



# Impact of energy spread on New Physics probes:

$B \rightarrow \phi K_s$  and  $\tau \rightarrow \mu \gamma$

**Nicola Neri**

Università di Pisa & INFN Pisa

**Maurizio Pierini**

University of Wisconsin, Madison



# The Idea

A super B factory is meaningful only if it has high potentiality in discovering New Physics.

We studied two of the available smoking guns

- + Time dependent CP measurement in  $b \rightarrow s$  decays ( $\phi K_s$ , but also  $\eta' K_s$  and  $K_s \pi^0$ )
  - ◆ We expect  $S \sim \sin 2\beta$  and  $C \sim 0$  in the SM (up to theoretical errors that will go down increasing experimental precision)
  - ◆ We can observe deviations, coming from NP contributions
- + Measurement of  $BR(\tau \rightarrow \mu \gamma)$ 
  - ◆ Small in the SM, enhanced in NP.  $\sim 10^{-10}$  is the goal.
- + These two “smoking guns” are strongly correlated in GUT models (see talk by M.Ciuchini)
- + We use a fast toy Monte Carlo simulation to extrapolate the improvement in the statistical error we will have



# WARNING

This is a fast simulation based on toys. We will cross check the results in a more detailed simulation in the next future

## For $b \rightarrow s$ time dependent analyses

- + We assume that present yields will scale with luminosity
- + We assume the same number of events for different energy spread values (so they do not correspond to the same lumi.)
- + We are (for the moment) ignoring any background rejection improvement that WILL come from looking at the D vertex on the tag side (work in progress)

## For $\tau \rightarrow \mu \gamma$

- + Work is **very preliminary**
- + We are assuming a **(too) conservative approach**, with worse energy spread, but same vertexing power (now vertexing improves  $\tau \gamma$  mass resolution by a factor 2)
- + With more time, we will provide more accurate simulation of the improvements coming from a better vertex resolution



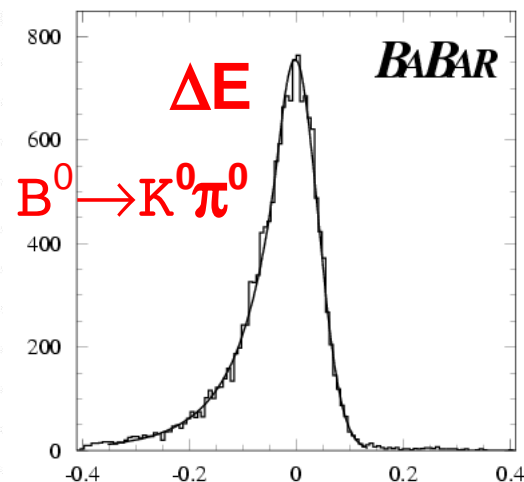
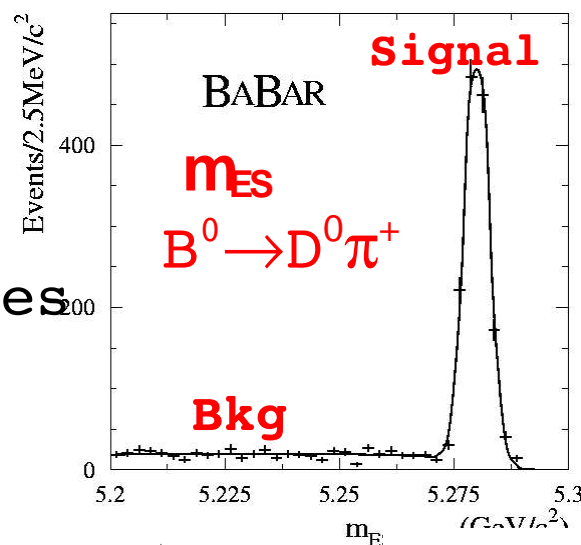
# A Fast Simulation of $b \rightarrow s$

The current analyses use three main ingredients

➤  $m_{ES} = \sqrt{(\sqrt{s}/2)^2 - p_B^{*2}}$

➤  $\Delta E = E_B^* - \sqrt{s}/2$

➤ Topological variables  
(Fisher, NN, ...)



We assume:

➤ Gaussian shapes for  $m_{ES}$  and  $\Delta E$

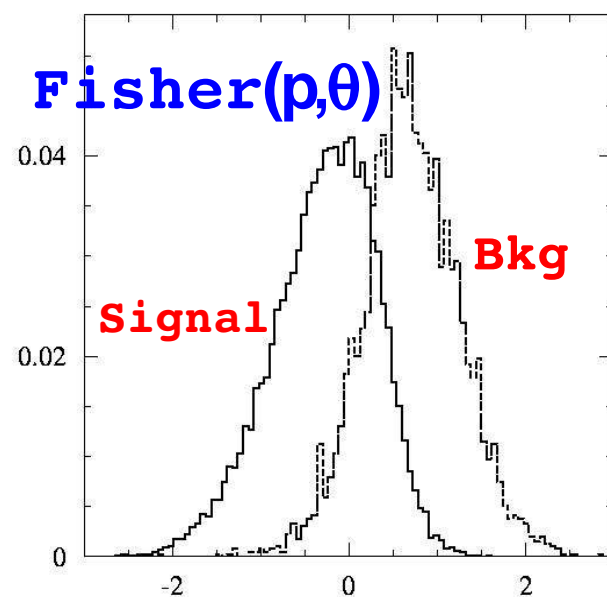
➤  $\sigma(m_{ES}) = \frac{1}{2}$  energy spread

➤  $\sigma(\Delta E) = \text{sqrt}((20\text{MeV})^2 + \sigma(m_{ES}))$

➤ We vary  $\sigma(\Delta t) \sim \sigma(\Delta z) / c\beta\gamma$

➤ Same bkg and signal Fisher shapes as now in BaBar

➤ No  $\bar{B}B$  background





# Time Dependence

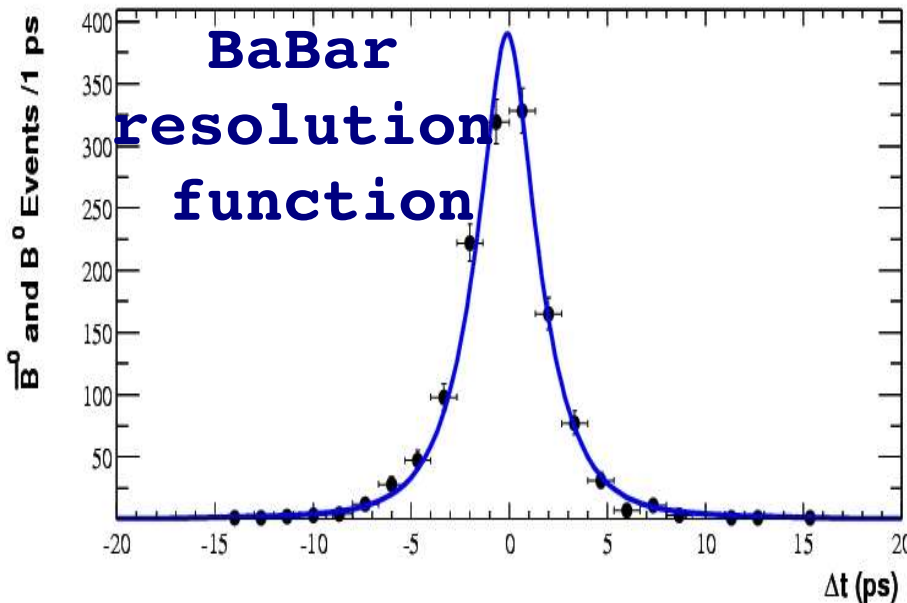
$\Delta t$  distribution comes from the convolution of the physics model and the Resolution Function (for signal) or from the Resolution Function alone (for background)

The **physics model is the same** (given by S and C)

Current BaBar RF is described by 3 Gaussians.

We assume a **single Gaussian**, centered in 0, for simplicity

We vary  $\sigma(\Delta t)$  from 0.1 to 0.9 ps<sup>-1</sup> (current value is 0.6)

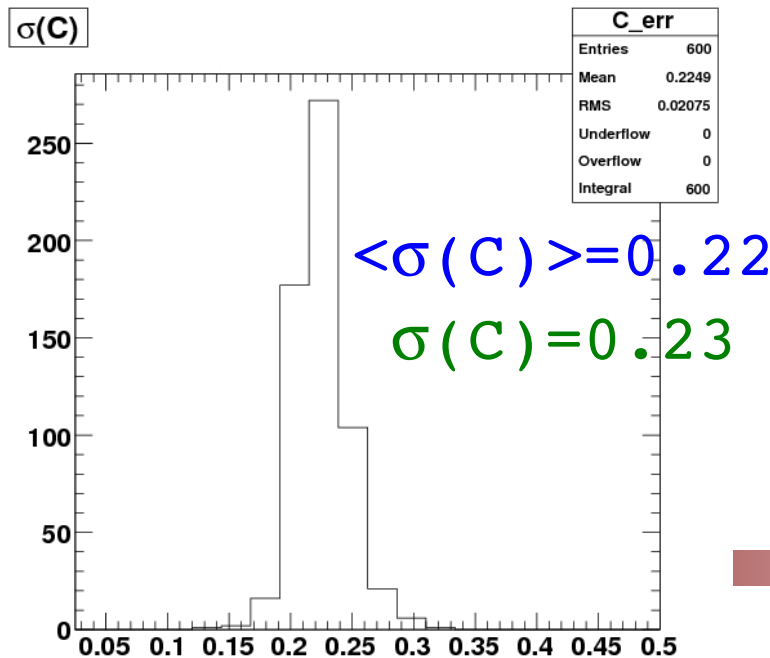
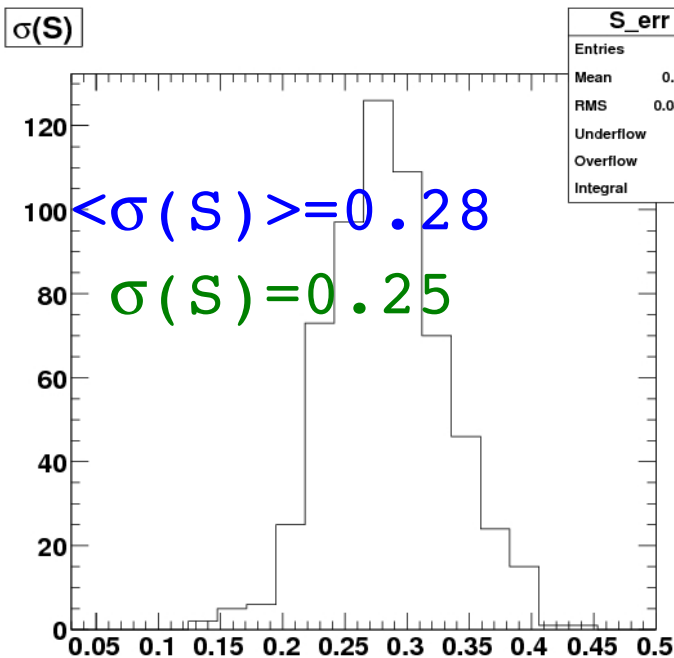
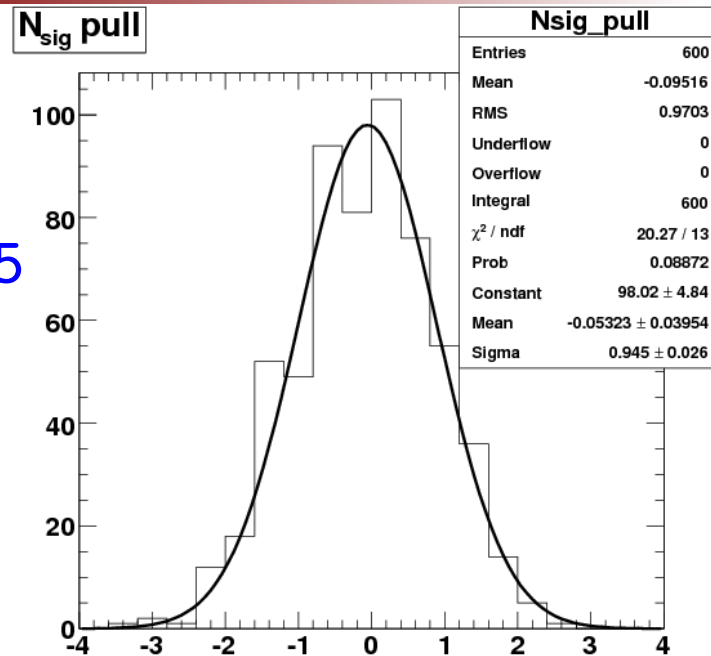
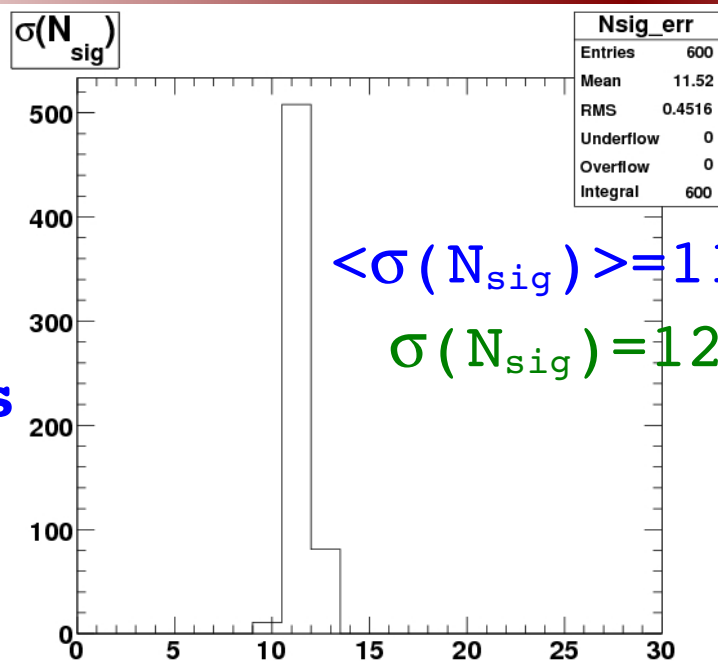


We are losing the accurate description of the tails, but we are retaining the core of the distribution

# Reproducing Current Analysis



Comparing our  
average results  
to BaBar  
analysis



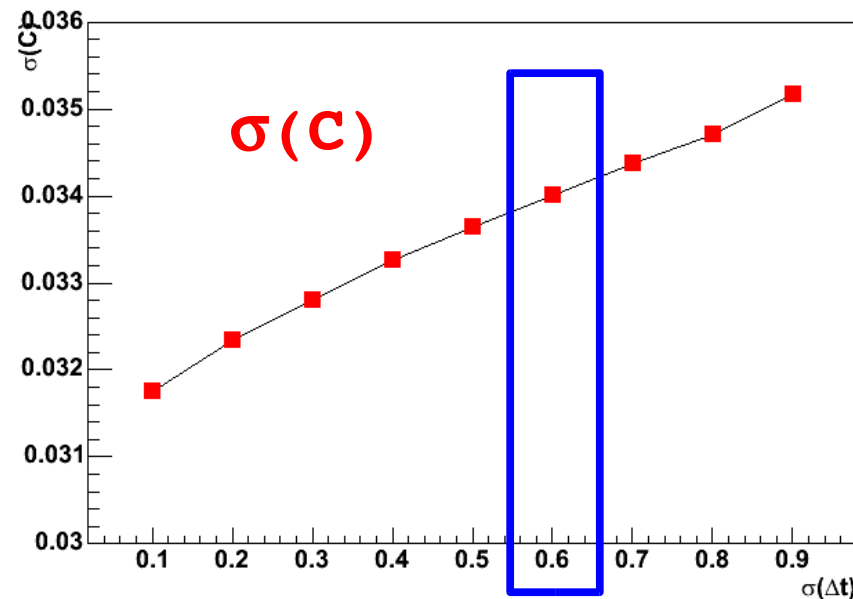
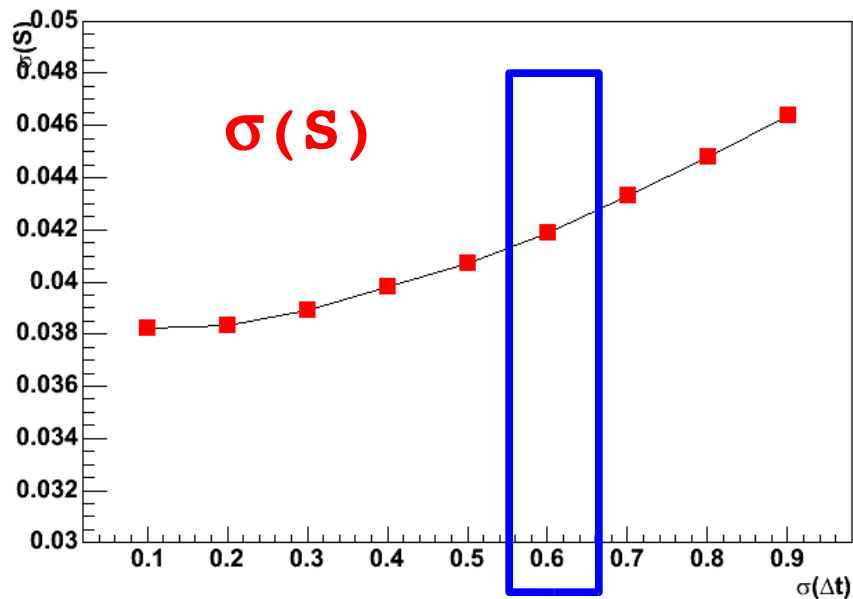
We are not  
simplifying  
things  
too much



# Going Up in Statistics

We repeat the previous exercise as a function of  $\sigma(\Delta t)$ , scaling signal and background yields to  $10 \text{ ab}^{-1}$  (we don't change energy spread for the moment).

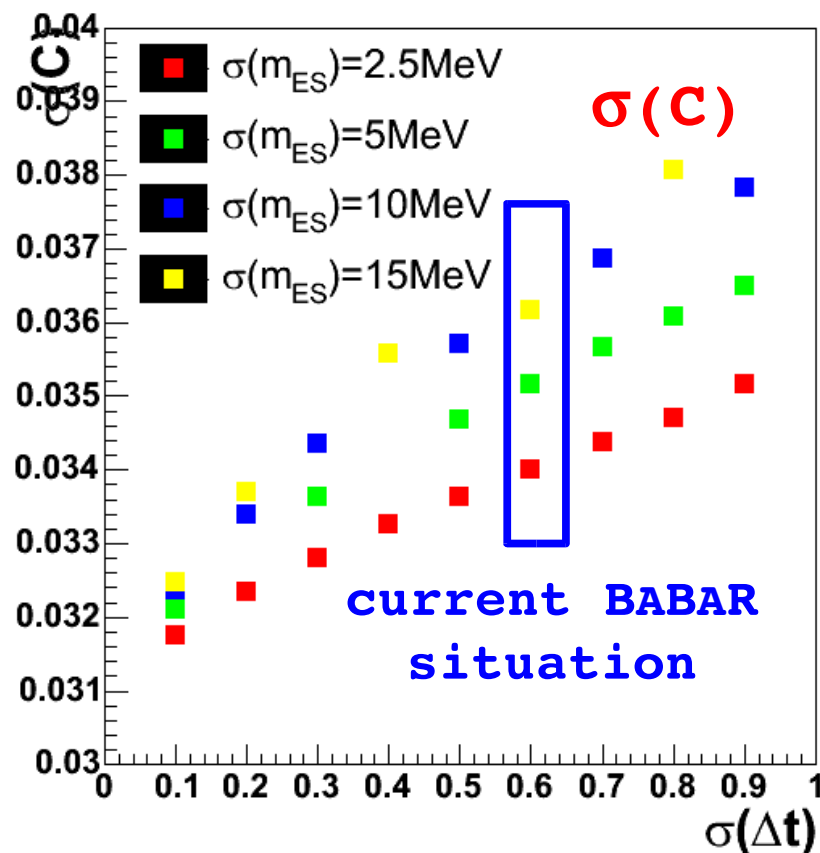
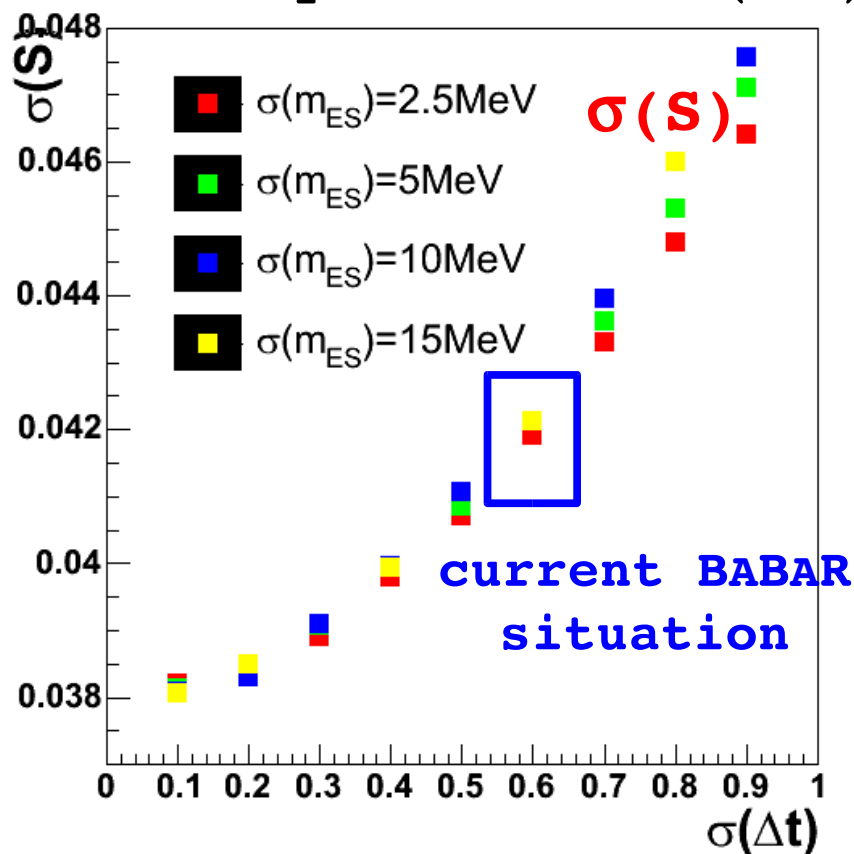
- ★ 114 signal events/  $211 \text{ fb}^{-1}$   $\rightarrow$  5K events/ $10 \text{ ab}^{-1}$
- ★ 4100 bkg events/ $211 \text{ fb}^{-1}$   $\rightarrow$  180K events/ $10 \text{ ab}^{-1}$





# The effect of Energy Spread

The test is redone increasing  $\sigma(m_{ES}) \sim \frac{1}{2}$  energy spread and  $\sigma(\Delta E) \sim \sqrt{(20\text{MeV})^2 + \sigma(m_{ES})^2}$



The good resolution on  $\Delta t$  compensates the reduction of background rejection because of increasing  $\sigma(m_{ES})$ .

It does not work for large  $\sigma(\Delta t)$





# Fast simulation of $\tau \rightarrow \mu \gamma$

The current analysis is based on a maximum likelihood fit on two kinematic variable

- +  $\tau$  mass (in the range [1.3, 2.3] GeV)
- +  $\Delta E$ , defined as before (in the range [-0.1, 0.2])

The  $\tau$  mass is evaluated after an energy-constrained vertexing

- ◆ The resolution is related to the energy spread
- ◆ The increase of vertex resolution can compensate the effect of increasing the energy spread

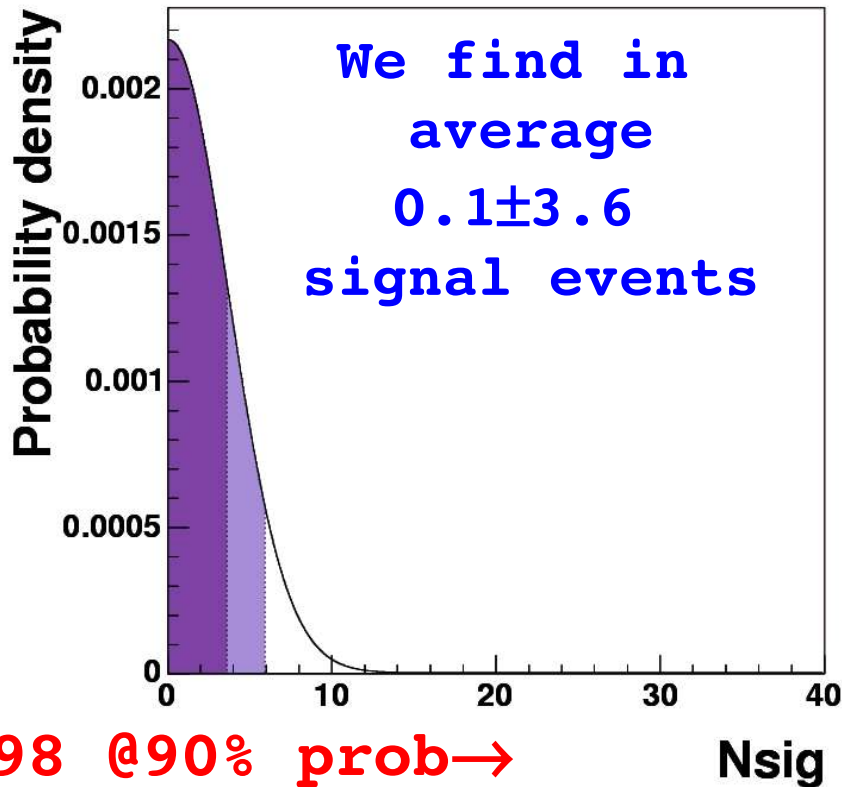
We simplify the analysis respect to what is done in BaBar

- No distinction between 3 prong and 1 prong events (now present because of poor statistics)
- We parameterize signal according to the outcome of Pravda
- We parameterize background according to the present analysis (forcing  $\Delta E$  pdf to be flat to simplify things)

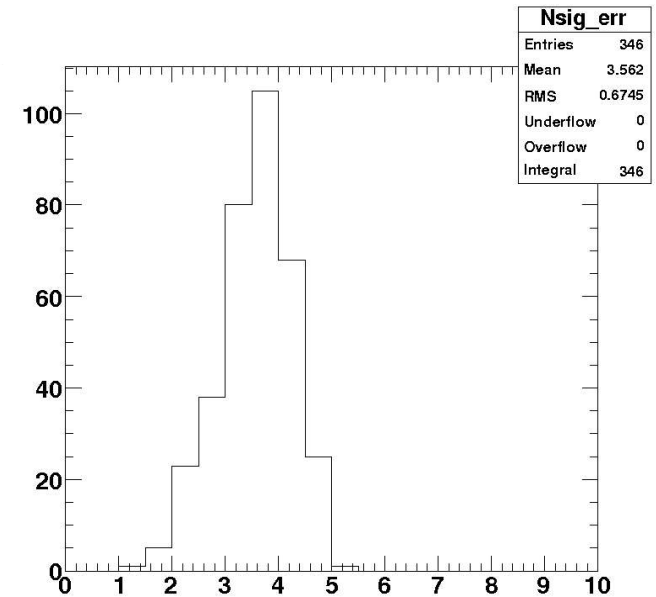
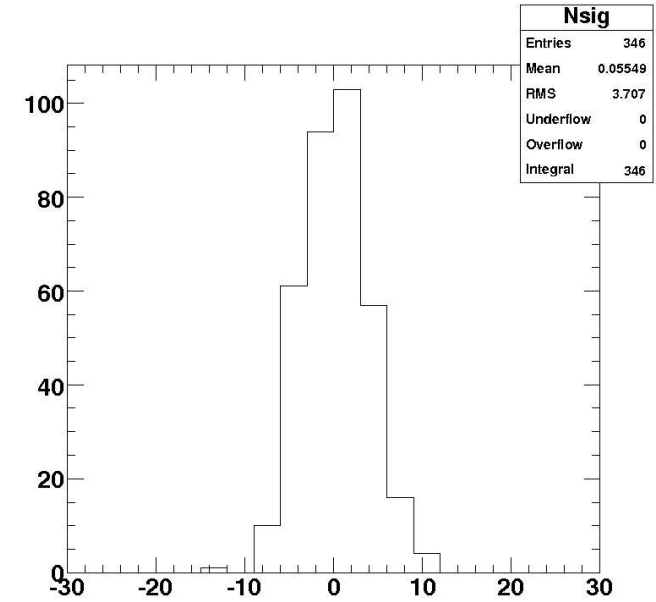


# Reproducing Current Analysis

We produce a set of Toy Monte Carlo experiments assuming 0 signal events and 442 background events



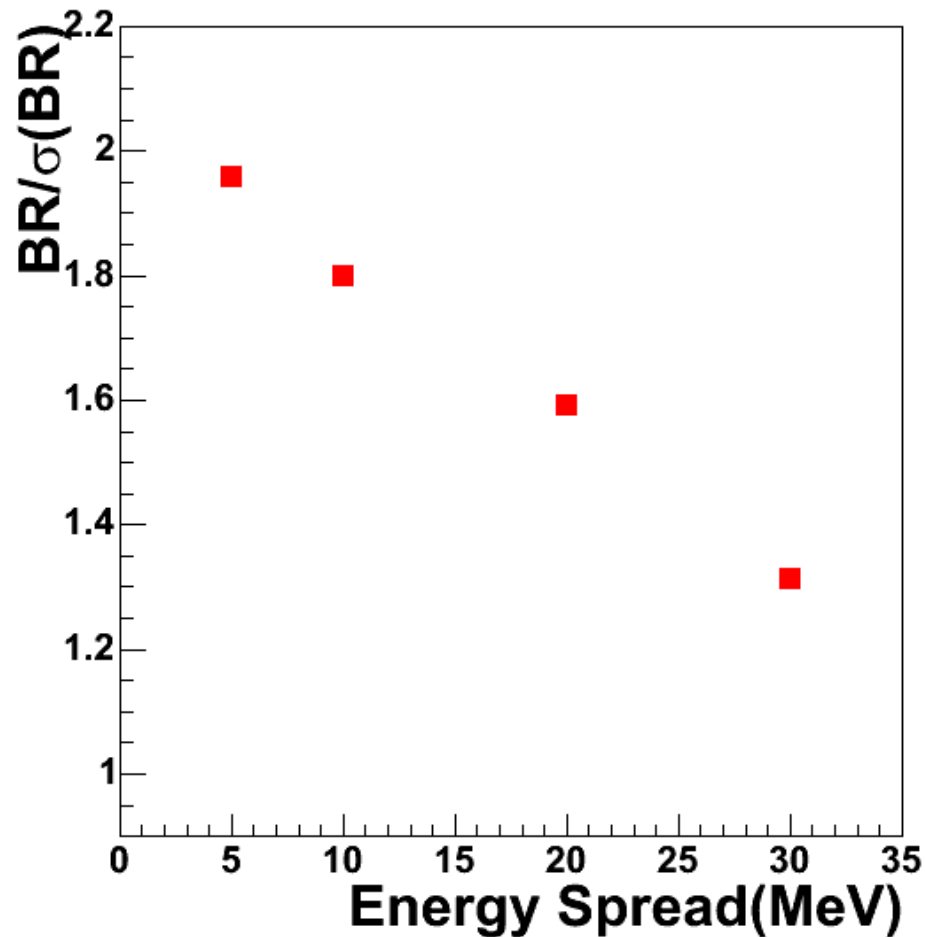
$N_{sig} < 5.98$  @ 90% prob  $\rightarrow$   
 $BR(\tau \rightarrow \mu\gamma) < 10^{-7}$  @ 90% prob  
(Reference analysis finds  $\sim 7 \pm 7$  events)





# Going Up in Statistics

Assuming 75 signal events and scaling bkg to  $50 \text{ ab}^{-1}$



- WARNING**: We are *very pessimistic*
- We make kinematic variables worse
  - But the actual resolution depends on a kinematic fit
  - We are neglecting (for the moment) the fact that even in this case the better vertex will reduce the impact of energy spread

**Work In Progress**



# Conclusion

Starting from the experience we have in BaBar, we can perform realistic fast simulations of superB factory performances using a Toy Monte Carlo approach.

This technique allows to simulate the effect of the energy spread on physics analyses

- In the case of  $\phi K_s$ , a good resolution on  $\Delta t$  can compensate the worse background rejection on kinematic variables.  $\sigma(S) \sim 0.039$ ,  $\sigma(C) \sim 0.033$  with  $10 \text{ ab}^{-1}$  (which means having an effect of  $\sim 5\sigma$  with the current central value!!!!). And we might have  $\sim 5$  times that /year.
- Simulation of  $\tau \rightarrow \mu \gamma$  still in progress. We need a detailed simulation to have a more accurate description of the impact of energy spread on non-Gaussian tails (present values are **PESSIMISTIC**)



# Backup Slides



# The effect of Energy Spread

The test is redone increasing  $\sigma(m_{ES}) = \frac{1}{2}$  energy spread and  $\sigma(\Delta E) = \sqrt{(20\text{MeV})^2 + \sigma(m_{ES})^2}$

