

# IHEP cosmics

Companion of LNF Runs 219,228

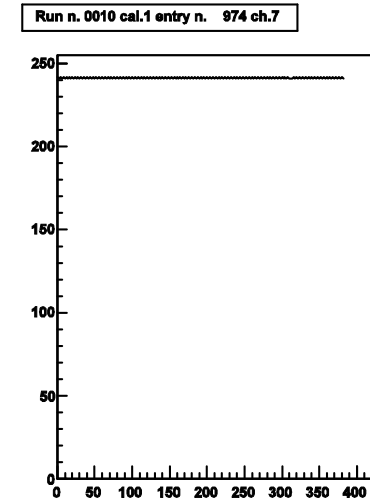
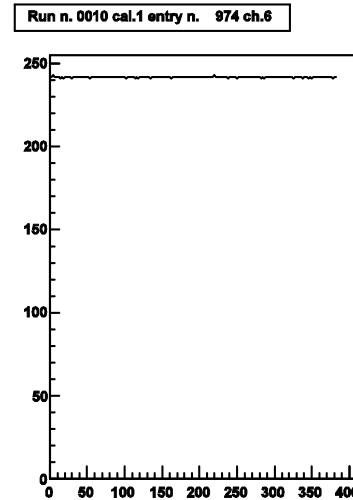
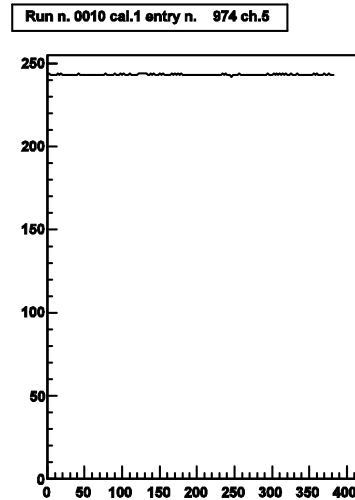
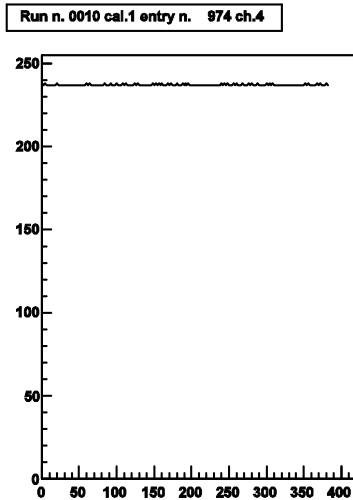
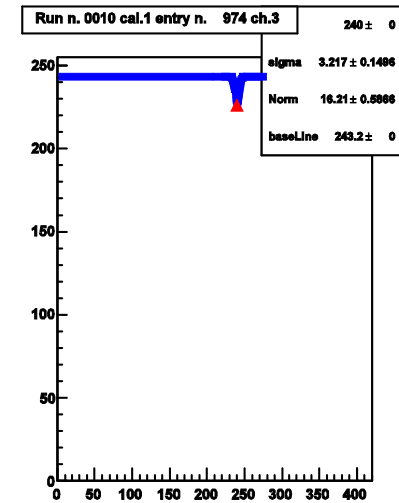
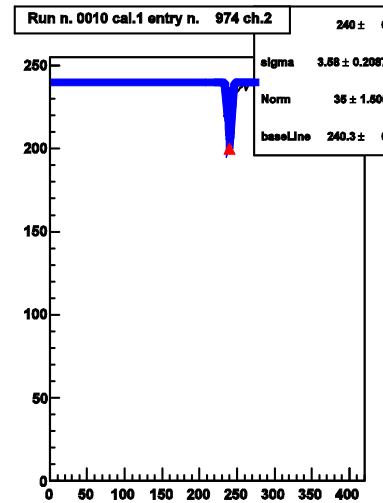
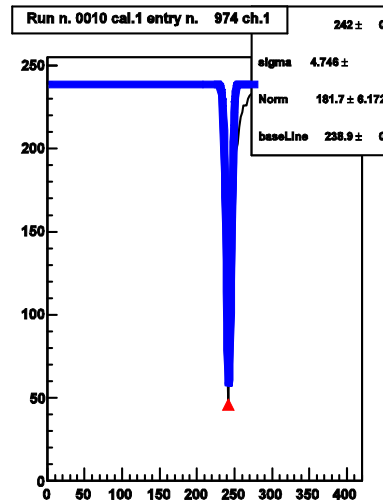
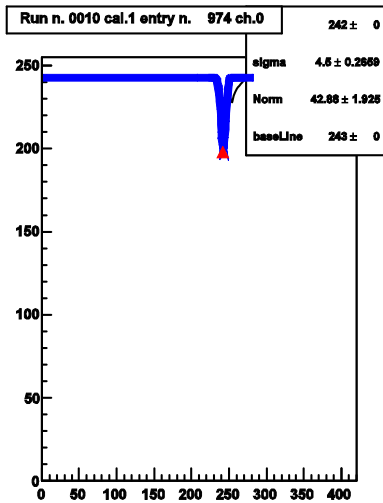
# Runs 10,11,12,13 at IHEP

- “up” and “down” minicalorimeters
- HV: “Frascati” 1.4 kV points, gain  $\approx 1.2 \cdot 10^6$ , Happy Box present
- Autotriggered, majority 3 channels
- One-channel trigger: baseline-4cts (16 mV)
- Independent trigger and DAQ on each minicalorimeter

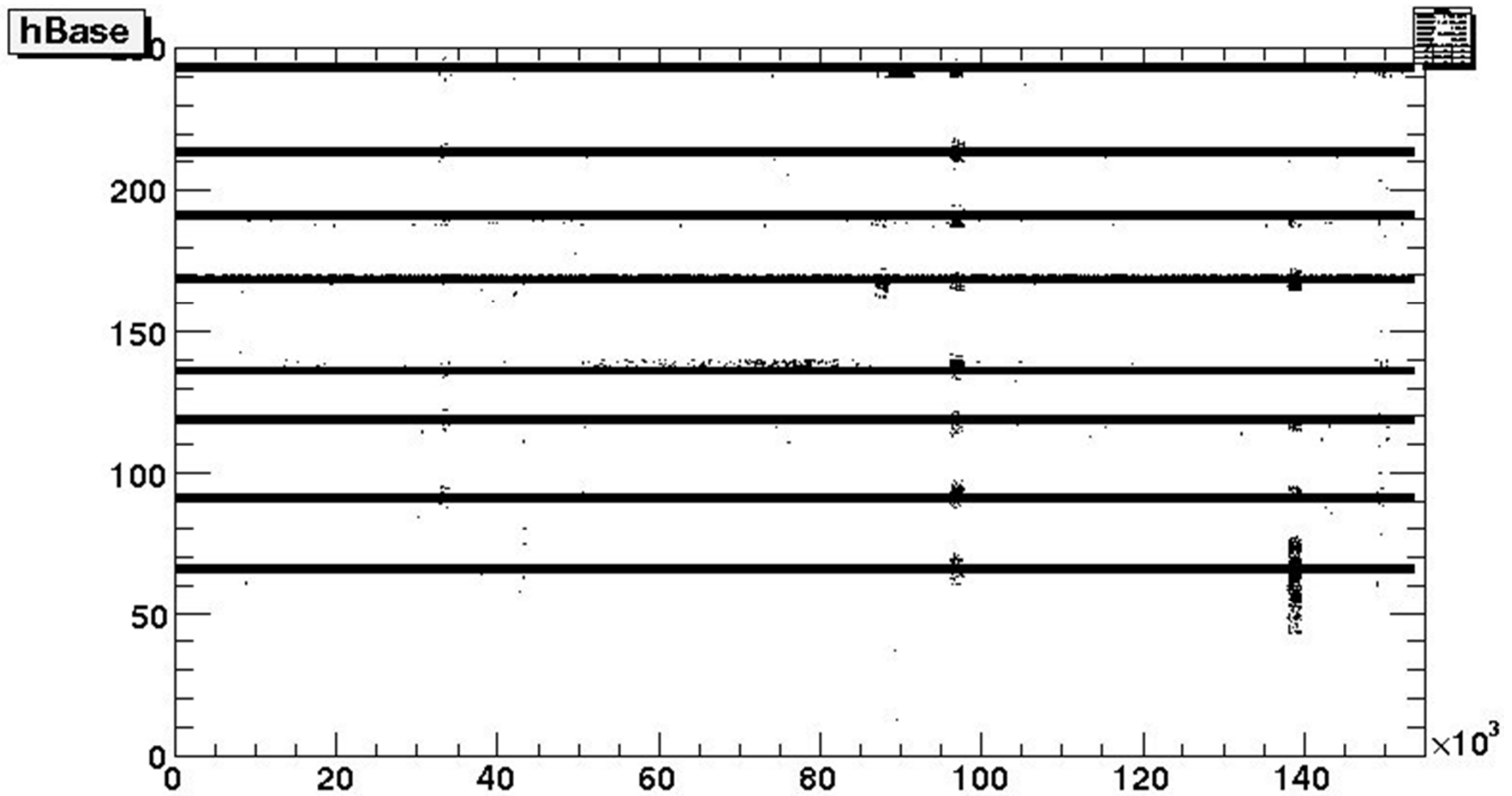
# Scanning waveforms

- First 80 ns data (40 samples) used to find the 8 baselines and baseline standard deviations on a per-event, per-waveform basis
- **Bad event**, rejected if one  $\sigma$  is  $> 0.5$  cts (2 mV)
- Samples 41 to 196 (82 to 384 ns) used to find the signal peak, and time of peak
- **Bad event**, rejected if one peak out-of-time (between 140 and 230 ns)

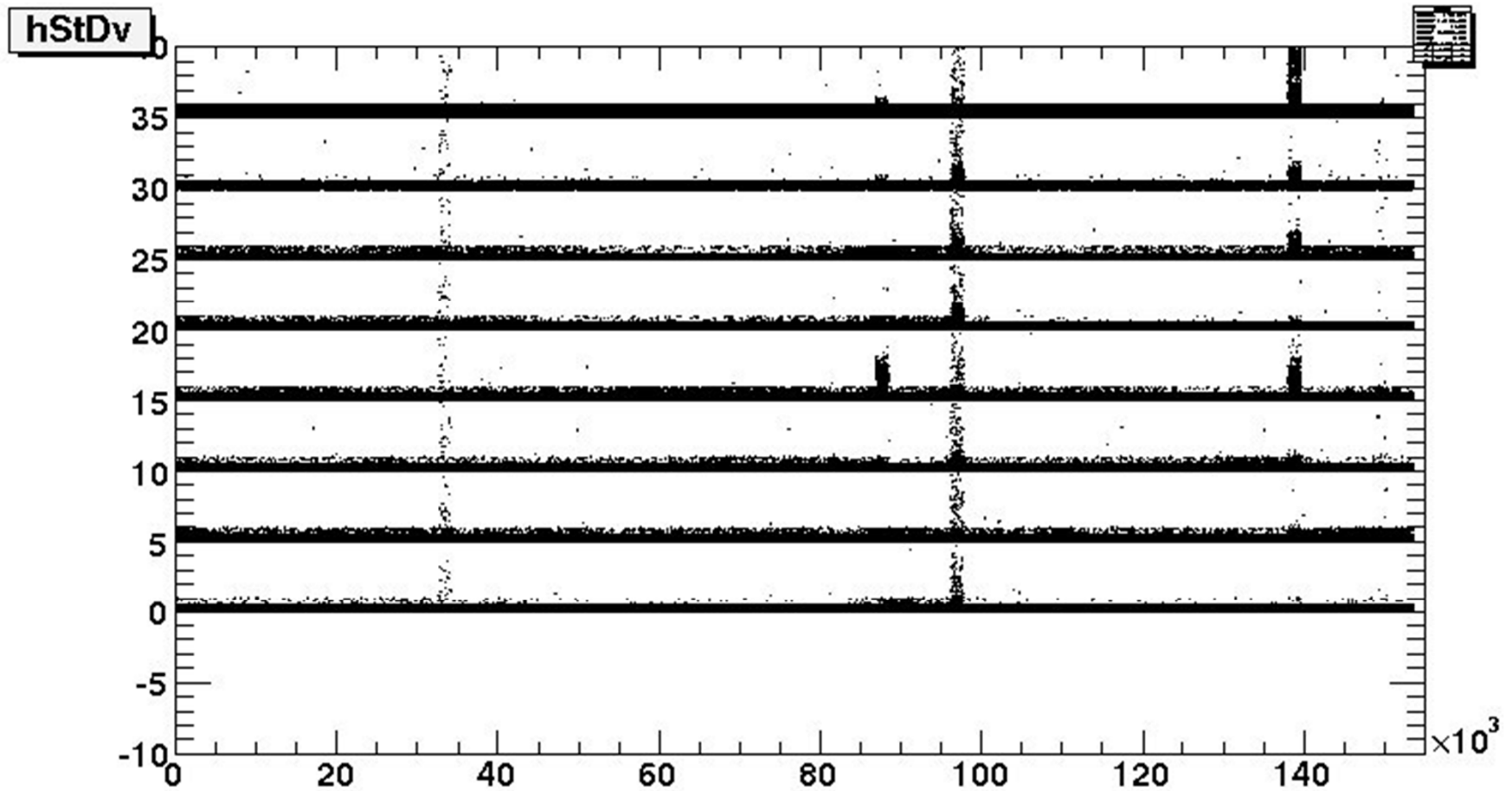
# One event from minical 1 (“up”)



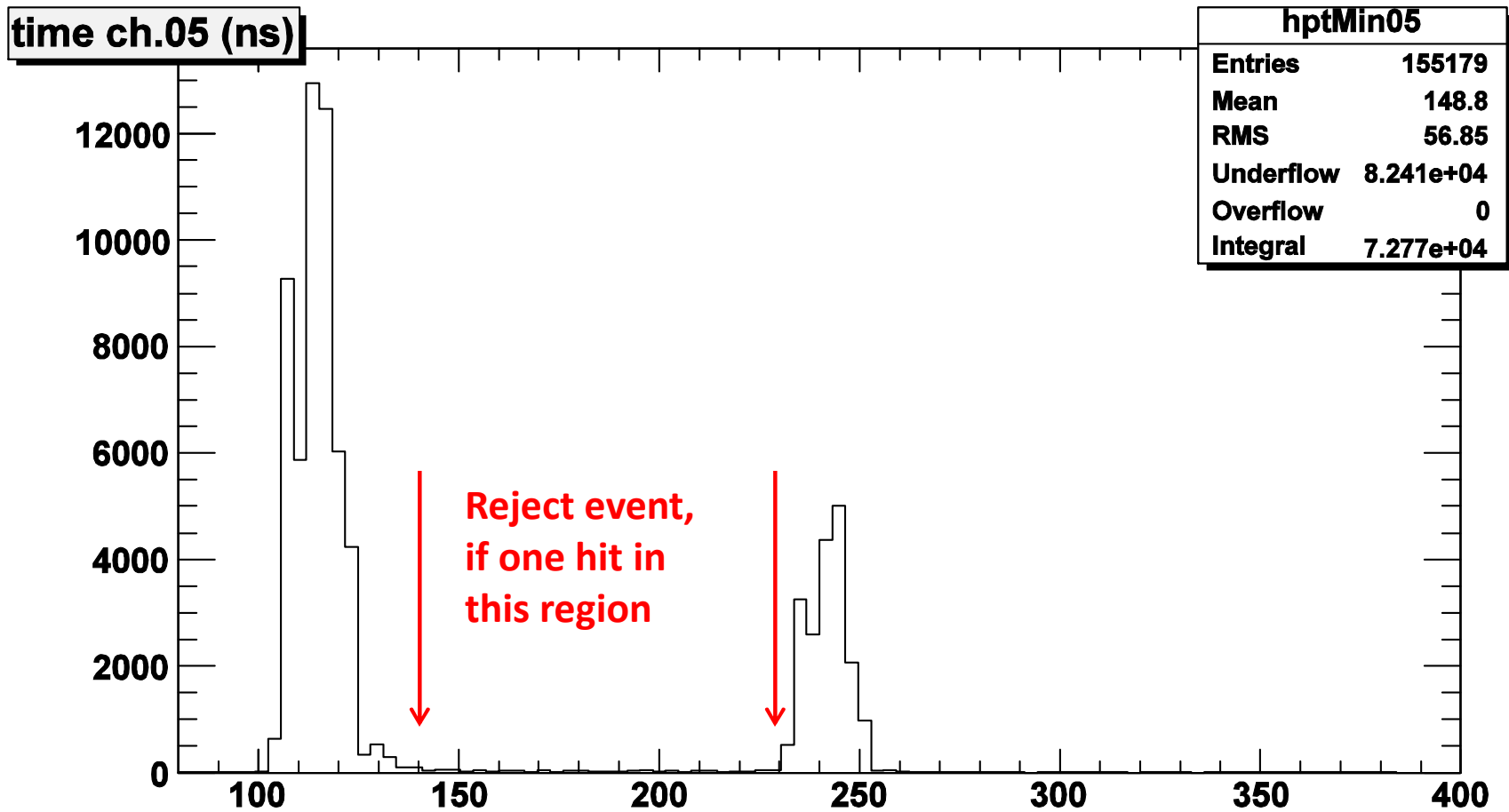
# Noise, seen on baselines



# Noise, seen in baseline $\sigma$ 's

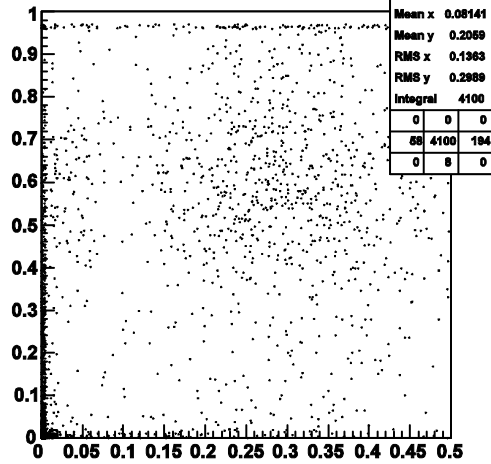


# Bad times...

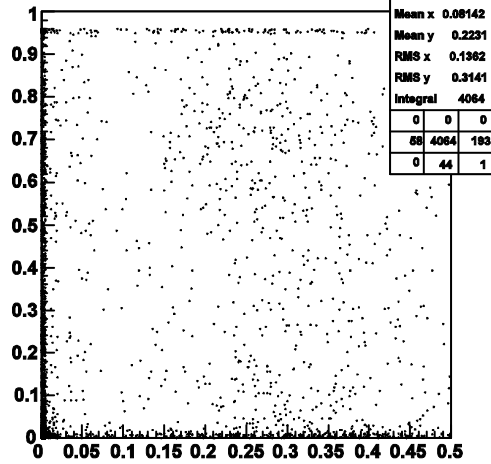


# Pulse height correlations (LNF)

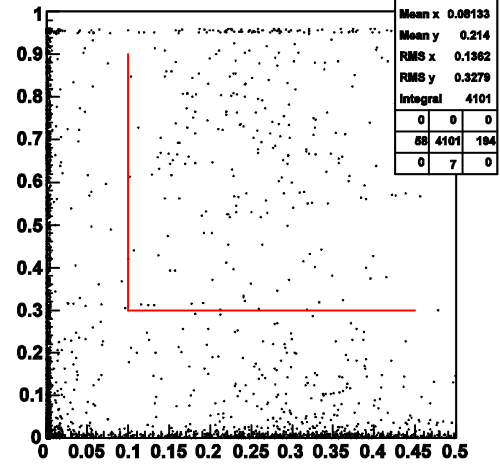
correlation ch.5 vs. ch. 4



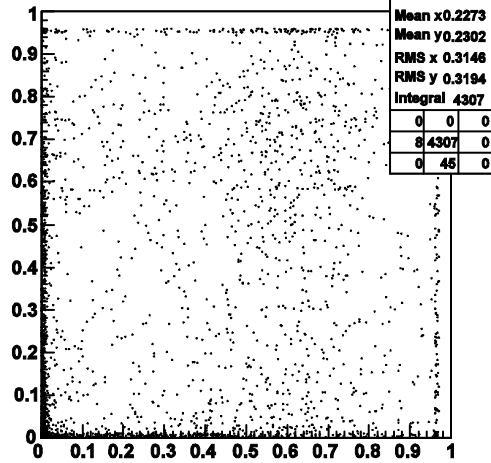
correlation ch.6 vs. ch. 4



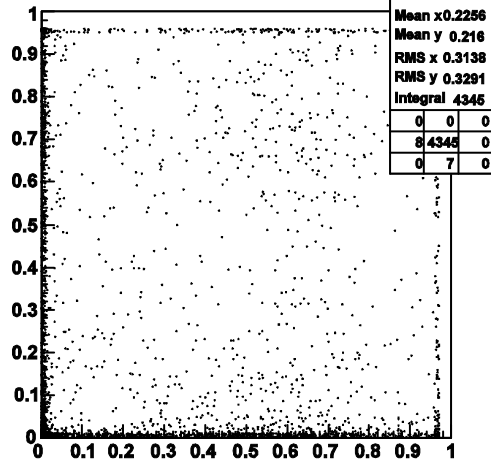
correlation ch.7 vs. ch. 4



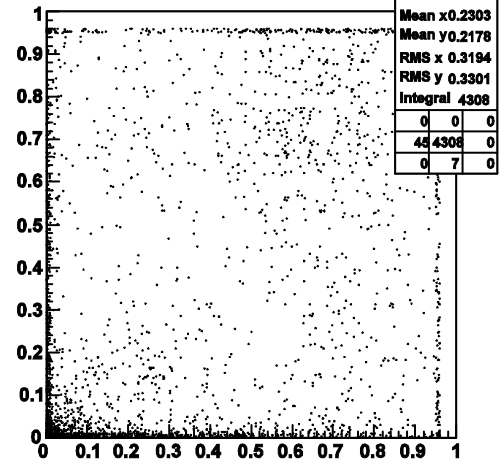
correlation ch.6 vs. ch. 5



correlation ch.7 vs. ch. 5

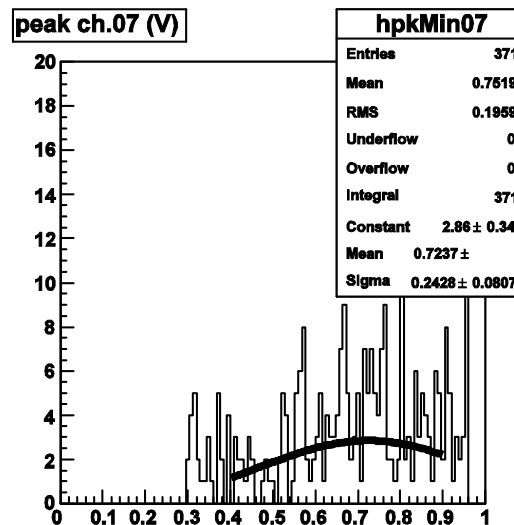
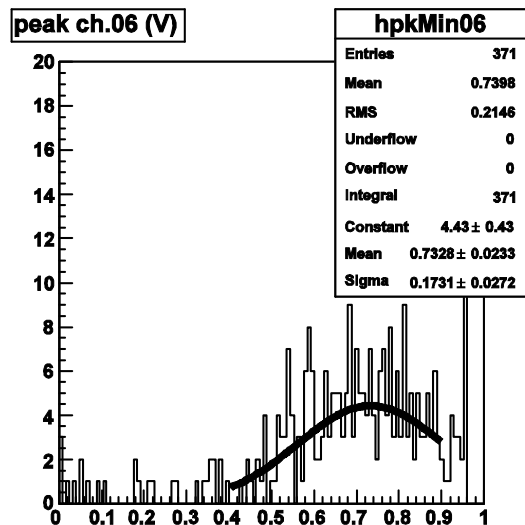
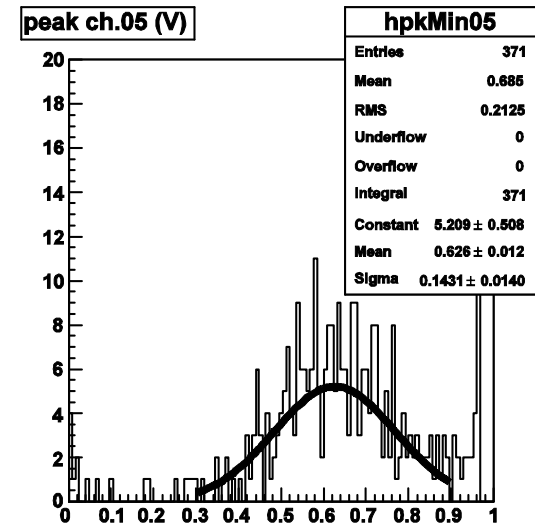
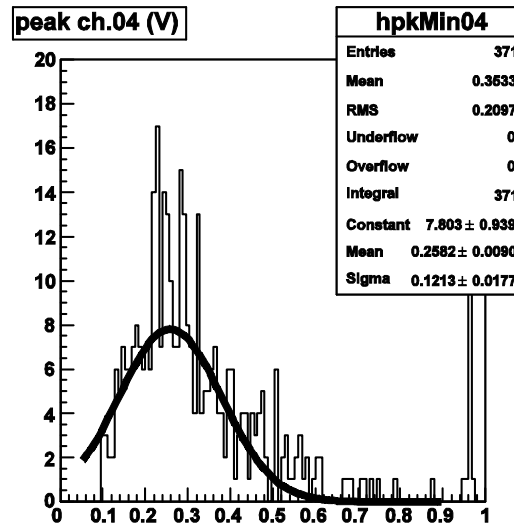
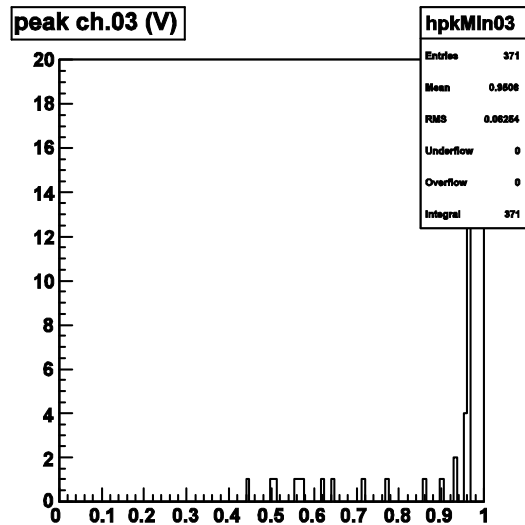


correlation ch.7 vs. ch. 6



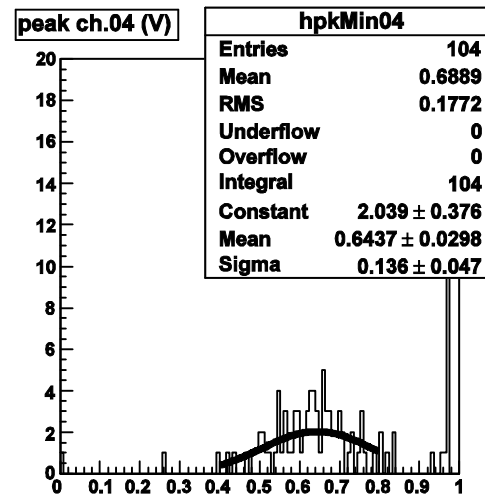
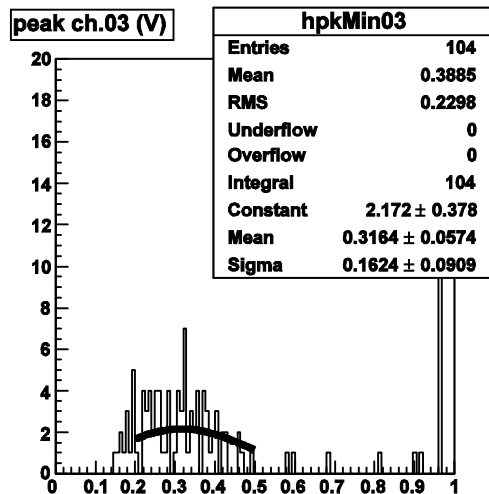
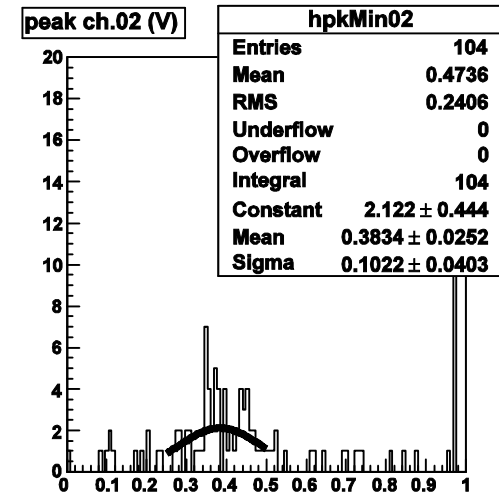
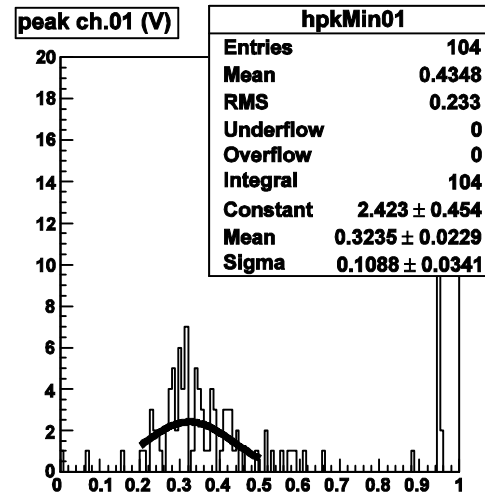
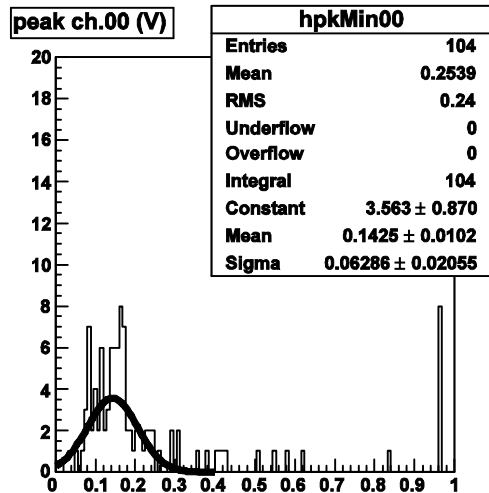


# LNF Run 219 after trigger cleanup



Peak in channels 5,6,7 is a factor 2 higher than in channel 4, just as expected. This is a good thing!

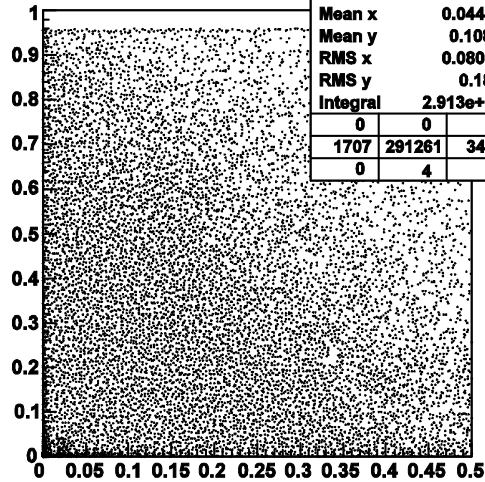
# LNF Run 228 after trigger cleanup



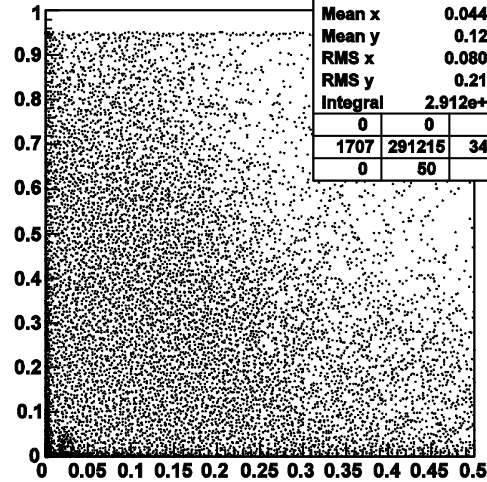
Peak in channels 1,2,3 is a factor 2 higher than in channel 0, just as expected. This is a good thing!

# Minical 0("dn") at IHEP

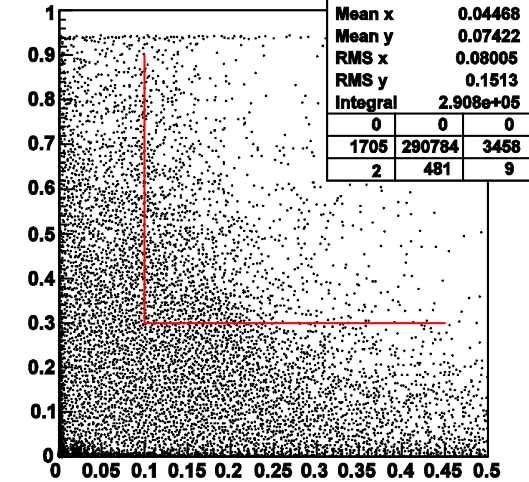
correlation ch.5 vs. ch. 4



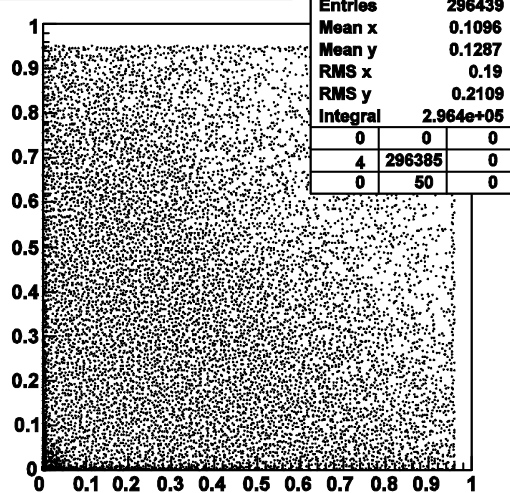
correlation ch.6 vs. ch. 4



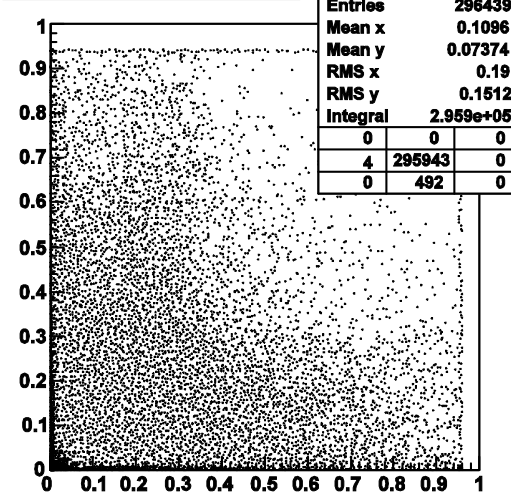
correlation ch.7 vs. ch. 4



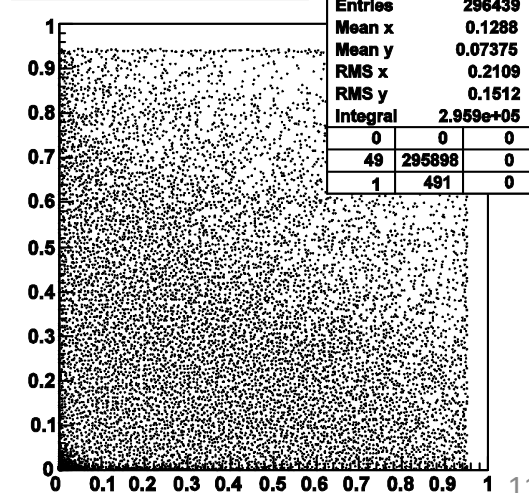
correlation ch.6 vs. ch. 5



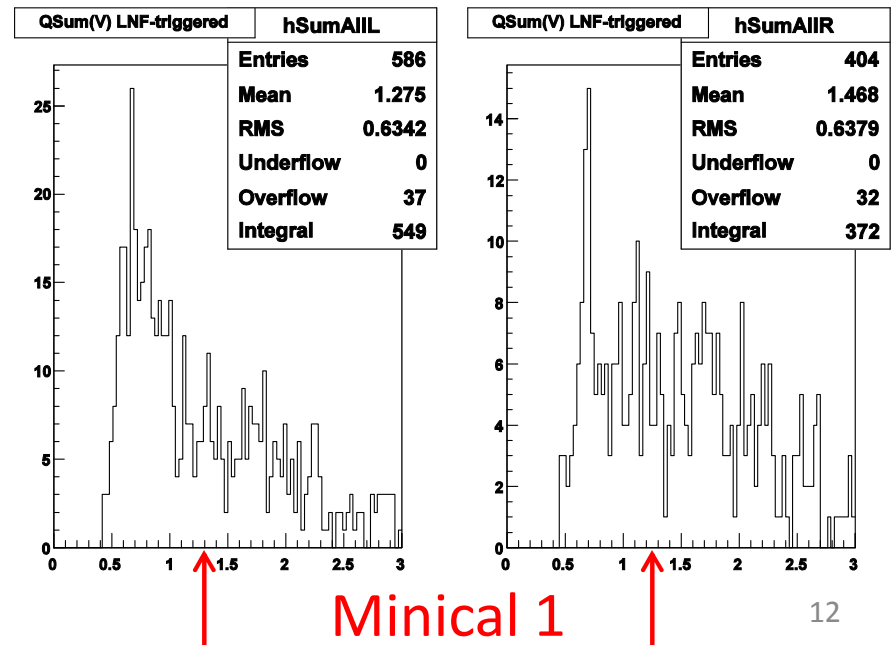
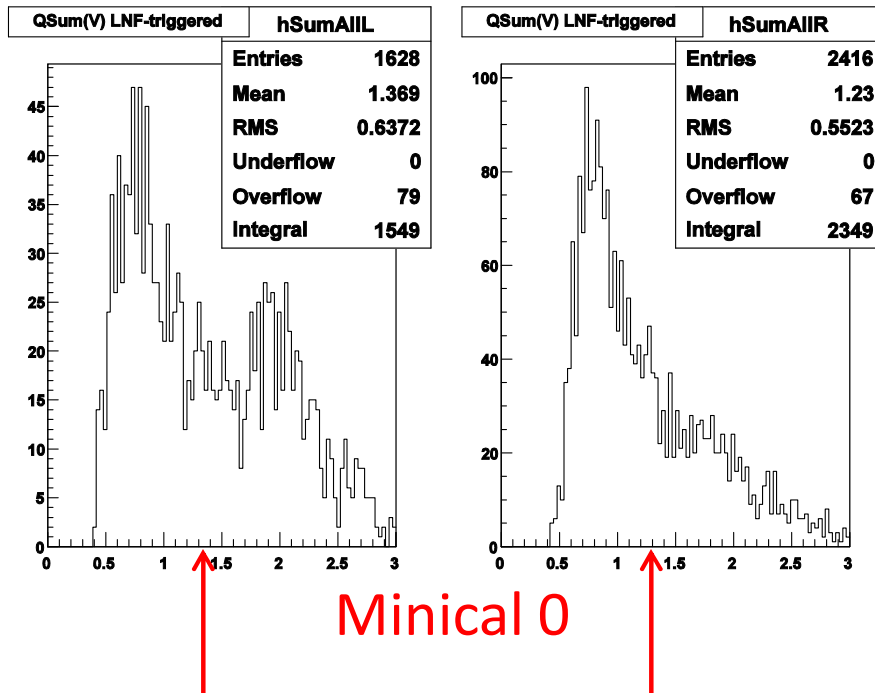
correlation ch.7 vs. ch. 5



correlation ch.7 vs. ch. 6



# “LNF” cuts on IHEP data



# Summary of facts

- The LNF-style criteria are useless at IHEP
- Cosmic rays are much rarer (factor 3 less if angular distribution  $\sin^2\vartheta$  taken seriously, possibly factor 10? *30-10 evts/day*)
- Very hard to identify and use cosmic rays using a simple cut on pulse heights
- Must investigate other techniques

# Use masks

- A “mask” is a configuration of contiguous hit cells

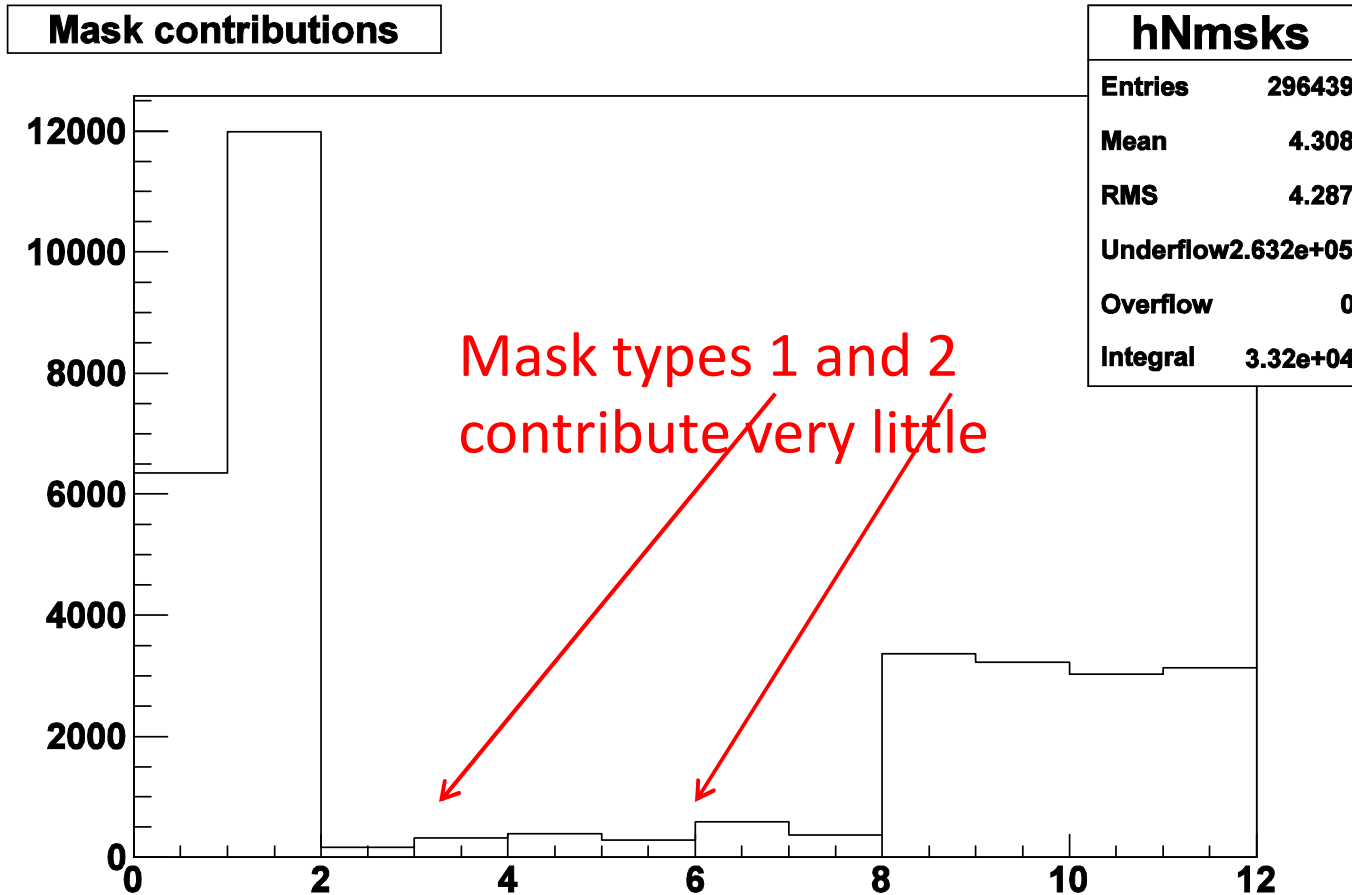
- 2 “type 0” masks: 

- 2 “type 1” masks: 

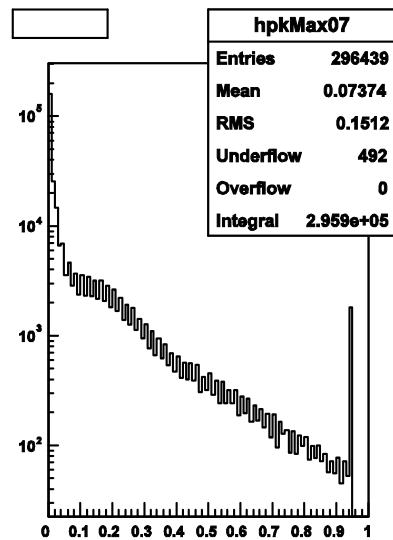
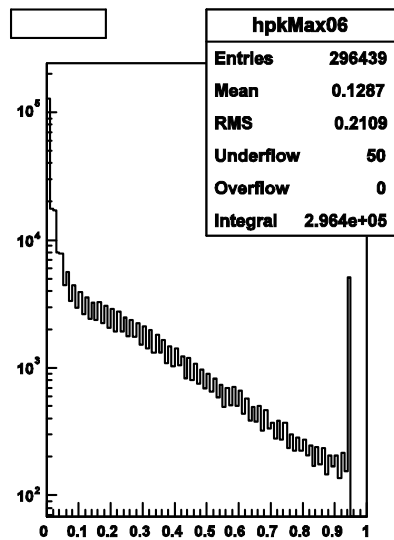
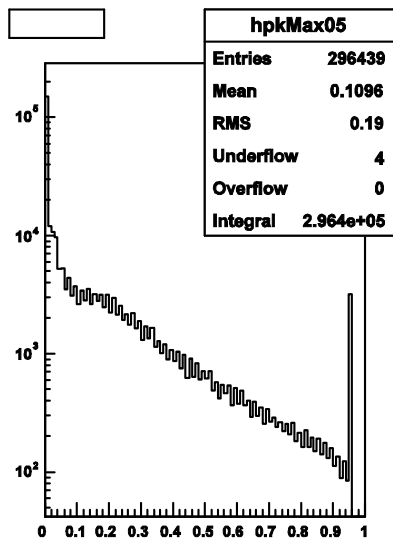
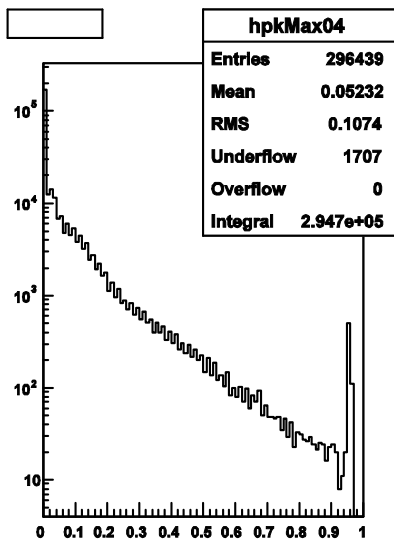
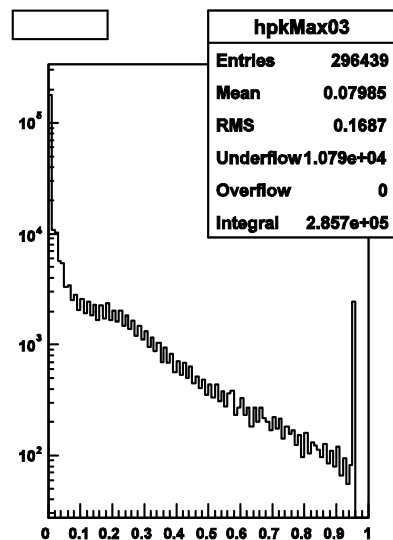
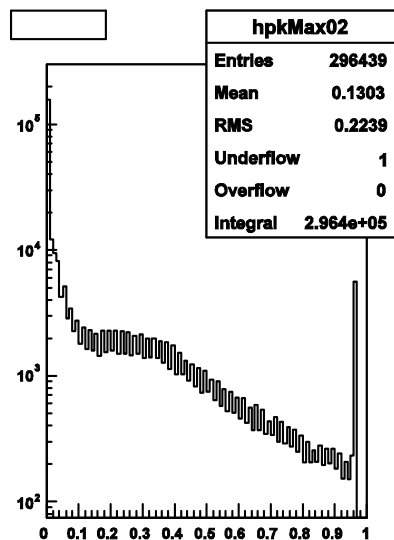
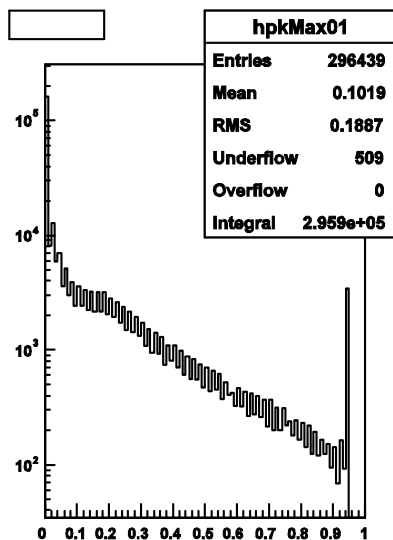
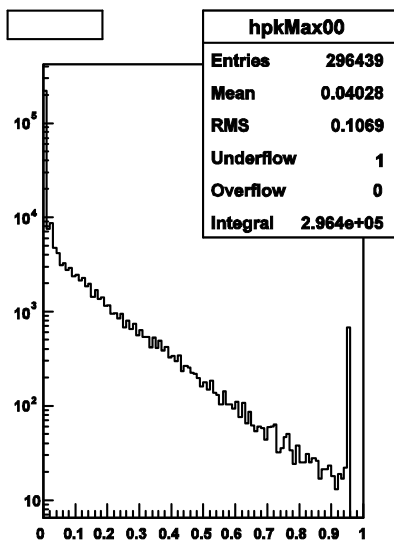
- 4 “type 2” masks: 

- 4 “type 3” masks: 

# Which masks are more frequent?

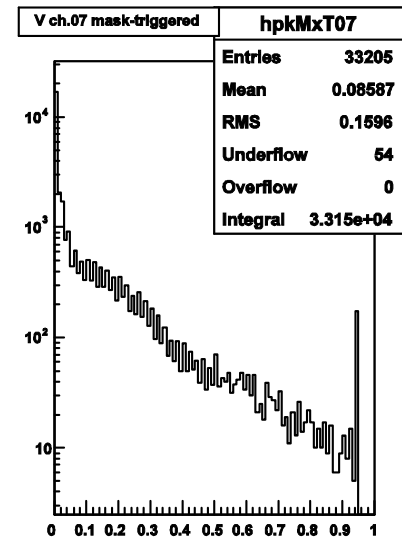
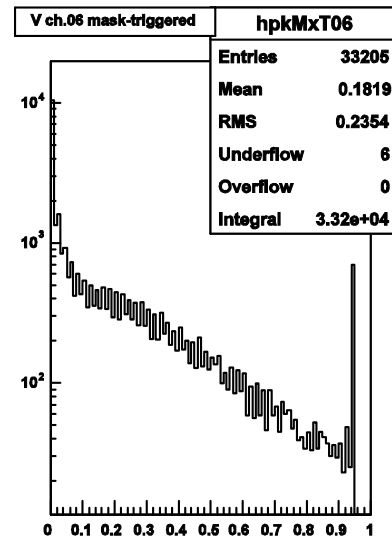
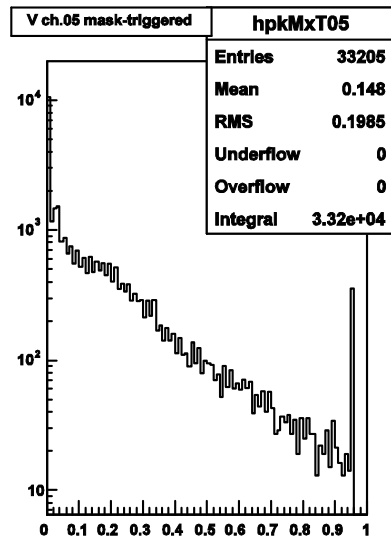
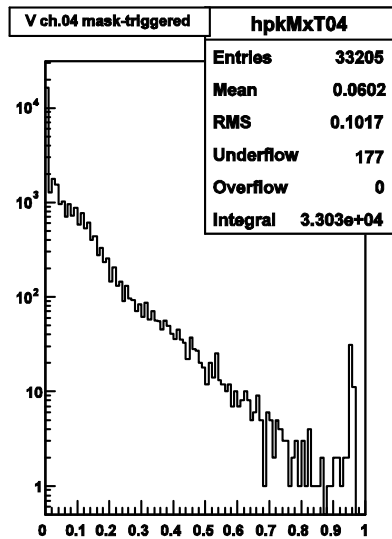
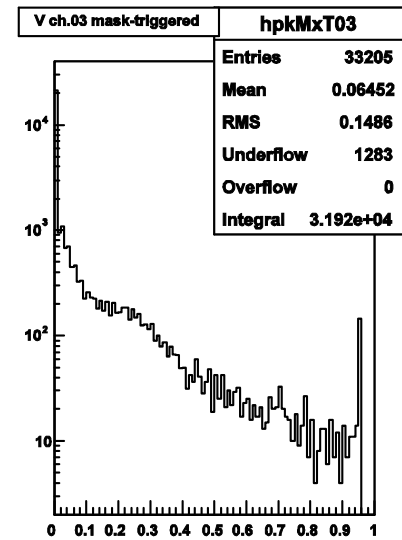
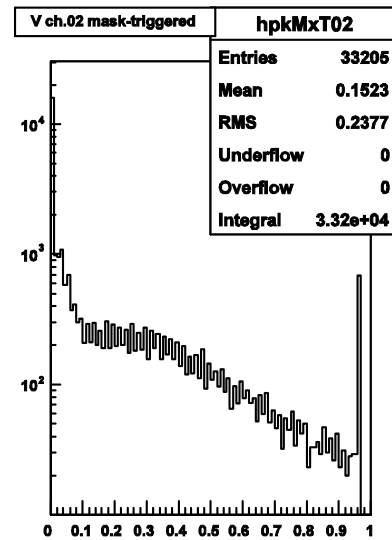
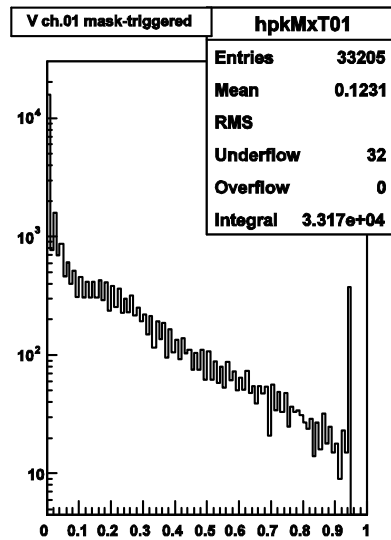
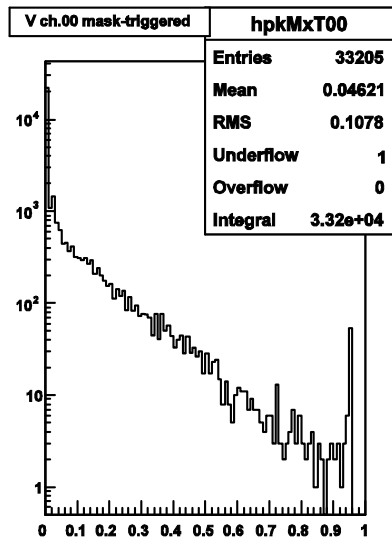


# Singles for minical 0

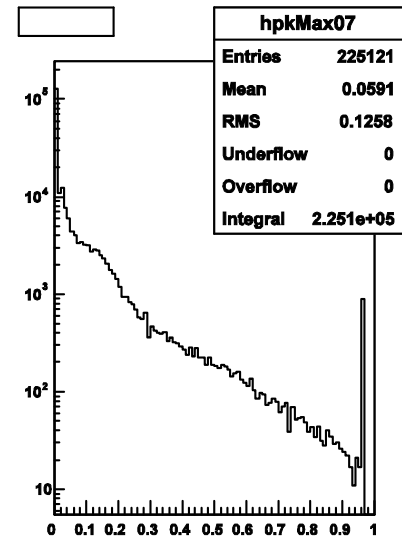
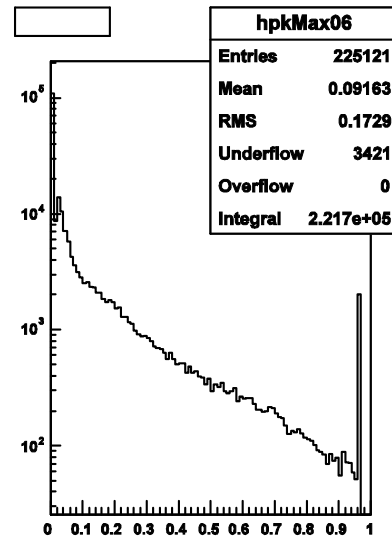
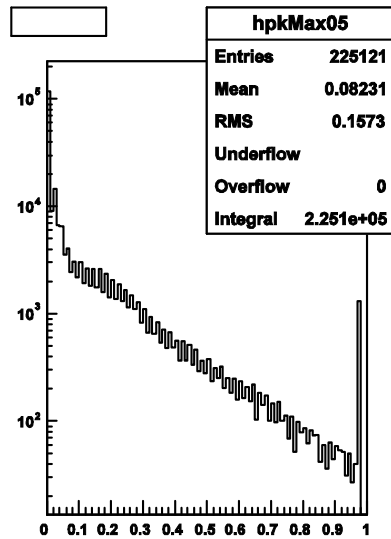
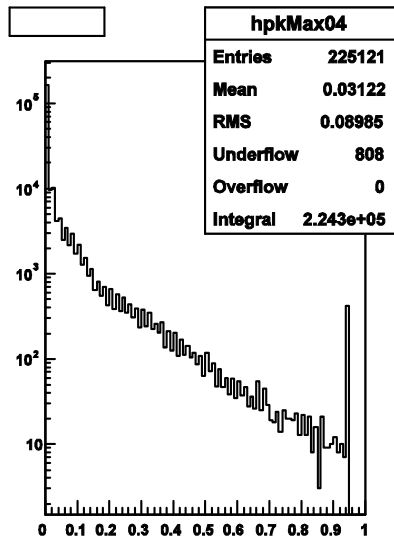
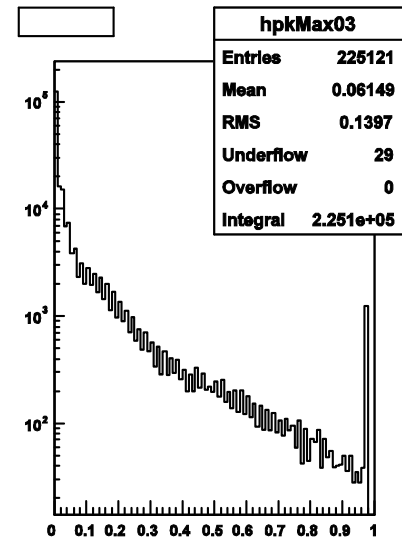
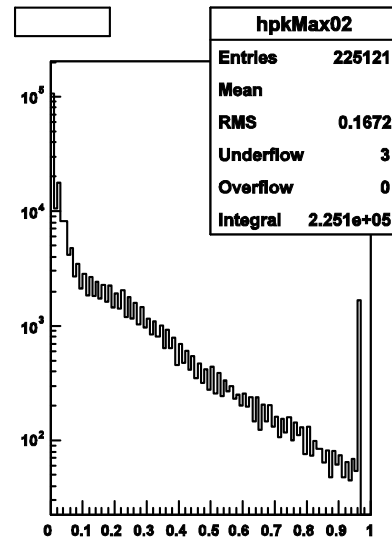
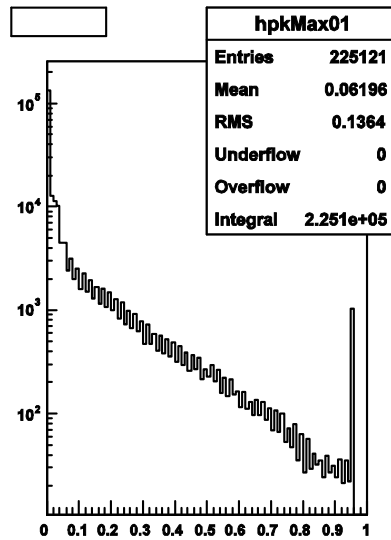
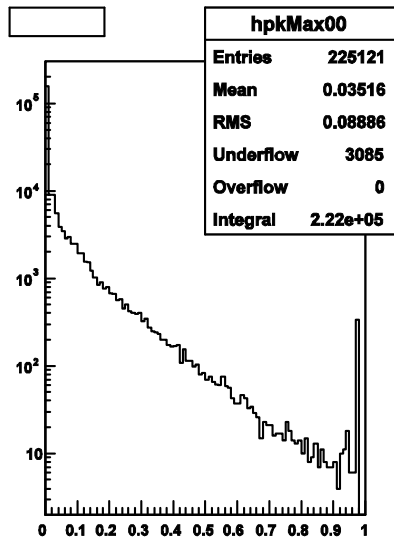




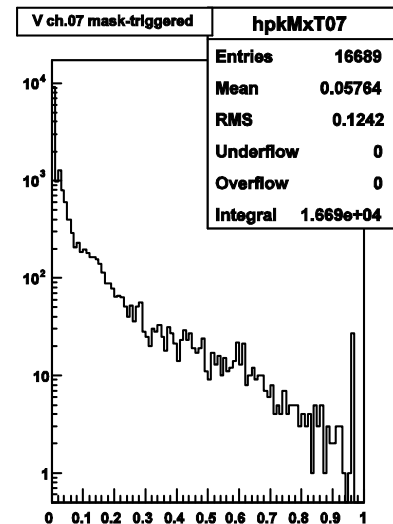
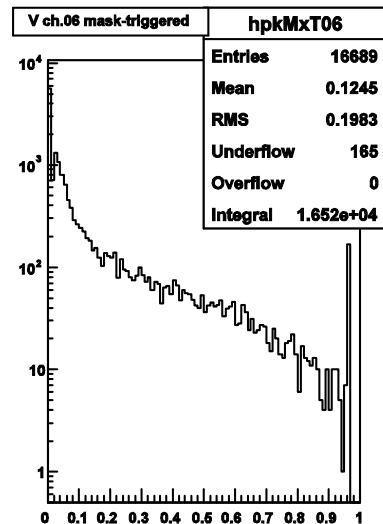
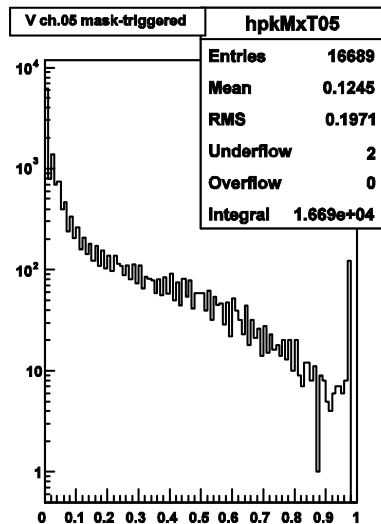
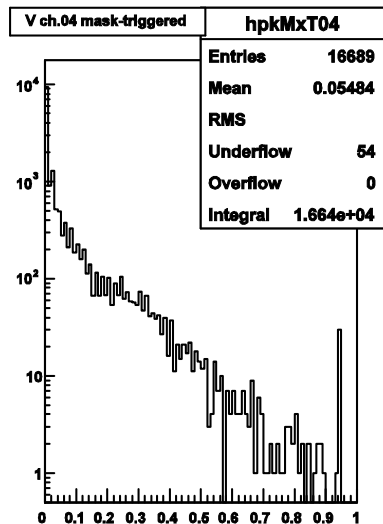
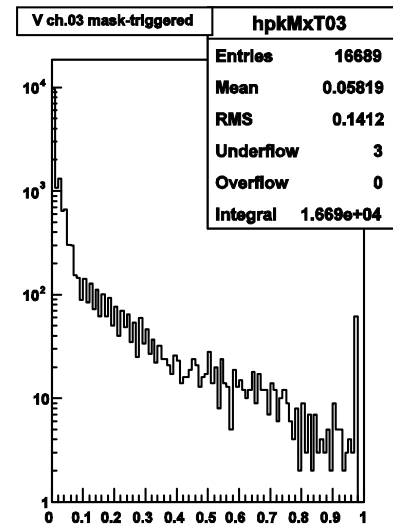
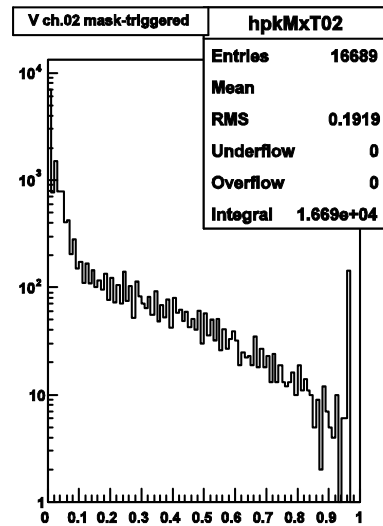
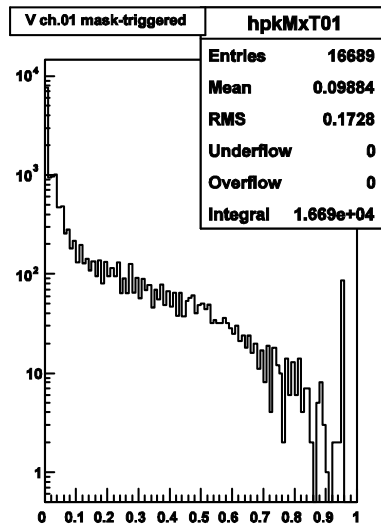
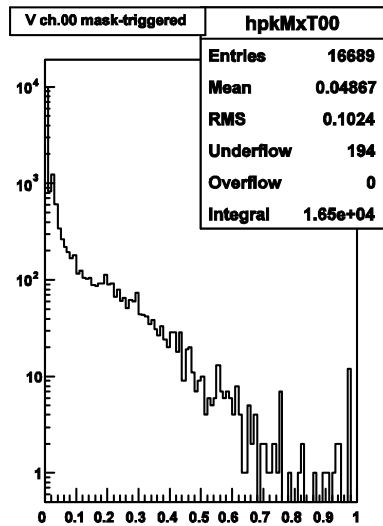
# Mask-trigger, minical 0



# Singles, minical 1

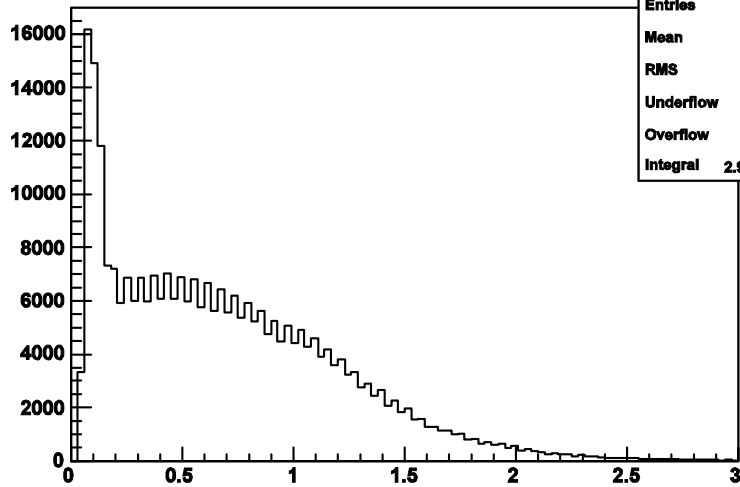


# Mask-triggered, minical 1



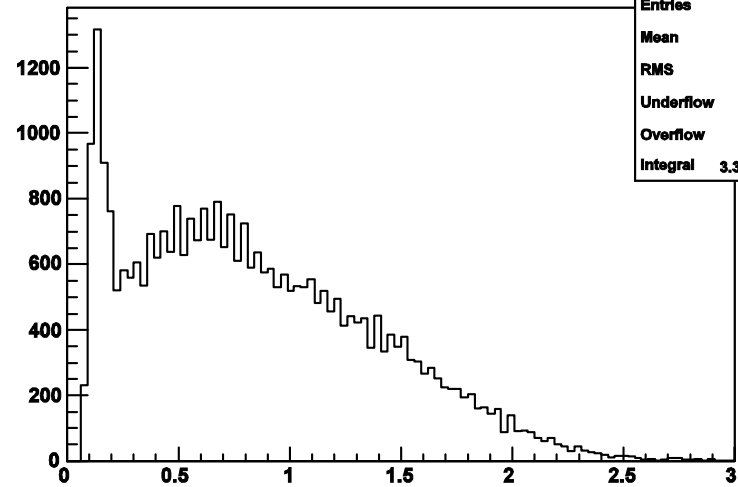
# Using masks, minical 0

QSum(V)



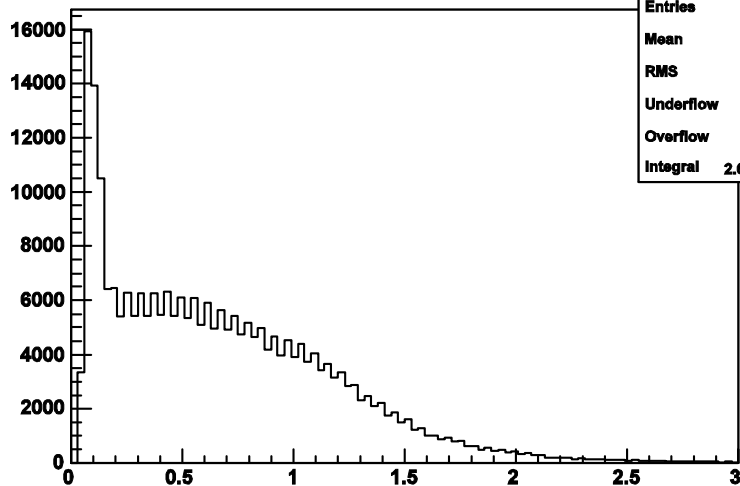
hSumAll	
Entries	296439
Mean	0.7002
RMS	0.5101
Underflow	0
Overflow	940
Integral	2.955e+05

QSum(V) mask-triggered



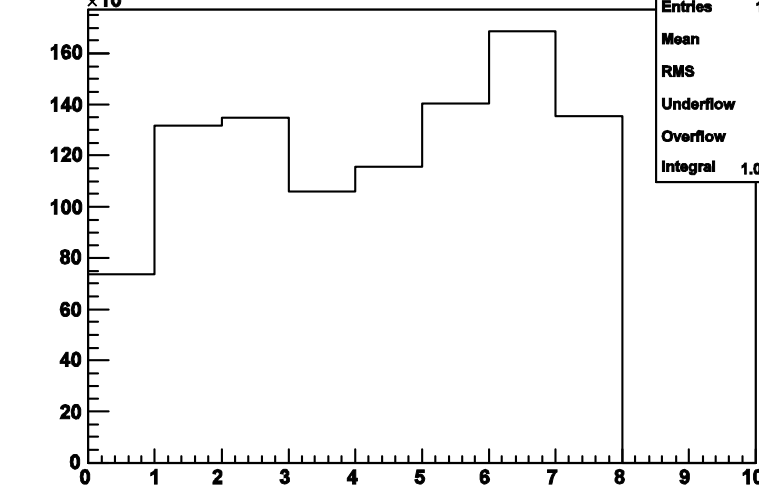
hSumAllIT	
Entries	33205
Mean	0.8537
RMS	0.5405
Underflow	0
Overflow	27
Integral	3.318e+04

QSum(V) noMask-triggered



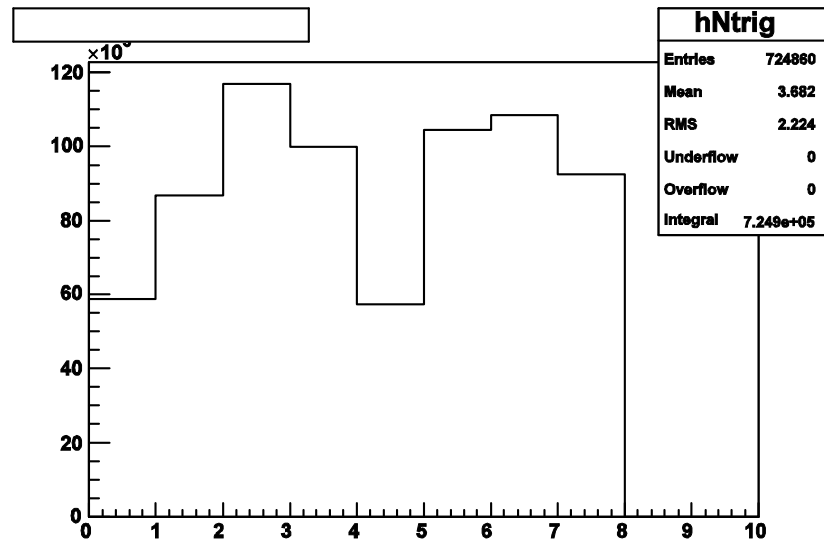
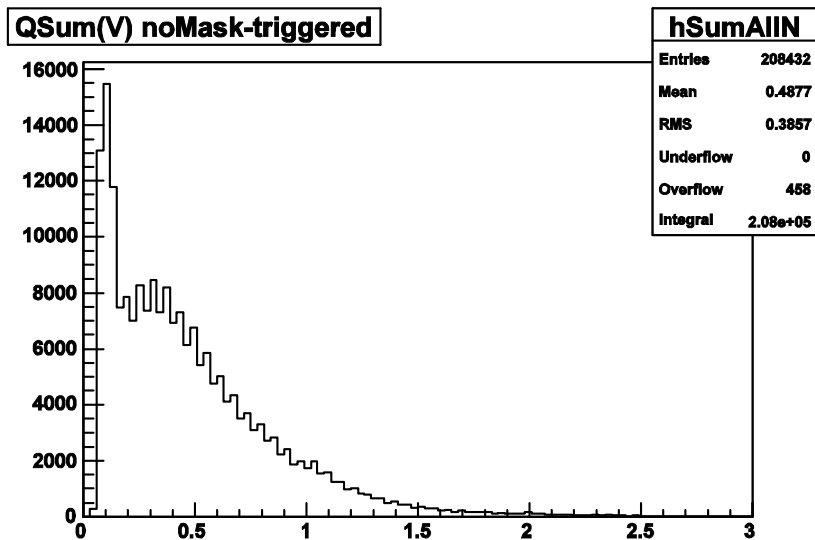
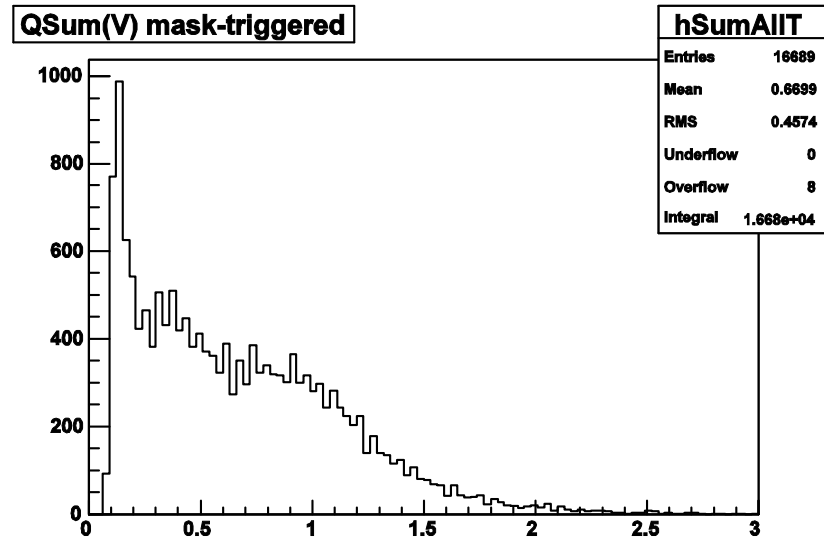
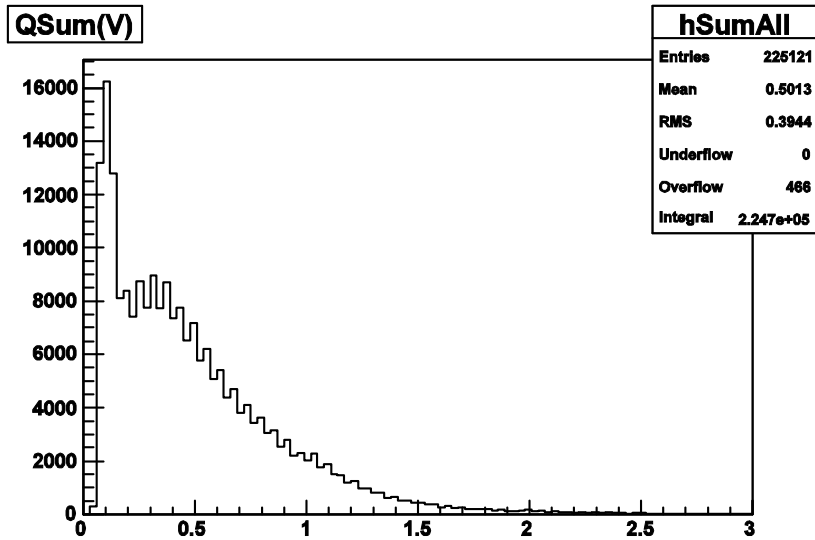
hSumAllIN	
Entries	263234
Mean	0.6808
RMS	0.5028
Underflow	0
Overflow	913
Integral	2.623e+05

hNtrig



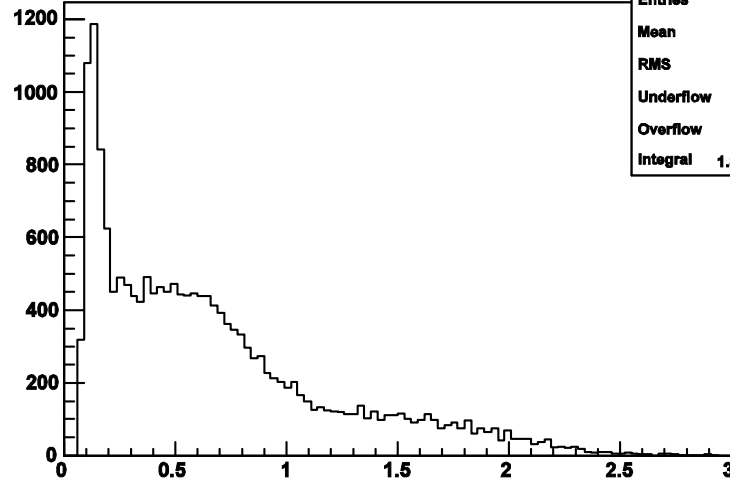
hNtrig	
Entries	1006447
Mean	3.82
RMS	2.231
Underflow	0
Overflow	0
Integral	1.006e+06

# Using masks, minical 1



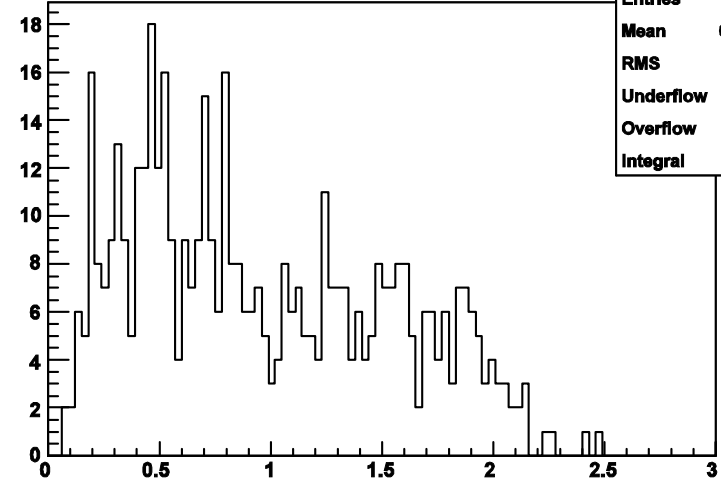
# Divide by mask type, minical 0

QSum(V) mask 0 (LNF)



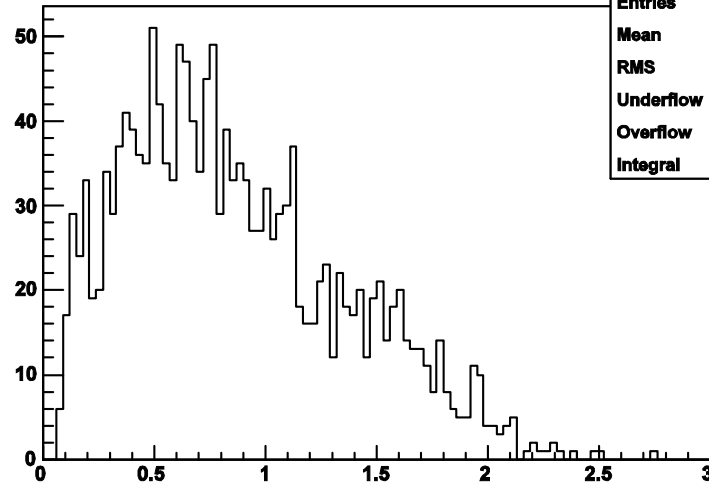
hSumAll0	
Entries	18332
Mean	0.678
RMS	0.5351
Underflow	0
Overflow	23
Integral	1.831e+04

QSum(V) mask 1 (4 2+2 long)



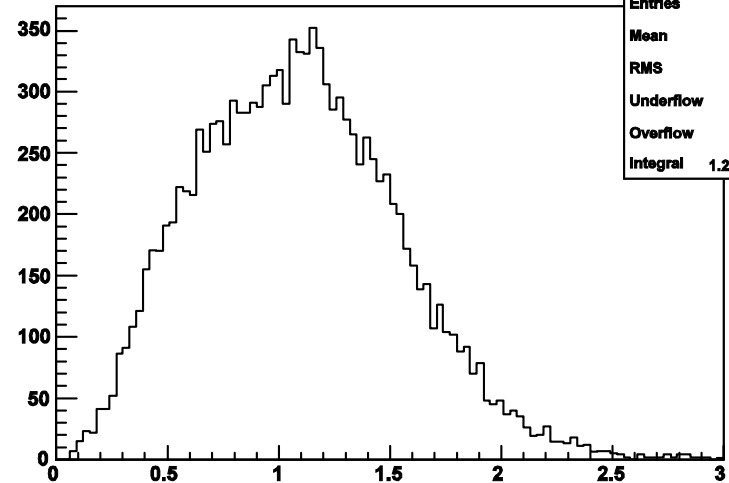
hSumAll1	
Entries	488
Mean	0.9686
RMS	0.571
Underflow	0
Overflow	0
Integral	488

QSum(V) mask 2 (4 1+3)



hSumAll2	
Entries	1638
Mean	0.8696
RMS	0.4954
Underflow	0
Overflow	0
Integral	1638

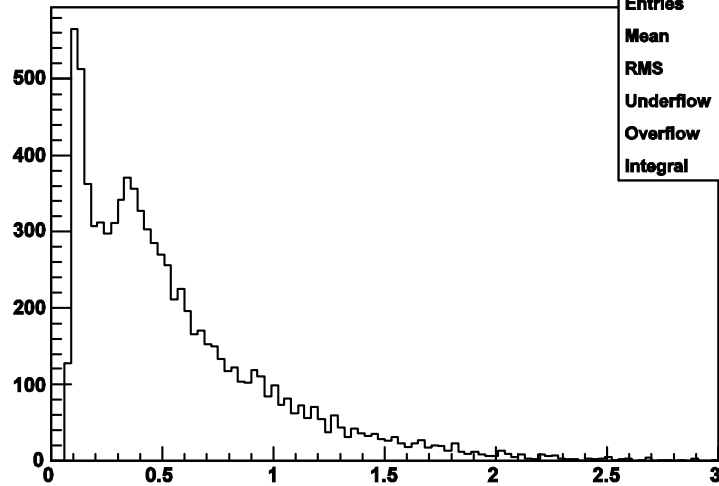
QSum(V) mask 3 (4 2+2 short)



hSumAll3	
Entries	12747
Mean	1.086
RMS	0.452
Underflow	0
Overflow	4
Integral	1.274e+04

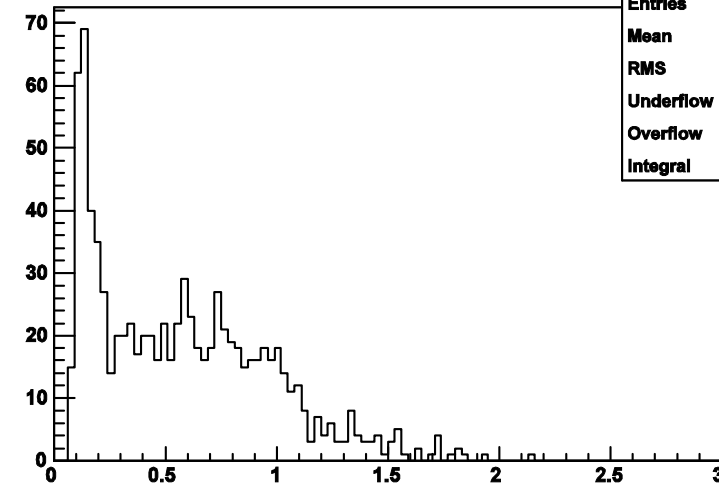
# Divide by mask type, minical 1

QSum(V) mask 0 (LNF)



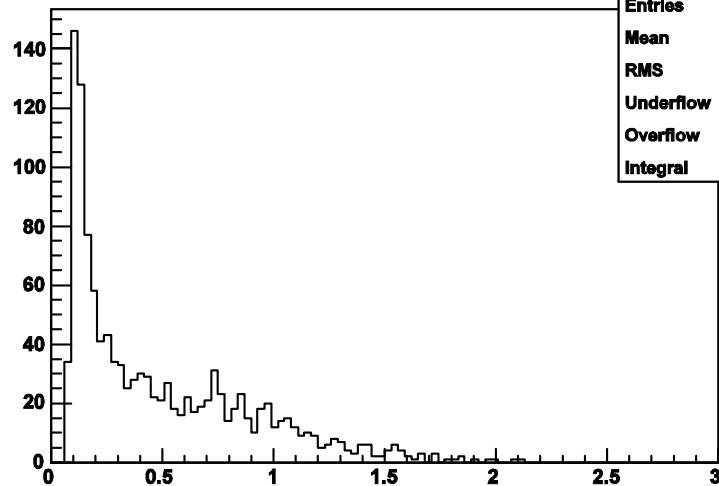
hSumAll0	
Entries	8773
Mean	0.5607
RMS	0.4354
Underflow	0
Overflow	7
Integral	8766

QSum(V) mask 1 (4 2+2 long)



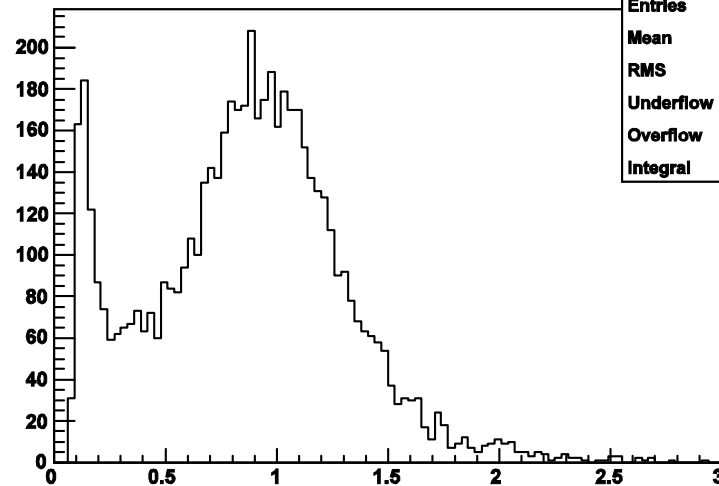
hSumAll1	
Entries	861
Mean	0.5752
RMS	0.3969
Underflow	0
Overflow	0
Integral	861

QSum(V) mask 2 (4 1+3)



hSumAll2	
Entries	1223
Mean	0.4922
RMS	0.406
Underflow	0
Overflow	0
Integral	1223

QSum(V) mask 3 (4 2+2 short)



hSumAll3	
Entries	5832
Mean	0.8648
RMS	0.4365
Underflow	0
Overflow	1
Integral	5831

# Conclusions? Ideas, anyone?

- There is little chance of selecting cosmic rays at IHEP like we did in Frascati
- At the level of “singles” differences between minical 0 and 1 are not much apparent
- Using mask-triggered events the sum of all charges is bigger in minical 0. PM gain changed? HV problem? Electronics?
- **Why mask type 3 shows such a big charge?**