



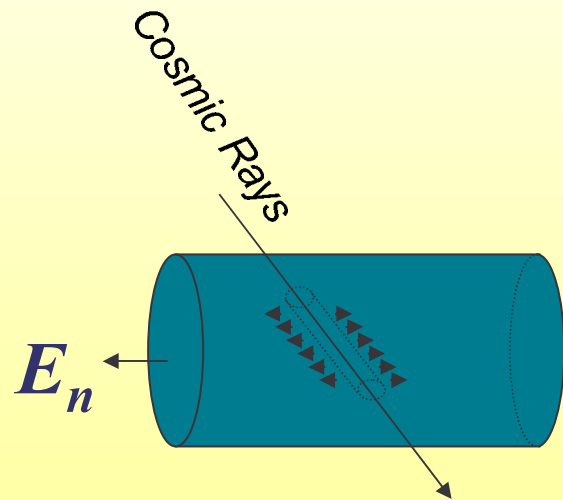

Rivelazione Acustica di Particelle

S. Bertolucci, E. Coccia, S. D'Antonio, A.C. Fauth, A. de Waard,
G. Delle Monache, D. Di Gioacchino, V. Fafone, G. Frossati, C. Ligi, A.
Marini, G. Mazzitelli, G. Modestino, G. Pizzella, L. Quintieri, G. Raffone,
F. Ronga, P. Tripodi, P. Valente

INFN – Laboratori Naz. di Frascati (RM)
INFN – sez. di Roma2 & Univ. Tor Vergata, Roma
Leiden University, The Netherlands
Universidade Estadual de Campinas, Brazil
ENEA Centro Ricerche Frascati (RM)

The NAUTILUS Gravitational Wave Antenna has recently recorded signals due to the passage of cosmic rays. *Large signals at higher rate than expected* has been observed in the superconductive state of the antenna.

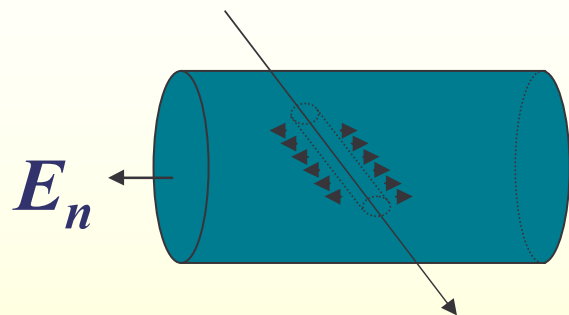
Thermo-acoustic model



CR crossing the antenna loss energy

- *warming up of the material*
- *local thermal expansion*
- *mechanical vibrations*

Thermo-acoustic model



The energy deposited by a particle is converted in a local increase of temperature

$$\delta T = \delta E / \rho C_V V$$

$$\delta P = \gamma \delta E / V$$

$\gamma \equiv \beta V / \chi_T C_V$ is the **Grüneisen parameter**

$$E_n \propto \gamma^2 (dE/dx)^2 F_n^2$$

thermal and mechanical

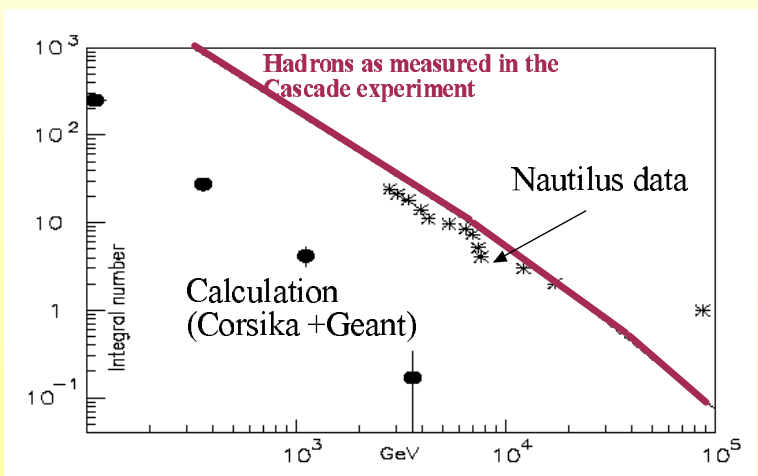
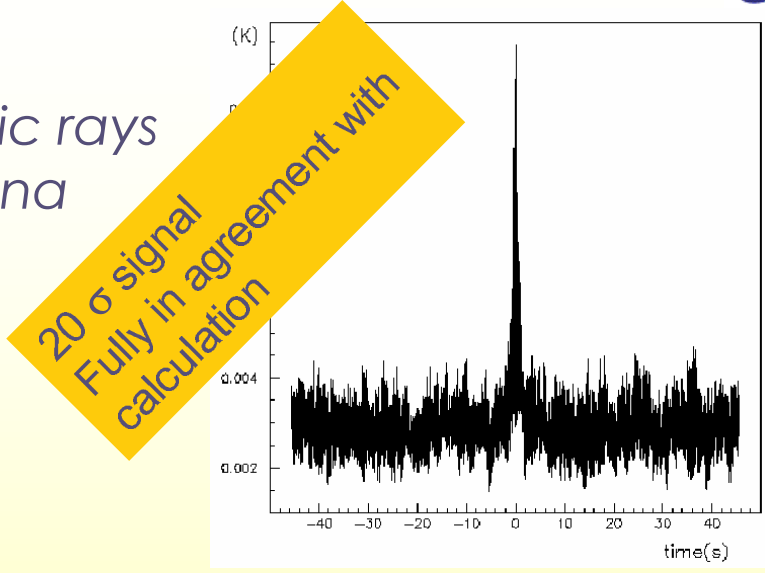
energy loss

Geometric factor

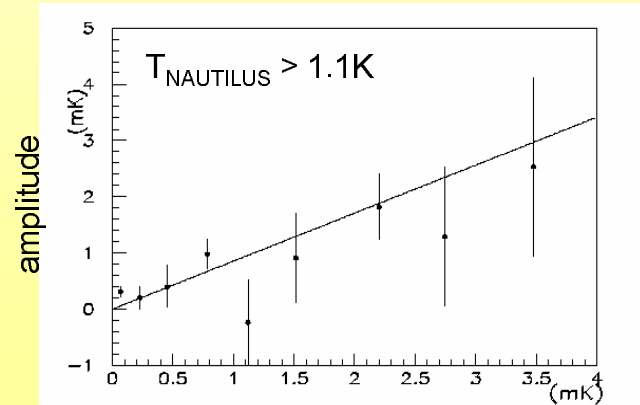
The Thermo-acoustic Model predicts **very small** signal for present resonant gravitational wave detector sensitivity. It has been proven effective at room temperature by previous experiments (Rev. Sci. Instrum. **71** (2000) 1345-1354 and previous papers)

1999 first measurement of cosmic rays signature in the NAUTILUS antenna
 Phys. Rev. Lett. **84** (2000) 14

2000 anomalous signal detected
 Phys. Lett. **B499** (2001) 16



2001 dependence from NAUTILUS thermodynamic temperature
 Phys. Lett. **B540** (2002) 179



T.A.M. expected signal Induced by measured shower
 SIF 2003 - Parma, 20 settembre 2003



(End of 2001)

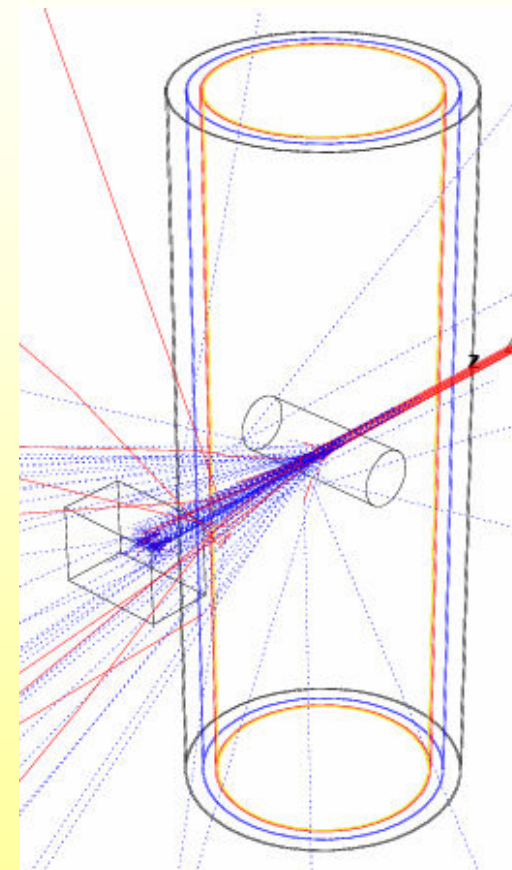
RAP Proposal

LNF-01-027(IR)

In order to understand:

$$E_n \propto \gamma^2 (dE/dx)^2 F_n^2$$

- enhancement of γ *Grüneisen factor* in super-conducting state
- enhancement of energy conversion (dE/dx) in super-conducting state
- exotic component of cosmic rays in (dE/dx) (nuclearites, monopoles)



GEANT simulation

$$\gamma \equiv \beta V / \chi_T C_V$$

Mechanical and thermal properties of the detector are contained in the Grüneisen parameter, commonly assumed almost constant with the temperature.

The extrapolated value at zero Kelvin from measurements done down to 4 K, gives: $\gamma = 1.6$

β cannot be directly measured at temperatures below T_c . Recent evaluation based on critical magnetic field $H_c(P,T)$ measurements and specific heat gives a different value of the Grüneisen parameter in the super-conducting state. In particular at $T = T_c$:

$$\gamma_s \sim -10$$

e^- / e^+

$n_{\text{average}} = 1 - 10^{10}$ particles

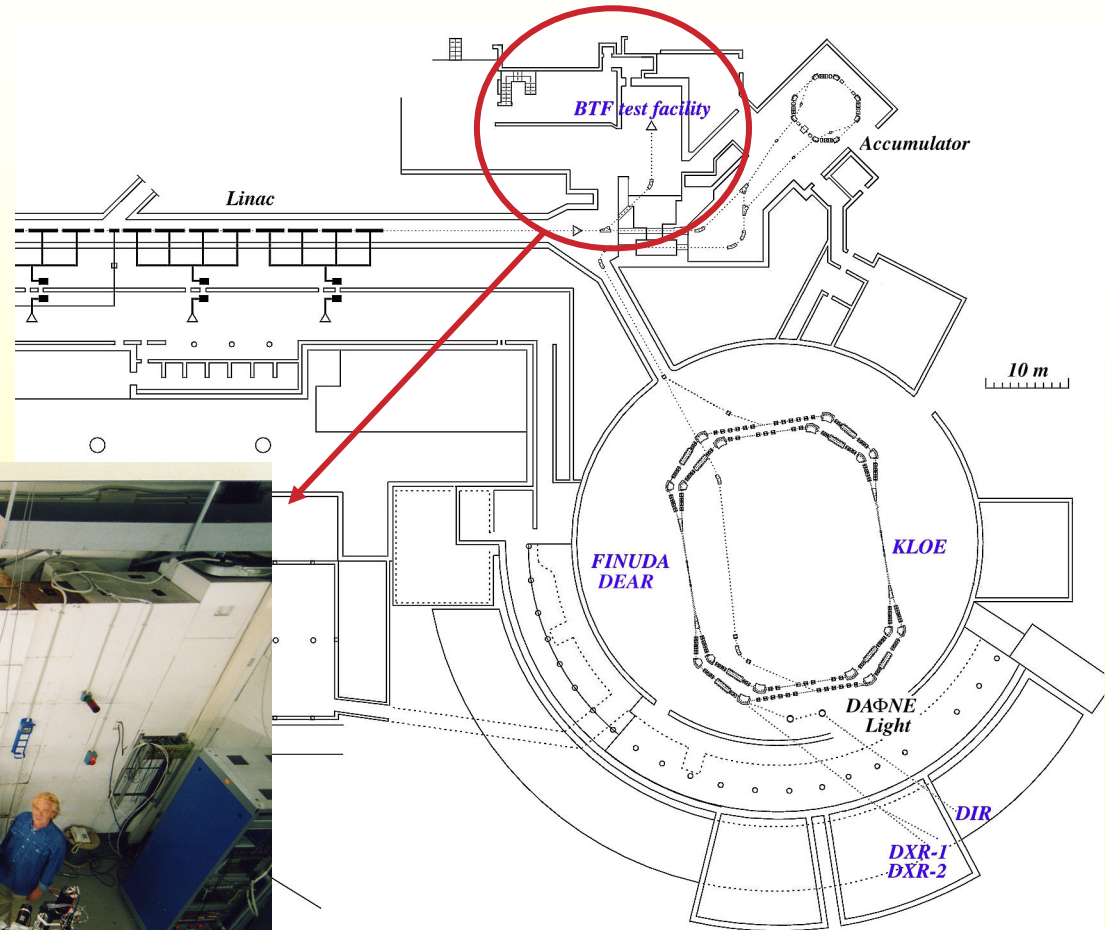
Energy: 20 – 800 MeV

Repetition rate: 50 Hz

Pulse Duration: 1 – 10 ns

1% energy selection

100 m² Experimental Hall



Carlo Ligi

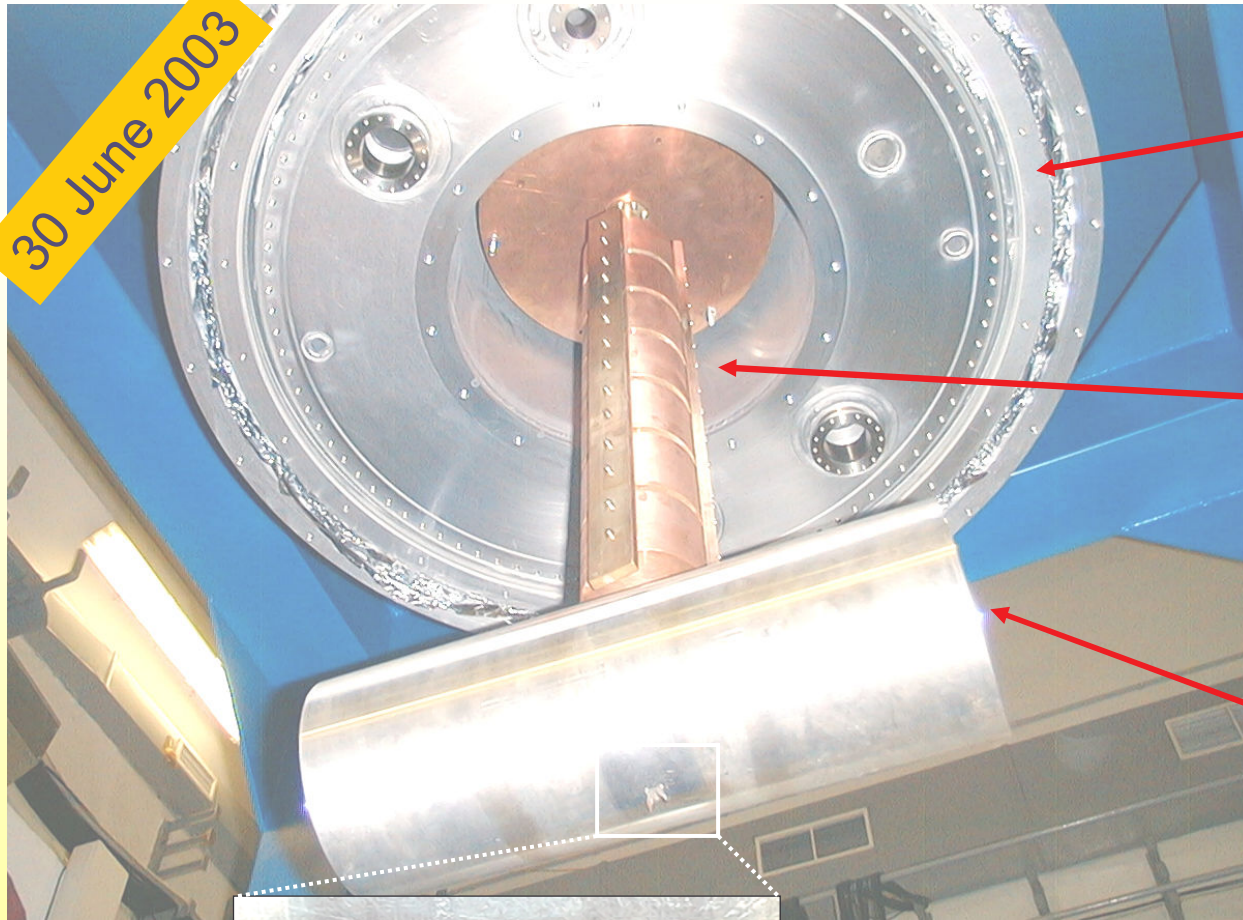
SIF 2003 – Parma, 20 settembre 2003



18 June 2003

Experimental Setup

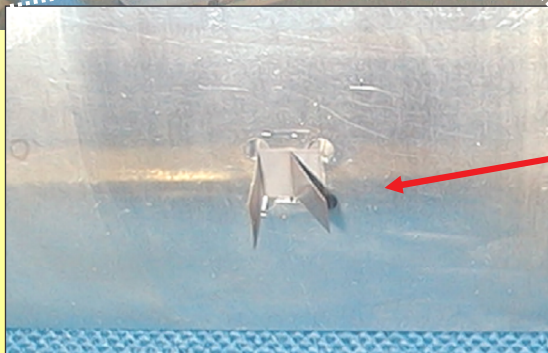
30 June 2003



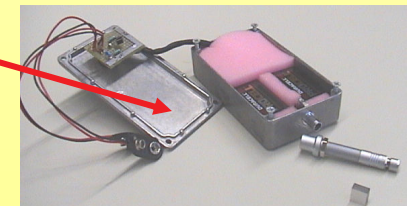
KADEL Liquid Helium Cryostat + Dilution Refrigerator
Working Temp: 100 mK

Suspension:
1 OFCH copper tube with
7 OFCH copper masses
Attenuation: -200db @ 5KHz

Antenna: Al 5056 bar
50x18 cm, 35 Kg
Res. Freq: 5096 Hz @ 300 K



Read-out: 2 piezo-electric
ceramics + JFET amplifier
 $1 \text{ nV Hz}^{-1/2}$ @ 5 KHz
bandwidth 25 KHz

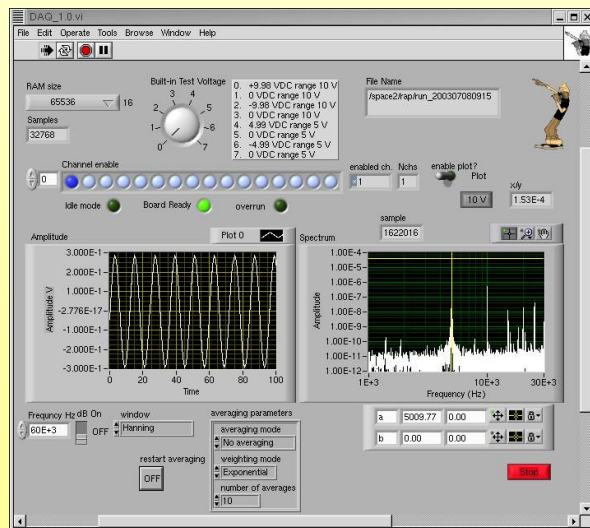
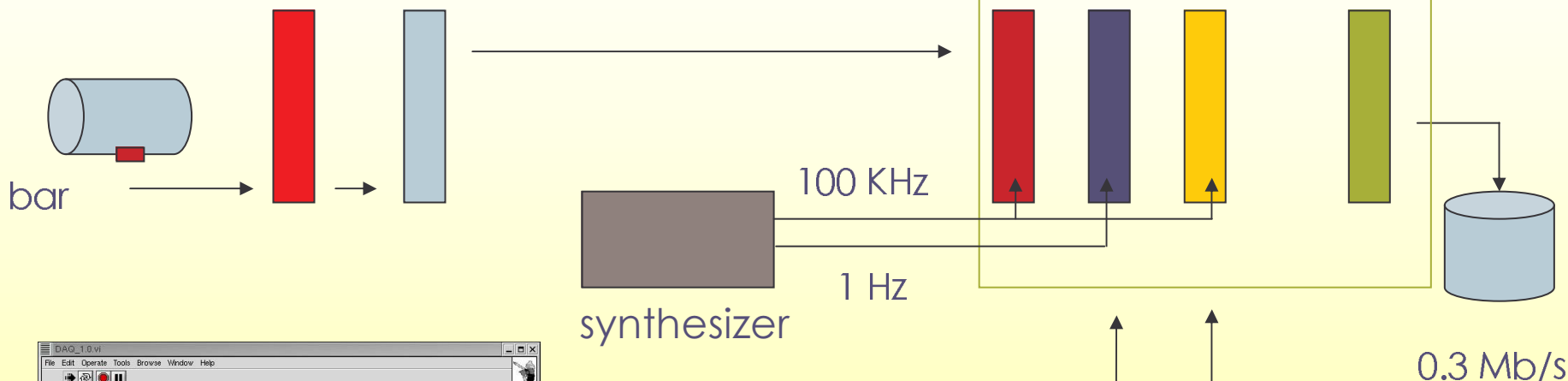


LN-JFET-AMP
1 nV Hz^{-1/2} @ 5 KHz
bandwidth 25 KHz

STANFORD
SR560 amplifier
4 nV Hz^{-1/2}

ADC INCAA 5758
ADC VMIC 3123
SCALER

LINUX VME
controller
VMIC 7740

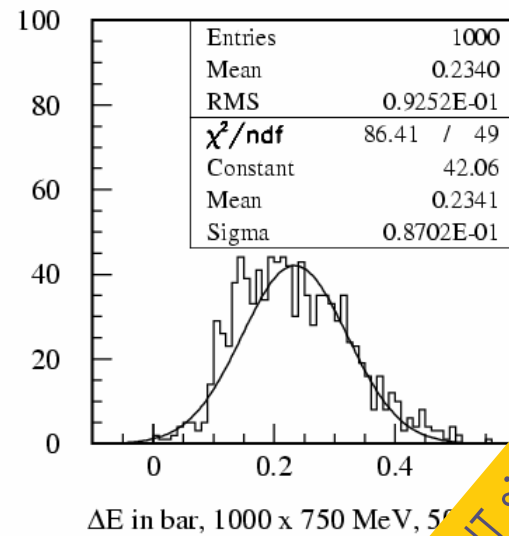
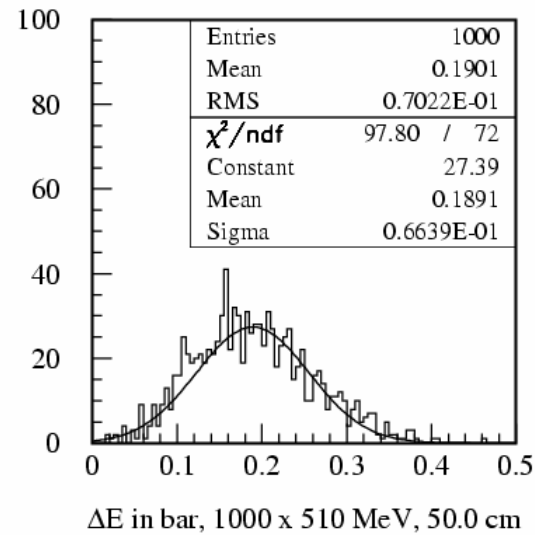


LabView™ run
controller & monitor

auxiliary channels

beam trigger

Expected Signal



GEANT simulation

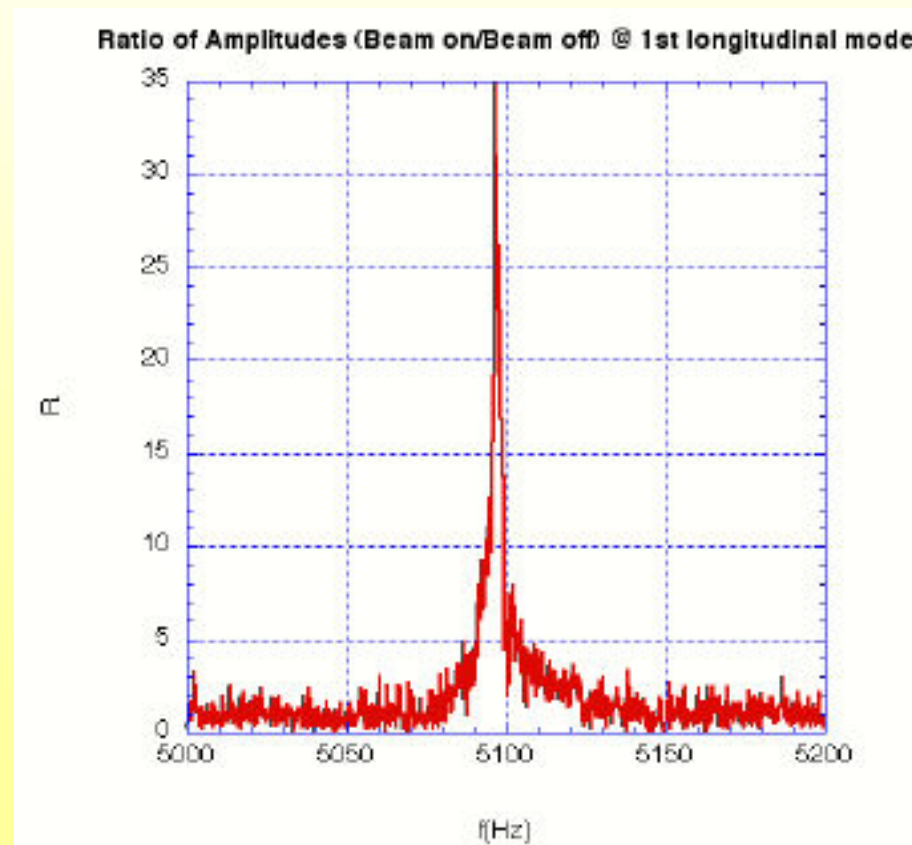
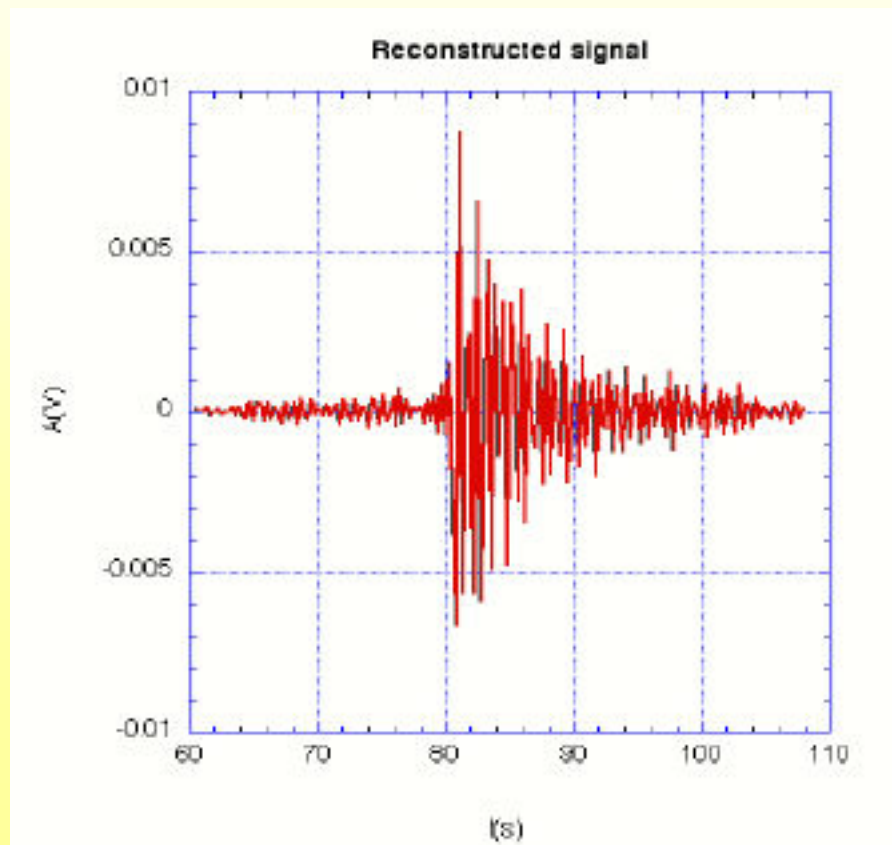
Phase 1: installation and start of full apparatus at room temperature, suspension, electronics, mechanical structure ready, first measurements at room temperature

DONE!

$9.4 \cdot 10^7$ electrons @ 510 MeV

Energy released in the bar = $1.8 \cdot 10^{-3}$ J

Expected vibration in the 1L mode = $4.99 \cdot 10^{-13}$ m



Phase 2:

- cryogenic test
- low temperature measurement in normal-conducting state (4.2 K)
(end 2003)

Phase 3:

- dilution refrigerator installation
- measurement in super-conducting state
(2004)



www.Inf.infn.it/esperimenti/rap/