+ > 1$, is rapidly decreasing with energy.

On Fig. 8 we show the behaviour with energy of the W in the "closure" approximation (3.23). The main difference with the exact calculation is in the backward direction, where the closure result is, for $p_2 = .65$ GeV/c, 10% higher than the "exact" result. This is due either to the cut of the spectator momentum $p_5 \leq 250$ MeV/c either to the exclusion from the integration region of the kinematically impossible values of \vec{p}_5 .

In Fig. 9 we show the behaviour of W^+ with the momentum cut at $p_2 = 1.51$ (- - - -; $p_5 \le 100 \text{ MeV/c}$; ----- $p_5 \le 150 \text{ MeV/c}$; _____ $p_5 \le 250 \text{ MeV/c}$).

Eventually in Fig. 10 we compare the results of W^+ obtained by the method, described above, (dashed lines) with the results of the calculation without making any approximation (continuous line), that is

1) keeping inside the integration $\frac{d\Omega^*}{dV(2)} \Phi^2$

2) determining the 3 body kinematics in terms of $\cos \vartheta_3^{*}$ taking in account the Fermi motion of the nucleon.

The result for the weight factors is very simple (3.24): however, we have to say that in the second term of formula (3.24) we cannot in principle determine the three body kinematics because, as from appendix B, we have to know the energy of the two body scattering which in this case has to be chosen between s_{34} and s_{35} . However we make the choice s_{34} and calculate it, remembering that the overlapping term is large only around the forward direction and in this region $s_{34} \sim s_{35}$.

The results for the two different procedure look quite similar, a part from the backward region, which is influenced by the different approximations for the phase space volume Fig. 10, 11). ACKNOWLEDGMENTS

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APPENDIX A

A practical method to determine the constant in the relation between the deuteron vertex and the non relativistic function is given in this appendix. The method consists in the calculation of the pole diagram and the comparison of its expression with the T matrix with one pole in potential scattering. The diagram is



where M is the mass of the deuteron.

Let us now consider this expression the C.M. system and the non relativistic limit: then

(A.2)

$$p^{2} = \left(2\sqrt{q^{2}+m^{2}}\right)^{2} = 4m^{2} + \frac{q^{4}}{m^{2}} + 4q^{2} \sim 4m^{2} + 4q^{2}$$

$$(A.2)$$

$$4m^{2} = \left(M + \frac{\chi^{2}}{m}\right)^{2} = M^{2} + \chi^{4} + \frac{2M}{m}\chi^{2} \sim M^{2} + 4\chi^{2}$$

where m is the mass of the nucleon, $\frac{\chi^2}{m}$ is the binding energy of the deuteron.

Extracting from the above expression the invariant T matrix, we obtain

$$(A.3) \qquad T = -\frac{G^2}{4(q^2+\chi^2)}$$

let us now calculate the scattering amplitude in potential scattering, assuming that it is simply a pole term in q^2 plane

(A.4)
$$F = \frac{1}{2iq} \quad \frac{f_0(q) - f_0(q)}{f_0(-q)} = -\frac{1}{N^2} \quad \frac{1}{q^2 + \chi^2}$$

where

$$\mathbb{N}^2 = \int_0^\infty \mathbf{f}_0^2 (-\mathbf{i} \ \chi, \mathbf{r}) d\mathbf{r}$$

and $f_0(q)$, $f_0(q,r)$ are respectively the Jost function and the irregular solution of the radial Schroedinger equation.

The relation between the invariant T matrix and the scattering amplitude in the C.M. system is

$$(A.5) \qquad \qquad T = 8 \pi \sqrt{s} f$$

The comparison between the two results gives the relation between G and N

(A.6)
$$G^2 = 64 \pi E N^{-2}$$

where E is the energy of the nucleon.

Let us now consider the deuteron vertex,



and assume that p_2 in on the mass shell. The expression of this vertex, apart from the delta function in energy and momentum and factors, is

(A.7)
$$\frac{G}{p_1^2 - m^2} \cong \frac{G}{(-2)(q^2 + \chi^2)}$$

The asymptotic wave function in momentum space is

(A.8)
$$\psi(q) = -\frac{1}{\pi \sqrt{2}N} - \frac{1}{q^2 + \chi^2}$$

which compared with (A.7), gives the following relation

(A.9)
$$\frac{G}{p_1^2 - m^2} = (32 \pi^3 E)^{\frac{1}{2}} \psi (q)$$

where E is the energy of the on-mass-shell particle.

We want to stress here that although this relation seems to be valid only for an asymptotic wave function, it turns out to be true even for a general deuteron wave function, as shown more rigorously by Gross (¹⁴). In this case G is thought as a function of q. APPENDIX B

In this appendix we consider in detail the procedure of calculating the full kinematics of the 3-body final state once 4 variables are fixed.

The method is based on the kinematical relation between the invariant quantities built with the energies and momenta of the 5 particles involved in the process



These relations are easily derived if we consider two of the 5 particles as a single one and reduce therefore the 3-body to a normal 2 body process; the relations are of the following type:

(B.1)
$$s_{12} + t_{23} + t_{13} = s_{45} + m_1^2 + m_2^2 + m_3^2$$

This is the usual relation between the Mandelstan variables for the two body case. The definition of the invariants is the usual one:

(B.2) $s_{ij} = (p_i + p_j)^2$ $t_{ij} = (p_i - p_j)^2$

Obviously we can consider as one body any pair from the 5 different particles and obtain therefore 10 different relations, relative to the 10 possible pairs of "incident" particles.

In the analysis of the experimental data and the successive phase shift analysis, the best choice of the independent variables is given by the three-momentum of one of the outgoing nucleons and the C.M. angle of the incident particle-nucleon process.

The definition of \vec{p}_5 and the complete knowledge of the initial state fixes the invariants s_{12} , t_{15} and t_{25} ; the energy and momentum conservation fixes $s_{34} = (p_2+p_1-p_5)^2$.

The knowledge of ϑ_3^* gives the possibility of determining the squared four momentum transfer tzs, once the energy of the two body process is known. There are now two different ways of choosing this energy:

1) taking the internal line off-mass-shell and keeping energy conservation on the upper vertex of diagram

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2) taking the internal line on the mass shell and violating energy conservation

S = S34

(B.4)
$$s = \left(\sqrt{p_2^2 + m_2^2} + \sqrt{p_5^2 + m_5^2}\right)^2 - \left(\vec{p}_2 - \vec{p}_5\right)^2$$

Once t_{23} is determined through one of these two methods, the procedure follows without further ambiguities.

The first step is the determination of t_{24} using equation (B.3) and the one of the same kind obtained considering as one body particles 1 and 3.

$$(B.5) t_{24} + s_{45} + t_{25} = t_{13} + m_2^2 + m_4^2 + m_5^2$$

Summing these two relations it follows.

(B.6)
$$t_{24} = -t_{23} - t_{25} - s_{12} + m_1^2 + 2m_2^2 + m_3^2 + m_4^2 + m_5^2$$

The second step is the derivation of linear equation for the invariants $s_{3\,5}$ and $t_{1\,3\,;}$ considering the two Mandelstan relations

 $t_{14}+t_{24}+s_{12} = s_{35}+m_1^2+m_2^2+m_4^2$ $t_{13}+t_{14}+s_{34} = t_{25}+m_1^2+m_3^2+m_4^2$

it follows

(B.7)

 $s_{35} = t_{25} + t_{24} + s_{12} - s_{34} - t_{13} + m_3^2 - m_2^2$

The invariant s35 can be also expressed using (B.2)

(B.8) $s_{35} = m_3^2 + m_5^2 + 2E_3E_5 - 2 \vec{p}_3\vec{p}_5$

where

 $\begin{array}{rcl} \mathbb{E}_{3} &= \left(\mathbb{m}_{1}^{2} + \mathbb{m}_{3}^{2} - \mathbb{t}_{1\,3}\right)/2\mathbb{m}_{1} \\ (\mathbb{B}.9) & & & & \\ \mathbb{p}_{3\,_{Z}} &= \left(\mathbb{t}_{2\,3} - \mathbb{m}_{2}^{2} - \mathbb{m}_{3}^{2} + 2\mathbb{E}_{2}\mathbb{E}_{3}\right)/2\mathbb{p}_{2} \\ & & & \\ \mathbb{p}_{3\,_{X}} &= \left(\mathbb{E}_{3}^{2} + \mathbb{m}_{3}^{2} - \mathbb{p}_{3\,_{Z}}^{2}\right)^{\frac{1}{2}} \end{array}$

By substitution of (B.9) in(B.8) we obtain another relation between s_{35} , t_{13} and t_{23} . Subtracting (B.8) from (B.7) we find a quadratic relation for t_{13} . The solution for t_{13} suffers for a sign ambiguity, which is solved comparing with the two body kinematics.

The knowledge of $t_{1\,3}\,allows$ the complete derivation of \vec{p}_3 and momentum conservation gives then \vec{p}_4 .

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FIGURE CAPTIONS

- Fig. 1 Two-body scattering diagram. The dashed lines correspond to mesons and the full lines to nucleons.
- Fig. 2 Feynman graphs describing the deuteron break-up in the impulse approximation. 1 is the deuteron, 2 and 3 are the incoming and outgoing mesons, 4 and 5 are the two nucleons in the final state.
- Fig. 3 Perspective representation of the function $E^{\dagger}dV(^{3})/d\Omega_{3}dp_{5}d\Omega_{5}$ for $\varphi_{5} = \pi$ and cos $\vartheta_{3}^{*} = 0$. The quasi elastic peak and the spectator moraine are apparent.
- Fig. 4 As Fig. 3 with $\cos \vartheta_3^* = 0.9$. In this case the overlapping of the quasi elastic peak and the spectator moraine clearly shows off.
- Fig. 5 Momentum distribution of the spectator nucleon for the $K^{\dagger}d \rightarrow K^{\circ}pp$ process at 0.98 GeV/c. Thick line refers to the experimental data and the thin line to the simplified nodel described in the text.
- Fig. 6 W^+ versus cos ϑ * for the $K^+d \rightarrow K^\circ pp$ process at 0.98 GeV/c, as calculated in ref.(⁵) for: cos P^Q = 0.0 dot-dashed

 $\cos P^Q = 0.5$ dashed and $\cos P Q = 1.0$ full line.

- Fig. 7 W^+ versus cos ϑ^* for the $K^+d \rightarrow K^\circ pp$ process at 0.65 (dashed curve) and 1.51 GeV/c (full line), calculating the three-fold integral on the volume element d^3_{p5} exactly.
- Fig. 8 As Fig. 7 but in closure approximation.
- Fig. 9 W⁺ versus cos θ* for the K⁺d → K^opp process at 1.51 GeV/c for three different cuts on p₅ : < 0.100 GeV/c dashed, < 0.150 dot-dashed and < 0.250 full line.</p>
- Fig.10 Comparison of W^+ versus cos ϑ^* for the $K^+d \rightarrow K^\circ pp$ process at 0.65 GeV/c with $p_5 < 0.100 \text{ GeV/c}$. The two curves are computed with the two different procedures explained in the text.

Fig. 11- As Fig. 10 with p5 < 0.250 GeV/c.









Fig. 7







Fig. 10

Fig. 11