NuFact05 workshop

Resolving parameter degeneracies in long-baseline experiments with atmospheric neutrino data

Thomas Schwetz SISSA, Trieste

based on:

P. Huber, M. Maltoni, TS, PRD 71 (2005) 053006 [hep-ph/0501037]

T. Schwetz, NuFact05, Frascati, Italy, 21–26 june 2005 – p.1

Introduction

Goals for future neutrino oscillation experiments:

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- Type of the neutrino mass ordering (sign of Δm^2_{31})



Parameter degeneracies in LBL experiments

G.L. Fogli, E. Lisi, Phys. Rev. D54 (1996) 3667
J. Burguet-Castell et al., Nucl. Phys. B608 (2001) 301
M. Koike, T. Ota, J. Sato, Phys. Rev. D65 (2002) 053015
H. Minakata, H. Nunokawa, JHEP 10 (2001) 001
V.Barger, D.Marfatia, K.Whisnant, Phys. Rev. D65 (2002) 073023; D66 (2002) 053007
P.Huber, M.Lindner, W.Winter, Nucl. Phys. B645 (2002) 3; Nucl. Phys. B654 (2003) 3
J. Burguet-Castell et al., Nucl.Phys. B646 (2002) 301
O. Yasuda, New J. Phys. 6 (2004) 83
A.Donini, D.Meloni, S.Rigolin, JHEP 0406 (2004) 011

and many more (I appologize for omissions)

• The intrinsic or $(\delta_{CP}, \theta_{13})$ degeneracy J. Burguet-Castell et al., Nucl. Phys. B608 (2001) 301 several solutions in the $(\delta_{CP}, \theta_{13})$ plane

 $P_{\mu e} \simeq \sin^2 2\theta_{13} \sin^2 \theta_{23} \sin^2 \Delta_{31} + \alpha^2 \sin^2 2\theta_{12} \cos^2 \theta_{23} \Delta_{31}^2$ $+ \alpha \sin 2\theta_{12} \sin 2\theta_{13} \sin 2\theta_{23} \Delta_{31} \sin \Delta_{31} \cos(\Delta_{31} \pm \delta_{CP})$

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- The hierarchy or sgn(Δm_{31}^2) degeneracy H. Minakata, H. Nunokawa, JHEP 10 (2001) 001

solutions for both signs of Δm_{31}^2 (affects mainly δ_{CP})

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 ν_{μ} -disappearance chanel gives only $\sin^2 2\theta_{23}$ solutions for θ_{23} and $\pi/2 - \theta_{23}$ (affects mainly $\sin^2 2\theta_{13}$)

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overall an 8-fold degeneracy

V.Barger, D.Marfatia, K.Whisnant, Phys. Rev. D65 (2002) 073023

The T2K-II long-baseline experiment

4 MW superbeam at JPARC mean neutrino energy: 0.76 GeV (2° OA) 1 Mt Cherenkov detector at Kamioka baseline: 295 km

| | ν (2 Mt yrs) | $\bar{\nu}$ (6 Mt yrs) |
|---|---------------------|------------------------|
| $ u_{\mu} \rightarrow \nu_{e} \text{ signal} $ | 21 300 | 16 000 |
| $ u_{\mu} \rightarrow u_{e} \text{ background} $ | 2 1 4 0 | 3 260 |
| $ u_{\mu} ightarrow u_{\mu}$ signal | 73 200 | 75 600 |
| $ u_{\mu} ightarrow u_{\mu}$ background | 340 | 320 |

 $\sin^2 2\theta_{13} = 0.05, \ \sin^2 \theta_{23} = 0.5, \ \sin^2 \theta_{12} = 0.3, \ \delta_{\rm CP} = 0,$ $\Delta m_{21}^2 = 8.1 \times 10^{-5} \ {\rm eV}^2, \ \Delta m_{31}^2 = 2.2 \times 10^{-3} \ {\rm eV}^2$

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GLoBES software P. Huber, M. Lindner, W. Winter, hep-ph/0407333

http://www.ph.tum.de/~globes/



allowed regions at 2σ , 99%, 3σ CL

true values: $\sin^2 2\theta_{13} = 0.03$ $\delta_{\rm CP} = -0.85\pi$ $\sin^2 \theta_{23} = 0.4$ $\Delta m_{31}^2 = 2.2 \times 10^{-3} {\rm eV}^2$

The (δ_{CP} , θ_{13}) degeneracy is not present for T2K-II because of spectral information



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ambiguities in θ_{13} and δ_{CP} no information on the hierarchy

several possibilities to resolve the degeneracies are known:

- combining information from detectors at different baselines
- using additional oscillation chanels ($\nu_e \rightarrow \nu_{\tau}$)
- spectral information (broadband beam)
- adding information on θ_{13} from a reactor experiment

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we propose a new method based on ...

3-flavour effects in atmospheric neutrinos

Petcov, Phys. Lett. B434, 321 (1998), hep-ph/9805262 Akhmedov, Nucl. Phys. B538, 25 (1999), hep-ph/9805272 Akhmedov, Dighe, Lipari, Smirnov, Nucl. Phys. B542, 3 (1999), hep-ph/9808270 Kim, Lee, Phys. Lett. B444, 204 (1998), hep-ph/9809491 Bernabeu, Palomares-Ruiz, Perez, Petcov, Phys. Lett. B531, 90 (2002), hep-ph/0110071 Bernabeu, Palomares-Ruiz, Petcov, Nucl. Phys. B669, 255 (2003), hep-ph/0305152 Peres, Smirnov, Phys. Lett. B456, 204 (1999), hep-ph/9902312 Peres, Smirnov, Nucl. Phys. B680, 479 (2004), hep-ph/0309312 Gonzalez-Garcia, Maltoni, Eur. Phys. J. C26, 417 (2003), hep-ph/0202218 Gonzalez-Garcia, Maltoni, Smirnov, Phys. Rev. D70, 093005 (2004), hep-ph/0408170

again I appologize for omissions

3-flavour effects in atmospheric neutrinos

excess of electron-like events:

$$\begin{split} \frac{N_e}{N_e^0} - 1 \simeq & (r \, s_{23}^2 - 1) \, P_{2\nu}(\Delta m_{31}^2, \theta_{13}) & \theta_{13} \text{-effects} \\ & + & (r \, c_{23}^2 - 1) \, P_{2\nu}(\Delta m_{21}^2, \theta_{12}) & \Delta m_{21}^2 \text{-effects} \\ & - & 2s_{13}s_{23}c_{23} \, r \, \text{Re}(A_{ee}^* A_{\mu e}) & \text{interference: } \delta_{\text{CP}} \end{split}$$

$$r = r(E_{\nu}) \equiv \frac{F_{\mu}^{0}(E_{\nu})}{F_{e}^{0}(E_{\nu})} \qquad r \approx 2 \quad \text{(sub-GeV)}$$
$$r \approx 2.6 - 4.5 \quad \text{(multi-GeV)}$$

 θ_{13} -effects

$$\frac{N_e}{N_e^0} - 1 \simeq (r \, s_{23}^2 - 1) \, P_{2\nu}(\Delta m_{31}^2, \theta_{13})$$

resonant matter effect in $P_{2\nu}(\Delta m_{31}^2, \theta_{13})$ for multi-GeV events ($r \approx 2.6 - 4.5$)

normal hierarchy: enhancement for neutrinos inverted hierarchy: enhancement for anti-neutrinos

detection cross sections are different for neutrinos and anti-neutrinos

sensitivity to the neutrino mass hierarchy

 θ_{13} -effects

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Bernabeu, Palomares-Ruiz, Petcov, Nucl. Phys. B669, 255 (2003), hep-ph/0305152

 Δm_{21}^2 -effects

 $\frac{N_e}{N^0} - 1 \simeq (r c_{23}^2 - 1) P_{2\nu}(\Delta m_{21}^2, \theta_{12})$

Peres, Smirnov, hep-ph/0309312

contours of $\frac{N_e}{N_e^0} - 1$

relevant for sub-GeV events

sensitivity to the octant of θ_{23}

3-flavour effects in atmospheric neutrinos

plot from T. Kajita

The HK atmospheric neutrino experiment

assume 9 Mt yrs ATM data (100 \times SK-I data)

| | zenith angle | ν | $ar{ u}$ |
|-----------------------|---|---------|----------|
| e-like sub-GeV | 10 bins | 239 000 | 58 000 |
| e-like multi-GeV | 10 bins | 52700 | 18100 |
| μ -like sub-GeV | 10 bins | 232 000 | 66 200 |
| μ -like multi-GeV | 10 bins | 108 000 | 49100 |
| upward going μ | 10_{thr} + 5_{st} | 127 000 | 65 400 |

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WARNING:

- same systematics as SK-I
- same binning (zenith angle, energy) as SK-I
- only single-ring events

The ATM analysis

• Full numerical three-flavour analysis

both \$\Delta m_{31}^2\$ and \$\Delta m_{21}^2\$ taken into account
realistic treatment of earth matter effects

based on:

Gonzalez-Garcia, Maltoni, Pena-Garay, Valle, Phys. Rev. D **63** (2001) 033005 Gonzalez-Garcia, Maltoni, Eur. Phys. J. C **26** (2003) 417 [hep-ph/0202218] Maltoni, TS, Tortola, Valle, Phys. Rev. D **67** (2003) 013011 [hep-ph/0207227] Gonzalez-Garcia, Maltoni, Smirnov, Phys. Rev. D **70** (2004) 093005 [hep-ph/0408170]

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Combined with LBL data by using a generalized version of the GLoBES software

solid curves: LBL only allowed regions at 2σ , 99% and 3σ CL (2 dof) true values: $\sin^2 2\theta_{13} = 0.03$, $\delta_{\rm CP} = -0.85\pi$, $\Delta m_{31}^2 = 2.2 \cdot 10^{-3} {\rm eV}^2$

solid curves: LBL only, colored regions: LBL+ATM allowed regions at 2σ , 99% and 3σ CL (2 dof) true values: $\sin^2 2\theta_{13} = 0.03$, $\delta_{\rm CP} = -0.85\pi$, $\Delta m_{31}^2 = 2.2 \cdot 10^{-3} {\rm eV}^2$

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Which are the relevant ATM data samples?

Do we really need a Mt experiment?

Luminosity scaling

CERN-Frejus LBL experiments (PRELIMINARY)

work in progress in collaboration with J.E. Campagne and M. Mezzetto

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Beta Beam

see talk of M. Mezzetto for details $\bar{\nu}$: ⁶He ($\gamma = 100$), ν : ¹⁸Ne ($\gamma = 100$), 10 yrs data

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SPL Superbeam

see talk of J.E. Campagne for details

3.5 GeV proton beam from 4 MW SPL, 2 yrs ν , 8 yrs $\bar{\nu}$

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see talk of J.E. Campagne for details 3.5 GeV proton beam from 4 MW SPL, 2 yrs ν , 8 yrs $\bar{\nu}$

 450 kt water Cherenkov detector at Frejus baseline 130 km

sensitivity to the mass hierarchy

Concluding remarks

 sensitivity to the neutrino mass ordering significantly increased

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- good sensitivity to the octant of θ_{23}

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given the Mt detector for the LBL experiment, ATM data come for free!

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Complementarity of LBL and ATM data:

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Complementarity of LBL and ATM data:

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Thank you for your attention!

P. Huber, M. Maltoni, TS, PRD 71 (2005) 053006 [hep-ph/0501037]

additional slides

 $\sin^2 2\theta_{13} = 0.01$ $\delta_{\rm CP} = \pi/4$ $\sin^2 \theta_{23} = 0.3$

$$R = \frac{N_i^{\rm tr} - N_i^{\rm deg}}{\sqrt{(N_i^{\rm tr} + N_i^{\rm deg})/2}}$$

P.Huber, M.Lindner, W.Winter, Nucl. Phys. B645 (2002) 3

The intrinsic degeneracy is absent for T2K-II

Additional slides

True $\theta_{13} = 0$

Resolving the octant-degeneracy:

True $\theta_{13} = 0$

The limit on $\sin^2 2\theta_{13}$:

dashed: LBL only, solid: LBL+ATM

$$P_{\mu e} \simeq \sin^2 2\theta_{13} \sin^2 \theta_{23} \sin^2 \Delta_{31} + \dots$$