# A proposal to tag ISR events at B E S III

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## **BESIII 2010 Spring Collaboration Meeting**



Liaoning University, Liaoning, June 3-7, 2010

# **Motivations**

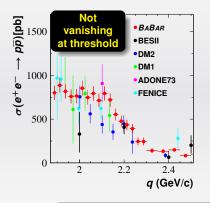


$$e^+e^- o p\overline{p}$$

$$\sigma(e^+e^-\!\to p\overline{p}) = \frac{4\pi\alpha^2\beta_p \frac{\textbf{C}}{\textbf{C}}}{3q^2} \left[ |\textbf{G}_{\textbf{M}}|^2 + \frac{2\textit{M}_p^2}{q^2} |\textbf{G}_{\textbf{E}}|^2 \right]$$

#### Coulomb factor

$$C_{eta_p o 0} rac{\pi lpha}{eta_p}$$



#### At the threshold

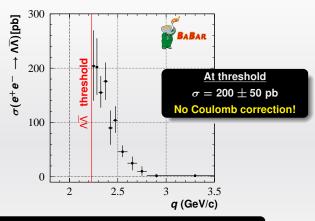
$$\sigma(e^+e^- \to p\overline{p})(4M_p^2) = \frac{\pi^2 \alpha^3}{2M_p^2} \frac{\beta_p}{\beta_p^2} |G^p(4M_p^2)|^2$$

$$\sigma(e^+e^- \to p\overline{p})(4M_p^2) = 850 |G^p(4M_p^2)|^2 \text{ pb}$$

 $|G^p(4M_p^2)| \equiv 1$ as pointlike fermion pairs!

Using the ISR technique with only few fb<sup>-1</sup> of integrated luminosity BESIII can easily achieve the BABAR statistics

PRD76, 092006 (BABAR)

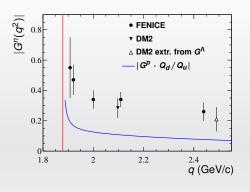


Only at the  $J/\psi$  mass BESIII can increase the BABAR statistics at least by a factor of two because of a better A reconstruction resolution (only one A reconstructed)

 $\Lambda$  polarization for free  $\Rightarrow G_F^{\Lambda} - G_M^{\Lambda}$  relative phase



#### $e^+e^-{ ightarrow} n\overline{n}$



- Measured only once by **FENICE at ADONE**
- $\mathcal{L} = 500 \text{ nb}^{-1} \text{ (15' at BESIII)}$
- $\sim$  100 candidates  $n\overline{n}$  events!
- $\sigma(n\overline{n}) > \sigma(p\overline{p})$ ?
- Not zero at threshold?

BESIII has the unique possibility to measure this cross section

No other experiments at present and in near future will be able to perform such a measurement





# Why Initial State Radiation



# ISR: Physics Motivations

 Existing results, obtained by BABAR (ISR), show interesting and unexpected behaviors, mainly at thresholds, for

$$e^+e^- o p\overline{p}$$
 and  $e^+e^- o \Lambda\overline{\Lambda}$ 

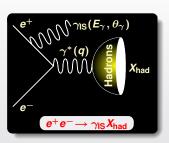
Only one measurement (FENICE with energy scan) for

$$e^+e^- o n\overline{n}$$

There are physical limits in reaching the threshold of many of these channels via energy scan (stable hadrons produced at rest can not be detected)

The Initial State Radiation technique provides a unique tool to access threshold regions working at higher resonances

#### Initial State Radiation



$$\frac{d^2\sigma}{dE_{\gamma}d\theta_{\gamma}} = W(E_{\gamma}, \theta_{\gamma}) \cdot \sigma_{e^+e^- \to X_{had}}(s)$$

$$W(E_{\gamma}, \theta_{\gamma}) = \frac{\alpha}{\pi x} \left( \frac{2 - 2x + x^2}{\sin^2 \theta_{\gamma}} \right)$$

$$W(E_{\gamma}, \theta_{\gamma}) = \frac{\alpha}{\pi x} \left( \frac{2 - 2x + x^2}{\sin^2 \theta_{\gamma}} \right)$$

- $s = q^2, q \dots X_{had}$  momentum
- $E_{\gamma}$ ,  $\theta_{\gamma}$ ...CM  $\gamma_{IS}$  energy, scatt. ang.
- $E_{\text{CM}} \dots \dots \dots \dots \text{CM } e^+e^-$  energy
- $x = E_{\gamma}/2E_{\rm CM}$

#### **Advantages**

- All energies (q²) at the same time Better control on systematics
  - (e.g. greatly reduced point to point)
- $\bigcirc$  Detected ISR  $\Rightarrow$  full  $X_{had}$  angular coverage



# Why ISR at zero degrees



# Proposal for a zero-degree detector

- $J/\psi$ ,  $\psi(2S)$ ,  $\psi(3770)$  resonances decay with high BR's to final states with  $\pi^0$  and  $\gamma_{\rm FS}$  (final state)
- At BESIII these decay channels represent severe backgrounds for typical ISR final states with  $\gamma_{\rm IS}$  detected at wide angle
  - $\bullet \ \pi^0$  and final  $\gamma$  angular distributions are isotropic
  - ISR angular distribution is peaked at small angles



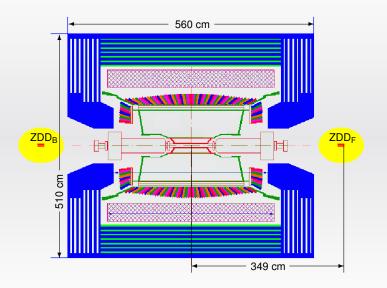
A zero-degree radiative photon tagger will suppress most of these backgrounds

We propose to upgrade (July 2011?) the present luminosity monitor with a new zero-degree detector (ZDD), with a better energy resolution, to tag ISR photons as well as to measure the luminosity

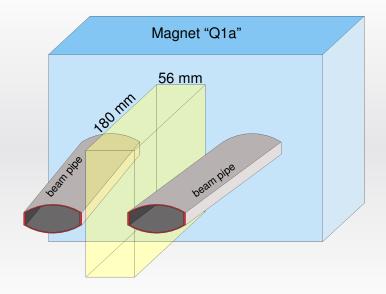
# Design and Installation two options: LYSO and Pb-Scint



## BESIII and ZDD

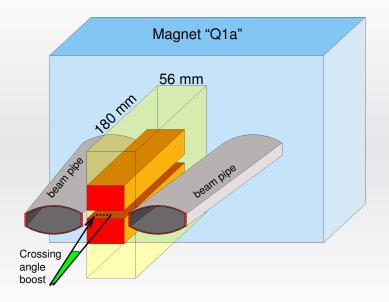


# Available space





## Available space



## LYSO design



#### Side view

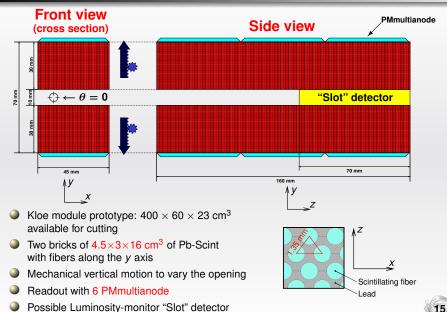


- Two  $3\times2$  matrices of  $1.5\times1.5\times16$  cm<sup>3</sup> of LYSO bars
- Mechanical vertical motion to vary the opening
- Total volume 432 cm<sup>3</sup>
- Readout with 2 PMmultianode
- Possible Luminosity-monitor "Slot" detector in the last 7 cm Identical structure but reduced size w.r.t. the actual one



**PMmultianode** 

# Pb-Scintillating fibers design à la Kloe



## Physical properties of materials

Material	LYSO	Pb-Scint (Kloe prototype)
Density (g/cm <sup>3</sup> )	7.4	5.3
Radiation Length (cm)	1.1	1.6
Molière Radius (cm)	1.9	2.9
Decay Constant (ns)	40-44	2.4
Peak Emission (nm)	428	460
Radiation Hardness (rad)	$\sim 10^8$	$\sim 10^6$

$$T = 1.5 \times 10^7 \text{ s}$$

$$\overline{\mathcal{L}} = 1.5 \times 10^{32} \text{ cm}^{-2} \text{s}^{-1}$$

$$\bullet$$
  $E_{c.m.} = 3.77 \text{ GeV}$ 

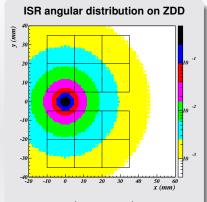
$$\Rightarrow$$

$$\frac{\text{dose absorbed}}{\text{year}} = \begin{cases} 0.4 \frac{\text{Mrad}}{\text{year}} & \text{LYSO} \\ \\ 1 \frac{\text{Mrad}}{\text{year}} & \text{Scint} \end{cases}$$

# **Energy Resolution**

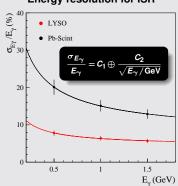


## Energy resolution, the ISR case



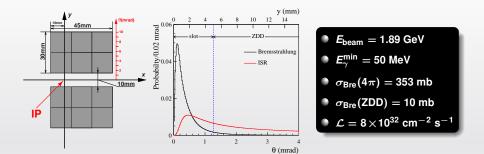
	LYSO	Pb-Scint
$ extbf{\emph{E}}_{\gamma}$ (GeV)	$\sigma_{ extbf{ extit{E}}_{\gamma}}/ extbf{ extit{E}}_{\gamma}$	$\sigma_{ extsf{ extsf{E}}_{\gamma}}/ extsf{ extsf{E}}_{\gamma}$
1.5	5.7%	12.9 %
1.0	6.4%	15.1 %
0.5	7.8%	20.1 %

#### **Energy resolution for ISR**



	LYSO	Pb-Scint
C <sub>1</sub>	4.3%	6.9 %
<b>C</b> <sub>2</sub>	4.6%	13.4 %

# $e^+e^- \rightarrow e^+e^- \gamma$ Bremsstrahlung simulation



- ISR in ZDD 13.7% of total solid angle
- Bremsstrahlung in ZDD 2.8% of total solid angle
- Bremsstrahlung rate in a single ZDD element (upper or lower):

800 kHz at 
$$\mathcal{L}=3\times 10^{32}~\text{cm}^{-2}~\text{s}^{-1}$$
 2.1 MHz at  $\mathcal{L}=8\times 10^{32}~\text{cm}^{-2}~\text{s}^{-1}$ 

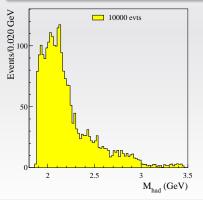


# $n\overline{n}$ physics case



# The $n\overline{n}\gamma_{\rm IS}$ physics case

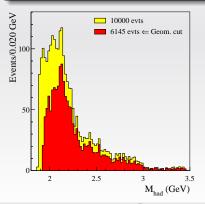
- $e^+e^- \rightarrow n\overline{n}\gamma_{1S}$  at a center of mass energy:  $E_{c,m} = 3.77$  GeV
- IS photon energy range: 50 MeV  $\leq E_{\gamma_{\rm IS}} \leq (E_{c.m.}/2) \left(1 4M_{\rm n}^2/E_{\rm c.m.}^2\right)$
- Beam pipe suppresses sinc. rad. bkg. and  $\gamma_{\rm IS}$  with  $E_{\gamma_{\rm IS}} < 50$  MeV
- $\gamma_{\rm IS}$  in ZDD and only antineutron detected in BESIII
- $\overline{n}$  annihilates in the scintillator with probability  $\sim 100\%$
- $\overline{n}$  annihilation star detected in TOF ( $\Delta t_{TOF} = 150 \text{ ps}$ )



- 10000 events with  $\gamma_{\rm IS} \rightarrow {\sf ZDD}$
- $M_{had} = E_{c.m.} \sqrt{1 2E_{\gamma_{IS}}/E_{c.m.}}$
- - No constraint in n

# The $n\overline{n}\gamma_{IS}$ physics case

- $e^+e^- \rightarrow n\overline{n}\gamma_{\rm IS}$  at a center of mass energy:  $E_{c.m.}=3.77~{\rm GeV}$
- IS photon energy range:  $50\,\mathrm{MeV} \le E_{\gamma_\mathrm{IS}} \le (E_{c.m.}/2) \left(1 4M_n^2/E_{c.m.}^2\right)$
- lacktriangle Beam pipe suppresses sinc. rad. bkg. and  $\gamma_{
  m IS}$  with  $E_{\gamma_{
  m IS}} <$  50 MeV
- $lacktriangleq \gamma_{
  m IS}$  in ZDD and  $\underline{
  m only}$  antineutron detected in BESIII
- $\overline{n}$  annihilates in the scintillator with probability $\sim$  100%
- $\overline{n}$  annihilation star detected in TOF ( $\Delta t_{TOF} = 150 \text{ ps}$ )



- lacksquare 10000 events with  $\gamma_{\mathsf{IS}} o \mathsf{ZDD}$
- Geometrical cut:
  - $\bullet$   $\overline{n} \rightarrow \mathsf{BESIII}$
  - No constraint in n

# $n\overline{n}\pi^0$ and $n\overline{n}\gamma_{\mathsf{FS}}$ backgrounds

- $e^+e^- \to n\overline{n}\pi^0$  and  $e^+e^- \to n\overline{n}\gamma_{FS}$  (with a final state photon) are important backgrounds for  $e^+e^- \to n\overline{n}\gamma_{IS}$
- lacksquare Assuming  $\sigma(e^+e^- o n\overline{n}\pi^0)\simeq\sigma(e^+e^- o p\overline{p}\pi^0)$ :

$$\frac{\mathrm{Ev}(n\overline{n}\pi^0)}{\mathrm{Ev}(n\overline{n}\gamma_{\mathrm{IS}})}\left[\mathit{M}_{\Upsilon(4S)}\right] \simeq \mathit{R}_{\mathrm{BABAR}} = \frac{\mathrm{Ev}(p\overline{p}\pi^0)}{\mathrm{Ev}(p\overline{p}\gamma_{\mathrm{IS}})}\left[\mathit{M}_{\Upsilon(4S)}\right] = 0.06$$

In **BESIII**, directly at the  $\psi$ (3770) mass:

$$R_{\text{BESIII}} = 0.06 \times \underbrace{\left( \frac{0.012}{3 \times 10^{-6}} \right)}_{p\bar{p}\pi^0 \text{ cross section ratio}} \times \underbrace{\left( \frac{1}{10.7} \right)}_{\text{Lum. ratio}} = 22.4$$

$$\gamma_{\text{IS}} \rightarrow \text{ZDD}$$

$$ZDD \text{ solid angle}$$

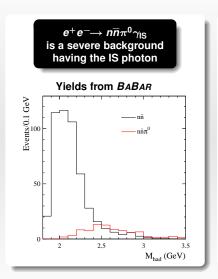
$$BESIII \text{ solid angle}$$

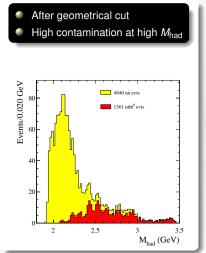
$$\frac{2 \cdot (2 \cdot 4.5 \cdot 3/349^2)}{4\pi \cos \theta_{\text{min}}} = 3.8 \cdot 10^{-5}$$

$$\frac{\mathsf{Ev}(n\overline{n}\pi^0,\ \pi^0 \rightarrow 0^o)}{\mathsf{Ev}(n\overline{n}\gamma_{\mathsf{IS}},\ \gamma_{\mathsf{IS}} \rightarrow 0^o)} = 0.0008$$

$$egin{aligned} rac{\mathsf{Ev}(n\overline{n}\gamma_{\mathsf{FS}},\gamma_{\mathsf{FS}}{
ightarrow}0^o)}{\mathsf{Ev}(n\overline{n}\gamma_{\mathsf{IS}},\gamma_{\mathsf{IS}}{
ightarrow}0^o)} \sim 0.0001 \end{aligned}$$

# The $n\overline{n}\pi^0\gamma_{\rm IS}$ background

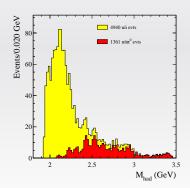




# The $n\overline{n}\pi^0\gamma_{\rm IS}$ background reduction

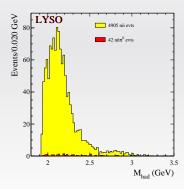
- Geometrical cut on  $\overline{n}$

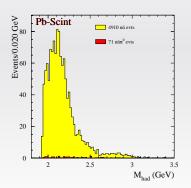
$$\chi^2 \le 10$$



# The $n\overline{n}\pi^0\gamma_{\rm IS}$ background reduction

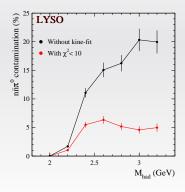
- Geometrical cut on  $\overline{n}$
- $\pi^0$  detection in BESIII: at least one of the  $\pi^0$  photons with  $E_{\gamma}>50$  MeV in BESIII not in a 200 mrad cone around  $\bar{n}$  direction
- Kinematic fit:  $\chi^2 < 10$

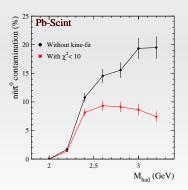




# The $n\overline{n}\pi^0\gamma_{\rm IS}$ background reduction

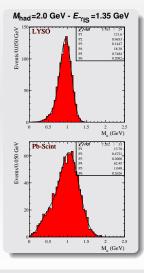
- Geometrical cut on  $\overline{n}$
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- $\chi^2 < 10$ Kinematic fit:

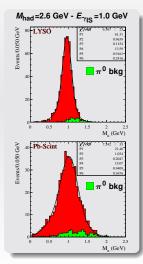


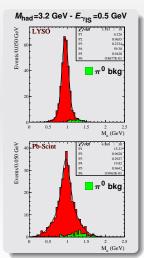


# Energy resolution in $\overline{n}\gamma_{1S}$ missing mass

- Events are generated with fixed value of  $M_{had} = E_{c.m.} \sqrt{1 2E_{\gamma_{|S}}/E_{c.m.}}$
- The  $\overline{n}\gamma_{\rm IS}$  missing mass is obtained only from experimental data

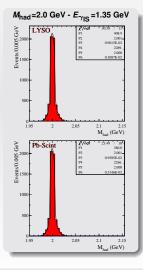


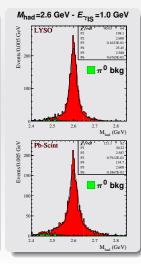


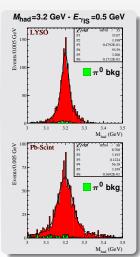


# Energy resolution in $M_{had}$ slices

- Events are generated with fixed value of  $M_{had} = E_{c.m.} \sqrt{1 2E_{\gamma_{|S}}/E_{c.m.}}$
- $M_{\rm had}$  is reconstructed using the kinematic fit procedure



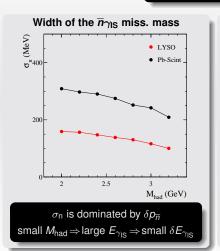


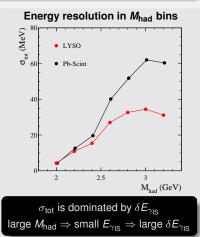


# Energy resolutions in summary

• Two-gaussian fit:  $\sigma$  =

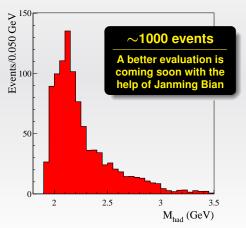
half width of the area, symmetric w.r.t. the center of mass of the distributions, which contains the 68% of events





### Expected events

- One year of data taking:  $T = 1.5 \times 10^7 \text{ s}$
- Average luminosity:  $\overline{\mathcal{L}} = 1.5 \times 10^{32} \text{ cm}^{-2} \text{s}^{-1}$ 
  - Detection efficiency:  $\epsilon \sim 0.5$
- Center of mass energy:  $E_{c.m.} = 3.77 \text{ GeV}$



# Other possible physics items $R_{\rm had}$ in the 1-3 GeV region

- Accessible had-CoM energy:  $M_{had} = \sqrt{E_{coll}^2 2E_{coll}E_{\gamma_{iS}}}$
- PDG:  $\gamma\gamma$ 2 and BESII (2-3 GeV) only
- ISR: small systematic error versus  $M_{had}$
- ISR on ZDD: negligible  $\pi^0$  background



 $|\Delta M_{\rm had}| = |\Delta E_{\gamma_{\rm IS}}|E_{\rm coll}/M_{\rm had}$ : feasible only if  $E_{\rm coll}/M_{\rm had}\sim 1$  (not for *B*-factories)



BESIII:  $E_{\text{coll}} \sim 3.5 \text{ GeV} \Rightarrow M_{\text{had}} \simeq 1 - 3 \text{ GeV}$ 

- LYSO:  $|\Delta M_{\rm had}| \simeq 150 \ {\rm MeV}$
- Pb-Scint:  $|\Delta M_{\rm had}| \simeq 300 \ {\rm MeV}$
- $|\Delta M_{\text{had}}|$  reduced by deconvolution techniques



## To sum up

We have presented the two-options proposal to INFN comittee for HEP, chaired by F. Ferroni:

- approved proposal with priority in the low cost Kloe option
- cut and test the Kloe-prototype funded
- in July we can ask for readout electronics budget
- We propose to install it in summer 2011 shutdown
- In the following years the crystals (LYSO?) option can be put ahead and eventually be installed during a following shutdown

## To sum up

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# 感谢您的关注!



# **Additional slides**



#### Radiation hardness

- Radiation damages mostly due to Bremsstrahlung:
- $\sigma_{\rm Bre}({\rm ZDD}/4)=2.6~{\rm mb}$

One year of data taking:

$$T = 1.5 \times 10^7 \, \mathrm{s}$$

Average luminosity:

$$\overline{\mathcal{L}} = 1.5 \times 10^{32}~\text{cm}^{-2}\text{s}^{-1}$$

Center of mass energy:

$$E_{c.m.} = 3.77 \text{ GeV}$$

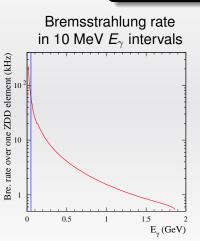
$$\frac{\text{Dose absorbed}}{\text{year}} = \frac{\text{energy deposited}}{\text{year} \cdot \text{mass}} = \begin{cases} &\frac{3 \times 10^{21} \text{ eV}}{0.12 \text{ kg}} = 4 \times 10^5 \frac{\text{rad}}{\text{year}} & \text{LYSO} \\ \\ &\frac{3 \times 10^{21} \cdot \frac{2}{13} \text{ eV}}{1.8 \times 10^{-2} \text{ kg}} = 10^6 \frac{\text{rad}}{\text{year}} & \text{Scint} \end{cases}$$

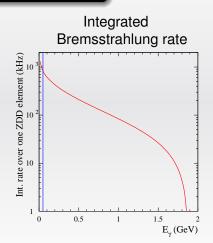
#### **Declared hardness**

- LYSO  $\sim 10^8$  rad
- Scint.  $\sim 10^6$  rad

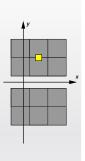
### Bremsstrahlung rate

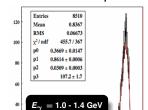
$${\cal L} = 3 \times 10^{32} \ \text{cm}^{-2} \text{s}^{-1}$$

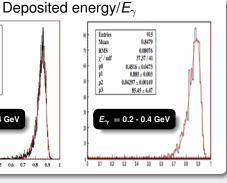




#### LYSO GEANT4 simulation<sub>1</sub>





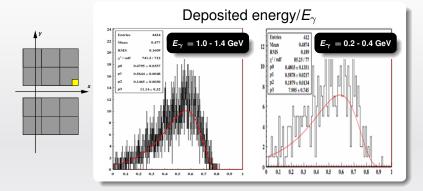


# $$\begin{split} & \underbrace{\frac{\text{df}}{\text{dE}} = \frac{\eta}{\sqrt{2\pi}\sigma_{E}\sigma_{0}} e^{-\frac{1}{2}\left[\frac{\ln^{2}\left(1-\frac{\eta(E-E_{0})}{\sigma_{E}}\right)}{\sigma_{0}^{2}} + \sigma_{0}^{2}\right]} \\ & \sigma_{0} = \frac{2}{2.35}\ln\left[\eta\frac{2.35}{2} + \sqrt{1+\left(\eta\frac{2.35}{2}\right)^{2}}\right], \quad \sigma_{E} = \frac{\text{FWHM}}{2.35} \end{split}$$

$m{\it E}_{\gamma}$ (GeV)	$\sigma_{E_{\gamma}}/E_{\gamma}$ Central (yellow square)	
1.0 - 1.4	3.6%	
0.2 - 0.4	4.9%	



### LYSO GEANT4 simulation<sub>2</sub>



$ extbf{\emph{E}}_{\gamma}$ (GeV)	$\sigma_{ extsf{E}_{\gamma}}/ extsf{E}_{\gamma}$ Central (yellow square)	
1.0 - 1.4	26.0%	
0.2 - 0.4	32.0%	

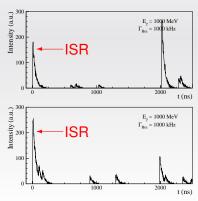
### Pileup effect: signal generation

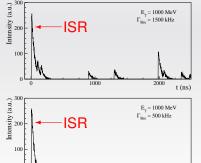
#### Maximum Bremsstrahlung rate expected 2.1 MHz (ZDD/4)

- Flash ADC: 500 MS/s, 8-bit resolution
- LYSO signal:

Intensity = 
$$e^{-t/\tau_r}(1-e^{-t/\tau_d})$$

rising time  $\tau_r = 2$  ns, decay time  $\tau_d = 40$  ns



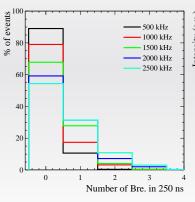


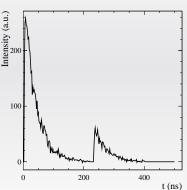
1000

2000

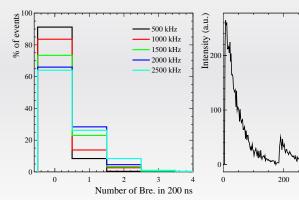
t (ns)

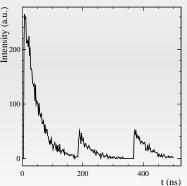
## Probability of pileup as a function of the Bremsstrahlung rate



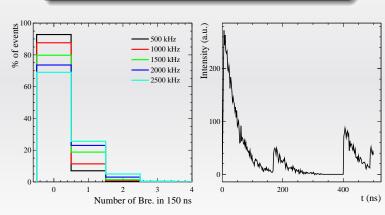


### Probability of pileup as a function of the Bremsstrahlung rate

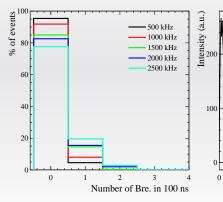


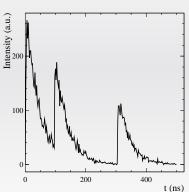


# Probability of pileup as a function of the Bremsstrahlung rate

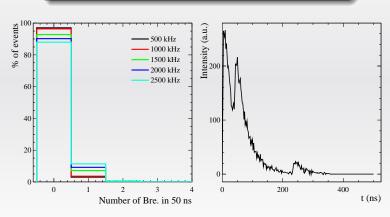


#### Probability of pileup as a function of the Bremsstrahlung rate



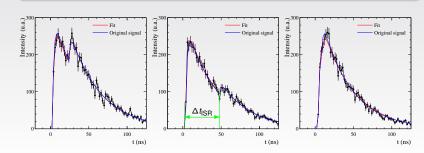


# Probability of pileup as a function of the Bremsstrahlung rate



#### Pileup effect: evaluation

- Signals have been generated at various rates with  $\text{Intensity} = \textit{\textbf{E}}_{\gamma_{\text{IS}}} \cdot e^{-t/\tau_{\text{decay}}} (1 e^{-t/\tau_{\text{raise}}})$
- E.g. at 2500 kHz: 31.6% has  $\Delta t_{\rm ISR} <$  160 ns  $\sim$  4 decay times
- We fit these signals to verify our capability to distinguish ISR and Bremsstrahlung contributions



#### Pileup effect<sub>2</sub> in T = 160 ns

■ The fit goodness is expressed as  $(\sigma_E/E)_{\text{fit}} = (E_{\text{gen}} - E_{\text{fit}})/E_{\text{gen}}$ , where  $E_{\text{gen}}$  is the generated ISR amplitude and  $E_{\text{fit}}$  is its fitted value

$$lacksquare$$
 We consider as a reference accuracies:  $\left\{egin{array}{ll} 7\% \sim rac{\sigma_E}{E} & ext{LYSO} \ 15\% \sim rac{\sigma_E}{E} & ext{Pb-Scint} \end{array}
ight.$ 

lacksquare  $E_{\gamma_{
m IS}}\in$  [0.5 GeV, 1.5 GeV], mild dependence on  $E_{\gamma_{
m IS}}$ 

rate	Pileup in 160 ns	$(\sigma_E/E)_{\mathrm{fit}} >$ 7%	$\sigma_{\it E}/E)_{ m fit} >$ 15%
(kHz)	(%)	(%)	(%)
2500	30	9.4	4.8
2100	26	8.1	4.2
1000	14	4.3	2.2
800	10	3.2	1.6

# The $n\overline{n}\gamma_{\rm IS}$ physics case: kinematic fit

#### Inputs (6)

- $\overline{n}$  3-momentum (TOF)
- $\gamma_{\rm IS}$  3-momentum (ZDD)

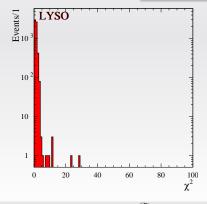
#### Constraints (4)

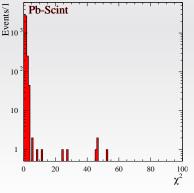
4-momentum cons.

#### Unknowns (3)

n 4-momentum

$$\chi^2 = \sum_{\text{tracks}} \sum_{i} \frac{\left(p_i^{\text{exp}} - p_i^{\text{fit}}\right)^2}{\sigma_{p_i}^2}$$





# The $n\overline{n}\gamma_{\rm IS}$ physics case: kinematic fit

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