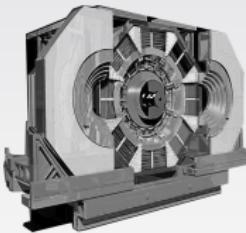


ZDD at BESIII

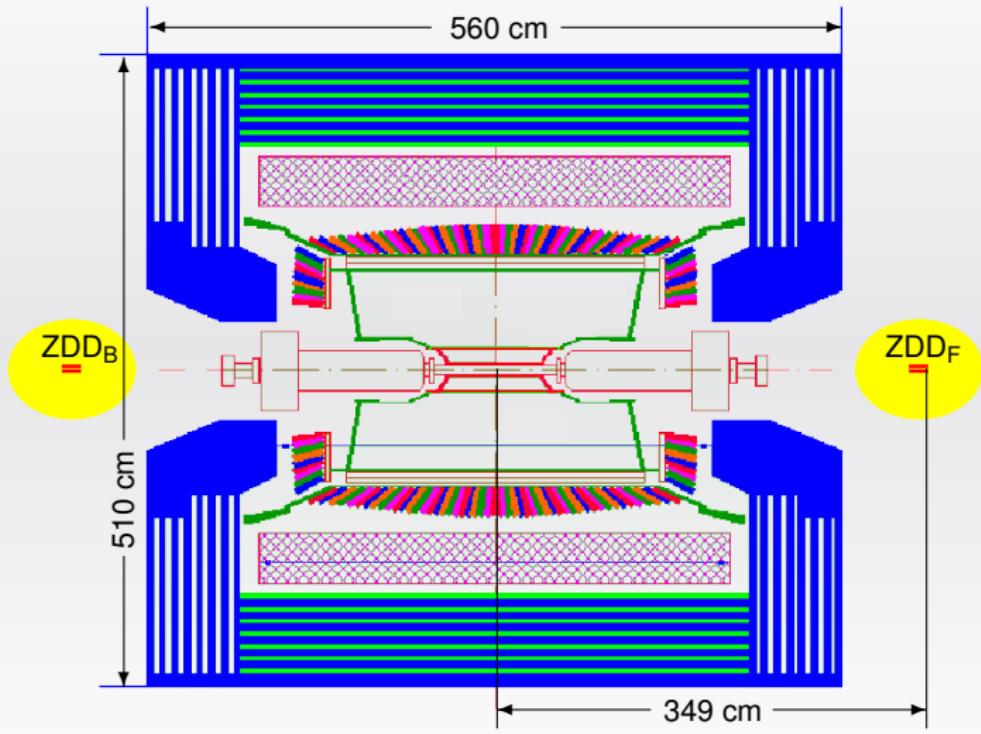
R. Baldini Ferroli and A. Zallo
for the Italian BESIII Group



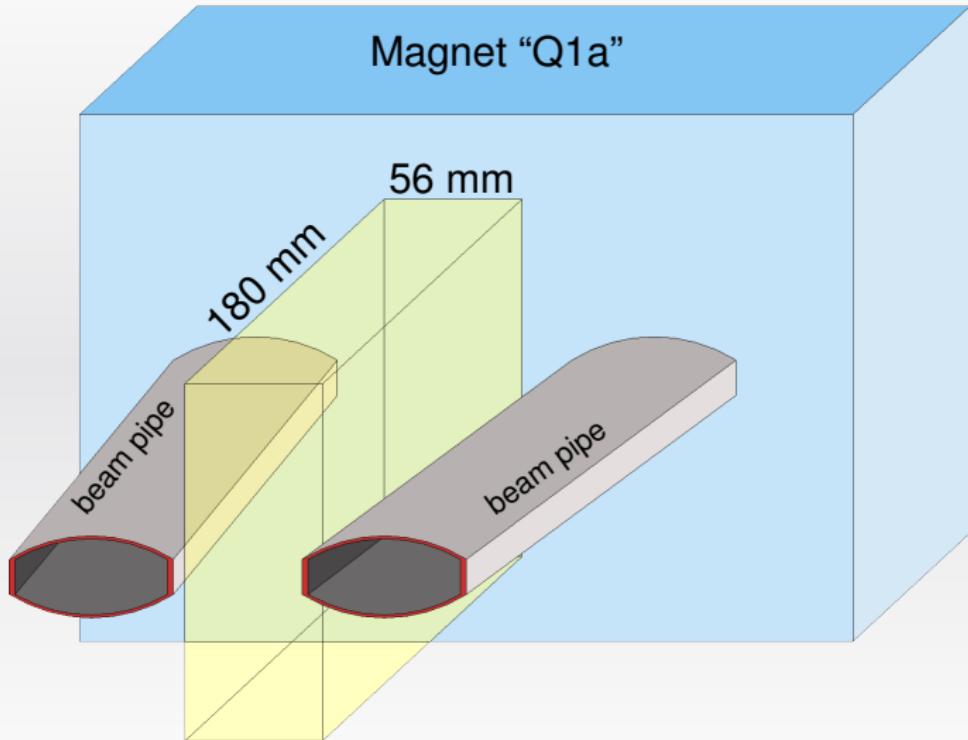
December 17, 2010 - Giessen, Germany

Design and Installation

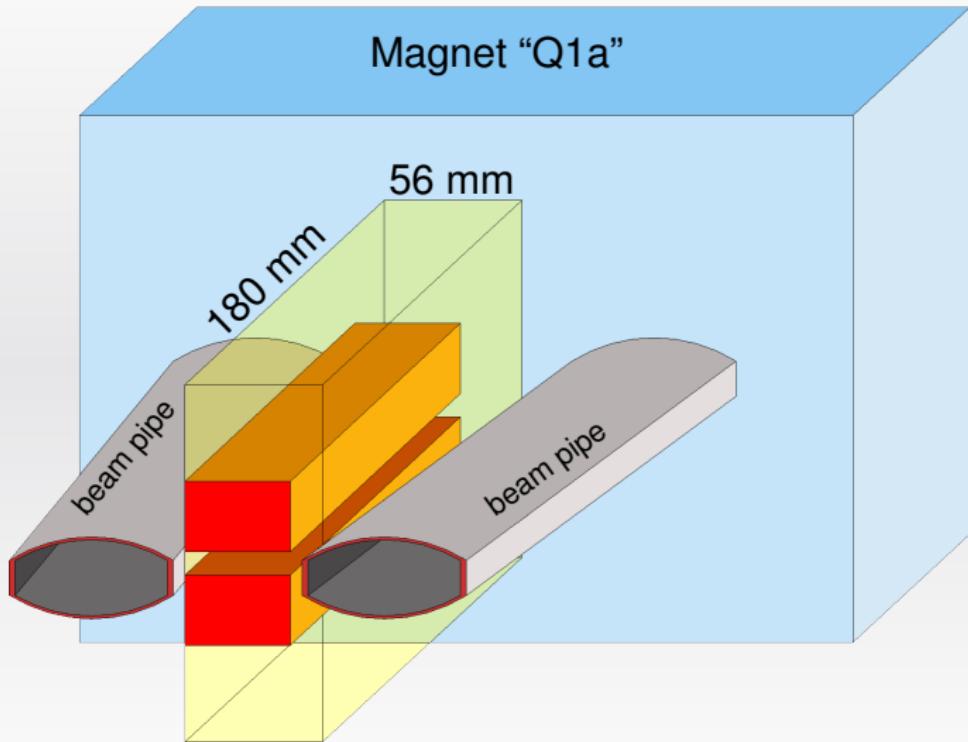
BESIII and ZDD



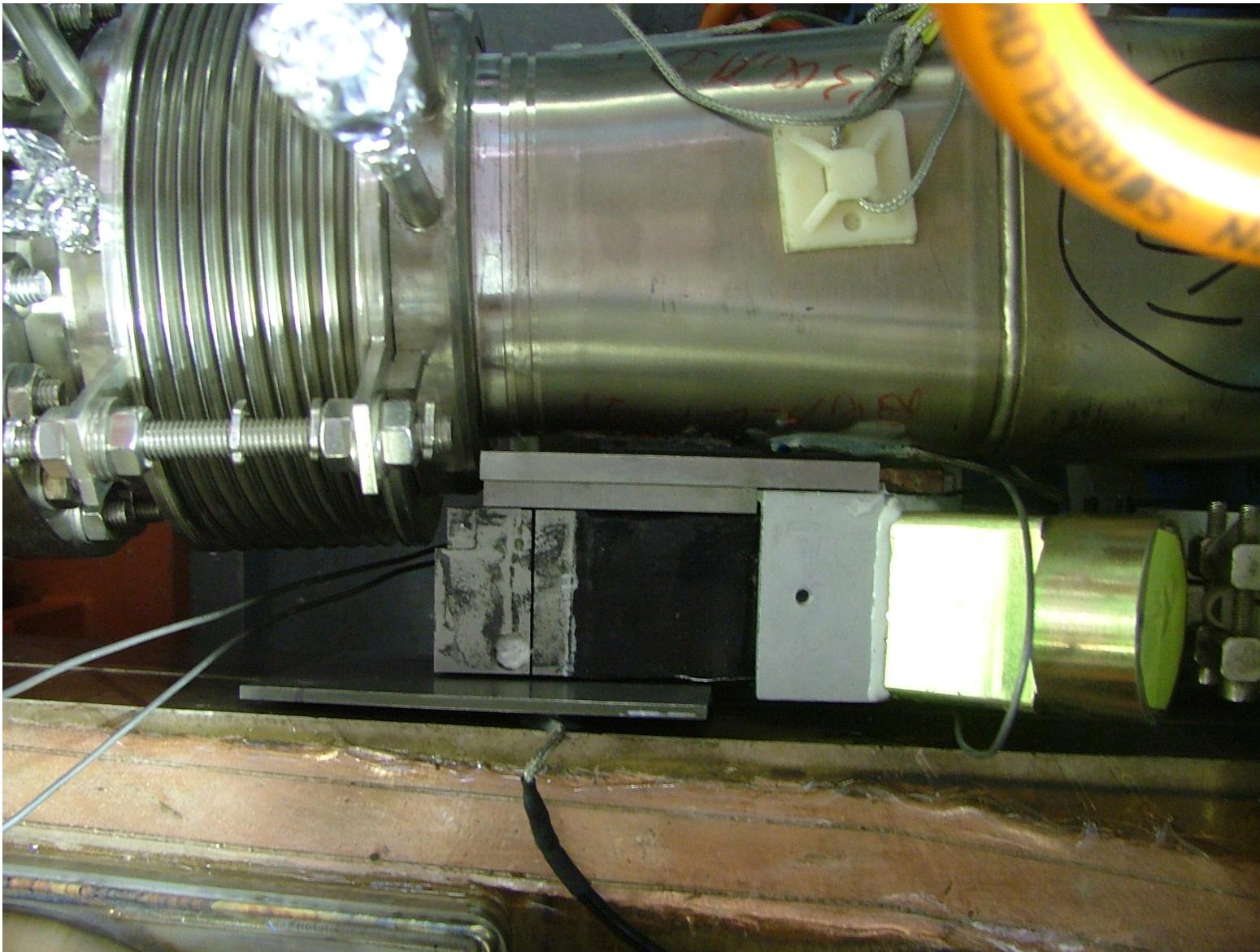
Available space



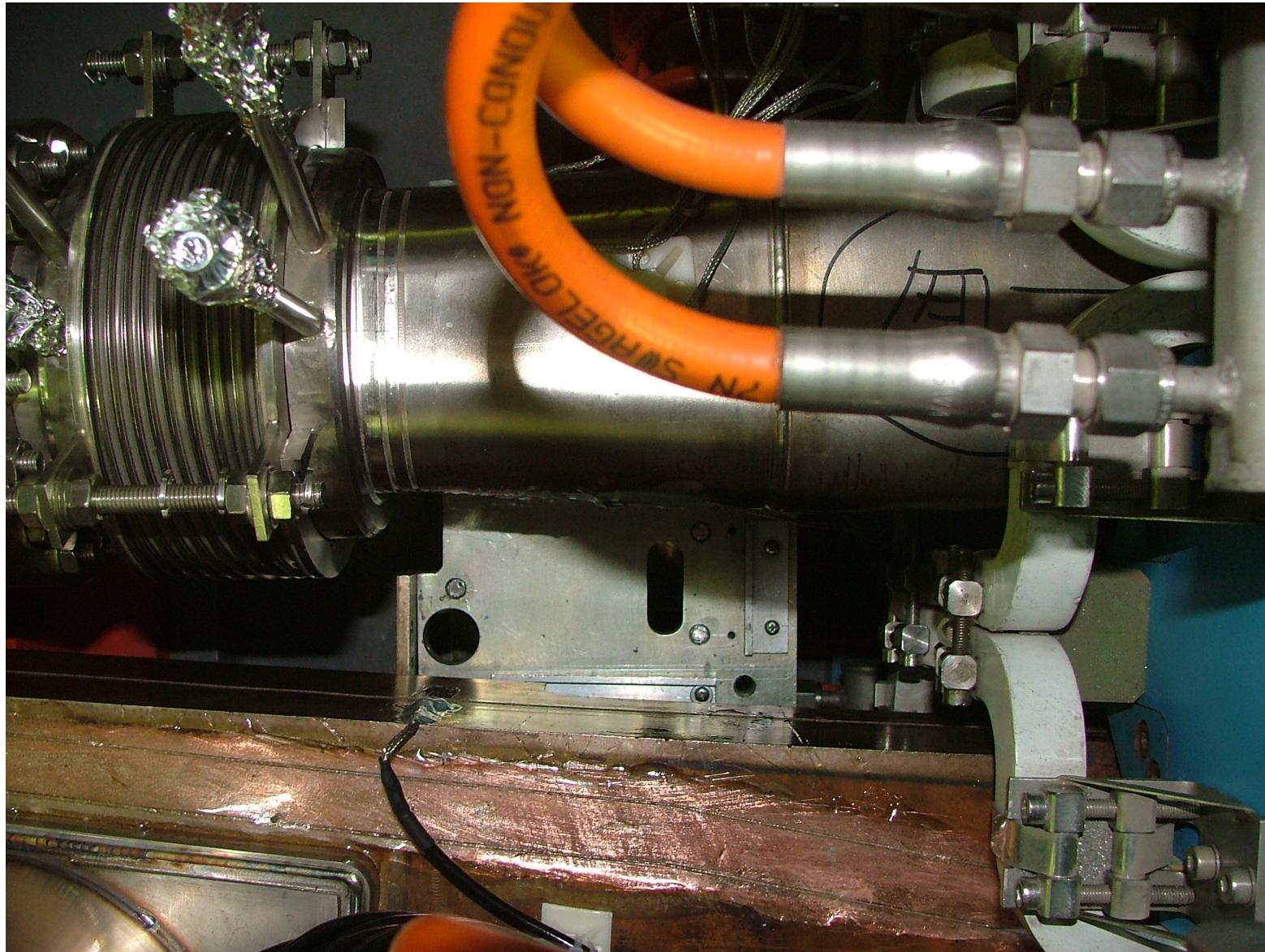
Available space



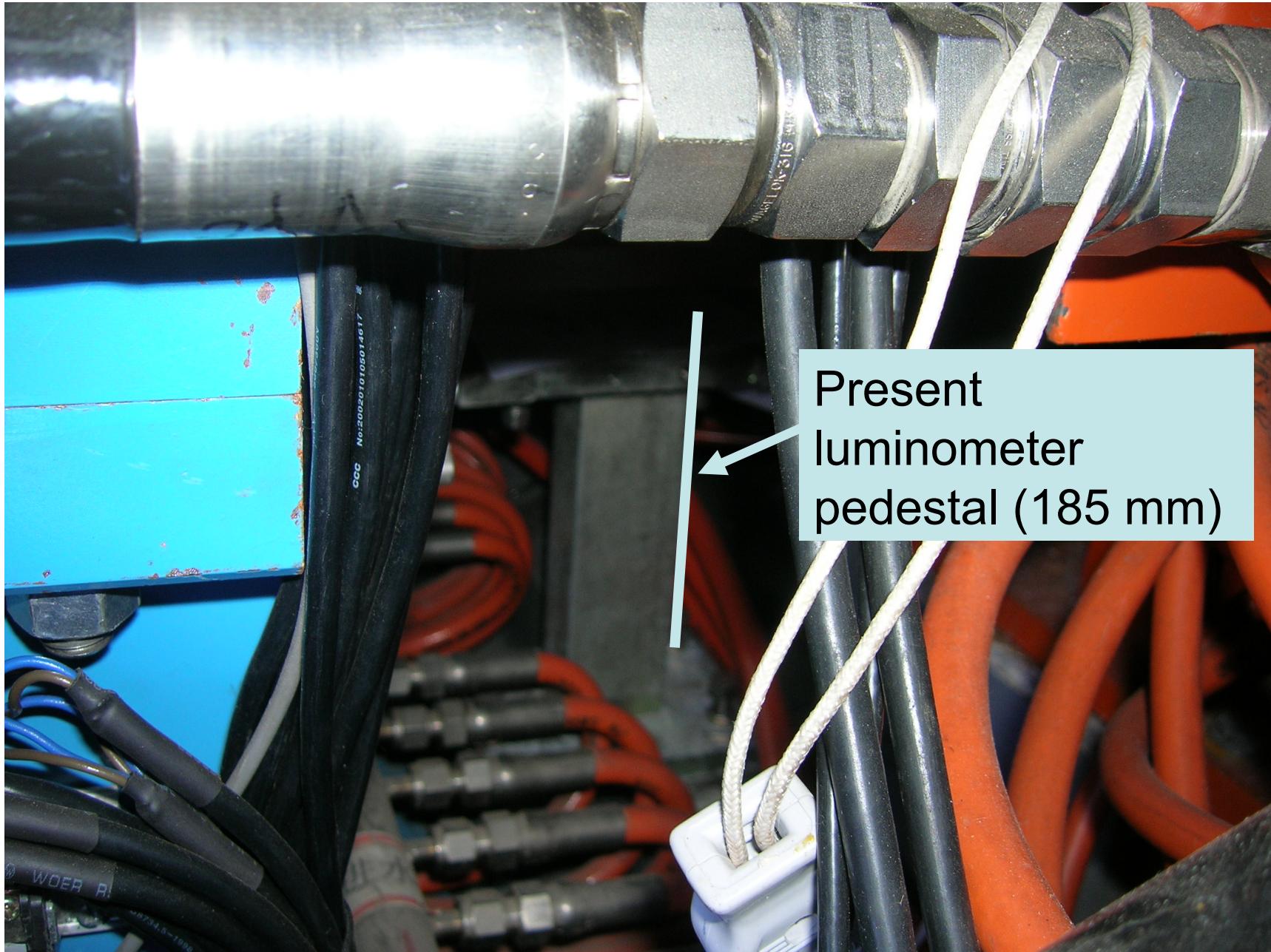
BESIII Lumi detector



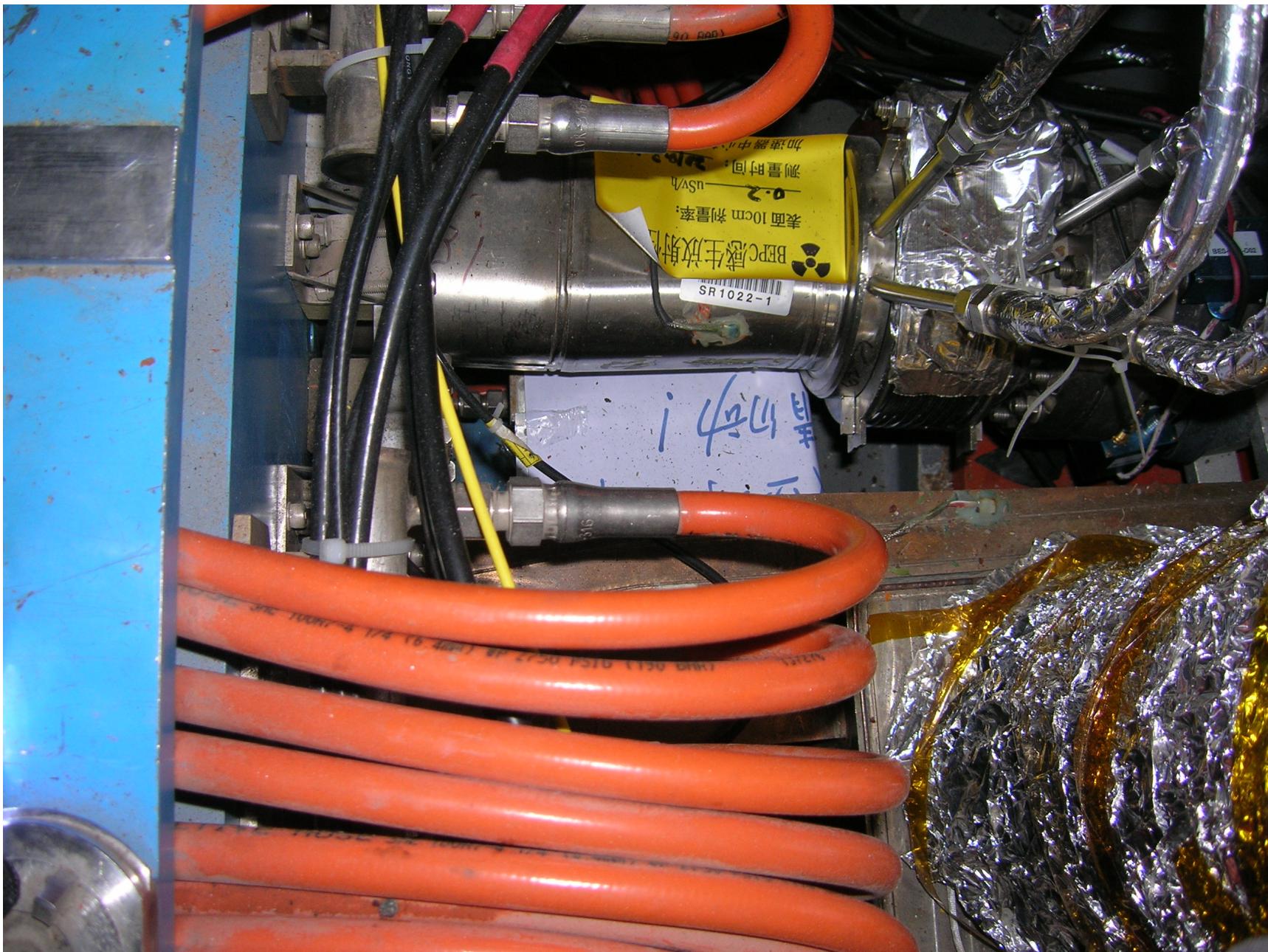
BESIII Lumi detector



August survey of the region, no Lumi detector



August survey of the region, no Lumi detector, top view

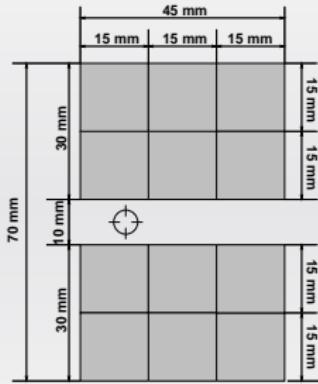


Two options:

LYSO and Pb-Scint

LYSO design

Front view
(cross section)

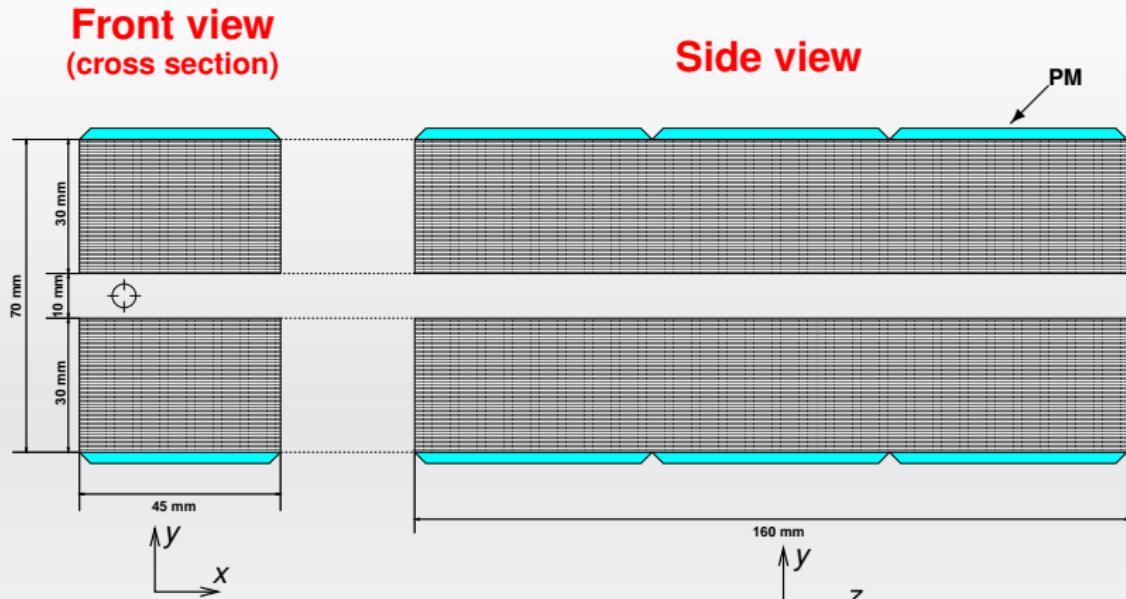


Side view

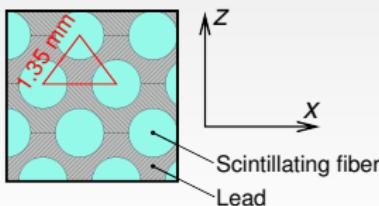


- Two 3×2 matrices of $1.5 \times 1.5 \times 16 \text{ cm}^3$ of LYSO bars
- Total volume 864 cm^3

Pb-Scintillator design à la Kloe



- Along z axis:
320 grooved 0.5 mm thick lead foils
alternated with layers of cladded
1 mm diameter scintillating fibers
- Readout with PM



Physical properties of materials

Material	LYSO	Pb-Scint
Density (g/cm ³)	7.4	5.3
Radiation Length (cm)	1.1	1.6
Molière Radius (cm)	1.9	2.9
Decay Constant (ns)	40-44	2.4
Peak Emission (nm)	428	460
Radiation Hardness (rad)	$\sim 10^8$	$\sim 10^6$

Radiation hardness

- Radiation damages mostly due to Bremsstrahlung: $\sigma_{\text{Bre}}(\text{ZDD}/4) = 2.6 \text{ mb}$
- One year of data taking: $T = 1.5 \times 10^7 \text{ s}$
- Average luminosity: $\bar{\mathcal{L}} = 1.5 \times 10^{32} \text{ cm}^{-2}\text{s}^{-1}$
- Center of mass energy: $E_{c.m.} = 3.77 \text{ GeV}$

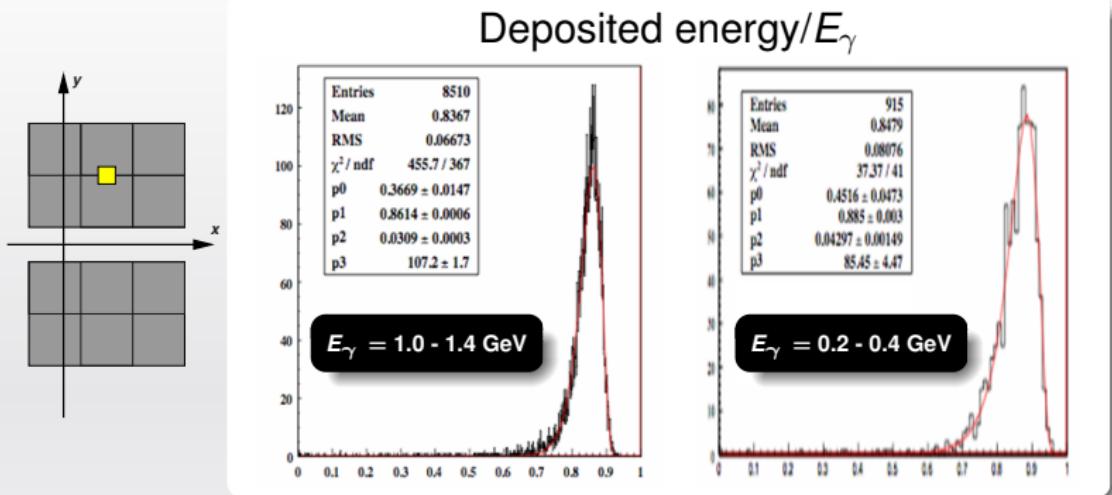
$$\frac{\text{Dose absorbed}}{\text{year}} = \frac{\text{energy deposited}}{\text{year} \cdot \text{mass}} = \begin{cases} \frac{3 \times 10^{21} \text{ eV}}{0.12 \text{ kg}} = 4 \times 10^5 \frac{\text{rad}}{\text{year}} & \text{LYSO} \\ \frac{3 \times 10^{21} \cdot \frac{2}{13} \text{ eV}}{1.8 \times 10^{-2} \text{ kg}} = 10^6 \frac{\text{rad}}{\text{year}} & \text{Scint} \end{cases}$$

Declared hardness

- LYSO $\sim 10^8 \text{ rad}$
- Scint. $\sim 10^6 \text{ rad}$

Energy Resolution

LYSO GEANT4 simulation₁



Log-normal distribution

$$\frac{df}{dE} = \frac{\eta}{\sqrt{2\pi}\sigma_E\sigma_0} e^{-\frac{1}{2} \left[\frac{\ln^2 \left(1 - \frac{\eta(E-E_0)}{\sigma_E} \right)}{\sigma_0^2} + \sigma_0^2 \right]}$$

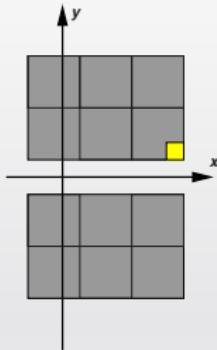
$$\sigma_0 = \frac{2}{2.35} \ln \left[\eta \frac{2.35}{2} + \sqrt{1 + \left(\eta \frac{2.35}{2} \right)^2} \right], \quad \sigma_E = \frac{\text{FWHM}}{2.35}$$

$E_{\gamma} (\text{GeV})$ $\sigma_{E_{\gamma}}/E_{\gamma}$ Central
(yellow square)

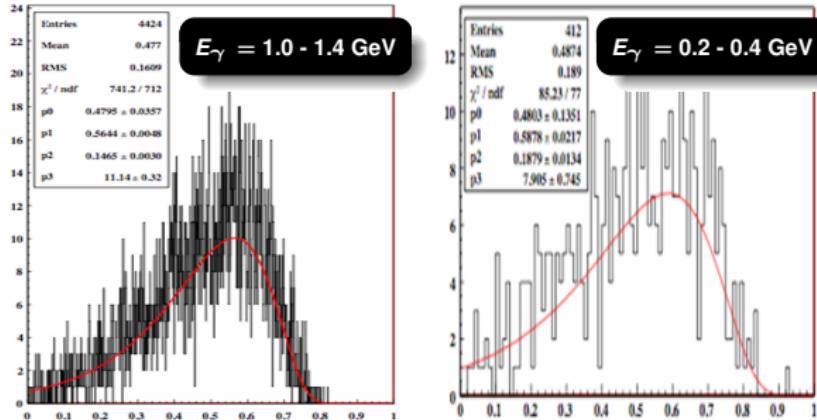
1.0 - 1.4 **3.6%**

0.2 - 0.4 **4.9%**

LYSO GEANT4 simulation₂



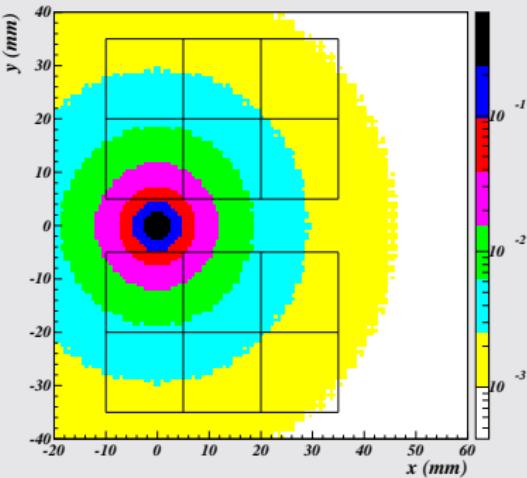
Deposited energy/ E_{γ}



$E_{\gamma} \text{ (GeV)}$	$\sigma_{E_{\gamma}}/E_{\gamma} \text{ Central}$ (yellow square)
1.0 - 1.4	26.0%
0.2 - 0.4	32.0%

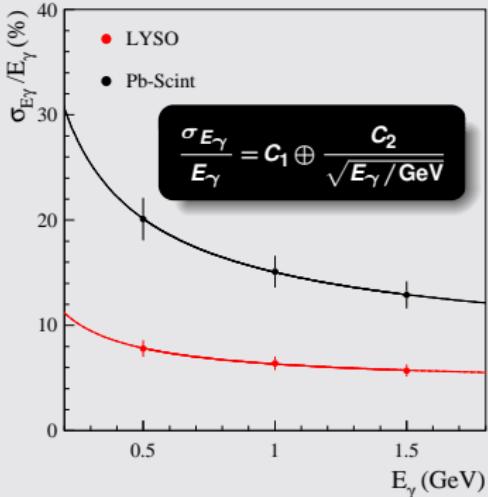
Energy resolution, the ISR case

ISR angular distribution on ZDD



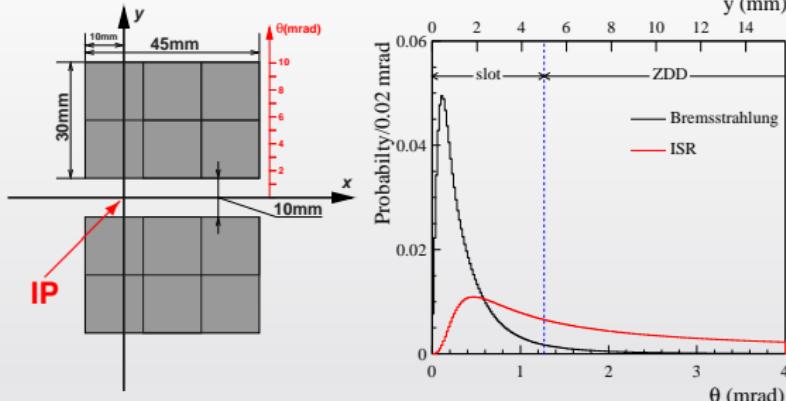
E_γ (GeV)	LYSO	Pb-Scint
$\sigma_{E_\gamma}/E_\gamma$	$\sigma_{E_\gamma}/E_\gamma$	$\sigma_{E_\gamma}/E_\gamma$
1.5	5.7%	12.9 %
1.0	6.4%	15.1 %
0.5	7.8%	20.1 %

Energy resolution for ISR



	LYSO	Pb-Scint
C_1	4.3%	6.9 %
C_2	4.6%	13.4 %

Bremsstrahlung simulation



- $E_{\text{beam}} = 1.89 \text{ GeV}$
- $E_{\gamma}^{\min} = 50 \text{ MeV}$
- $\sigma_{\text{Bre}}(4\pi) = 353 \text{ mb}$
- $\sigma_{\text{Bre}}(\text{ZDD}) = 10 \text{ mb}$
- $\mathcal{L} = 8 \times 10^{32} \text{ cm}^{-2} \text{ s}^{-1}$

- ISR in ZDD **13.7%** of total solid angle
- Bremsstrahlung in ZDD **2.8%** of total solid angle
- Bremsstrahlung rate in a single ZDD element (upper or lower):

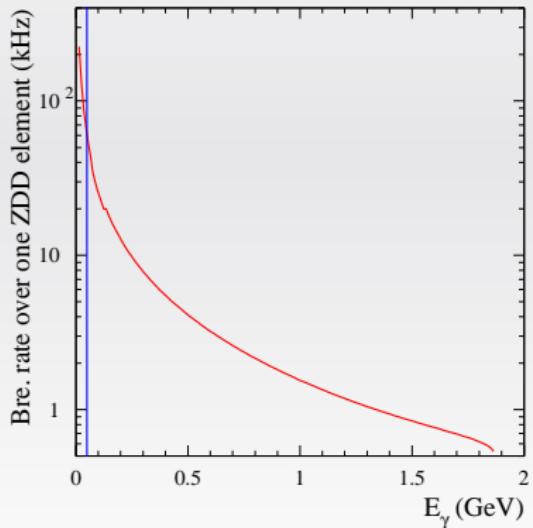
800 kHz at $\mathcal{L} = 3 \times 10^{32} \text{ cm}^{-2} \text{ s}^{-1}$

2.1 MHz at $\mathcal{L} = 8 \times 10^{32} \text{ cm}^{-2} \text{ s}^{-1}$

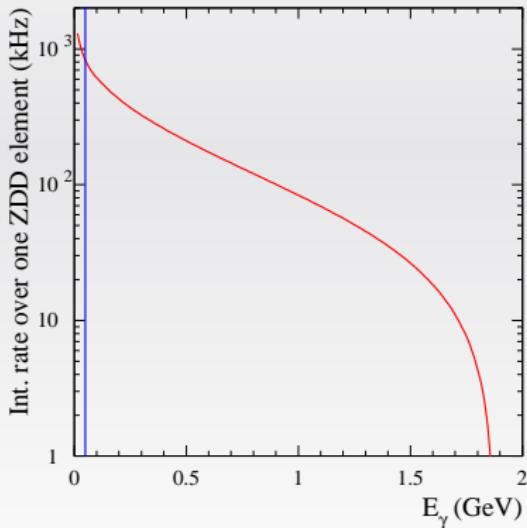
Bremsstrahlung rate

$$\mathcal{L} = 3 \times 10^{32} \text{ cm}^{-2}\text{s}^{-1}$$

Bremsstrahlung rate
in 10 MeV E_γ intervals



Integrated
Bremsstrahlung rate



Pileup effect₁: signal generation

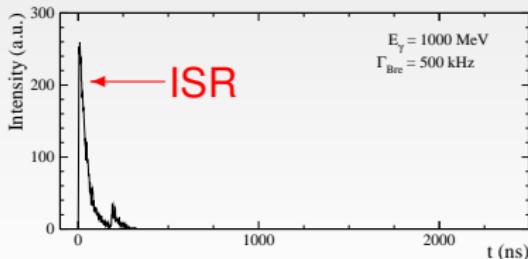
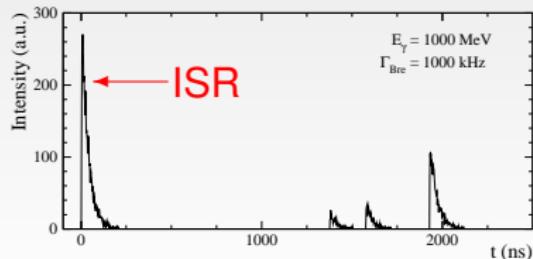
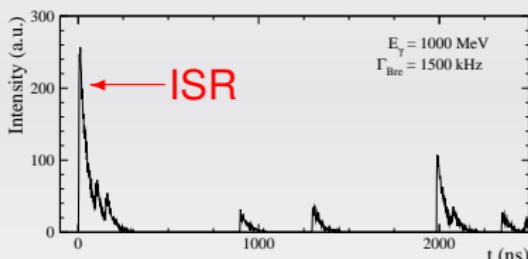
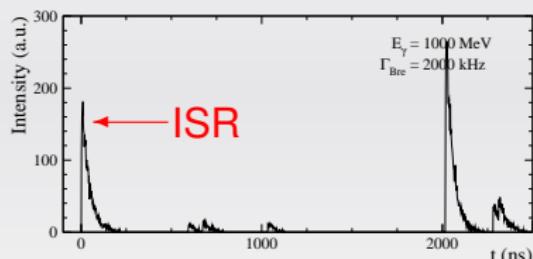
Maximum Bremsstrahlung rate expected 2.1 MHz (ZDD/4)

- Flash ADC: 500 MS/s, 8-bit resolution

- LYSO signal:

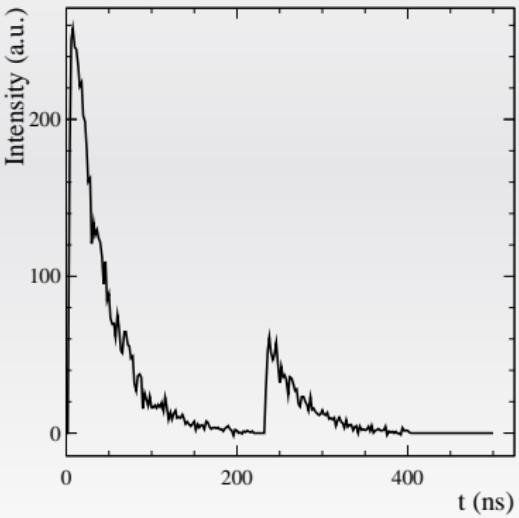
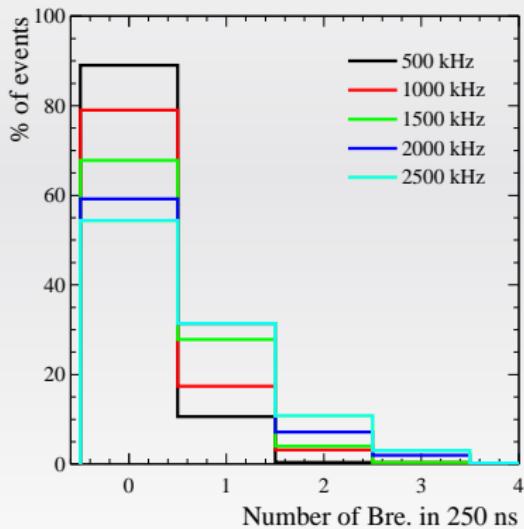
$$\text{Intensity} = e^{-t/\tau_d} (1 - e^{-t/\tau_r})$$

rising time $\tau_r = 2 \text{ ns}$, decay time $\tau_d = 40 \text{ ns}$

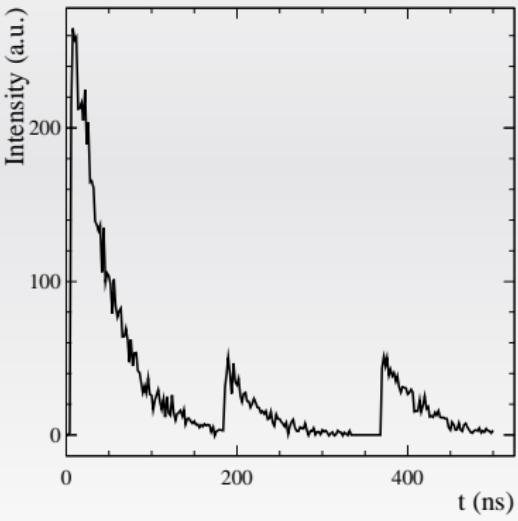
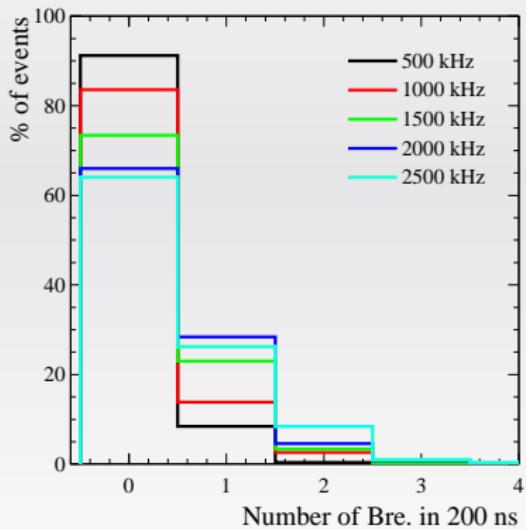


Pileup effect₂: probability

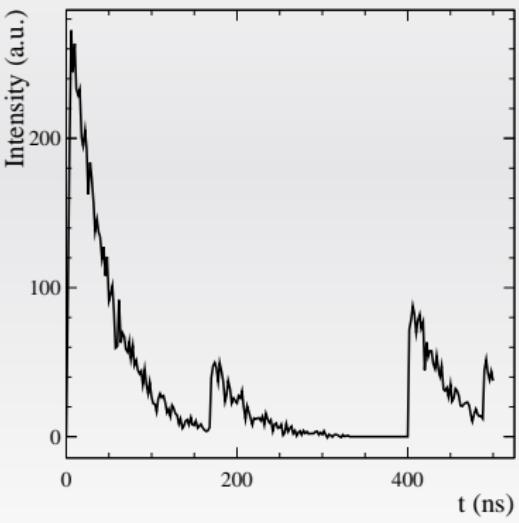
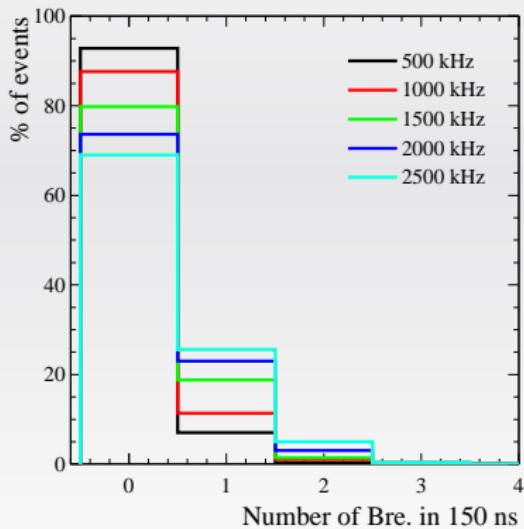
Probability of pileup as a function of the Bremsstrahlung rate



Probability of pileup as a function of the Bremsstrahlung rate

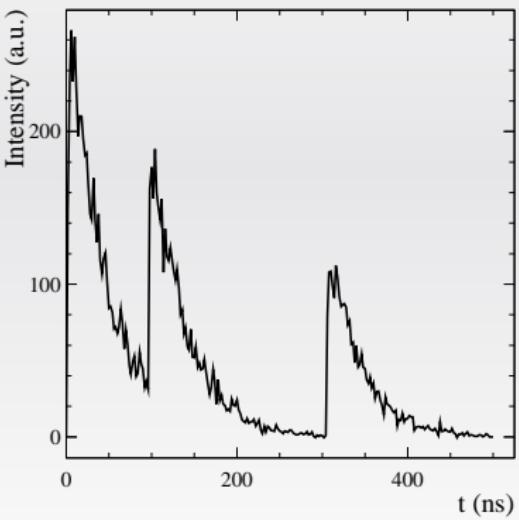
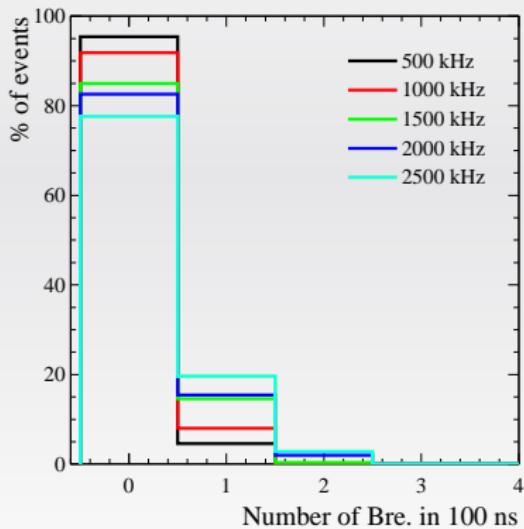


Probability of pileup as a function of the Bremsstrahlung rate

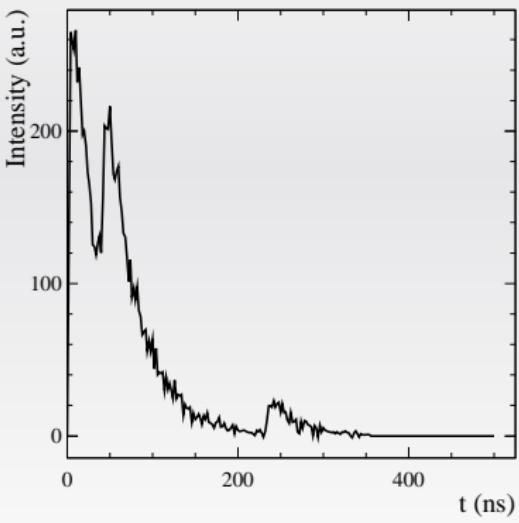
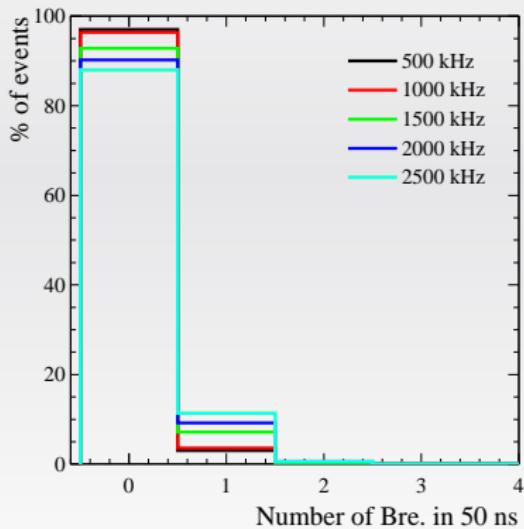


Pileup effect₂: probability

Probability of pileup as a function of the Bremsstrahlung rate

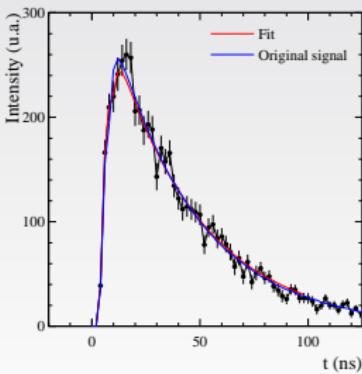
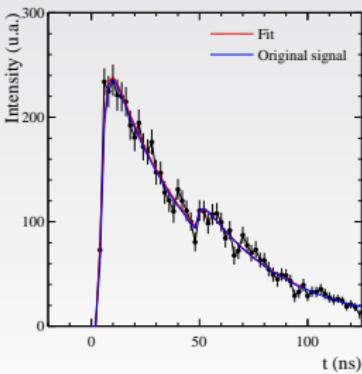
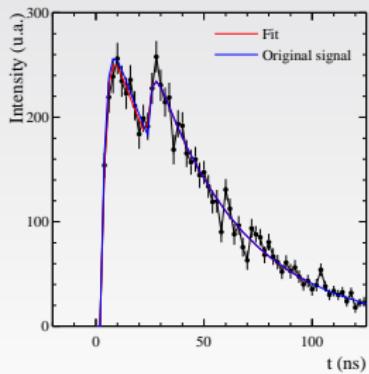


Probability of pileup as a function of the Bremsstrahlung rate



Pileup effect₃: evaluation

- 500 events have been generated at various rates
- E.g. at 2500 kHz:
158 (31.6%) have $\Delta t_{\text{ISR}} < 160 \text{ ns} \sim 4 \text{ decay times}$
- We fit these signals to verify our capability to distinguish ISR and Bremsstrahlung contributions

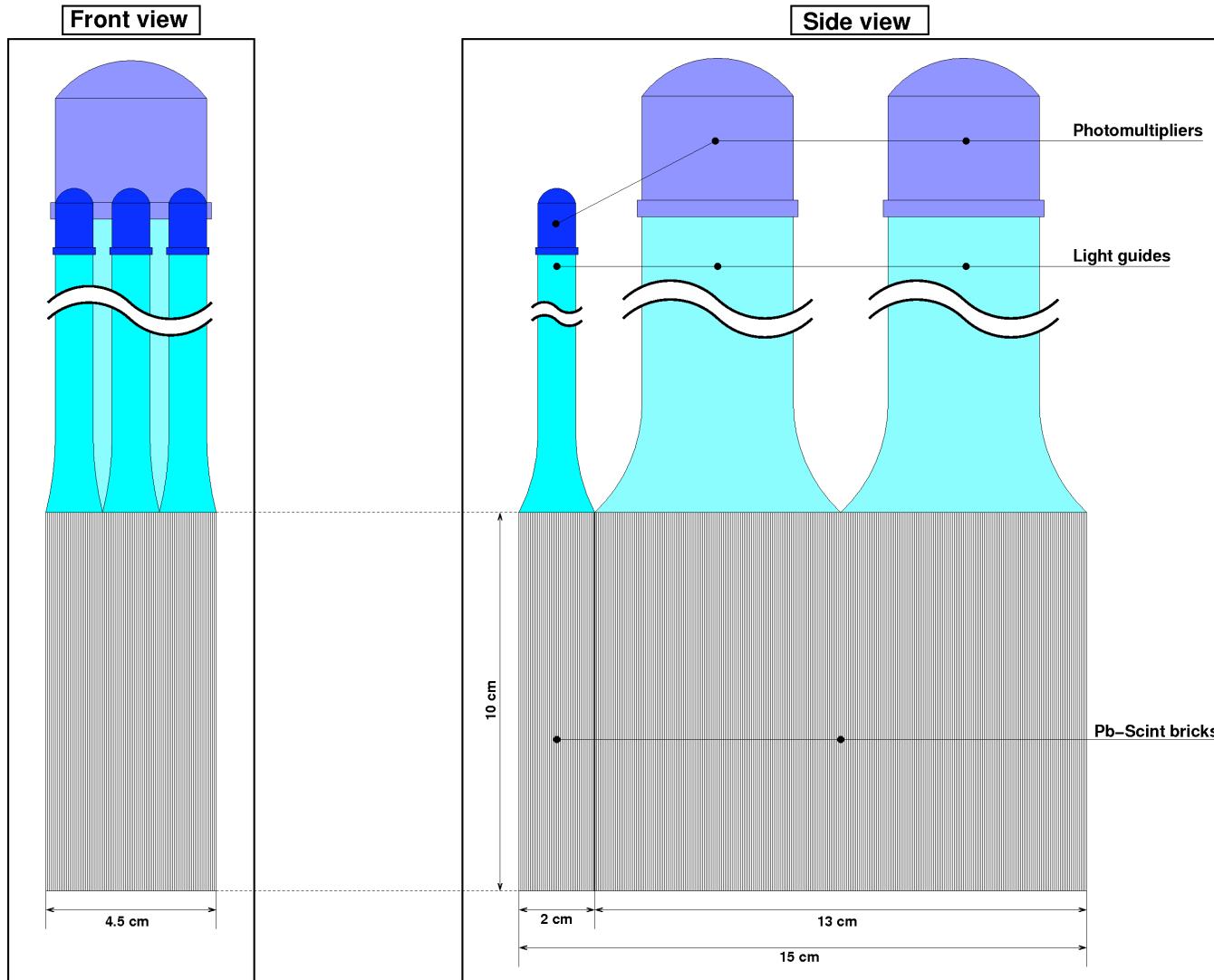


Pileup effect₄ in $T = 160$ ns

- The fit goodness is expressed as $(\sigma_E/E)_{\text{fit}} = (E_{\text{gen}} - E_{\text{fit}})/E_{\text{gen}}$, where E_{gen} is the generated ISR amplitude and E_{fit} is its fitted value
- We consider as a reference accuracies:
$$\begin{cases} 7\% \sim \frac{\sigma_E}{E} & \text{LYSO} \\ 15\% \sim \frac{\sigma_E}{E} & \text{Pb-Scint} \end{cases}$$
- $E_{\gamma_{\text{IS}}} \in [0.5 \text{ GeV}, 1.5 \text{ GeV}], \text{mild dependence on } E_{\gamma_{\text{IS}}}$

rate (kHz)	Pileup in 160 ns (%)	$(\sigma_E/E)_{\text{fit}} > 7\%$ (%)	$(\sigma_E/E)_{\text{fit}} > 15\%$ (%)
2500	30	9.4	4.8
2100	26	8.1	4.2
1000	14	4.3	2.2
800	10	3.2	1.6

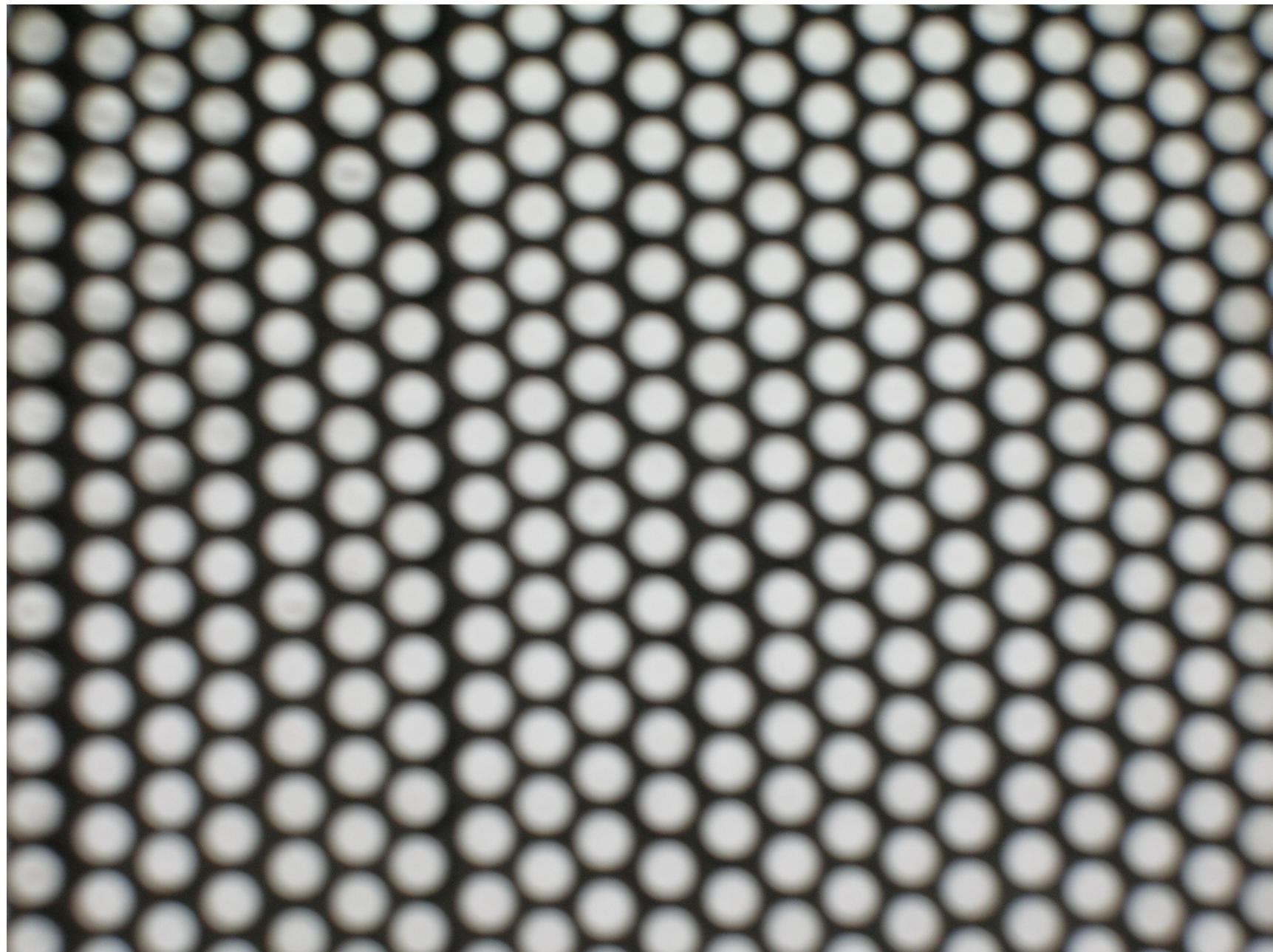
Present ZDD Design



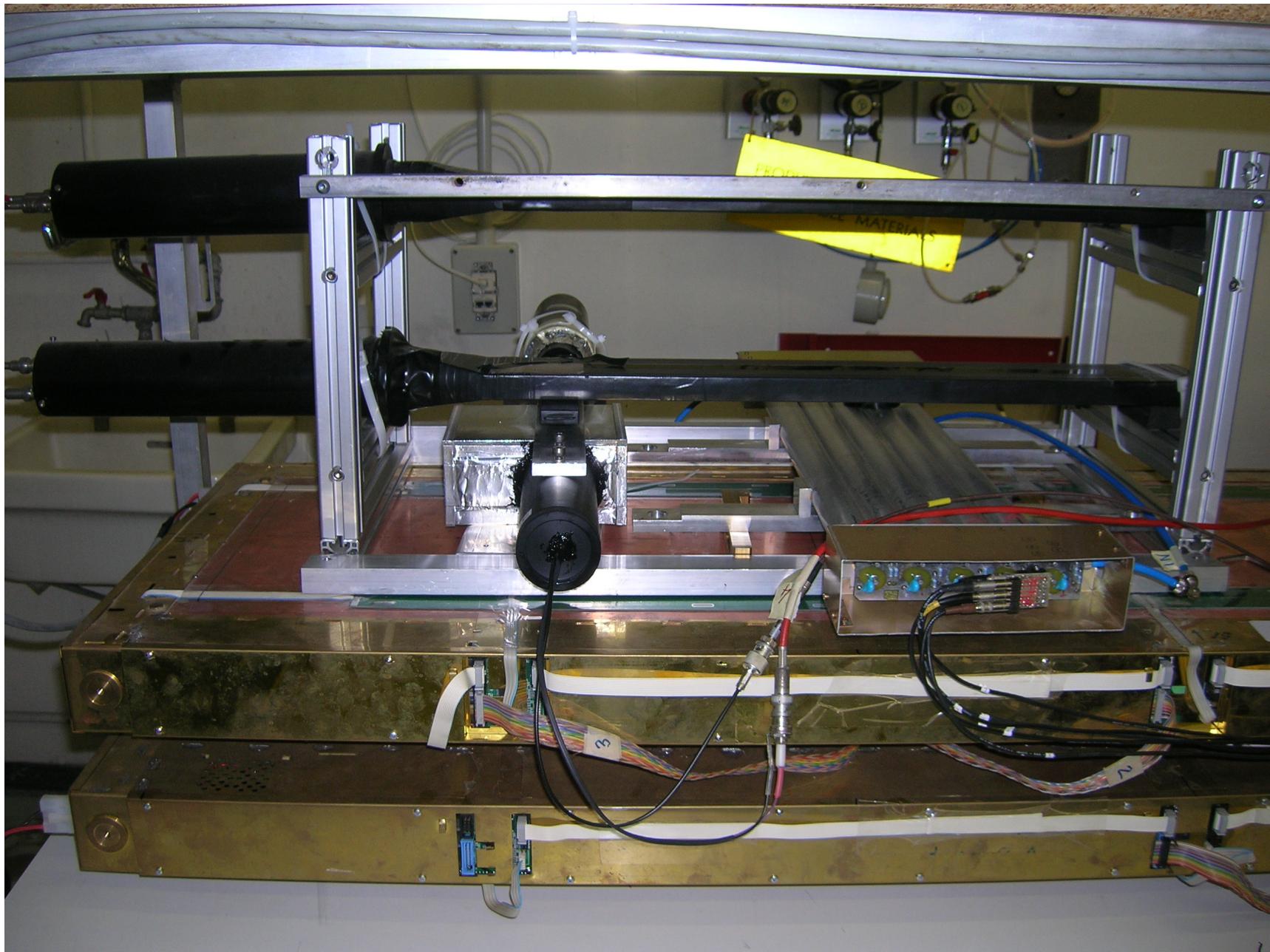
First ZDD prototype, September 2010



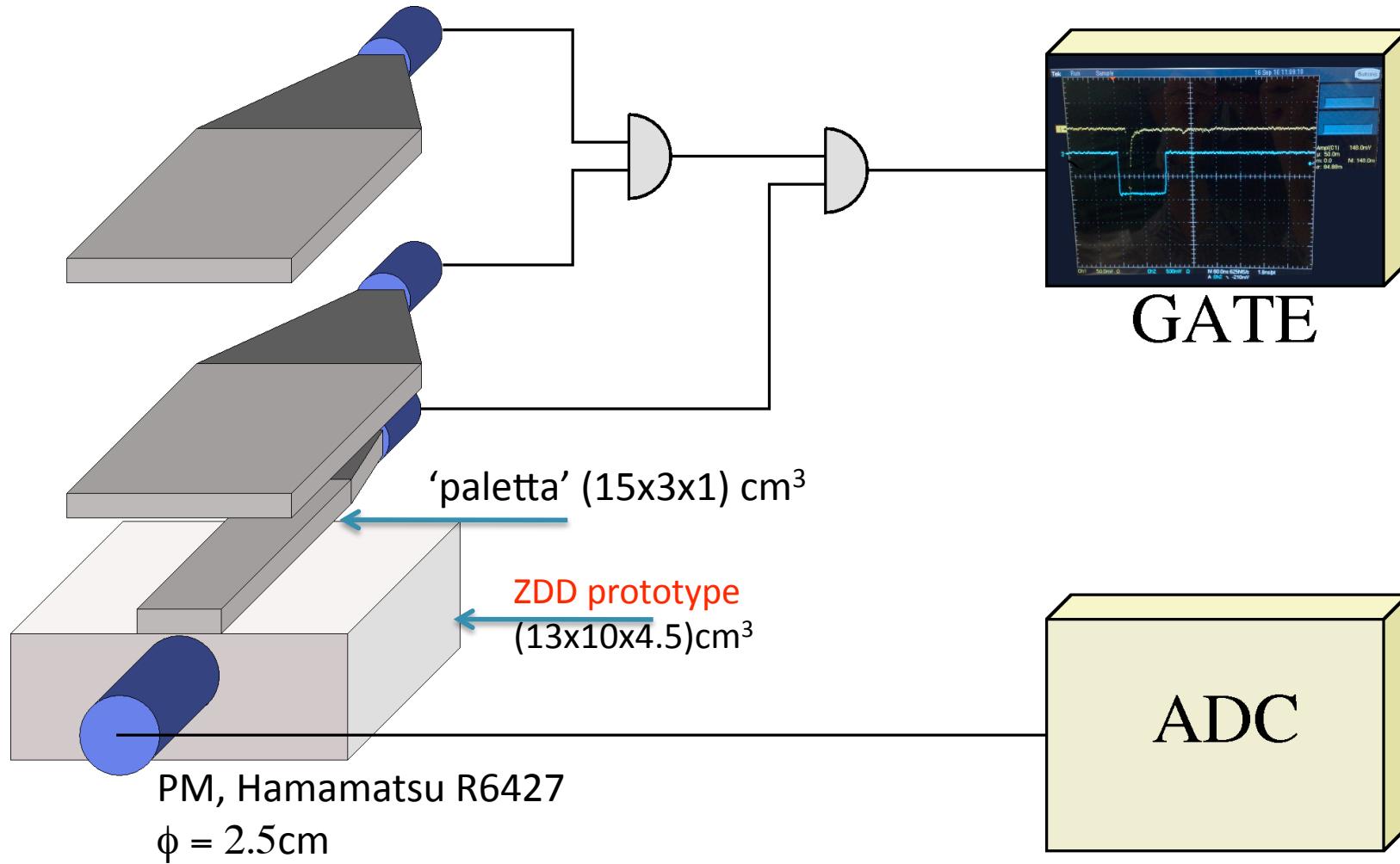
First prototype, September 2010



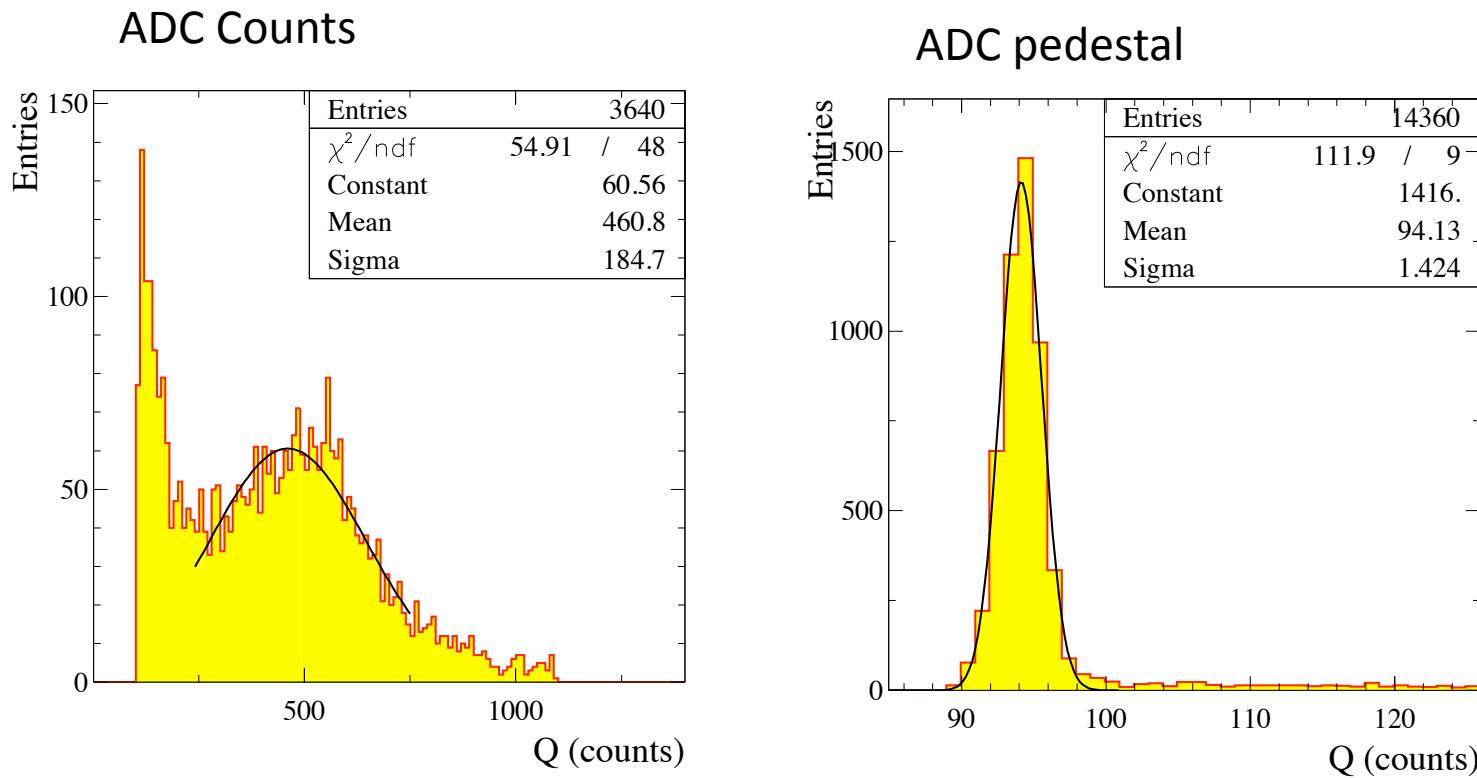
Cosmic rays test for first prototype



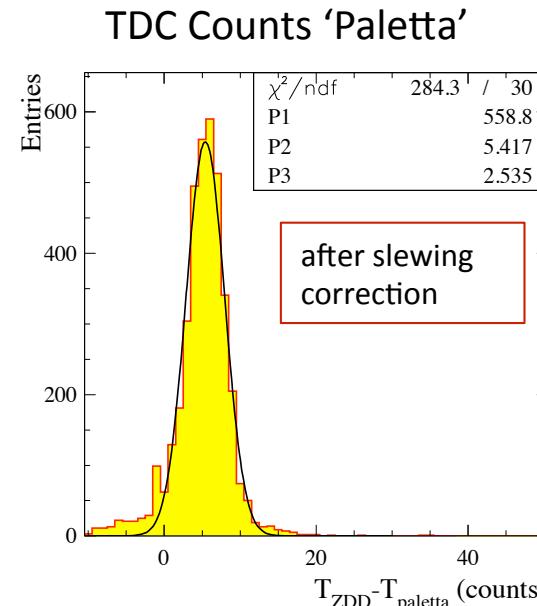
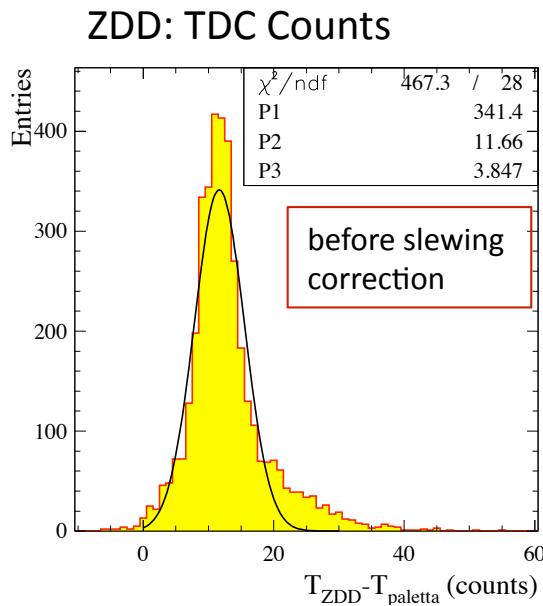
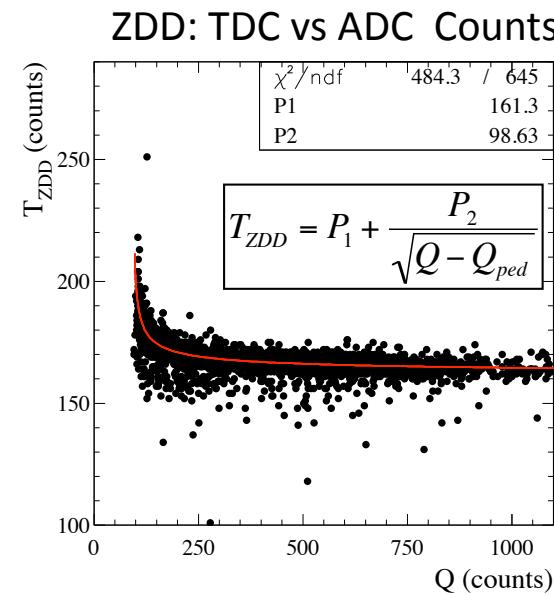
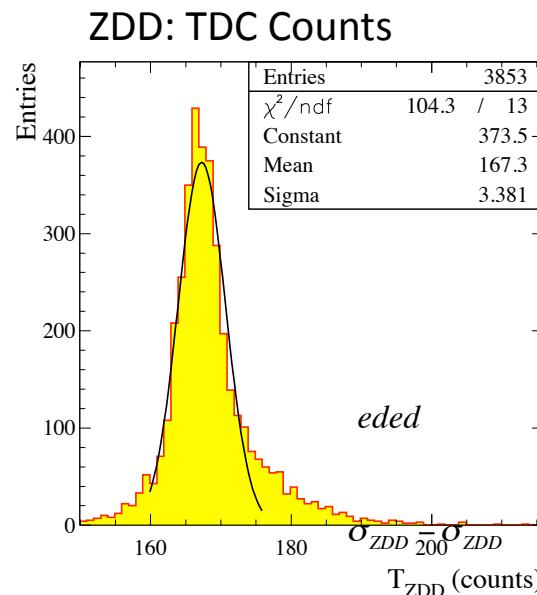
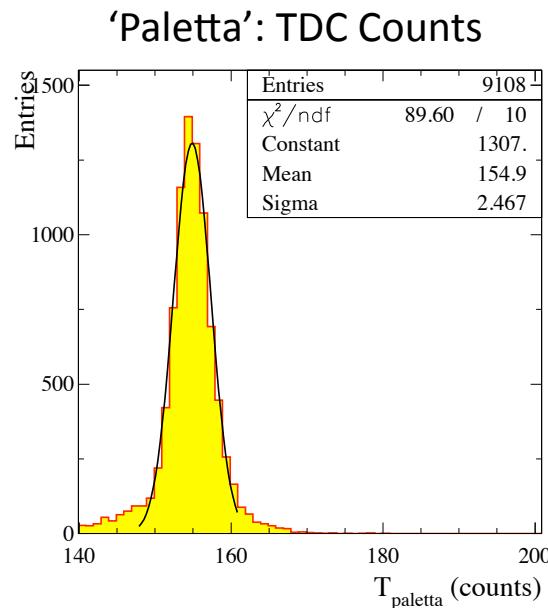
Cosmic ray setup @ LNF



Cosmic ray test: preliminary results



Cosmic ray test: preliminary results



$$\sigma_{ZDD} - \sigma_{paletta} = \frac{\tau_{ZDD}}{\sqrt{N_{p.e.}}}$$

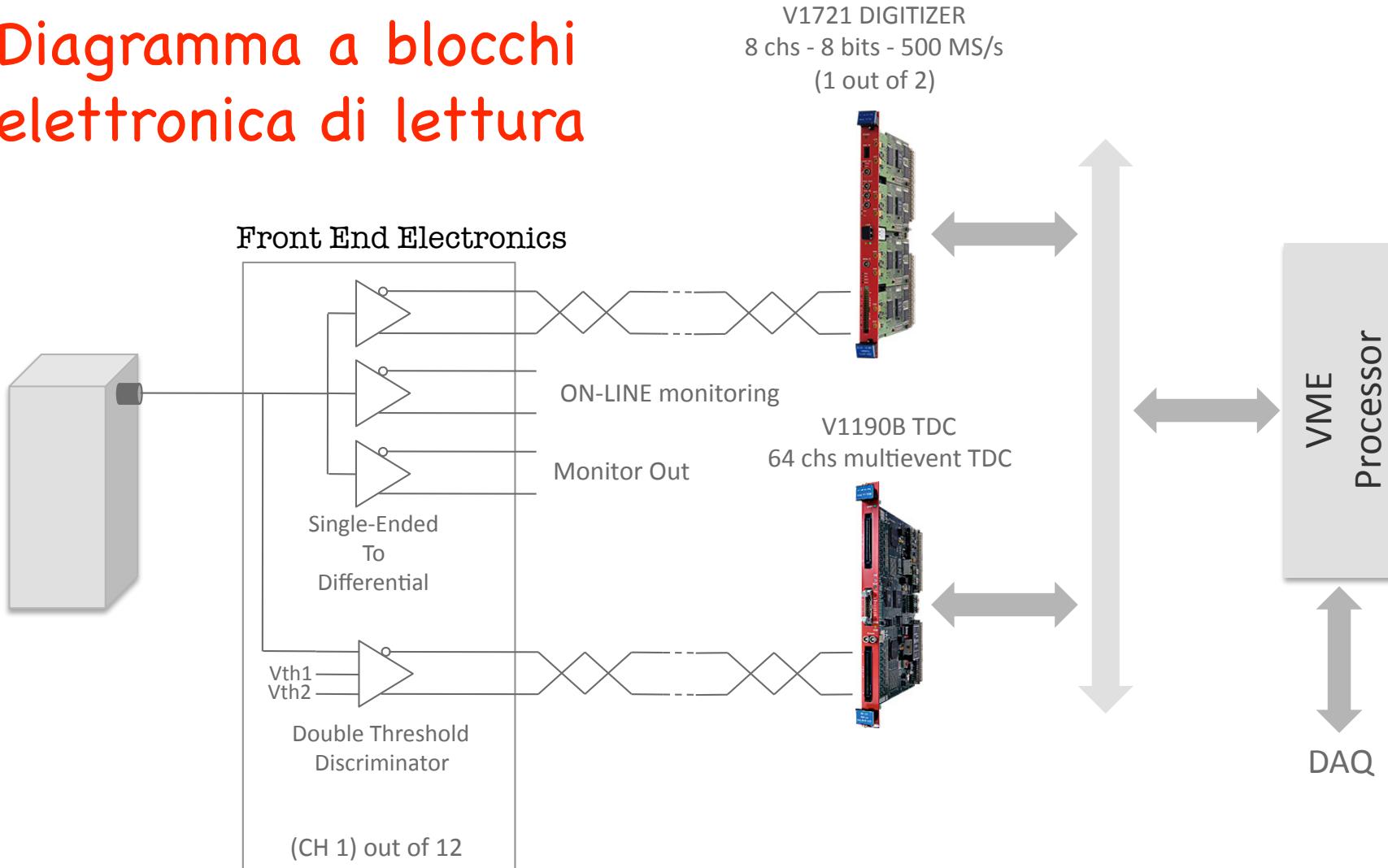
$$\tau_{ZDD} = 2.2 \text{ ns}$$

@17MeV:

$$N_{p.e.} = \left(\frac{\tau_{ZDD}}{\sigma_{ZDD} - \sigma_{paletta}} \right)^2 = 20$$

$$N_{p.e.} = 20/17 \geq 1 \text{ p.e./MeV}$$

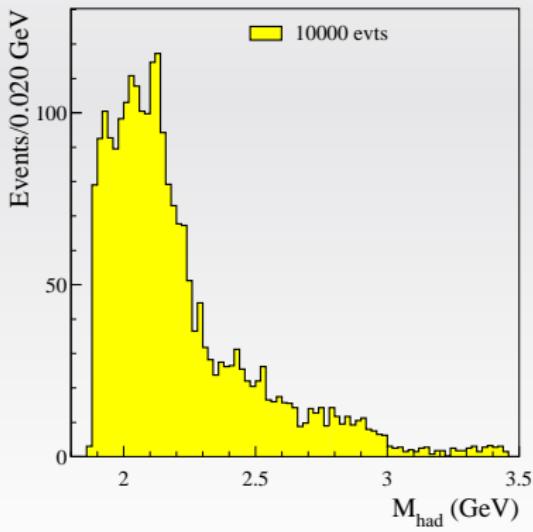
Diagramma a blocchi elettronica di lettura



Physics

The $n\bar{n}\gamma_{\text{IS}}$ physics case

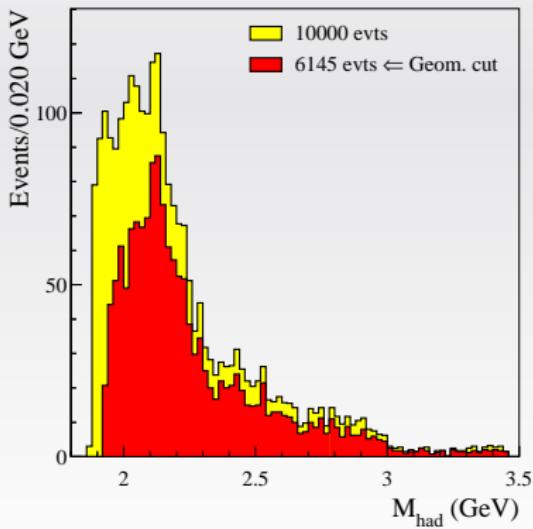
- $e^+e^- \rightarrow n\bar{n}\gamma_{\text{IS}}$ at a center of mass energy: $E_{c.m.} = 3.77 \text{ GeV}$
- Initial state photon energy range: $50 \text{ MeV} \leq E_{\gamma_{\text{IS}}} \leq \frac{E_{c.m.}}{2} \left(1 - \frac{4M_n^2}{E_{c.m.}^2}\right)$
- Beam pipe suppresses sinc. rad. bkg. and γ_{IS} with $E_{\gamma_{\text{IS}}} < 50 \text{ MeV}$
- γ_{IS} in ZDD and only antineutron detected in BESIII



- 10000 events with $\gamma_{\text{IS}} \rightarrow \text{ZDD}$
- $M_{\text{had}} = E_{c.m.} \sqrt{1 - \frac{2E_{\gamma_{\text{IS}}}}{E_{c.m.}}}$
- Geometrical cut:
 - $\bar{n} \rightarrow \text{BESIII}$
 - No constraint in n

The $n\bar{n}\gamma_{\text{IS}}$ physics case

- $e^+e^- \rightarrow n\bar{n}\gamma_{\text{IS}}$ at a center of mass energy: $E_{c.m.} = 3.77 \text{ GeV}$
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- 10000 events with $\gamma_{\text{IS}} \rightarrow \text{ZDD}$
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- Geometrical cut:
 - $\bar{n} \rightarrow \text{BESIII}$
 - No constraint in n

The $n\bar{n}\gamma_{IS}$ physics case: kinematic fit

Inputs (6)

- \bar{n} 3-momentum (TOF)
- γ_{IS} 3-momentum (ZDD)

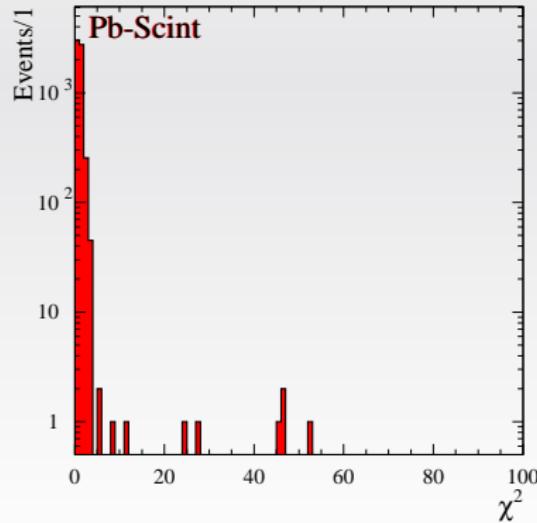
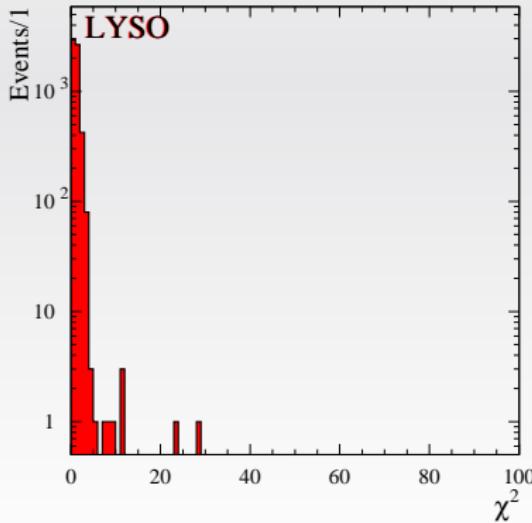
Constraints (4)

- 4-momentum cons.

Unknowns (3)

- n 4-momentum

$$\chi^2 = \sum_{\text{tracks}} \sum_i \frac{(p_i^{\text{exp}} - p_i^{\text{fit}})^2}{\sigma_{p_i}^2}$$



The $n\bar{n}\gamma_{IS}$ physics case: kinematic fit

Inputs (6)

- \bar{n} 3-momentum (TOF)
- γ_{IS} 3-momentum (ZDD)

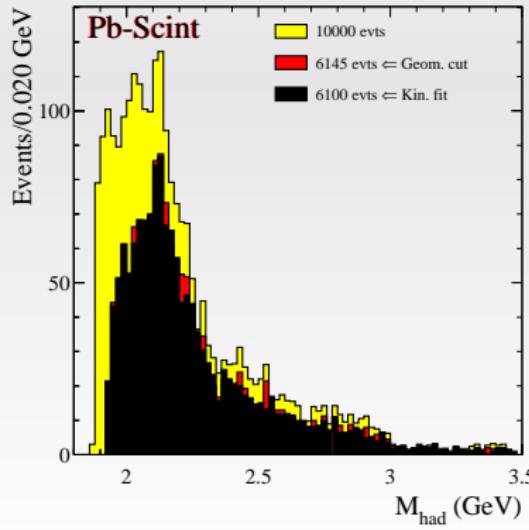
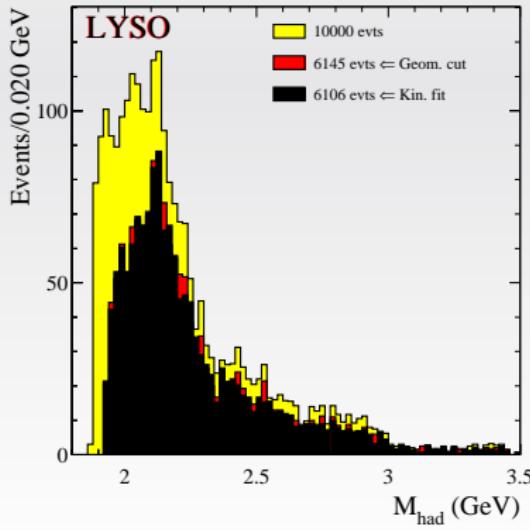
Constraints (4)

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Unknowns (3)

- n 4-momentum

$$\chi^2 = \sum_{\text{tracks}} \sum_i \frac{(p_i^{\text{exp}} - p_i^{\text{fit}})^2}{\sigma_{p_i}^2}$$



The $n\bar{n}\pi^0$ background

- $e^+e^- \rightarrow n\bar{n}\pi^0$ is one of the main backgrounds
- Assuming $\sigma(e^+e^- \rightarrow n\bar{n}\pi^0) \simeq \sigma(e^+e^- \rightarrow p\bar{p}\pi^0)$:

$$\frac{\text{Ev}(n\bar{n}\pi^0)}{\text{Ev}(n\bar{n}\gamma)} [M_{\Upsilon(4S)}] \simeq R_{BABAR} = \frac{\text{Ev}(p\bar{p}\pi^0)}{\text{Ev}(p\bar{p}\gamma)} [M_{\Upsilon(4S)}] = 0.06$$

- In **BESIII**, directly at the $\psi(3770)$ mass:

$$R_{BESIII} = 0.06 \times \underbrace{\left(\frac{0.012}{3 \times 10^{-6}} \right)}_{p\bar{p}\pi^0 \text{ cross section ratio}} \times \underbrace{\left(\frac{1}{10.7} \right)}_{\text{Lum. ratio}} = 22.4$$

$\gamma_{IS} \rightarrow ZDD$

$$\frac{\text{ZDD solid angle}}{\text{BESIII solid angle}}$$

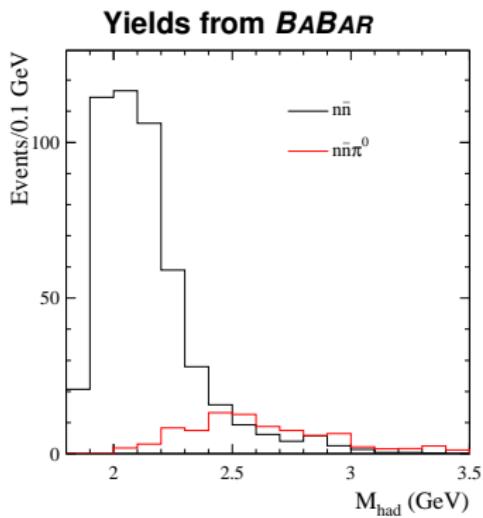
\Rightarrow

$$\frac{\text{Ev}(n\bar{n}\pi^0, \pi^0 \rightarrow 0^\circ)}{\text{Ev}(n\bar{n}\gamma, \gamma \rightarrow 0^\circ)} = 0.0008$$

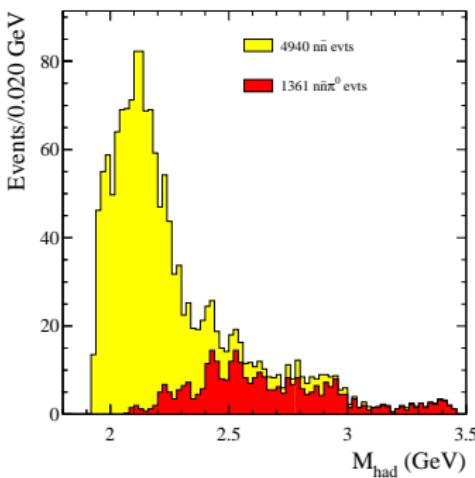
$$\frac{2 \cdot (2 \cdot 4.5 \cdot 3 / 349^2)}{4\pi \cos \theta_{\min}} = 3.8 \cdot 10^{-5}$$

The $n\bar{n}\pi^0\gamma_{IS}$ background

$e^+e^- \rightarrow n\bar{n}\pi^0\gamma_{IS}$
is a severe background
having the IS photon

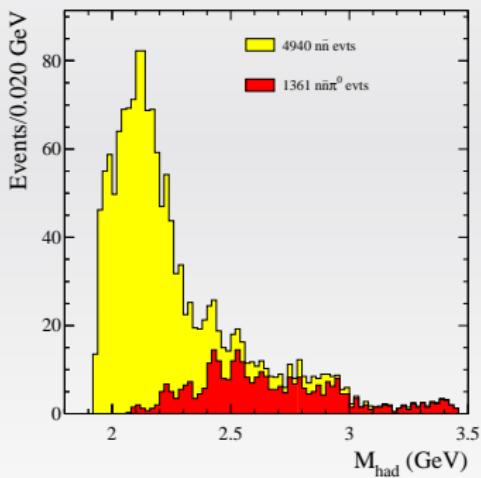


- After geometrical cut
- High contamination at high M_{had}



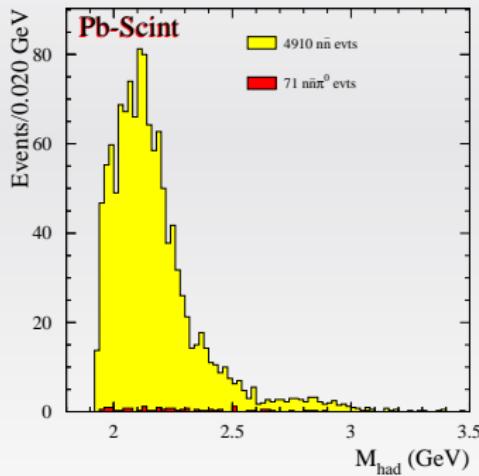
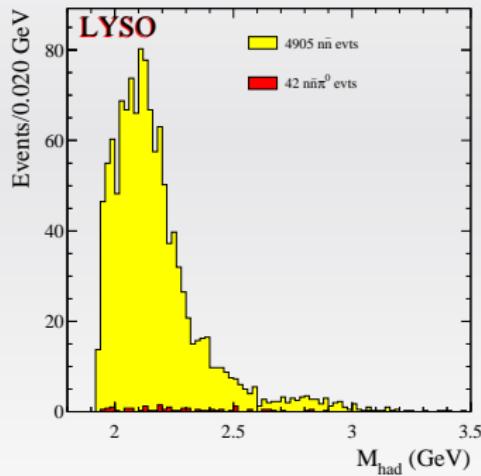
The $n\bar{n}\pi^0\gamma$ background reduction

- **π^0 detection in BESIII:** at least one of the π^0 photons with $E_\gamma > 50$ MeV in BESIII
not in a 200 mrad cone around \bar{n} direction
- **Kinematic fit:** $\chi^2 \leq 10$



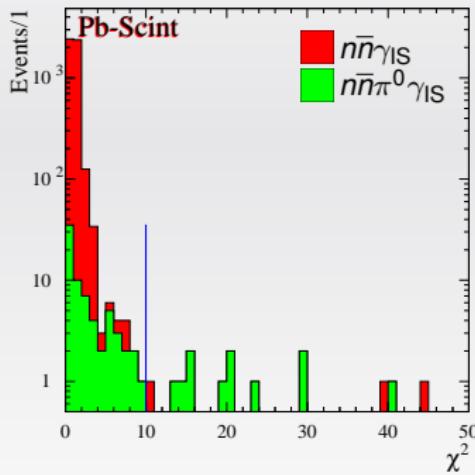
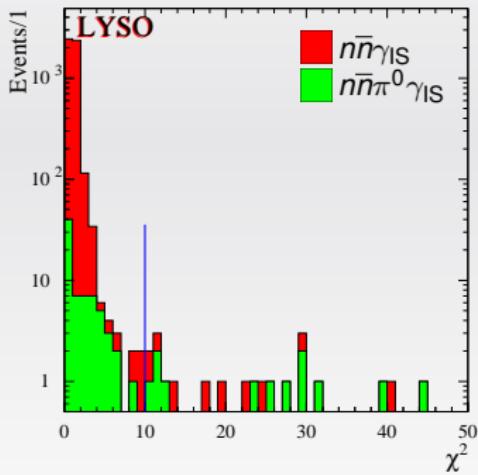
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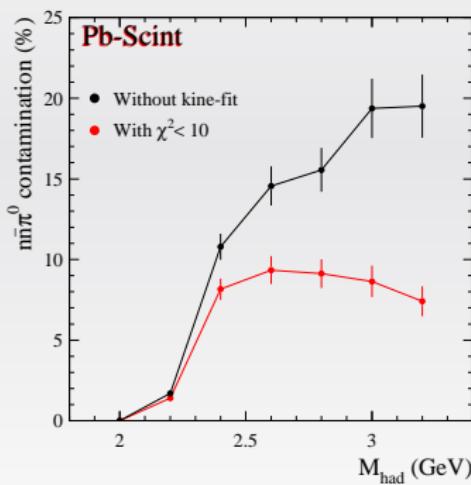
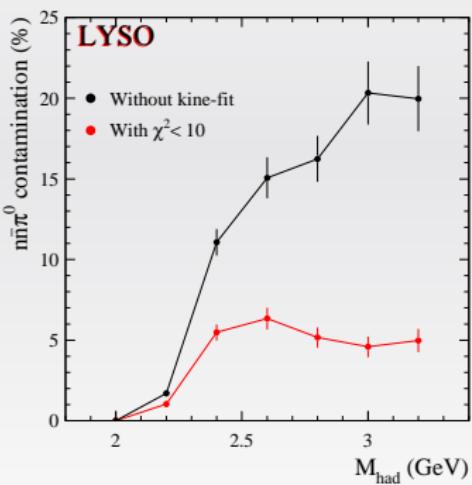
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The $n\bar{n}\pi^0\gamma_{IS}$ background reduction

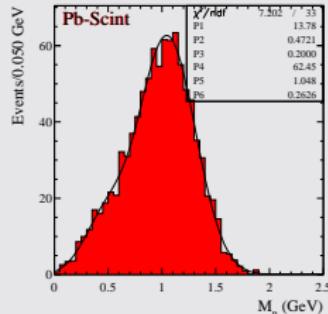
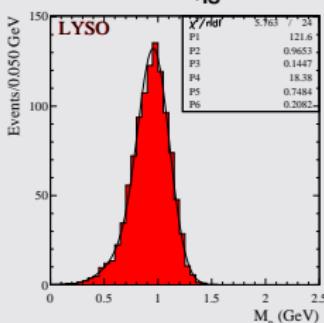
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not in a 200 mrad cone around \bar{n} direction
- **Kinematic fit:** $\chi^2 \leq 10$



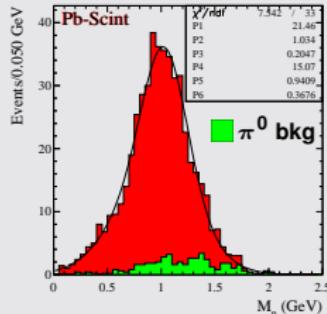
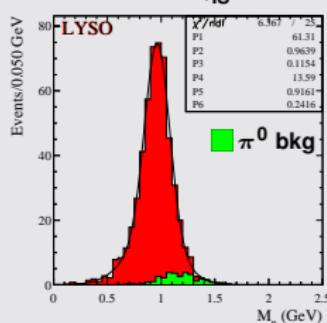
Energy resolution in $\bar{n}_{\gamma IS}$ missing mass

- Events are generated with fixed value of $M_{had} = E_{c.m.} \sqrt{1 - 2E_{\gamma IS}/E_{c.m.}}$.
- The $\bar{n}_{\gamma IS}$ missing mass is obtained only from experimental data

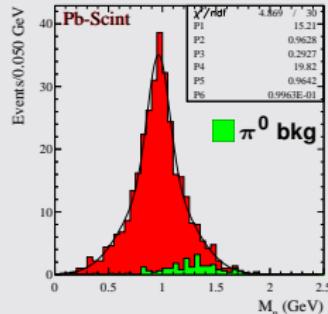
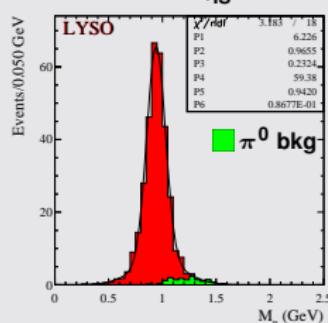
$M_{had}=2.0$ GeV - $E_{\gamma IS}=1.35$ GeV



$M_{had}=2.6$ GeV - $E_{\gamma IS}=1.0$ GeV



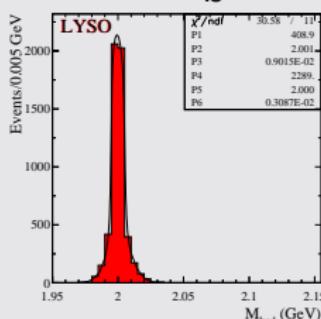
$M_{had}=3.2$ GeV - $E_{\gamma IS}=0.5$ GeV



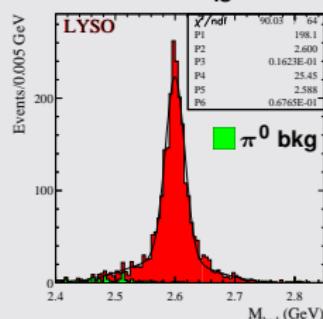
Energy resolution in M_{had} slices

- Events are generated with fixed value of $M_{\text{had}} = E_{c.m.} \sqrt{1 - 2E_{\gamma\text{IS}}/E_{c.m.}}$.
- M_{had} is reconstructed using the kinematic fit procedure

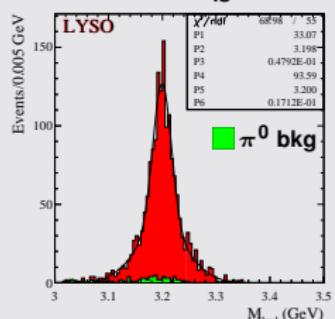
$M_{\text{had}}=2.0 \text{ GeV} - E_{\gamma\text{IS}}=1.35 \text{ GeV}$



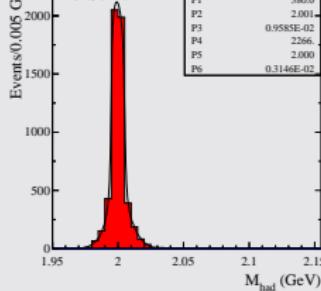
$M_{\text{had}}=2.6 \text{ GeV} - E_{\gamma\text{IS}}=1.0 \text{ GeV}$



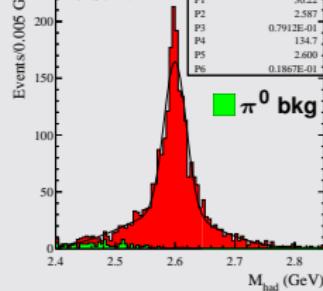
$M_{\text{had}}=3.2 \text{ GeV} - E_{\gamma\text{IS}}=0.5 \text{ GeV}$



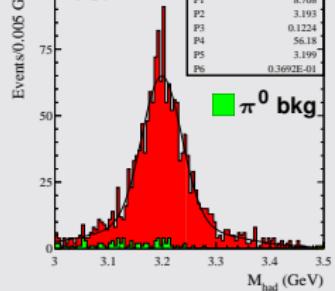
Pb-Scint



Pb-Scint



Pb-Scint

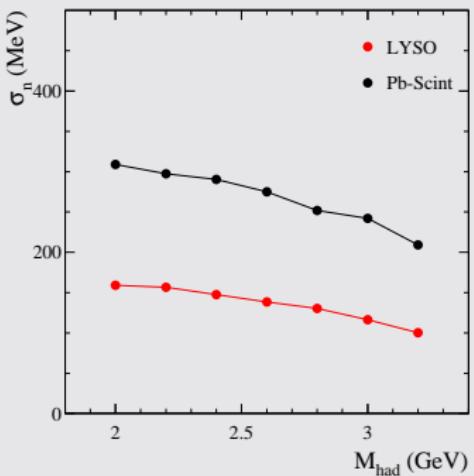


Energy resolutions

- Two-gaussian fit: $\sigma =$

half width of the area, symmetric w.r.t. the center of mass of the distributions, which contains the 68% of events

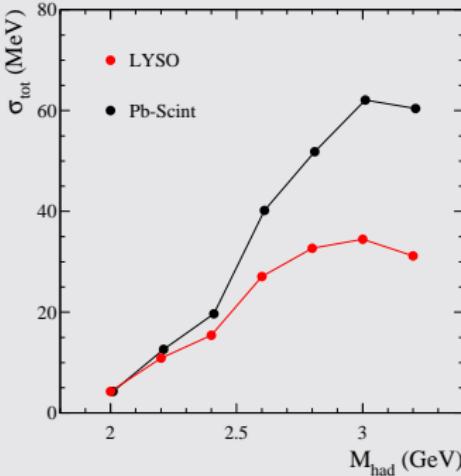
Width of the $\bar{n}\gamma_{IS}$ miss. mass



σ_n is dominated by $\delta p_{\bar{n}}$

$small M_{had} \Rightarrow large E_{\gamma_{IS}} \Rightarrow small \delta E_{\gamma_{IS}}$

Energy resolution in M_{had} bins

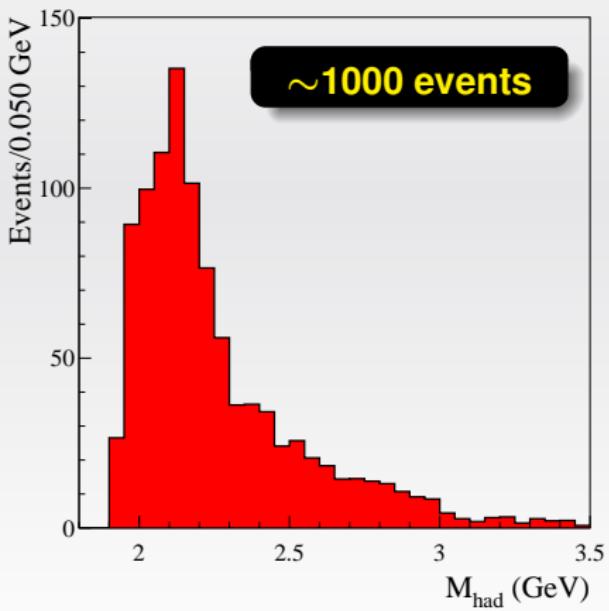


σ_{tot} is dominated by $\delta E_{\gamma_{IS}}$

$large M_{had} \Rightarrow small E_{\gamma_{IS}} \Rightarrow large \delta E_{\gamma_{IS}}$

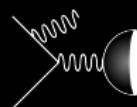
Expected events

- One year of data taking: $T = 1.5 \times 10^7 \text{ s}$
- Average luminosity: $\mathcal{L} = 1.5 \times 10^{32} \text{ cm}^{-2}\text{s}^{-1}$
- Detection efficiency: $\epsilon = 0.5$
- Center of mass energy: $E_{c.m.} = 3.77 \text{ GeV}$

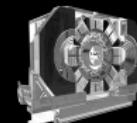


Other possible physics items R_{had} in the 1-3 GeV region

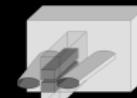
- Accessible had-CoM energy: $E_{\text{had}} = \sqrt{E_{\text{coll}}^2 - 2E_{\text{coll}}E_{\gamma\text{IS}}}$
- PDG: $\gamma\gamma 2$ and BESII (2-3 GeV) only
- ISR: small systematic error versus E_{had}
- ISR on ZDD: negligible π^0 background



- $|\Delta E_{\text{had}}| = |\Delta E_{\gamma\text{IS}}| E_{\text{coll}} / E_{\text{had}}$:
feasible only if $E_{\text{coll}} / E_{\text{had}} \sim 1$ (not for B -factories)
- BESIII: $E_{\text{coll}} \sim 3.5 \text{ GeV} \Rightarrow E_{\text{had}} \simeq 1 - 3 \text{ GeV}$



- LYSO: $|\Delta E_{\text{had}}| \simeq 150 \text{ MeV}$
- Pb-Scint: $|\Delta E_{\text{had}}| \simeq 300 \text{ MeV}$
- $|\Delta E_{\text{had}}|$ reduced by deconvolution techniques



Radiation hardness

- Radiation damages mostly due to Bremsstrahlung: $\sigma_{\text{Bre}}(\text{ZDD}/4) = 2.6 \text{ mb}$
- One year of data taking: $T = 1.5 \times 10^7 \text{ s}$
- Average luminosity: $\bar{\mathcal{L}} = 1.5 \times 10^{32} \text{ cm}^{-2}\text{s}^{-1}$
- Center of mass energy: $E_{c.m.} = 3.77 \text{ GeV}$

$$\frac{\text{Dose absorbed}}{\text{year}} = \frac{\text{energy deposited}}{\text{year} \cdot \text{mass}} = \begin{cases} \frac{3 \times 10^{21} \text{ eV}}{0.12 \text{ kg}} = 4 \times 10^5 \frac{\text{rad}}{\text{year}} & \text{LYSO} \\ \frac{3 \times 10^{21} \cdot \frac{2}{13} \text{ eV}}{1.8 \times 10^{-2} \text{ kg}} = 10^6 \frac{\text{rad}}{\text{year}} & \text{Scint} \end{cases}$$

Declared hardness

- LYSO $\sim 10^8 \text{ rad}$
- Scint. $\sim 10^6 \text{ rad}$