



# $\nu_\mu$ Disappearance at SPL, T2K-I and the NuFactory

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- Introduction
  - Oscillation parameters
  - Experiments description
- SPL vs T2K-I
- Subleading effects in  $\nu_\mu$  disappearance
  - $\Delta m^2_{\text{atm}}$  and the sign degeneracy
  - $\theta_{23}$  and the octant degeneracy
  - The effects of  $\theta_{13}$  and  $\delta$
- T2K-I bounds revised
- The Neutrino Factory
- Conclusions



# The oscillation parameters

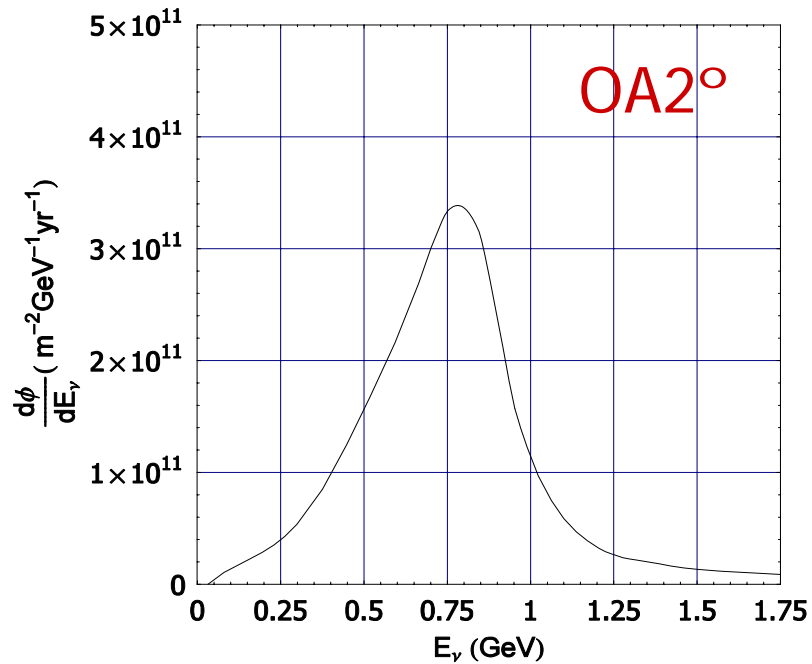
- What we already know (at  $3\sigma$ )

- Solar sector  $\begin{cases} \Delta m_{12}^2 = 8.2_{-0.9}^{+1.1} \cdot 10^{-5} \text{ eV}^2 \\ \tan^2 \theta_{12} = 0.39_{-0.11}^{+0.21} \end{cases} \quad \theta_{12} = 28^\circ - 38^\circ$
- Atm sector  $\begin{cases} |\Delta m_{23}^2| = 2.2_{-0.6}^{+1.4} \cdot 10^{-3} \text{ eV}^2 \\ \tan^2 \theta_{23} = 1_{-0.5}^{+1.1} \end{cases} \quad \theta_{23} = 35^\circ - 55^\circ$

- What we still do not know

- $\sin^2 2\theta_{13} < 0.16$   $\theta_{13} < 11.5^\circ$
- $\delta_{cp}$
- Mass hierarchy  $s_{atm} = \text{sign}(\Delta m_{23}^2)$
- Octant of  $\theta_{23}$   $s_{oct} = \text{sign}[\tan(2\theta_{23})]$

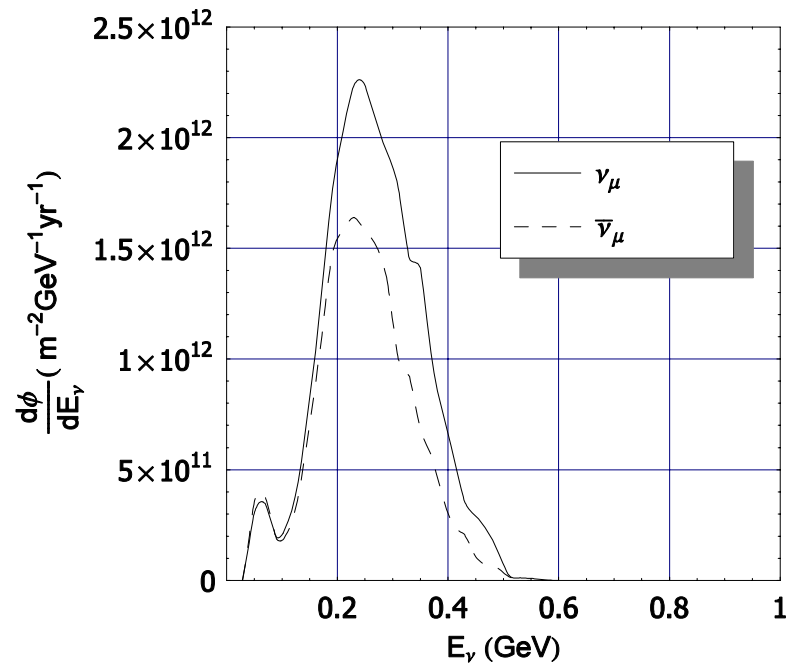
T2K-I



$L=295\text{Km}$

$\nu_{\mu}$  flux from  $\pi^+$  decay at  $\langle E_{\nu} \rangle = 0.75\text{GeV}$

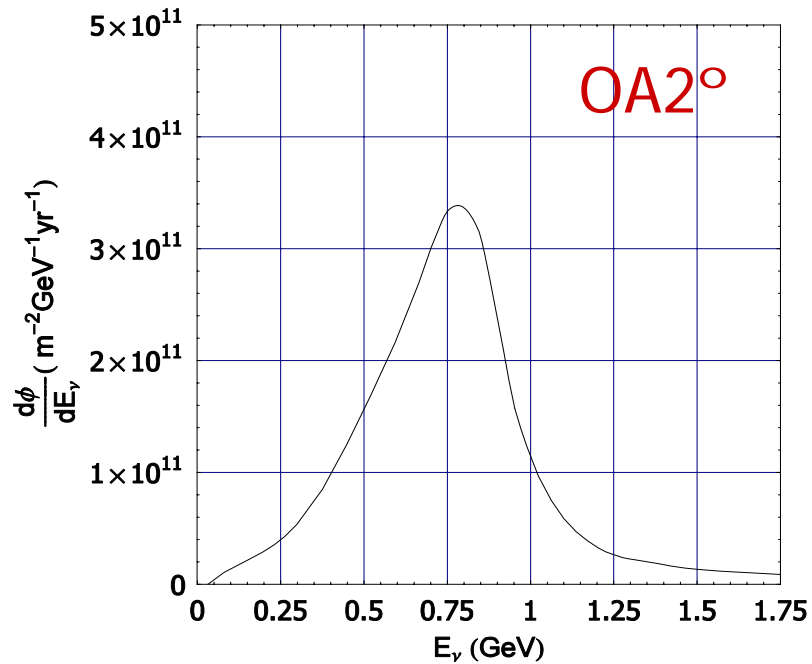
SPL



$L=130\text{Km}$

$\nu_{\mu}$  flux from  $\pi^+$  decay  $\langle E_{\nu} \rangle = 0.27\text{GeV}$   
 $\bar{\nu}_{\mu}$  flux from  $\pi^-$  decay  $\langle E_{\bar{\nu}} \rangle = 0.25\text{GeV}$

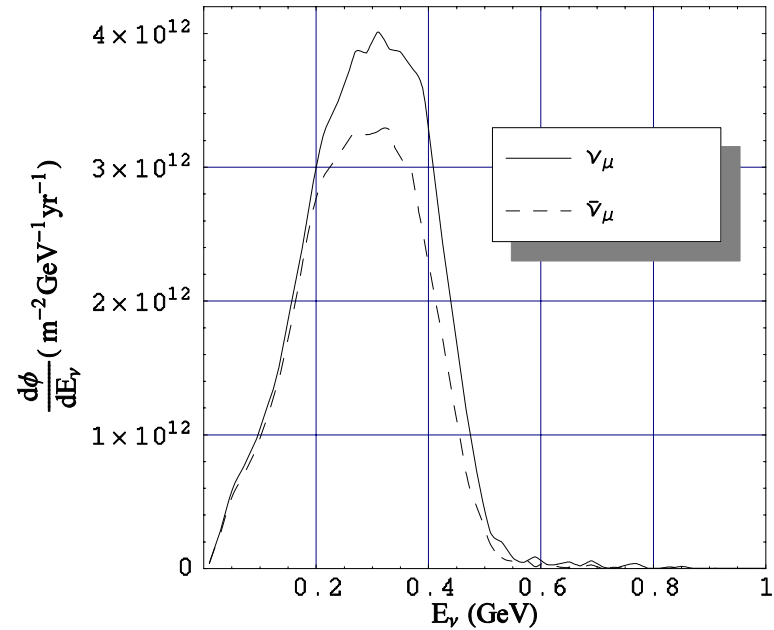
T2K-I



L=295Km

$\nu_{\mu}$  flux from  $\pi^+$  decay at  $\langle E_{\nu} \rangle = 0.75 GeV$

SPL



L=130Km

$\nu_{\mu}$  flux from  $\pi^+$  decay  $\langle E_{\nu} \rangle = 0.29 GeV$   
 $\bar{\nu}_{\mu}$  flux from  $\pi^-$  decay  $\langle E_{\bar{\nu}} \rangle = 0.28 GeV$



# Event Rates

T2K-I	B1	B2	B3	B4
No osc. $N_{\mu}$	753	2228	2273	757
Signal $N_{\mu}$	46	101	381	239

SPL	$\mu^{-}$	$\mu^{+}$
No osc. $N_{\mu}$	24245	25467
Signal $N_{\mu}$	1746	1614

4 energy bins of 200MeV  
Between 0.4 – 1.2GeV

L=295Km

Statistics dominated

L=130Km

Systematic dominated

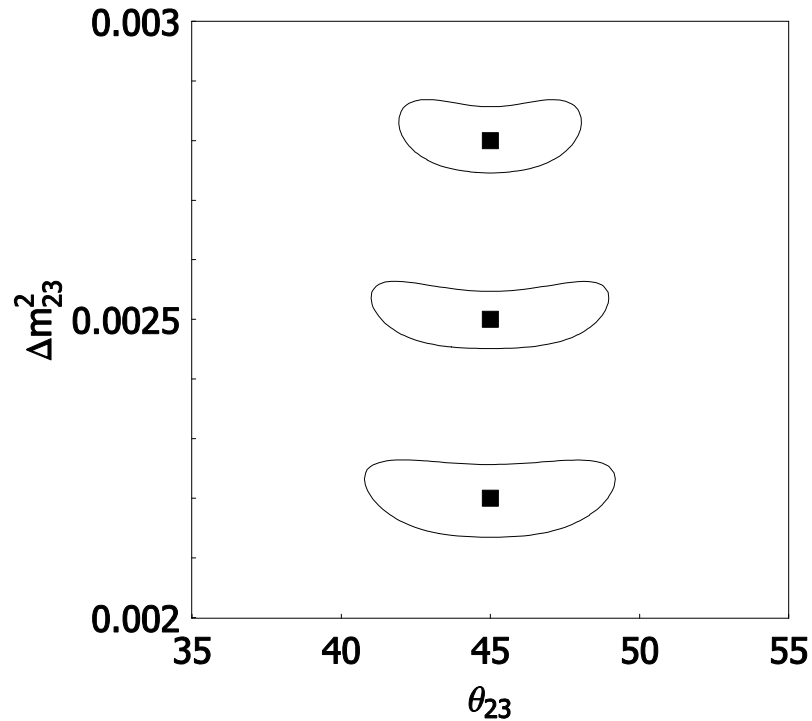
5yr  $\nu_{\mu}$  exposure with a 22.5Kt water cerenkov detector for T2K-I

2yr  $\nu_{\mu}$  + 8yr  $\bar{\nu}_{\mu}$  exposure with a 440Kt water cerenkov detector for the SPL

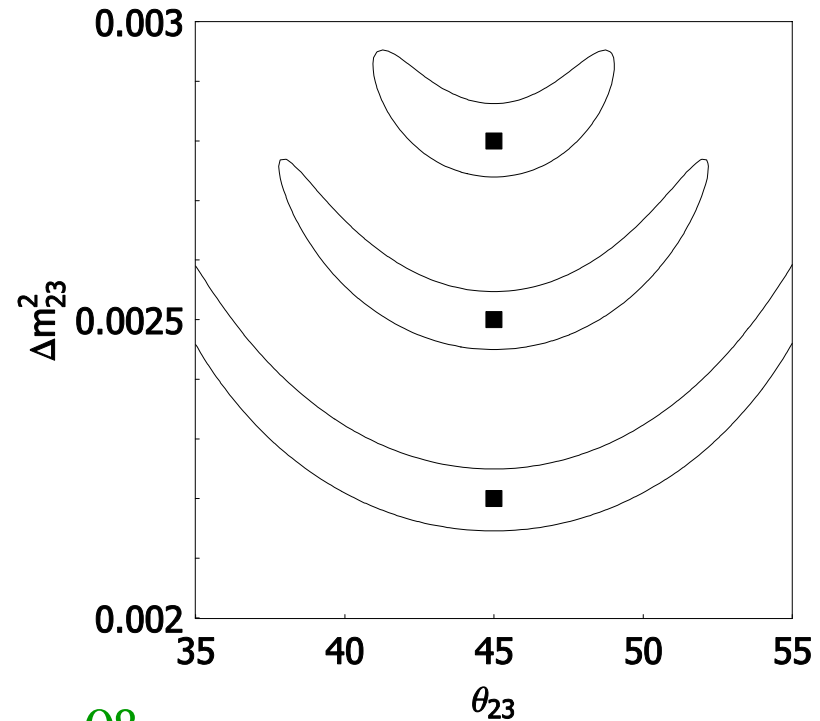


# The importance of energy resolution

T2K-I



SPL



$$\theta_{13} = 0^\circ$$

$$\delta = 0^\circ$$

90% CL contours, 2 dofs

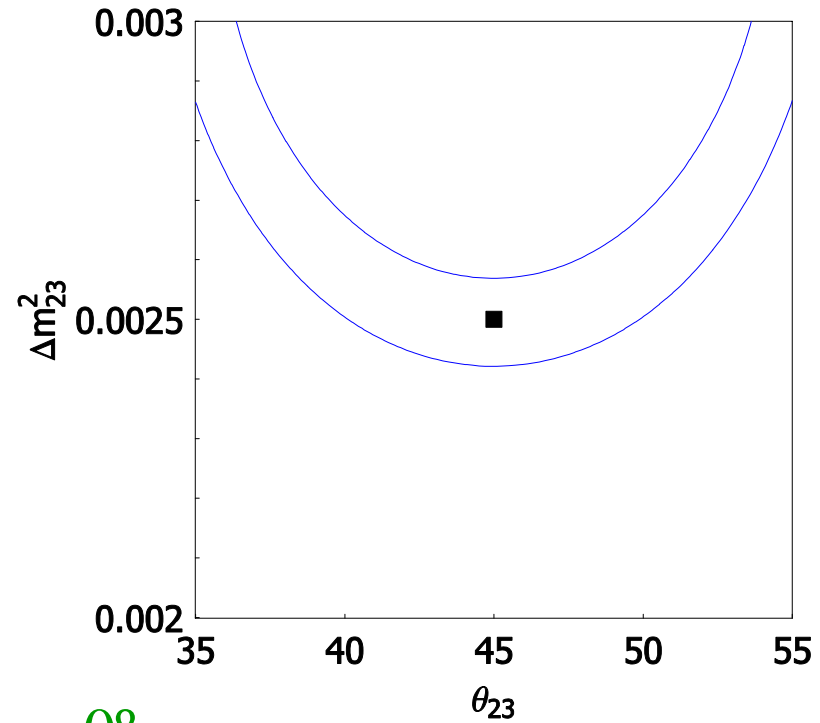
5% systematic error and backgrounds taken into account



# The importance of energy resolution

T2K-I

SPL



$$\theta_{13} = 0^\circ$$

$$\delta = 0^\circ$$

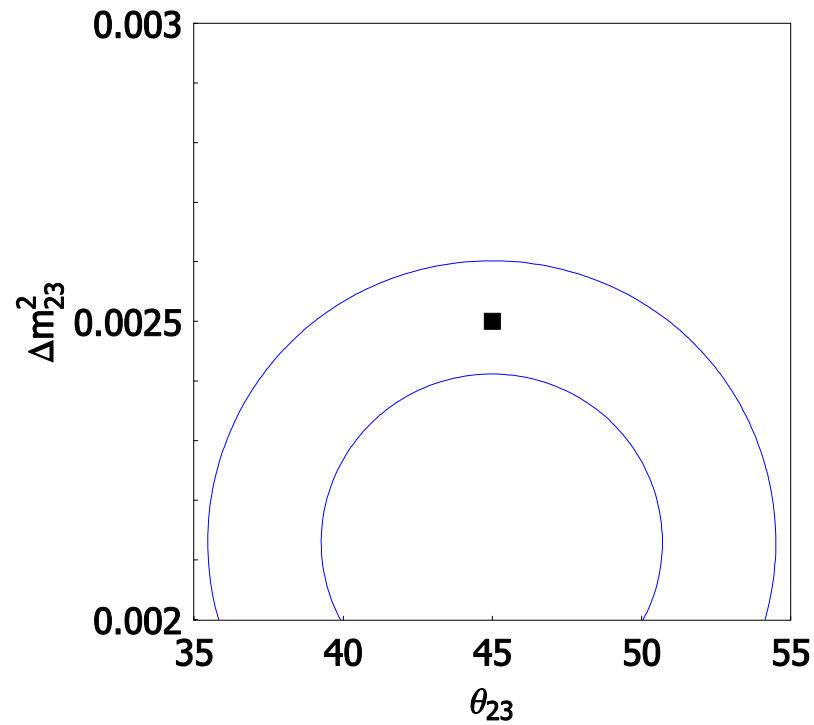
$$\langle E_\nu \rangle = 0.27 \text{ GeV}$$





# The importance of energy resolution

T2K-I

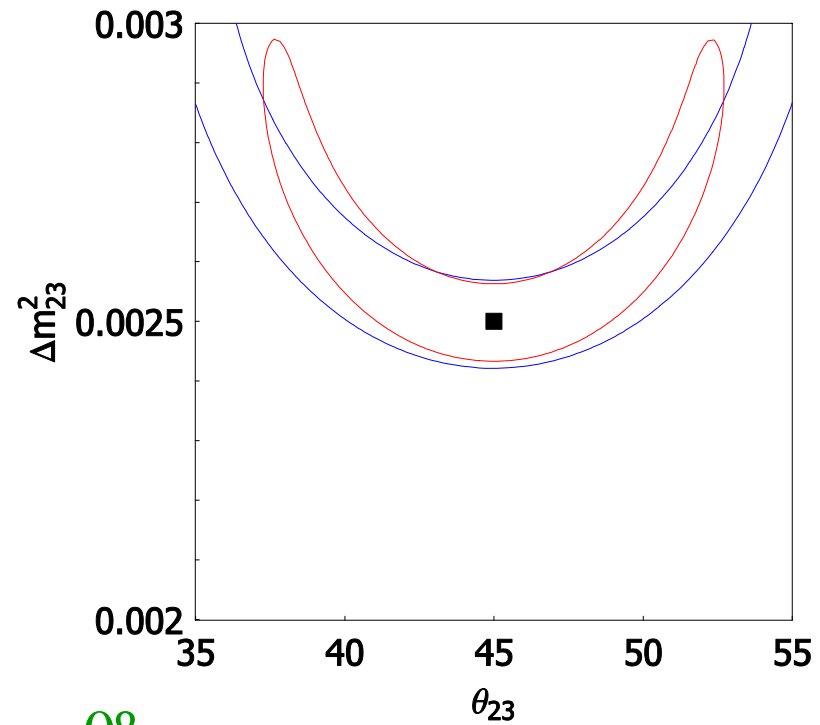


$E_1 = 0.4 - 0.6 \text{ GeV}$

$$\theta_{13} = 0^\circ$$

$$\delta = 0^\circ$$

SPL



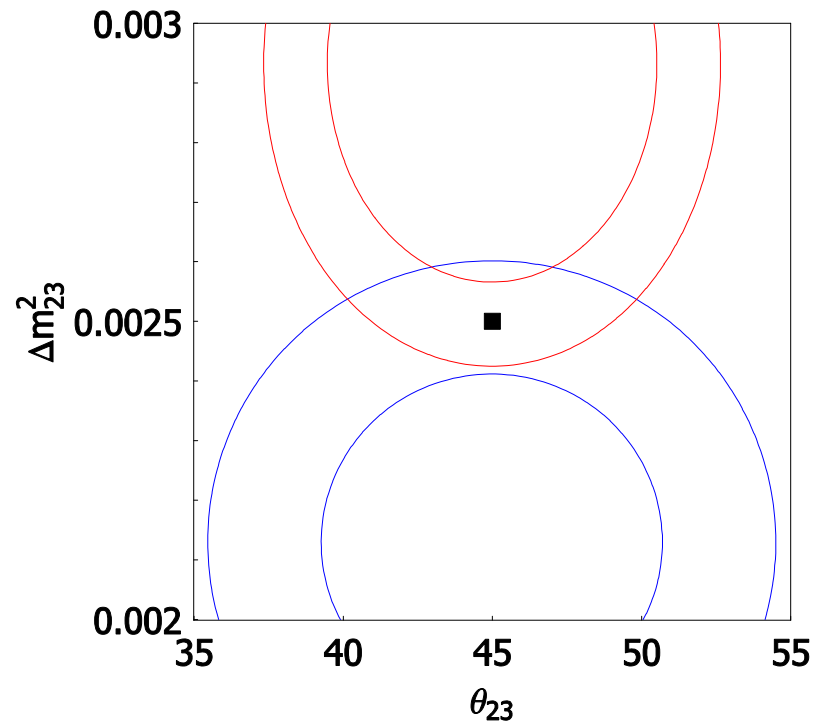
$$\langle E_\nu \rangle = 0.27 \text{ GeV}$$

$$\langle E_\nu \rangle = 0.25 \text{ GeV}$$



# The importance of energy resolution

T2K-I



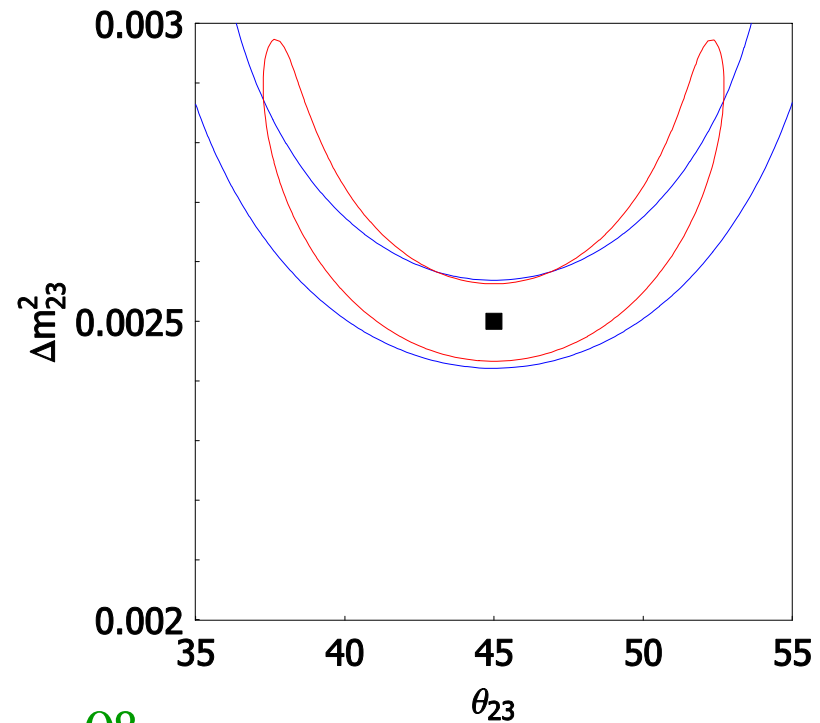
$E1 = 0.4 - 0.6 \text{ GeV}$

$E2 = 0.6 - 0.8 \text{ GeV}$

$\theta_{13} = 0^\circ$

$\delta = 0^\circ$

SPL



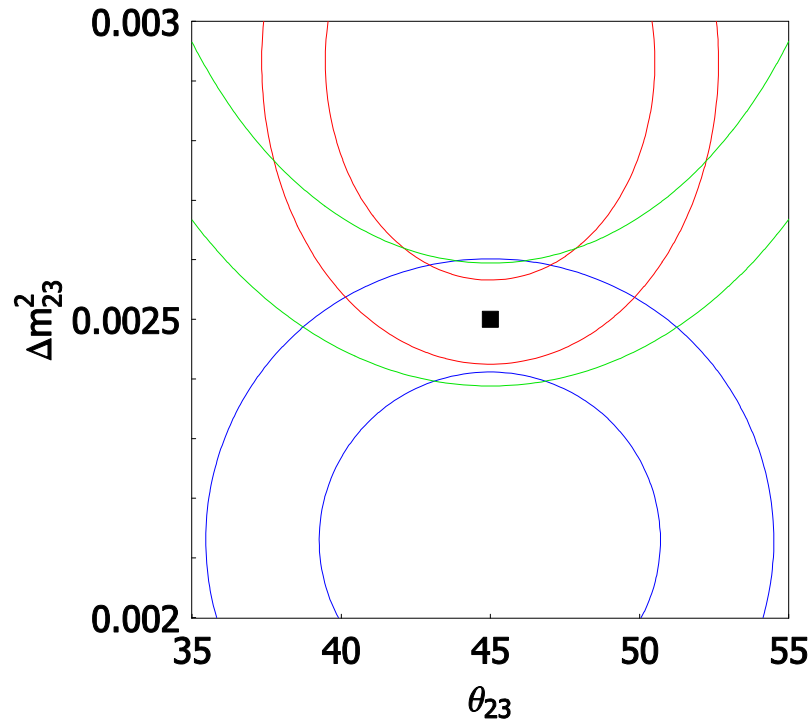
$\langle E_{\nu} \rangle = 0.27 \text{ GeV}$

$\langle E_{\bar{\nu}} \rangle = 0.25 \text{ GeV}$



# The importance of energy resolution

T2K-I



E1 = 0.4 - 0.6 GeV

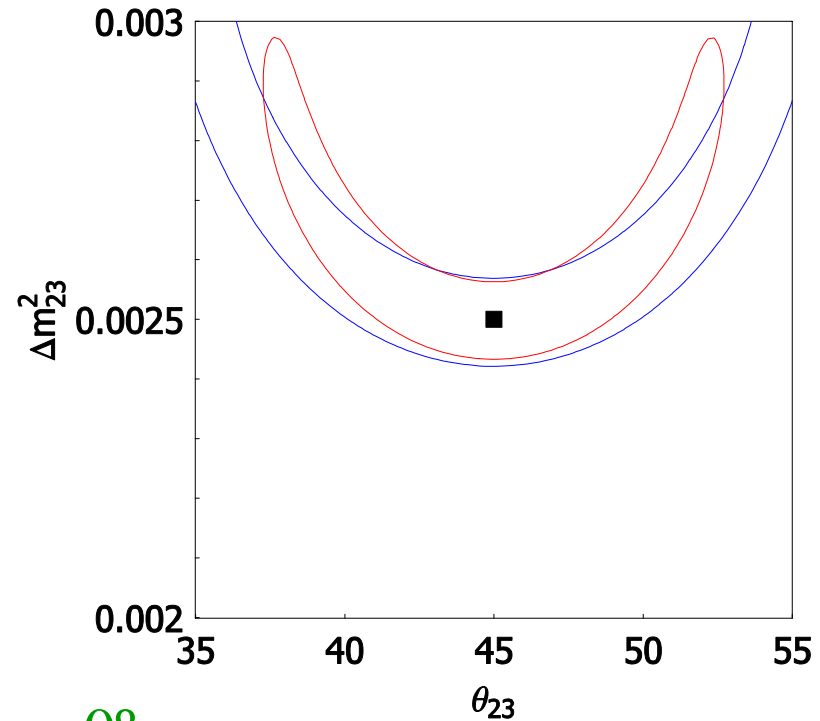
E2 = 0.6 - 0.8 GeV

E3 = 0.8 - 1.0 GeV

$\theta_{13} = 0^\circ$

$\delta = 0^\circ$

SPL



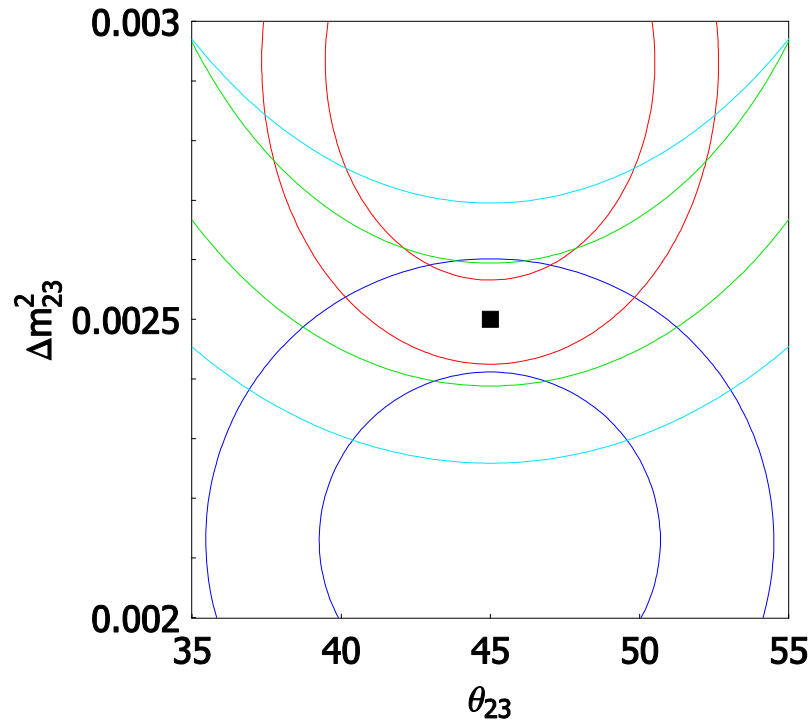
$\langle E_\nu \rangle = 0.27 \text{ GeV}$

$\langle E_{\bar{\nu}} \rangle = 0.25 \text{ GeV}$



# The importance of energy resolution

T2K-I



E1 = 0.4 - 0.6 GeV

E2 = 0.6 - 0.8 GeV

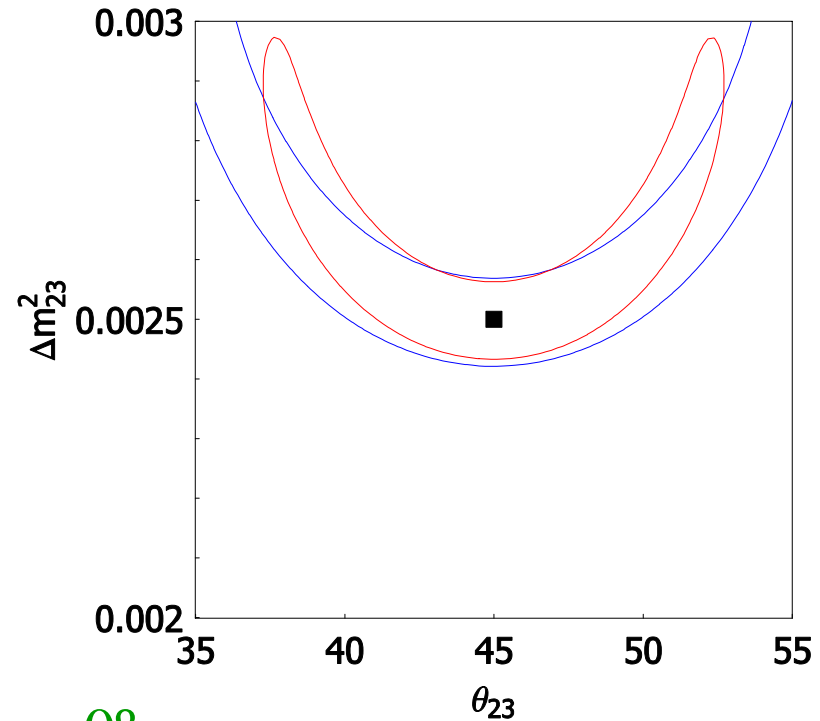
E3 = 0.8 - 1.0 GeV

E4 = 1.0 - 1.2 GeV

$\theta_{13} = 0^\circ$

$\delta = 0^\circ$

SPL



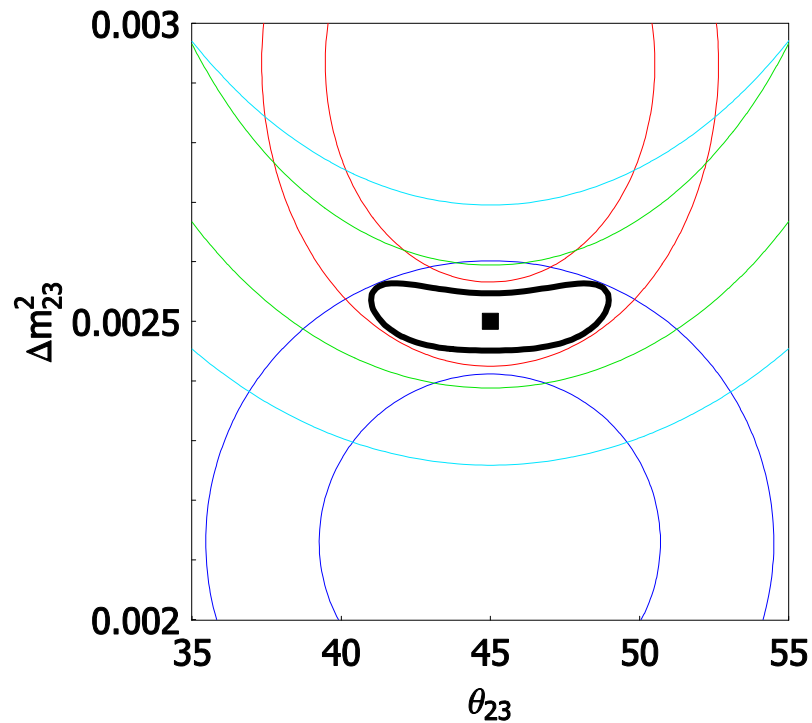
$\langle E_\nu \rangle = 0.27 \text{ GeV}$

$\langle E_{\bar{\nu}} \rangle = 0.25 \text{ GeV}$



# The importance of energy resolution

T2K-I



E1 = 0.4 - 0.6 GeV

E2 = 0.6 - 0.8 GeV

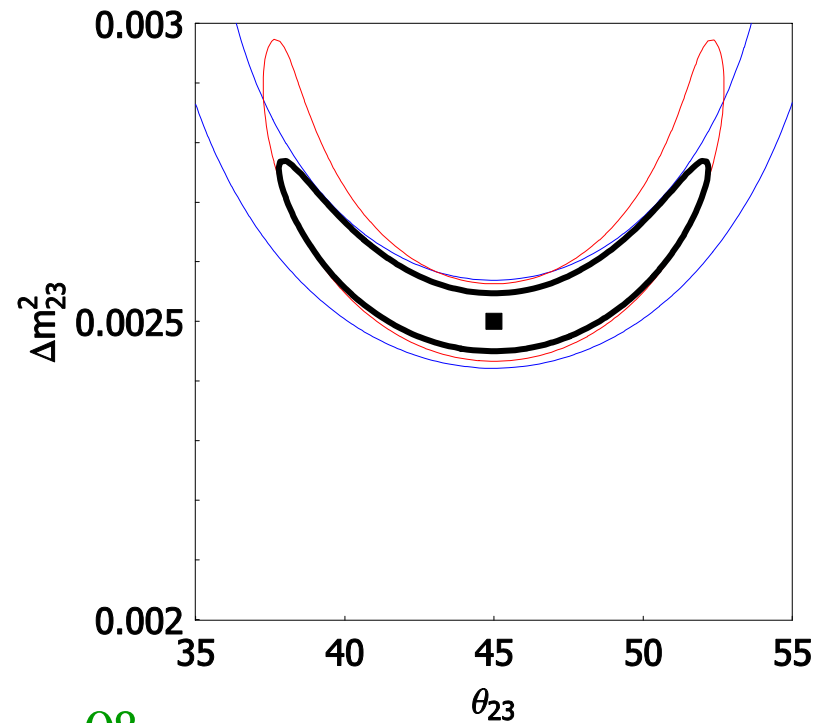
E3 = 0.8 - 1.0 GeV

E4 = 1.0 - 1.2 GeV

$\theta_{13} = 0^\circ$

$\delta = 0^\circ$

SPL



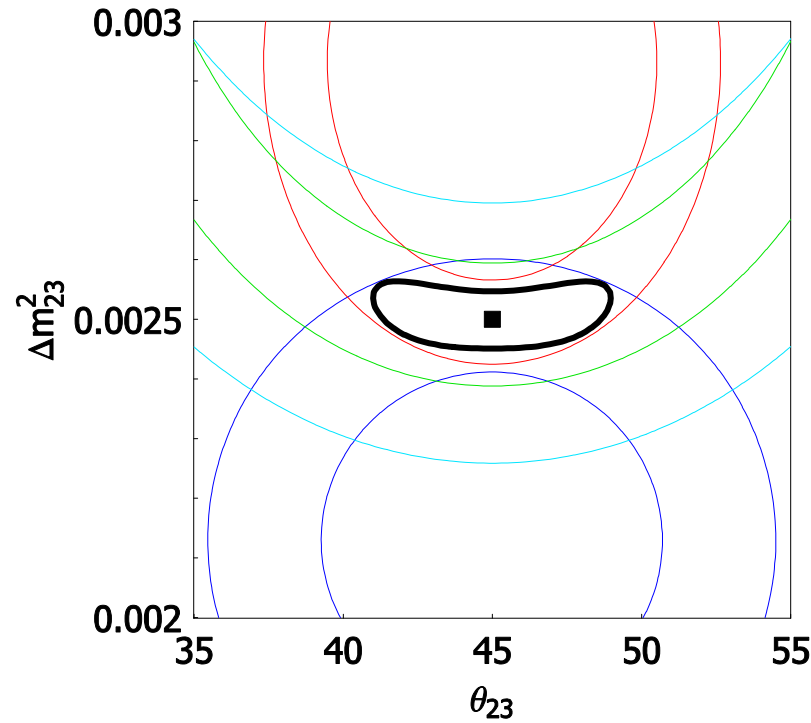
$\langle E_\nu \rangle = 0.27 \text{ GeV}$

$\langle E_{\bar{\nu}} \rangle = 0.25 \text{ GeV}$



# The importance of energy resolution

T2K-I



E1 = 0.4 - 0.6 GeV

E2 = 0.6 - 0.8 GeV

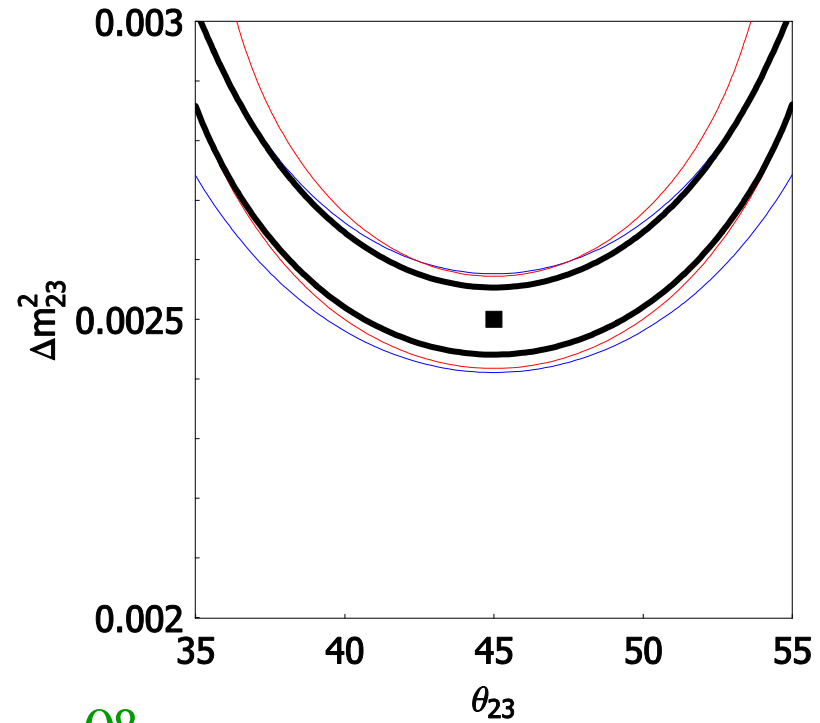
E3 = 0.8 - 1.0 GeV

E4 = 1.0 - 1.2 GeV

$\theta_{13} = 0^\circ$

$\delta = 0^\circ$

SPL-new



$\langle E_\nu \rangle = 0.29 \text{ GeV}$

$\langle E_{\bar{\nu}} \rangle = 0.28 \text{ GeV}$



## The $\nu_\mu$ disappearance channel

$$P(\nu_\mu \rightarrow \nu_\mu) = 1 - (\sin^2 2\theta_{23} - s_{23}^2 \sin^2 2\theta_{13} \cos^2 2\theta_{23}) \sin^2\left(\frac{\Delta_{atm} L}{2}\right) \\ - \left(\frac{\Delta_{sol} L}{2}\right) [s_{12}^2 \sin^2 2\theta_{23} + \tilde{J} s_{23}^2 \cos \delta] \sin(\Delta_{atm} L) \\ - \left(\frac{\Delta_{sol} L}{2}\right)^2 [c_{23}^4 \sin^2 2\theta_{12} + s_{12}^2 \sin^2 2\theta_{23} \cos(\Delta_{atm} L)]$$

Where

$$\tilde{J} = \cos \theta_{13} \sin 2\theta_{13} \sin 2\theta_{12} \sin 2\theta_{23} \qquad \Delta_{sol} = \frac{\Delta m_{12}^2}{2E}$$
$$\sin 2\theta_{13} < 0.4 \qquad \Delta_{atm} = \frac{\Delta m_{23}^2}{2E} \qquad \left(\frac{\Delta_{sol} L}{2}\right) \cong 0.05$$

E. K. Akhmedov *et al.* hep-ph/0402175  
A. Donini *et al.* hep-ph/0411402

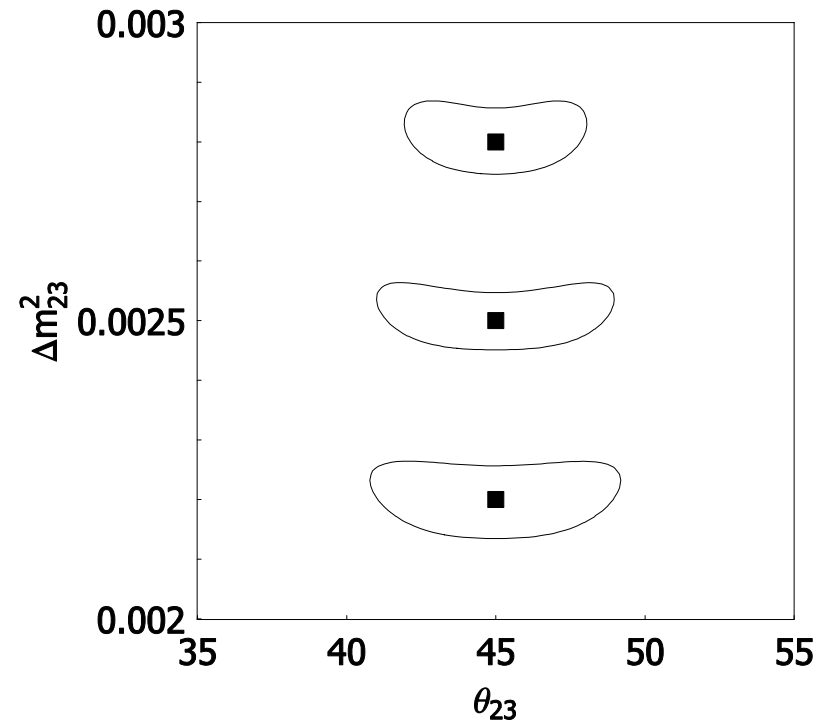


# The sign degeneracy

Input:

$$\theta_{13} = 0^\circ$$

$$\delta = 0^\circ$$



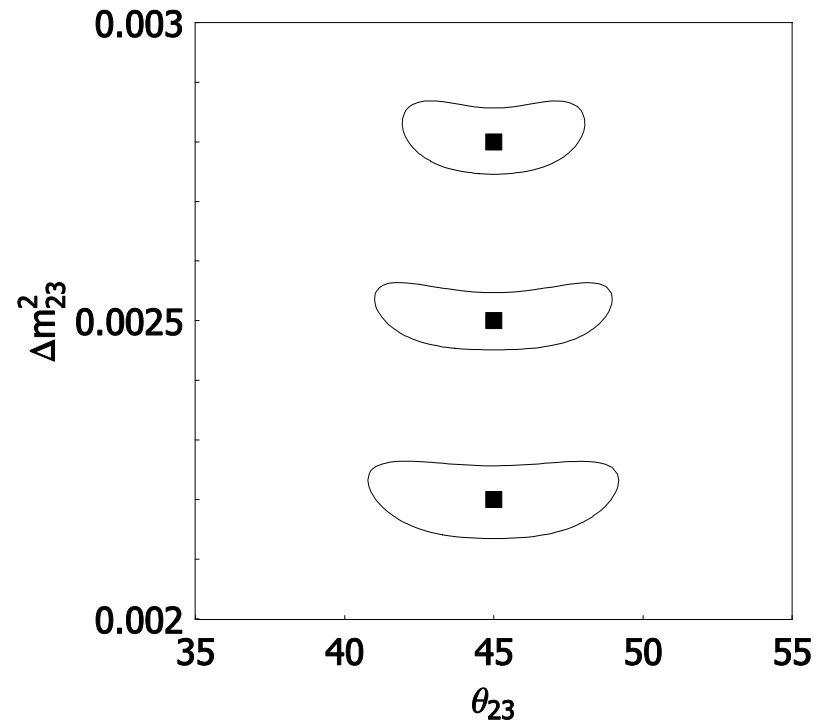
$$P(\nu_\mu \rightarrow \nu_\mu) = 1 - (\sin^2 2\theta_{23}) \sin^2\left(\frac{\Delta_{atm} L}{2}\right) \\ + \mathcal{O}\left(\frac{\Delta_{sol} L}{2}\right) \\ + \mathcal{O}\left(\frac{\Delta_{sol} L}{2}\right)^2$$





# The sign degeneracy

Input:  
 $\theta_{13} = 0^\circ$   
 $\delta = 0^\circ$

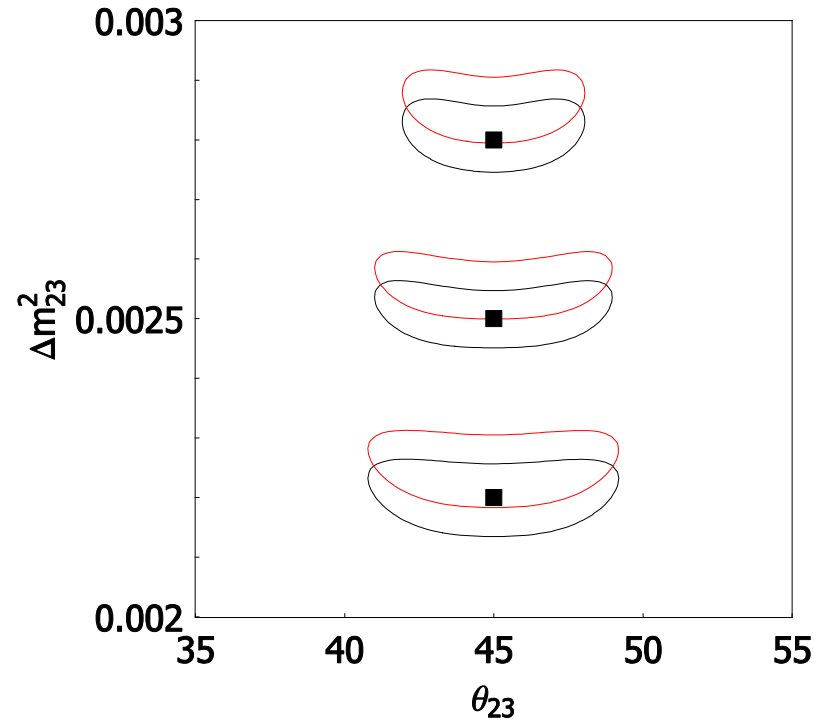


$$\begin{aligned}
 P(\nu_\mu \rightarrow \nu_\mu) = & 1 - (\sin^2 2\theta_{23}) \sin^2\left(\frac{\Delta_{atm} L}{2}\right) \\
 & - \left(\frac{\Delta_{sol} L}{2}\right) [s_{12}^2 \sin^2 2\theta_{23}] \sin(\Delta_{atm} L) \\
 & + \mathcal{O}\left(\frac{\Delta_{sol} L}{2}\right)^2
 \end{aligned}$$



# The sign degeneracy

Input:  
 $\theta_{13} = 0^\circ$   
 $\delta = 0^\circ$

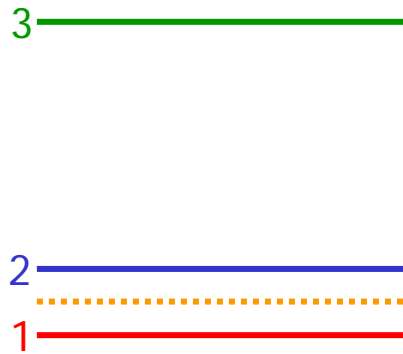
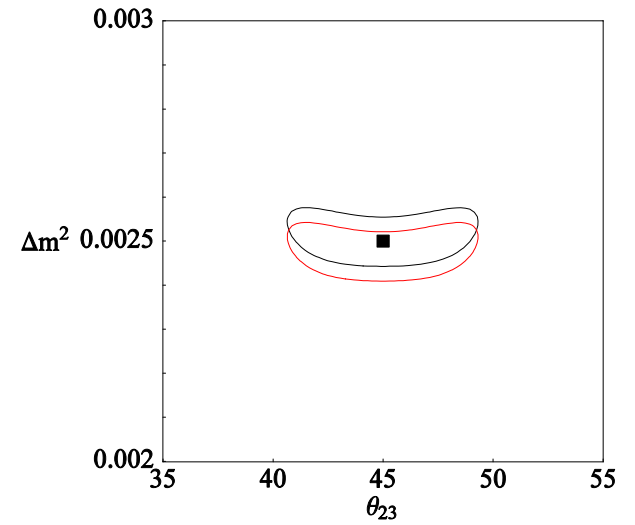
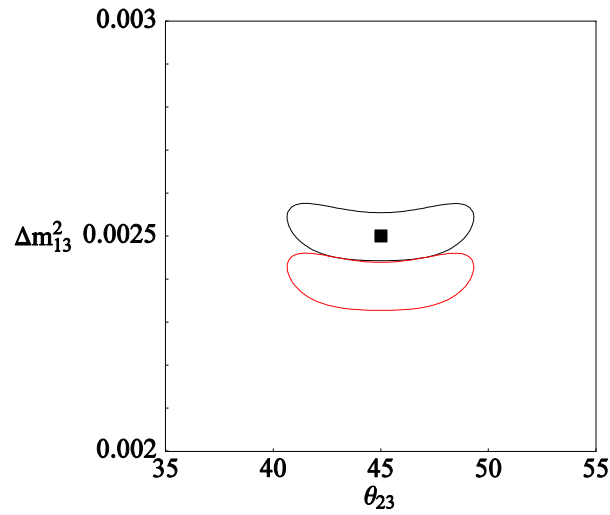
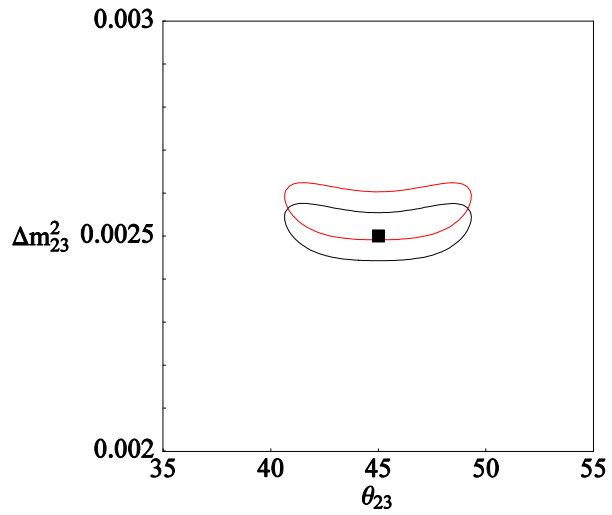


Fit assuming  
 inverted hierarchy

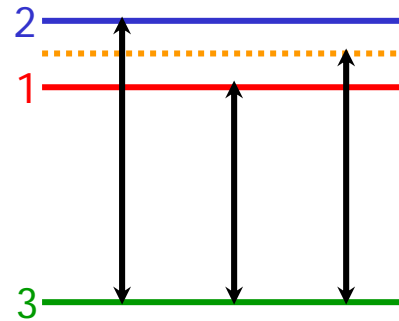
$$\begin{aligned}
 P(\nu_\mu \rightarrow \nu_\mu) = & 1 - (\sin^2 2\theta_{23}) \sin^2\left(\frac{\Delta_{atm} L}{2}\right) \\
 & - \left(\frac{\Delta_{sol} L}{2}\right) [s_{12}^2 \sin^2 2\theta_{23}] \sin(\Delta_{atm} L) \\
 & + \mathcal{O}\left(\frac{\Delta_{sol} L}{2}\right)^2
 \end{aligned}$$



# The sign degeneracy



Normal



Inverted

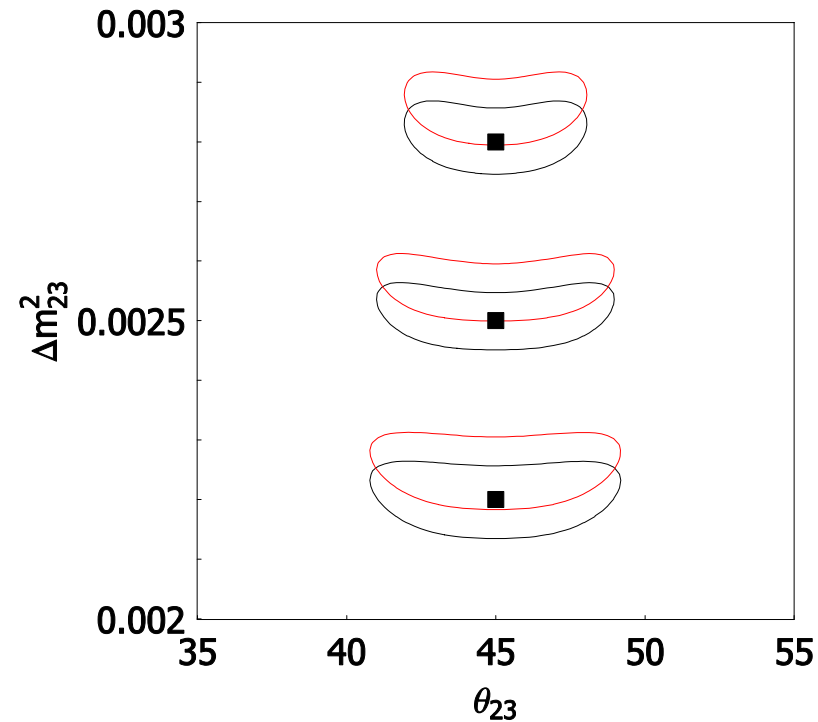


# The octant degeneracy

Input:

$$\theta_{13} = 0^\circ$$

$$\delta = 0^\circ$$

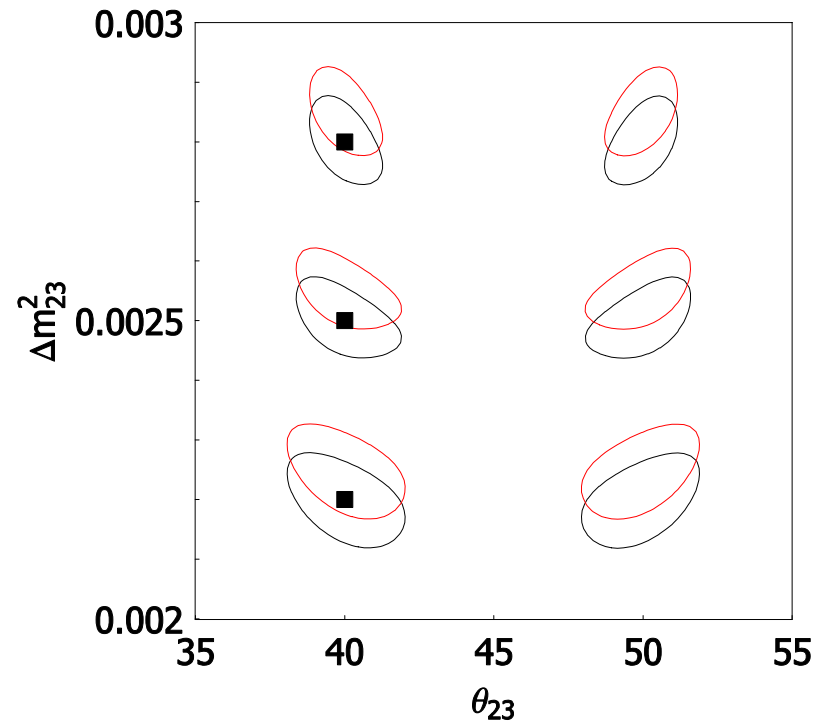


$$\begin{aligned}
 P(\nu_\mu \rightarrow \nu_\mu) = & 1 - \underbrace{(\sin^2 2\theta_{23})}_{\text{orange}} \sin^2\left(\frac{\Delta_{atm} L}{2}\right) \\
 & - \left(\frac{\Delta_{sol} L}{2}\right) \left[ s_{12}^2 \underbrace{\sin^2 2\theta_{23}}_{\text{orange}} \right] \sin(\Delta_{atm} L) \\
 & + \mathcal{O}\left(\frac{\Delta_{sol} L}{2}\right)^2
 \end{aligned}$$



# The octant degeneracy

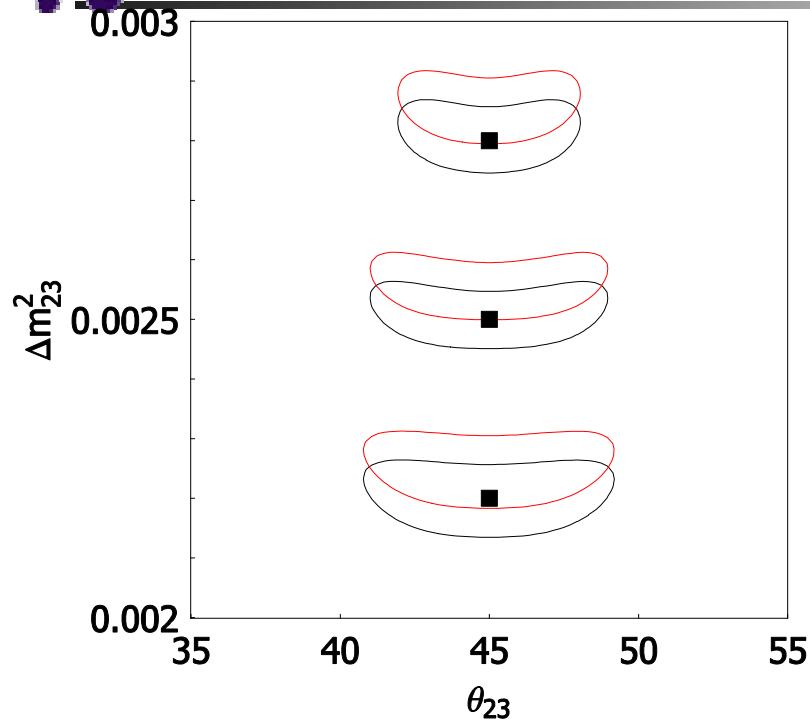
Input:  
 $\theta_{13} = 0^\circ$   
 $\delta = 0^\circ$



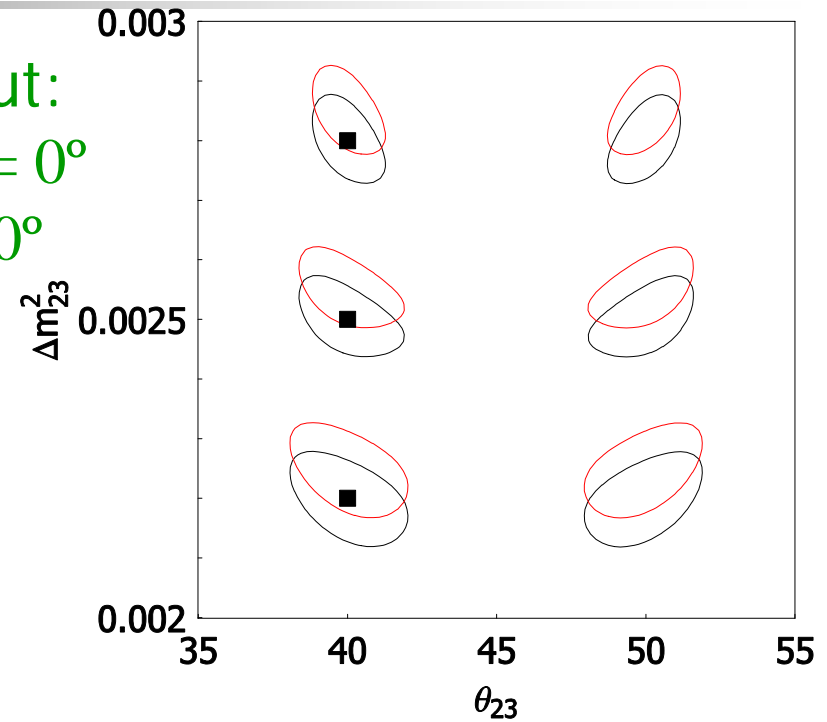
$$\begin{aligned}
 P(\nu_\mu \rightarrow \nu_\mu) = & 1 - \sin^2 2\theta_{23} \sin^2\left(\frac{\Delta_{atm} L}{2}\right) \\
 & - \left(\frac{\Delta_{sol} L}{2}\right) [s_{12}^2 \sin^2 2\theta_{23}] \sin(\Delta_{atm} L) \\
 & + \mathcal{O}\left(\frac{\Delta_{sol} L}{2}\right)^2
 \end{aligned}$$



# The effect of $\theta_{13}$



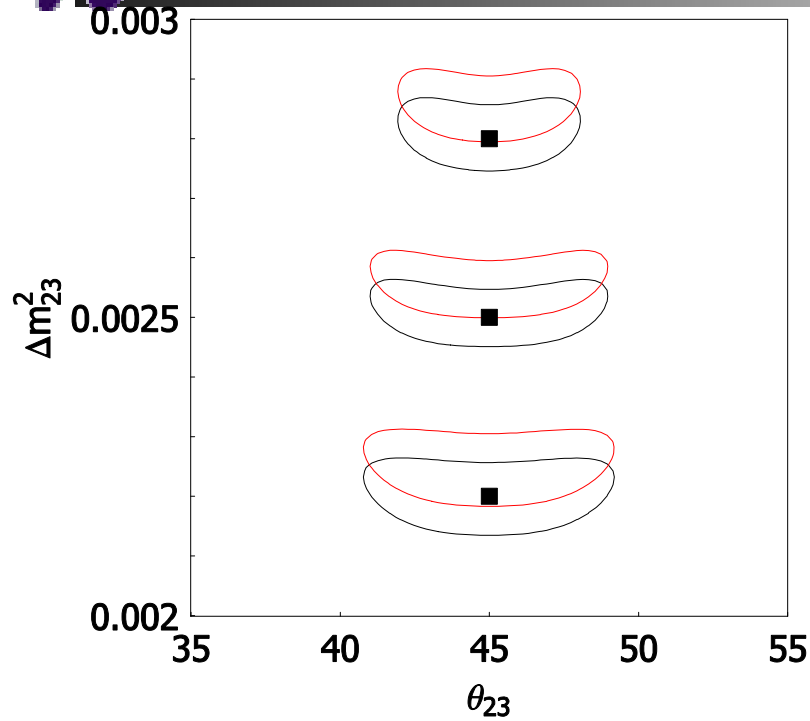
Input:  
 $\theta_{13} = 0^\circ$   
 $\delta = 0^\circ$



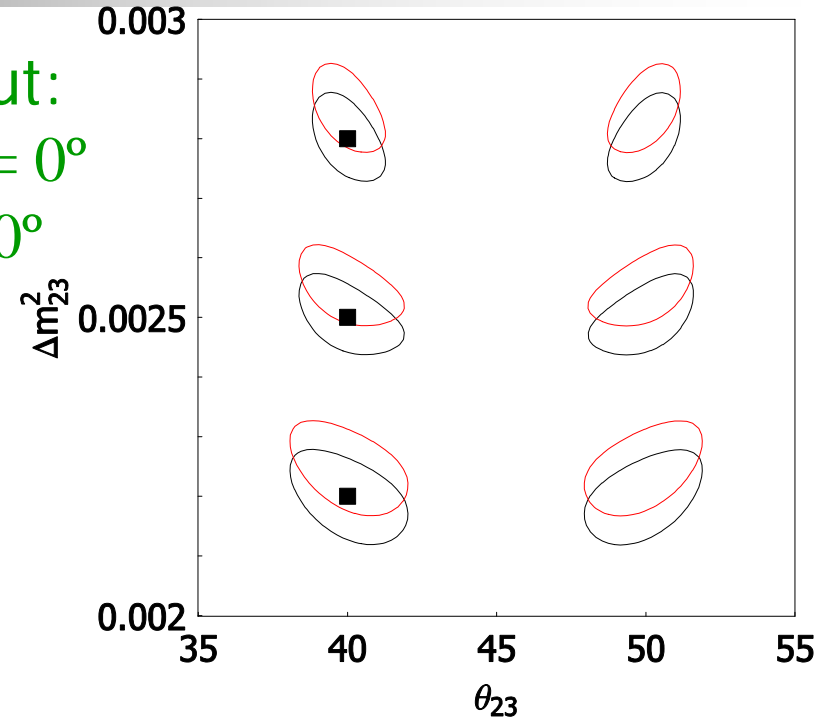
$$\begin{aligned}
 P(\nu_\mu \rightarrow \nu_\mu) = & 1 - (\sin^2 2\theta_{23}) \sin^2\left(\frac{\Delta_{atm} L}{2}\right) \\
 & - \left(\frac{\Delta_{sol} L}{2}\right) [s_{12}^2 \sin^2 2\theta_{23}] \sin(\Delta_{atm} L) \\
 & + \mathcal{O}\left(\frac{\Delta_{sol} L}{2}\right)^2
 \end{aligned}$$



# The effect of $\theta_{13}$



Input:  
 $\theta_{13} = 0^\circ$   
 $\delta = 0^\circ$

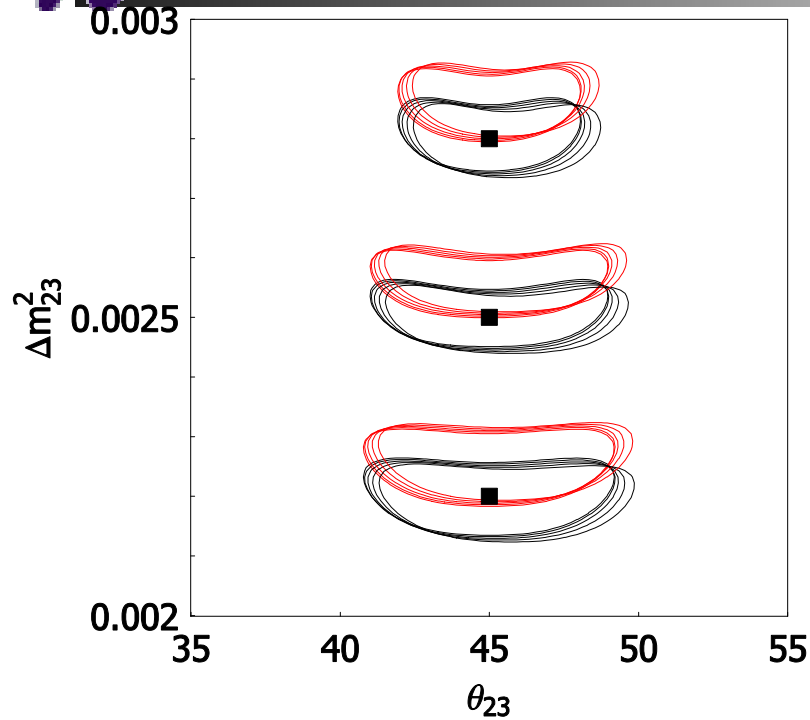


Assuming  $\theta_{13} = 0^\circ$

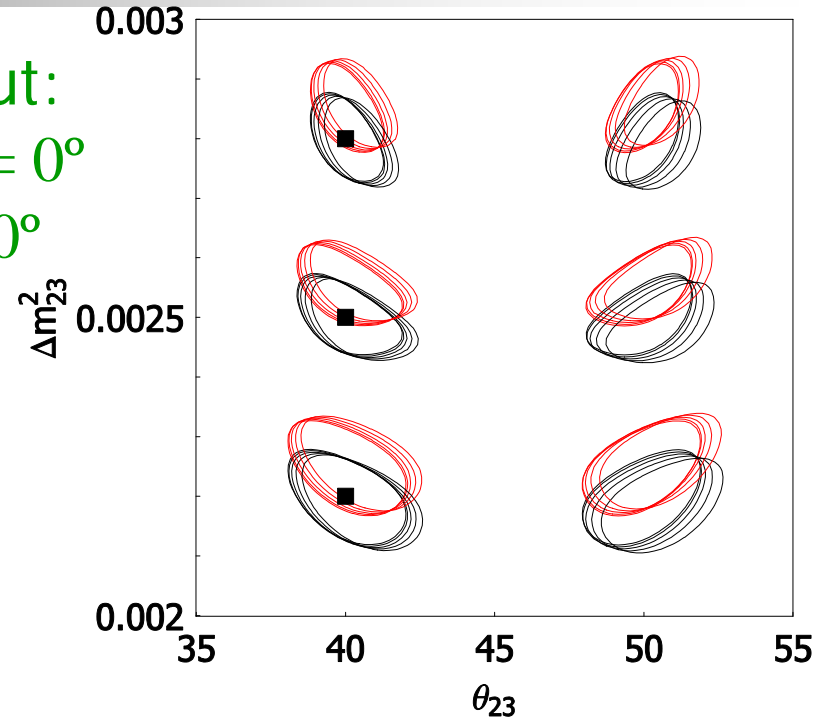
$$\begin{aligned}
 P(\nu_\mu \rightarrow \nu_\mu) = & 1 - (\sin^2 2\theta_{23} - \underbrace{s_{23}^2}_{\text{orange}} \sin^2 2\theta_{13} \cos^2 2\theta_{23}) \sin^2\left(\frac{\Delta_{atm} L}{2}\right) \\
 & - \left(\frac{\Delta_{sol} L}{2}\right) [s_{12}^2 \sin^2 2\theta_{23} + \underbrace{c_{23}^2}_{\text{orange}} \cos \delta] \sin(\Delta_{atm} L) \\
 & + \mathcal{O}\left(\frac{\Delta_{sol} L}{2}\right)^2
 \end{aligned}$$



# The effect of $\theta_{13}$



Input:  
 $\theta_{13} = 0^\circ$   
 $\delta = 0^\circ$



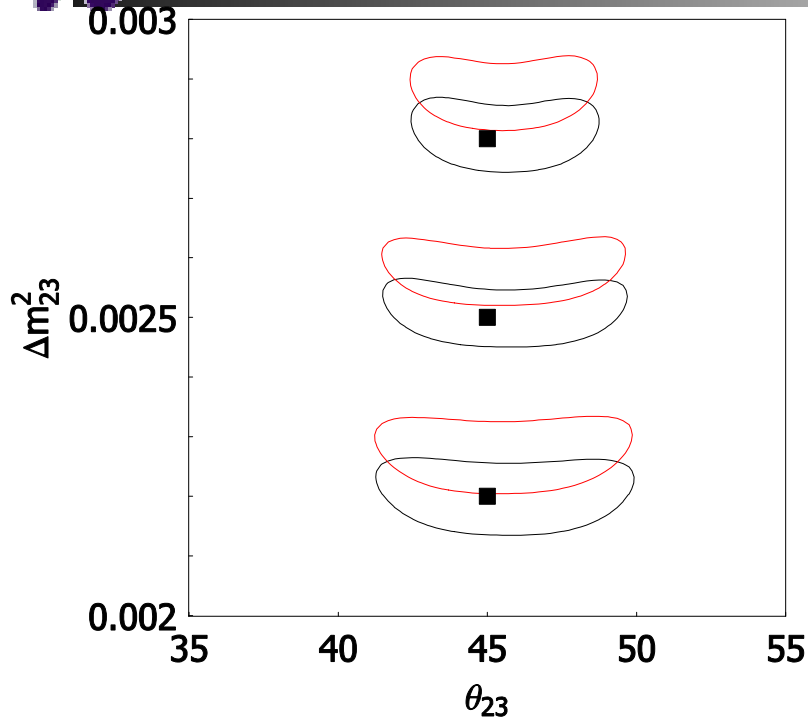
Assuming  $\theta_{13} = 0^\circ, 2^\circ, 4^\circ, 6^\circ, 8^\circ$

$$\begin{aligned}
 P(\nu_\mu \rightarrow \nu_\mu) = & 1 - (\sin^2 2\theta_{23} - s_{23}^2 \sin^2 2\theta_{13} \cos^2 2\theta_{23}) \sin^2\left(\frac{\Delta_{atm} L}{2}\right) \\
 & - \left(\frac{\Delta_{sol} L}{2}\right) [s_{12}^2 \sin^2 2\theta_{23} + s_{23}^2 \cos \delta] \sin(\Delta_{atm} L) \\
 & + \mathcal{O}\left(\frac{\Delta_{sol} L}{2}\right)^2
 \end{aligned}$$

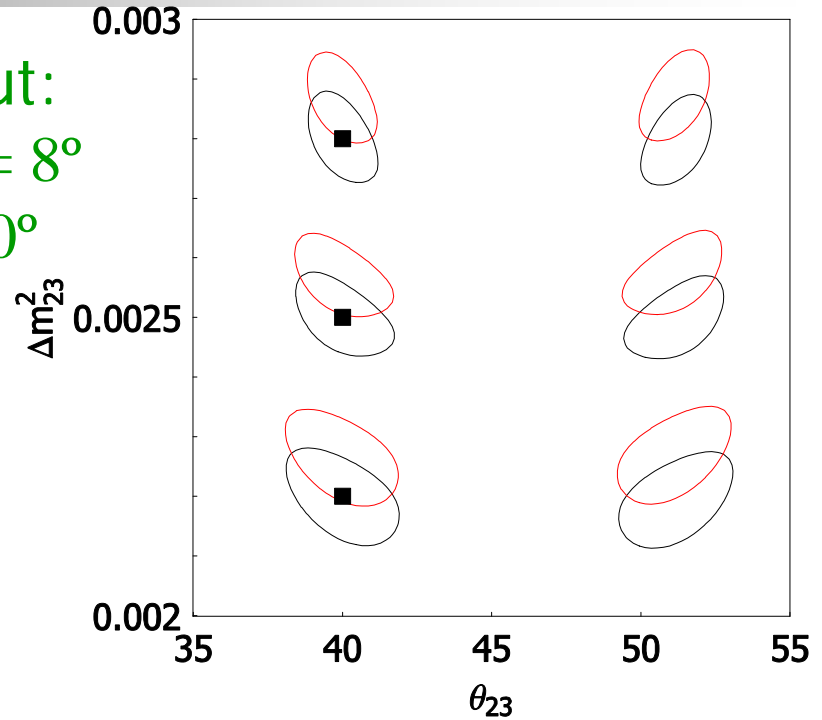




# The effect of $\delta$



Input:  
 $\theta_{13} = 8^\circ$   
 $\delta = 0^\circ$

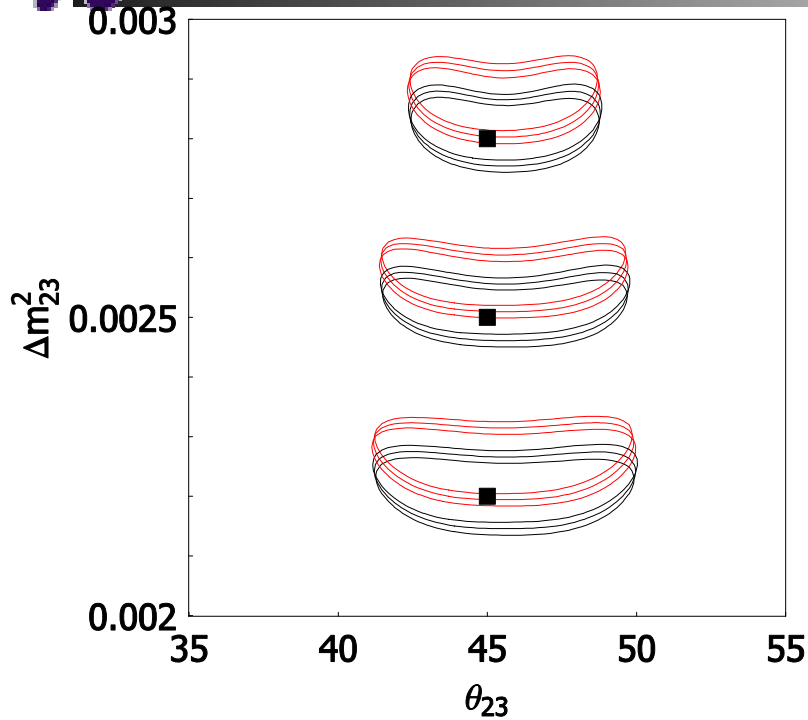


Assuming  $\delta = 0^\circ$

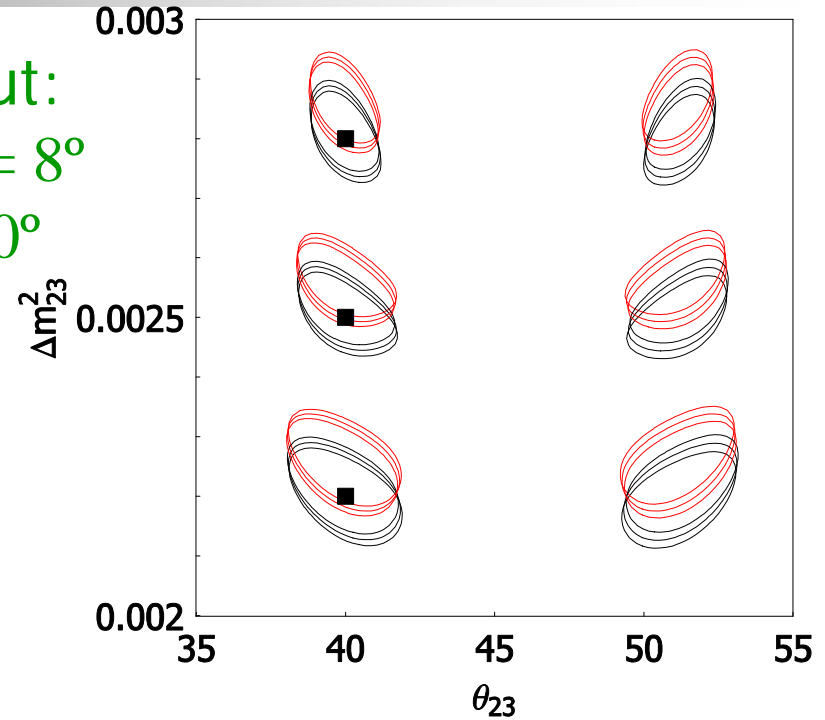
$$\begin{aligned}
 P(\nu_\mu \rightarrow \nu_\mu) = & 1 - (\sin^2 2\theta_{23} - s_{23}^2 \sin^2 2\theta_{13} \cos^2 2\theta_{23}) \sin^2\left(\frac{\Delta_{atm} L}{2}\right) \\
 & - \left(\frac{\Delta_{sol} L}{2}\right) [s_{12}^2 \sin^2 2\theta_{23} + \tilde{J} s_{23}^2 \cos \delta] \sin(\Delta_{atm} L) \\
 & + \mathcal{O}\left(\frac{\Delta_{sol} L}{2}\right)^2
 \end{aligned}$$



# The effect of $\delta$



Input:  
 $\theta_{13} = 8^\circ$   
 $\delta = 0^\circ$

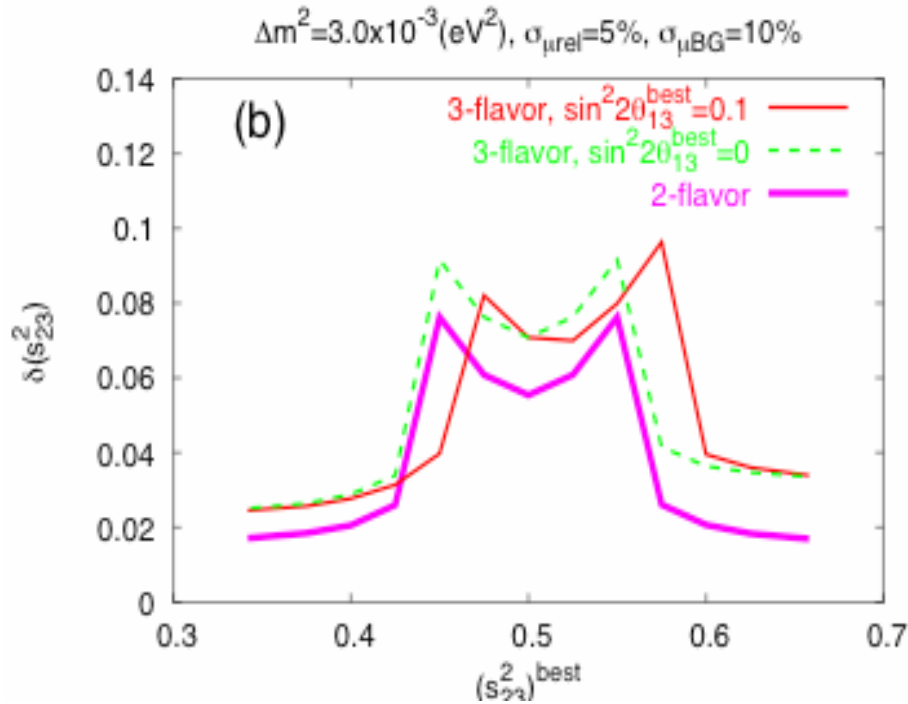


Assuming  $\delta = 0^\circ, 90^\circ, 180^\circ$

$$\begin{aligned}
 P(\nu_\mu \rightarrow \nu_\mu) = & 1 - (\sin^2 2\theta_{23} - s_{23}^2 \sin^2 2\theta_{13} \cos^2 2\theta_{23}) \sin^2\left(\frac{\Delta_{atm} L}{2}\right) \\
 & - \left(\frac{\Delta_{sol} L}{2}\right) [s_{12}^2 \sin^2 2\theta_{23} + \tilde{J} s_{23}^2 \cos \delta] \sin(\Delta_{atm} L) \\
 & + \mathcal{O}\left(\frac{\Delta_{sol} L}{2}\right)^2
 \end{aligned}$$



# T2K-I errors revised



Minakata *et al.* hep-ph/0406073

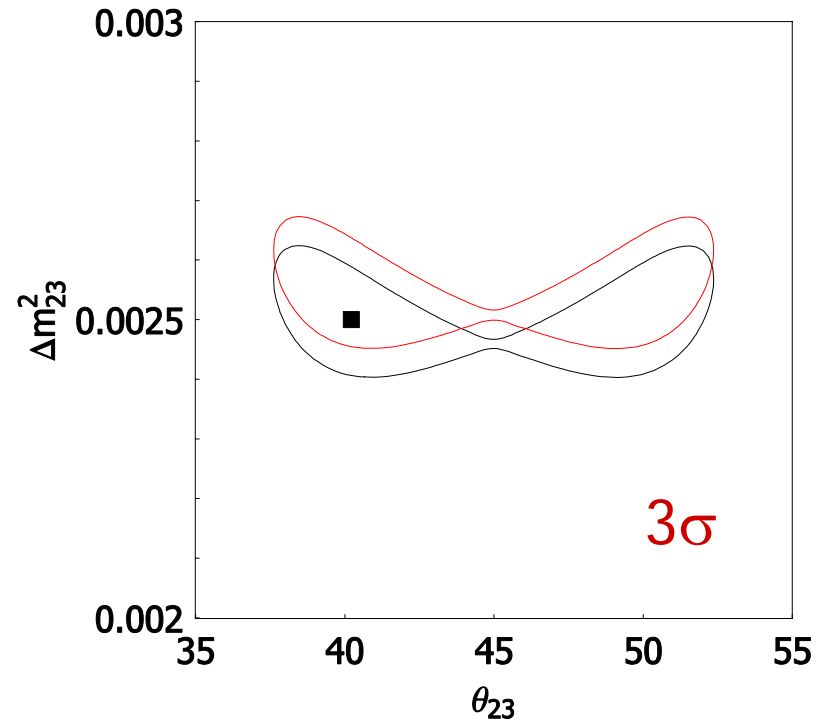
$$\Delta m^2 = (1.7 - 3.5) \cdot 10^{-3} \text{ eV}^2$$

$$\sin^2 2\theta > 0.9 \quad \tan^2 \theta = 0.53 - 2.04$$

$$\Delta m^2 = (2.43 - 2.60) \cdot 10^{-3} \text{ eV}^2$$

$$\sin^2 2\theta > 0.97 \quad \tan^2 \theta = 0.73 - 1.39$$

$$\Delta m^2 = (-2.63 - -2.49) \cdot 10^{-3} \text{ eV}^2$$



Antusch *et al.* hep-ph/0404268

$$\Delta m^2 = (2.42 - 2.61) \cdot 10^{-3} \text{ eV}^2$$

$$\sin^2 2\theta = 0.94 - 0.99$$

$$\tan^2 \theta = 0.62 - 0.85, 1.21 - 1.66$$

$$\Delta m^2 = (-2.64 - -2.47) \cdot 10^{-3} \text{ eV}^2$$

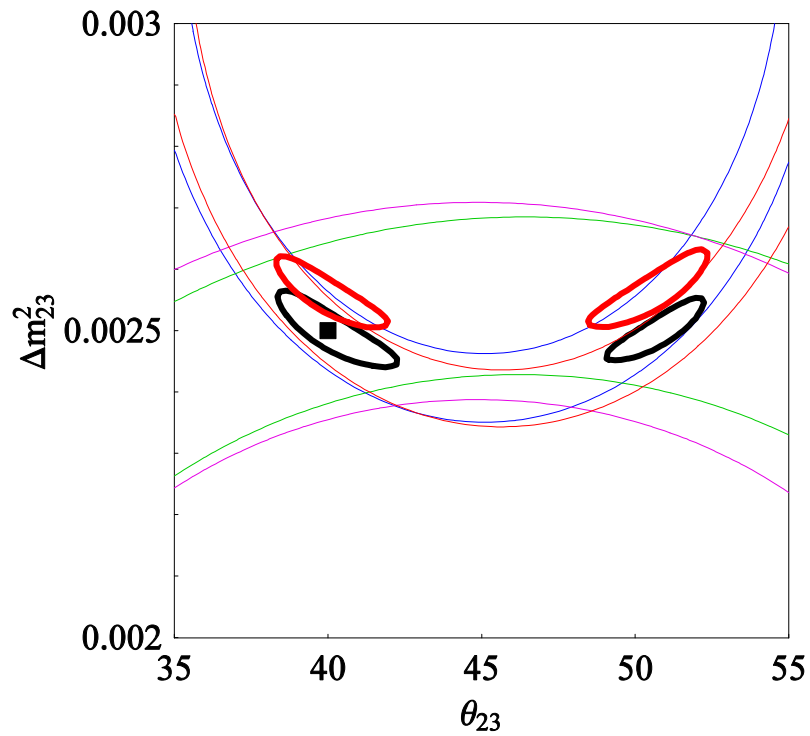


# Preliminary analysis at the NuFactory

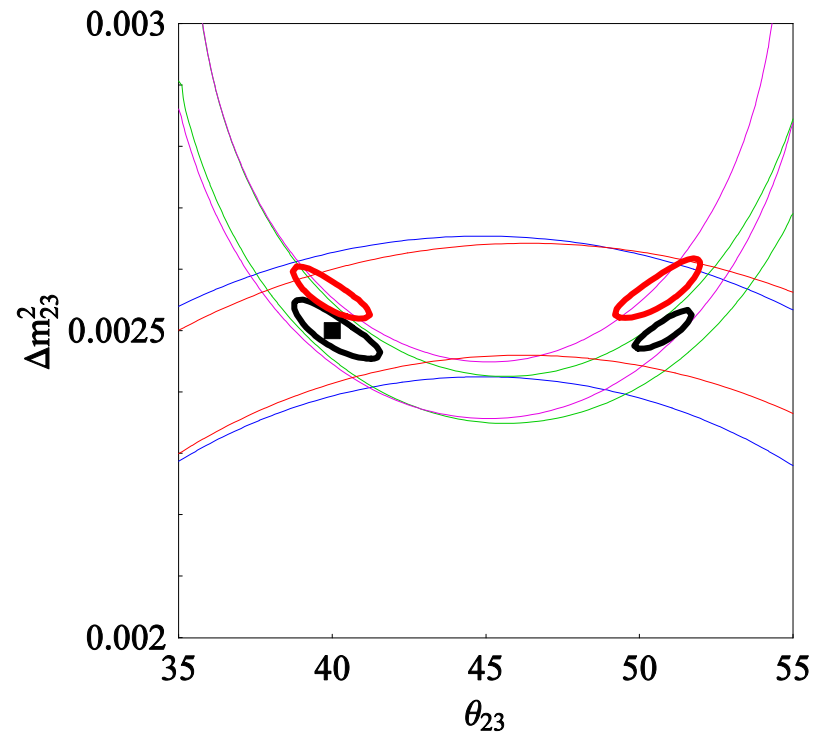
5yr  $\nu_\mu$  + 5yr  $\bar{\nu}_\mu$  exposure with a 40Kt iron calorimeter for the NF

- Possible Setups:
  - L = 3000Km E = 20, 50 GeV
  - L = 7000Km E = 50 GeV
- 5 GeV bins considered
- Efficiency:
  - $\varepsilon_\mu = 0.5$  for neutrinos “Cervera *et al.* hep-ph/0002108”
  - $\varepsilon_\mu = 0.33$  for antineutrinos
- Systematics = 2%

See e.g. Bueno *et al.* hep-ph/0005007 for an Icarus analysis

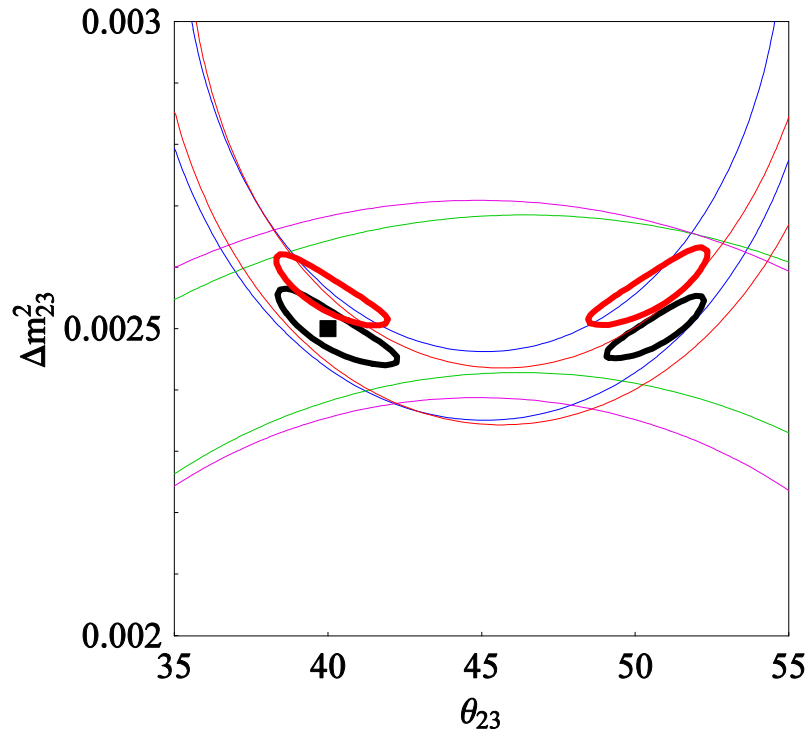


$E = 50$  GeV  
 $L = 3000$  Km

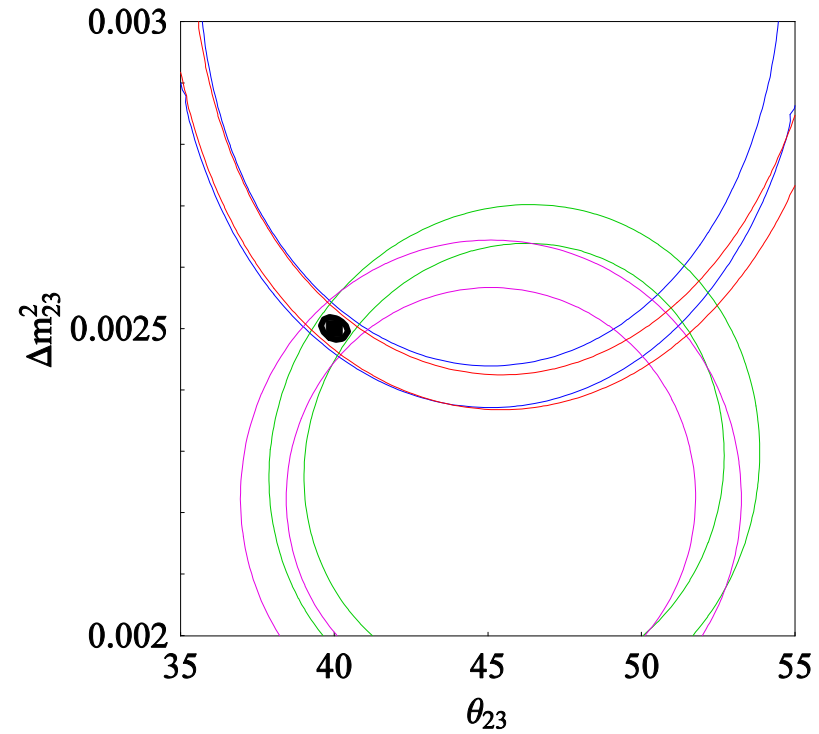


$E = 20$  GeV  
 $L = 3000$  Km

Input:  $\theta_{23} = 40^\circ$ ,  $\theta_{13} = 6^\circ$ ,  $\delta = 0^\circ$

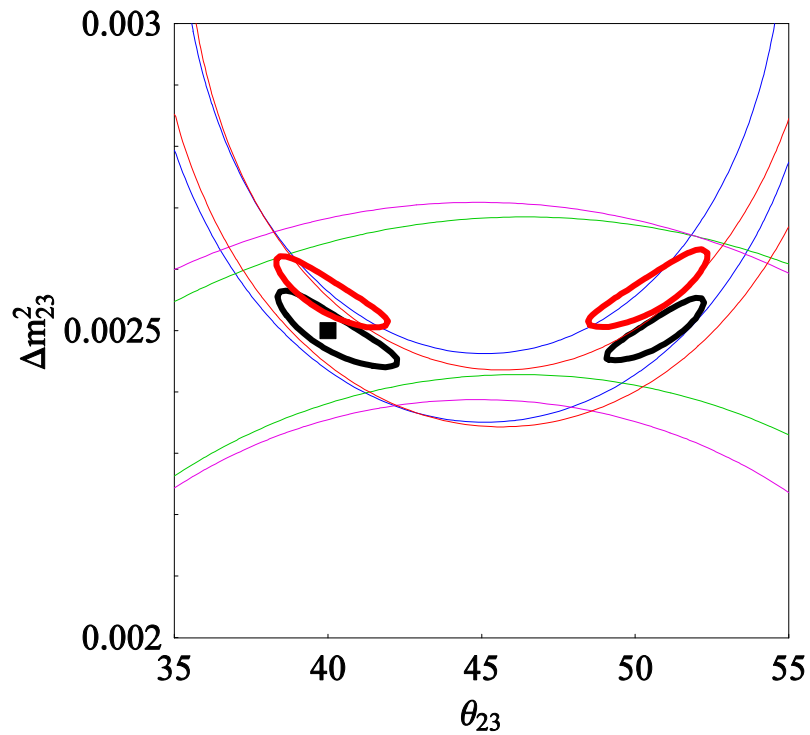


$E = 50 \text{ GeV}$   
 $L = 3000 \text{ Km}$

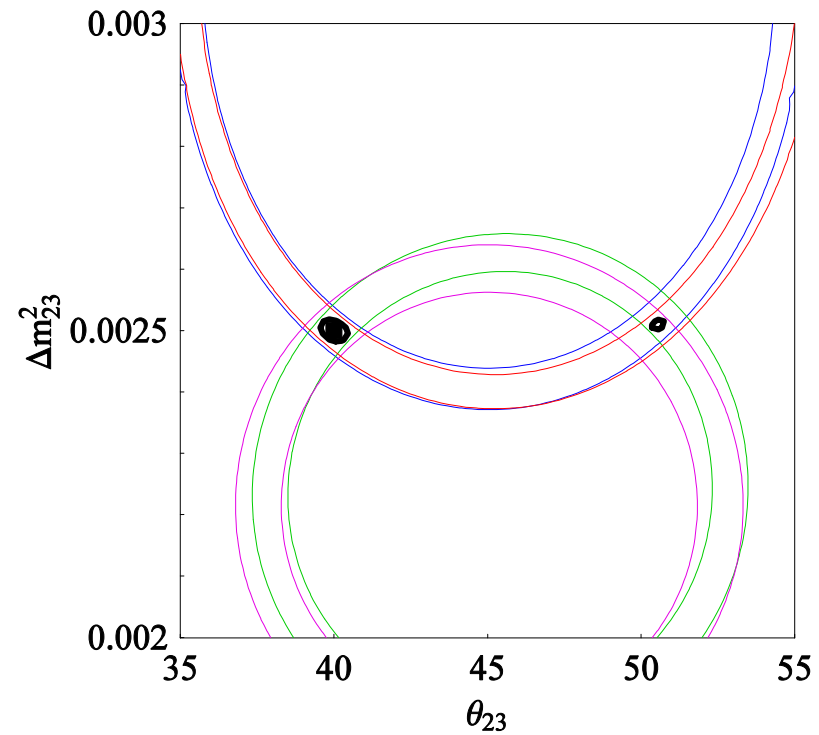


$E = 50 \text{ GeV}$   
 $L = 7000 \text{ Km}$

Input:  $\theta_{23} = 40^\circ$ ,  $\theta_{13} = 6^\circ$ ,  $\delta = 0^\circ$

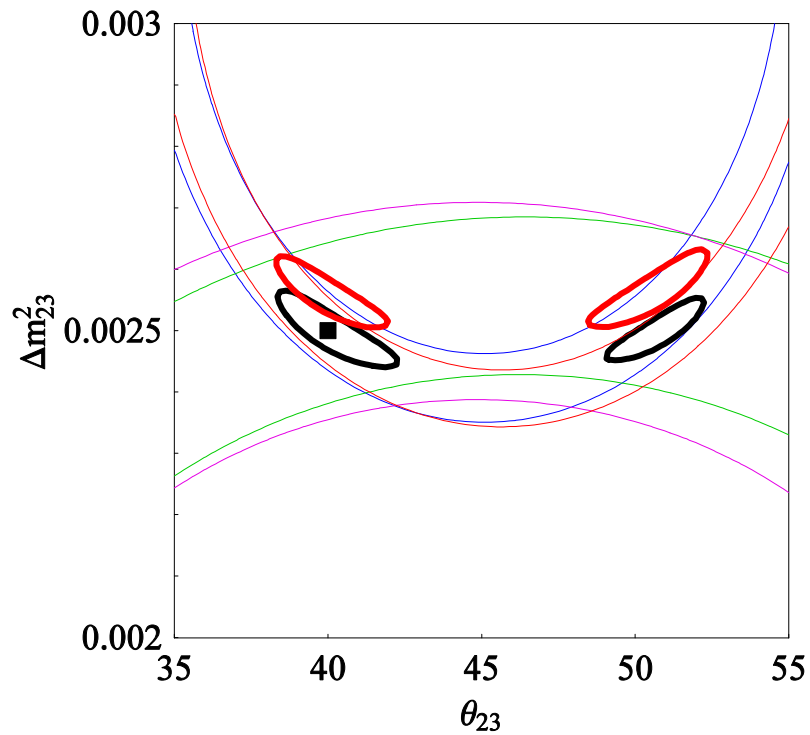


$E = 50$  GeV  
 $L = 3000$  Km

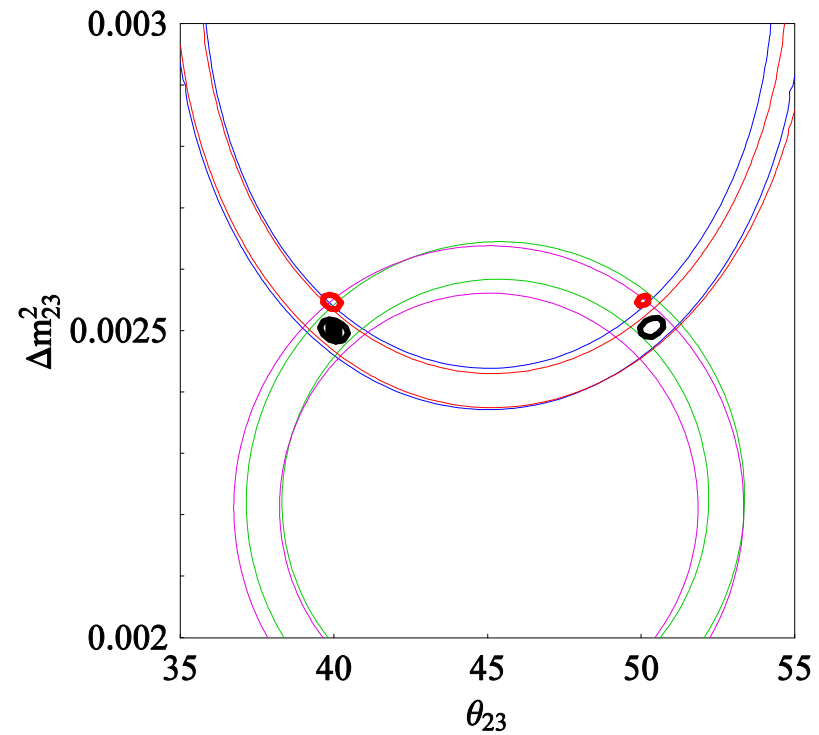


$E = 50$  GeV  
 $L = 7000$  Km

Input:  $\theta_{23} = 40^\circ$ ,  $\theta_{13} = 4^\circ$ ,  $\delta = 0^\circ$



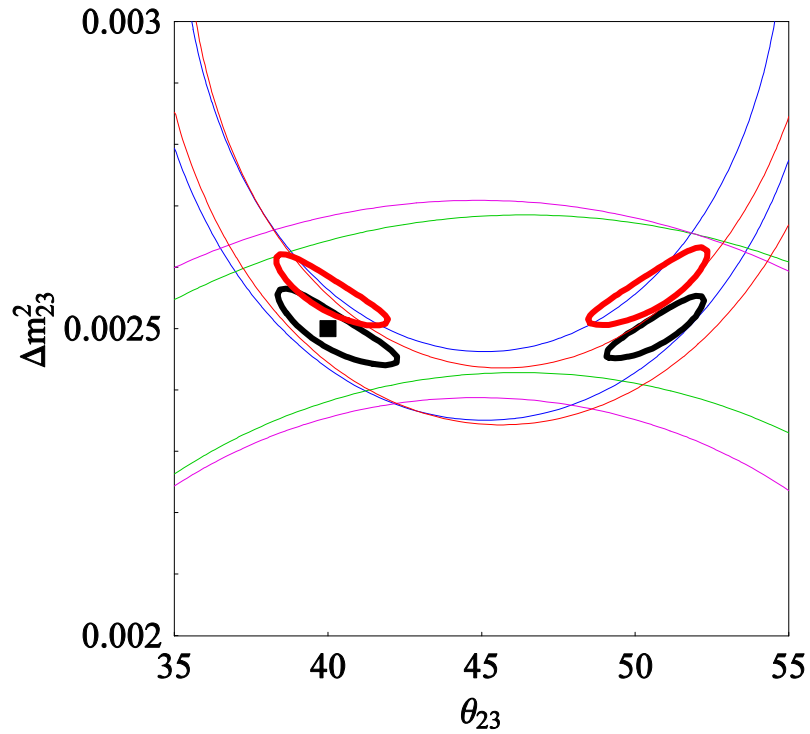
$E = 50 \text{ GeV}$   
 $L = 3000 \text{ Km}$



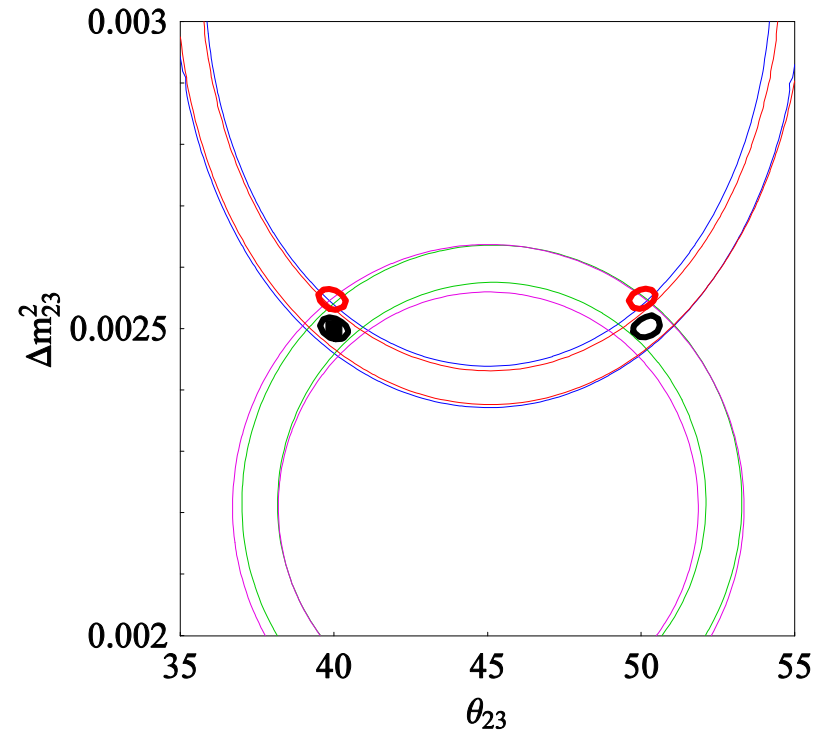
$E = 50 \text{ GeV}$   
 $L = 7000 \text{ Km}$

Input:  $\theta_{23} = 40^\circ$ ,  $\theta_{13} = 3^\circ$ ,  $\delta = 0^\circ$





$E = 50 \text{ GeV}$   
 $L = 3000 \text{ Km}$



$E = 50 \text{ GeV}$   
 $L = 7000 \text{ Km}$

Input:  $\theta_{23} = 40^\circ$ ,  $\theta_{13} = 2^\circ$ ,  $\delta = 0^\circ$



## Conclusions

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- The measurement of  $\theta_{13}$  and  $\delta$  will rely heavily on an improvement of the measure of  $\theta_{23}$  and  $\Delta m^2_{23}$

See Meloni's talk

- Precision measurements of  $\theta_{23}$  and  $\Delta m^2_{23}$  need energy resolution and events above and below the oscillation peak
- SPL is clearly inadequate for the task. T2K-I is very good due to energy resolution and it can exclude maximal mixing for  $\theta_{23} < 41^\circ$
- The NuFactory seems extremely promising but more study is needed (a **very long baseline?**)
- $\nu_\mu$  disappearance can be combined with the appearance channel to solve degeneracies



# Event Rates

NF 50GeV	B1	B2	B3	B4
No osc. $N_\mu$	1137	15390	60590	147987
Signal $\mu^-$	266	1297	16150	54128
Signal $\mu^+$	153	751	9032	28635

New SPL	$\mu^-$	$\mu^+$
No osc. $N_\mu$	79365	95511
Signal $N_\mu$	8811	11347

4 energy bins of 5GeV  
Between 0 – 20GeV

L=3000Km

L=130Km

Systematic dominated

Systematic dominated

5yr  $\nu_\mu$  + 5yr  $\bar{\nu}_\mu$  exposure with a 40Kt iron calorimeter for the NF

2yr  $\nu_\mu$  + 8yr  $\bar{\nu}_\mu$  exposure with a 440Kt water cerenkov detector for the SPL

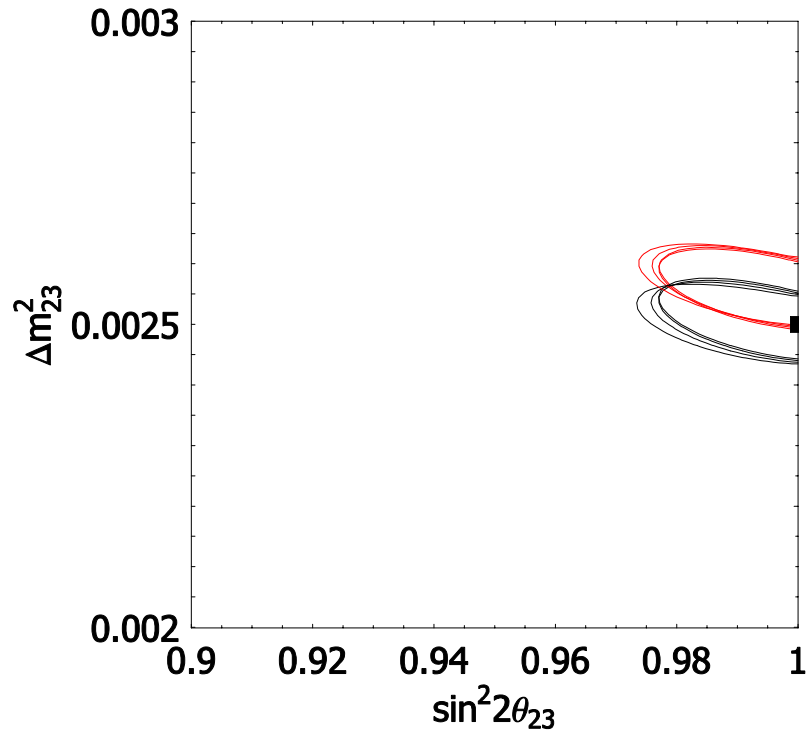


# Event Rates

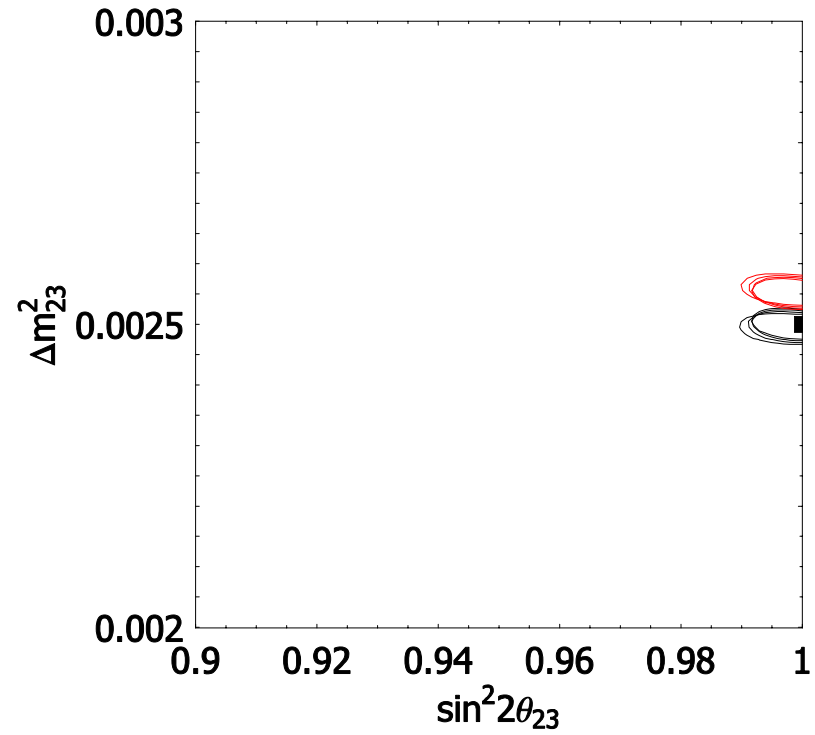
NF 20GeV	B1	B2	B3	B4
No osc. $N_\mu$	2598	31286	103756	196373
Signal $\mu^-$	614	2574	27399	71472
Signal $\mu^+$	337	1257	10415	11557

NF 50GeV 7000km	B1	B2	B3	B4
No osc. $N_\mu$	209	2827	11129	27281
Signal $\mu^-$	37	980	380	1328
Signal $\mu^+$	21	569	215	697

5yr  $\nu_\mu$  + 5yr  $\bar{\nu}_\mu$  exposure with a 40Kt iron calorimeter for the NF

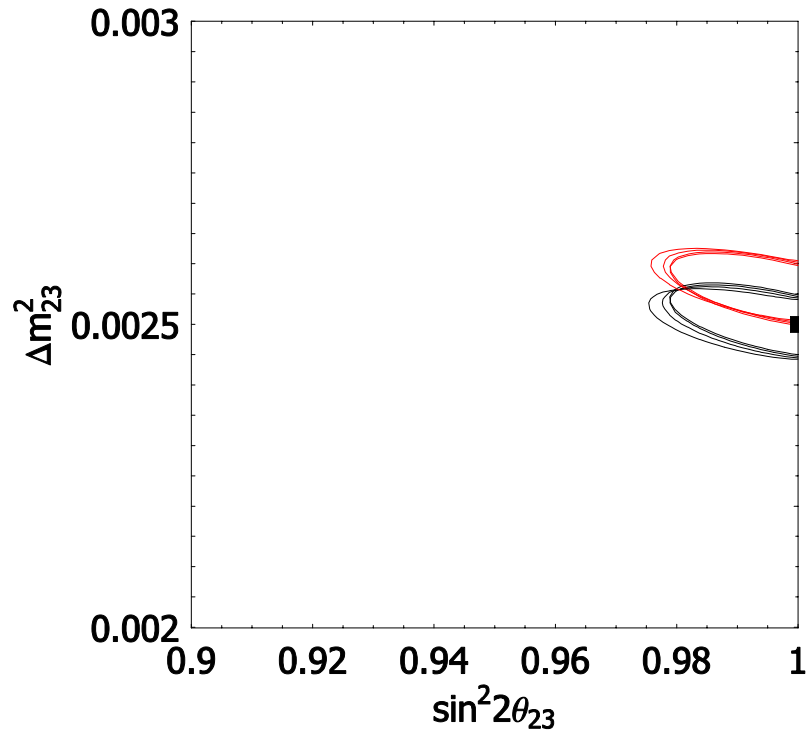


T2K-I

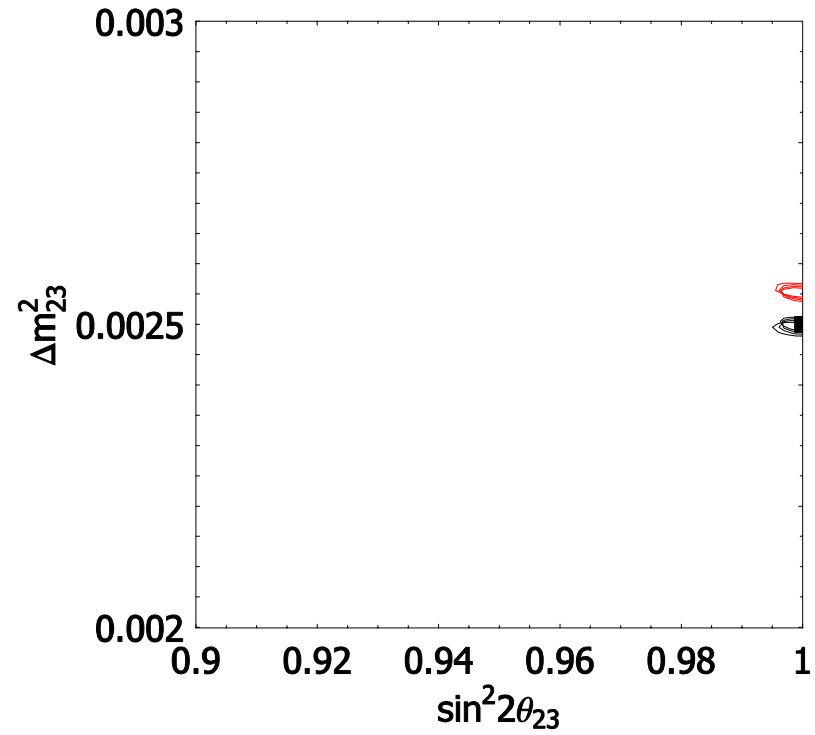


T2K-2

5% systematic error



T2K-I

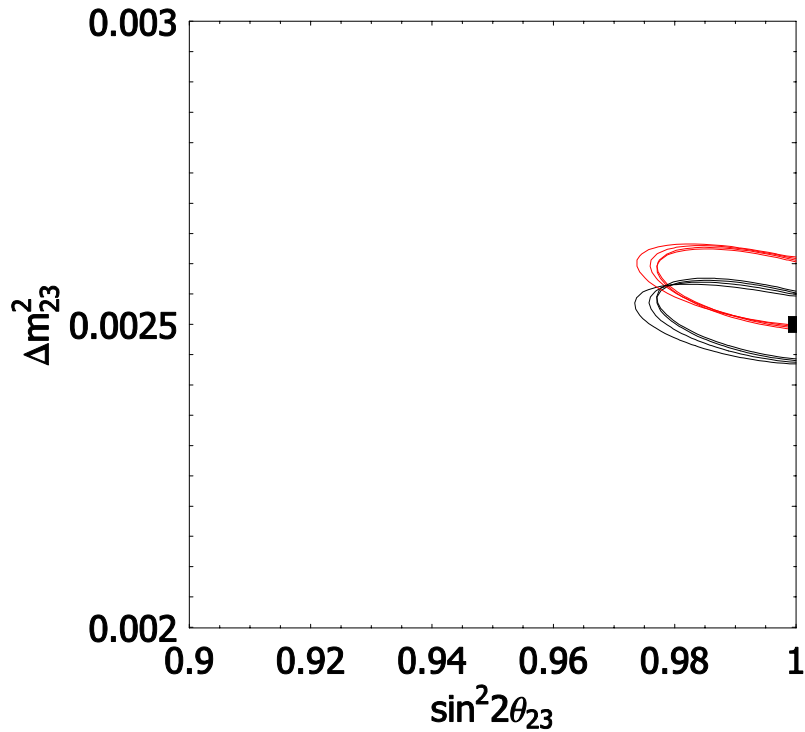


T2K-2

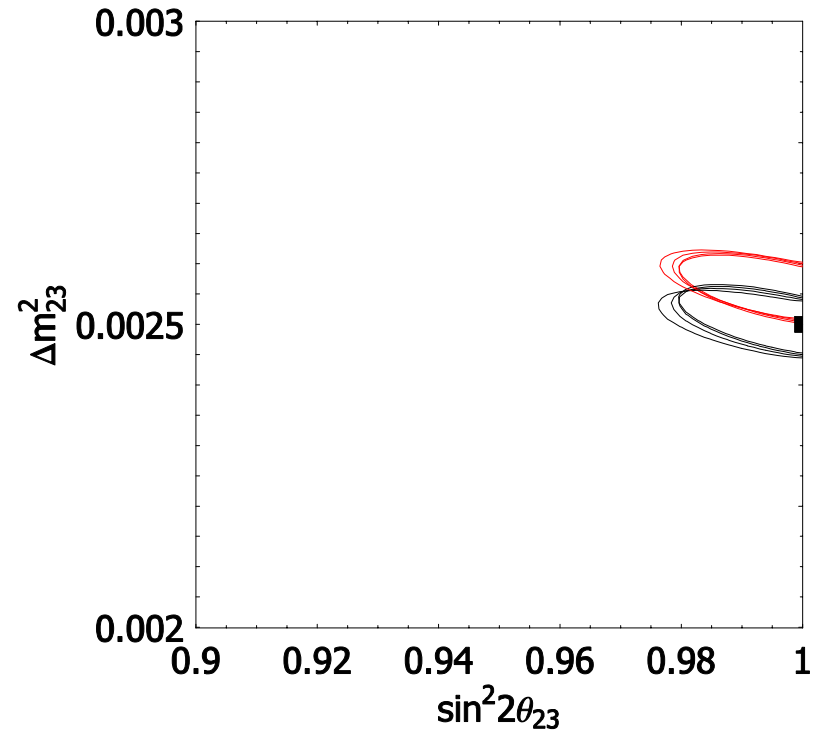
2% systematic error



# No background and no systematic



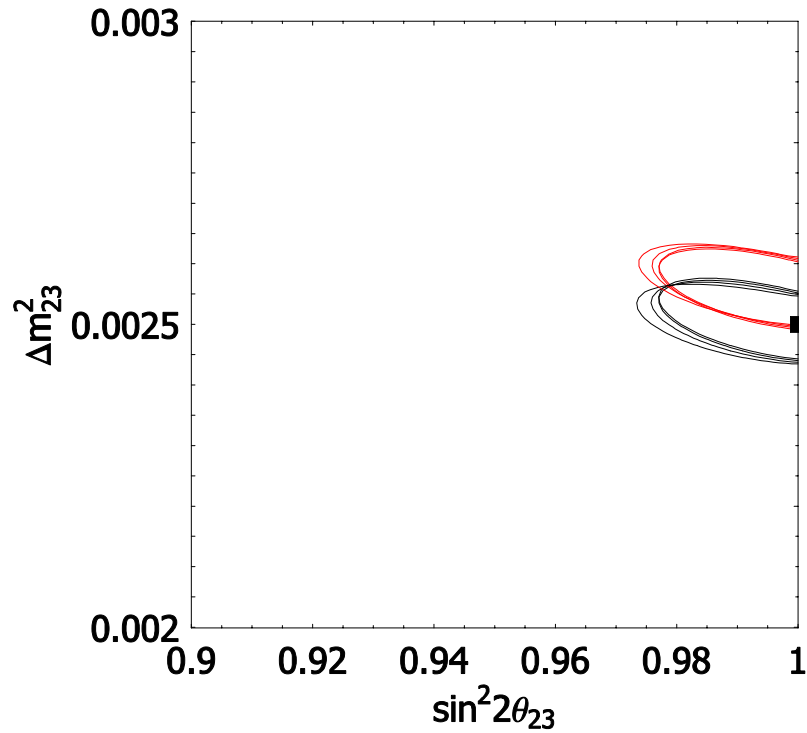
Systematic 5%  
With Background



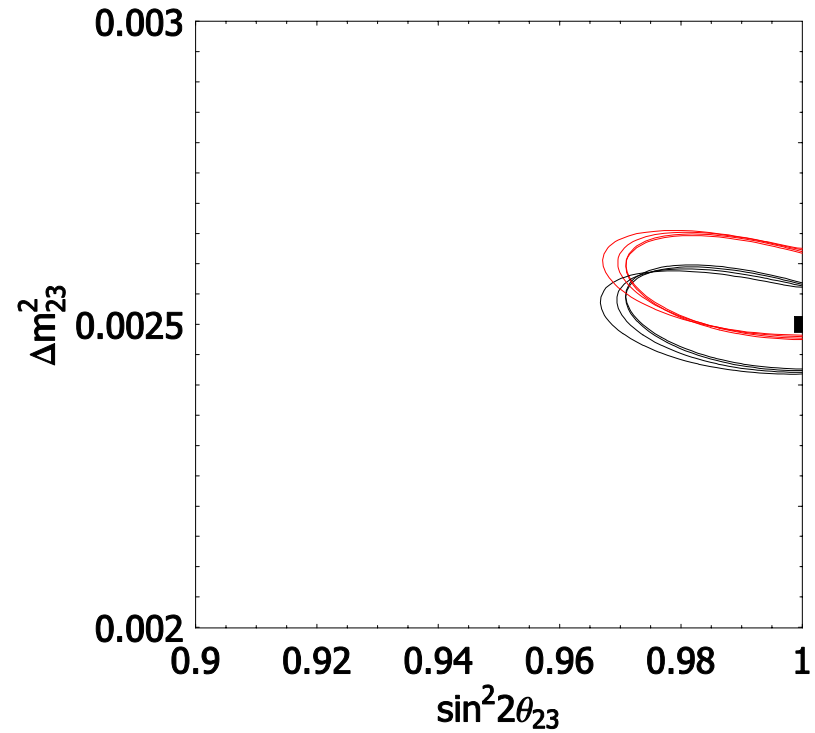
Systematic 0%  
No Background

Errors dominated by statistics

fact  $\pi$   $\nu$   $\nu$  10% systematic



Systematic 5%



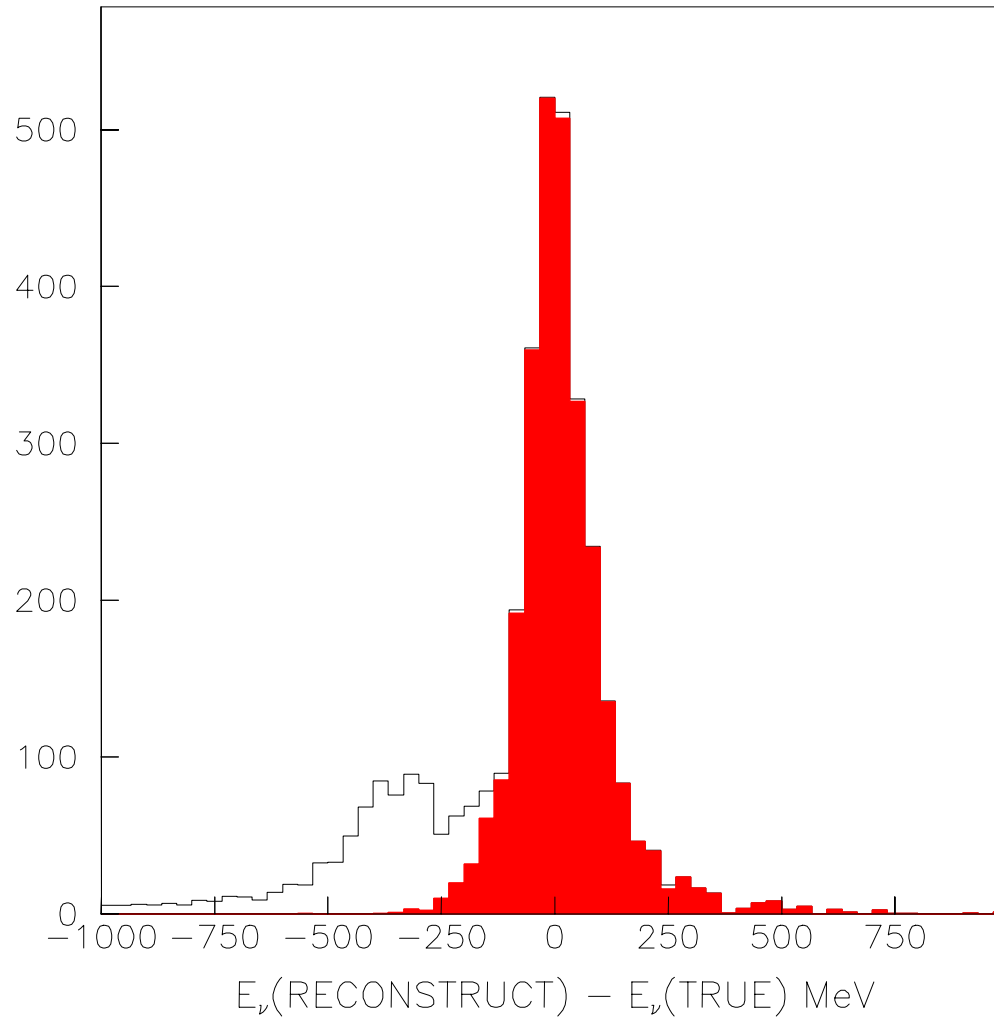
Systematic 10%

Errors dominated by statistics





# Energy Resolution

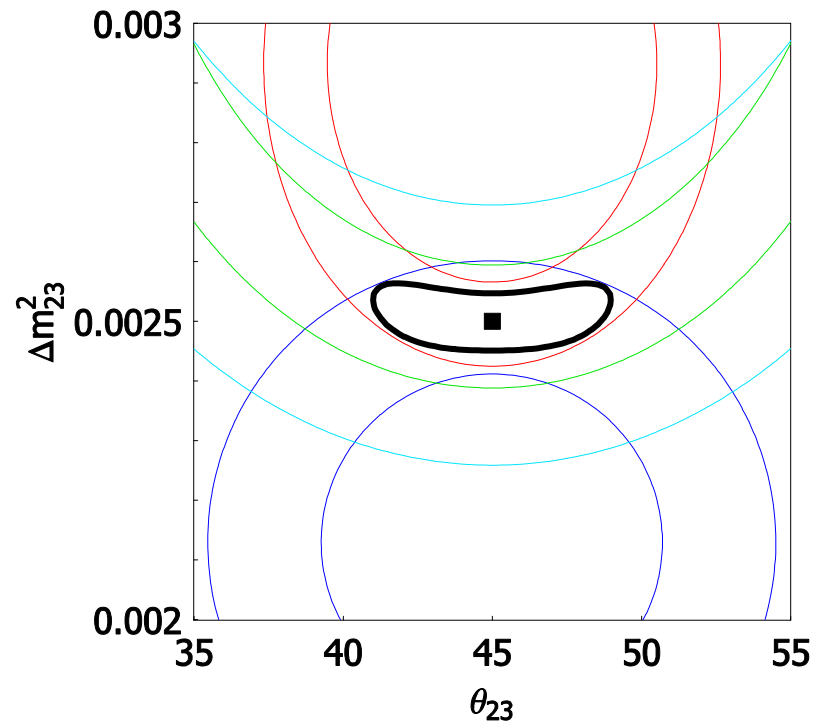


Red histogram for true QE events

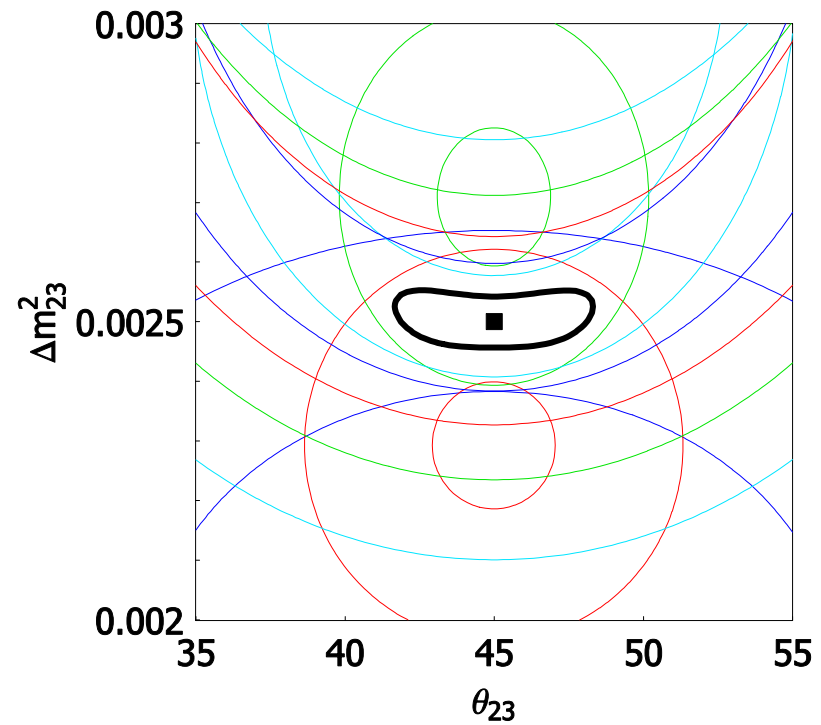
Figure taken from  
Y. Itow *et al.* hep-ex/0106019



# Double energy resolution



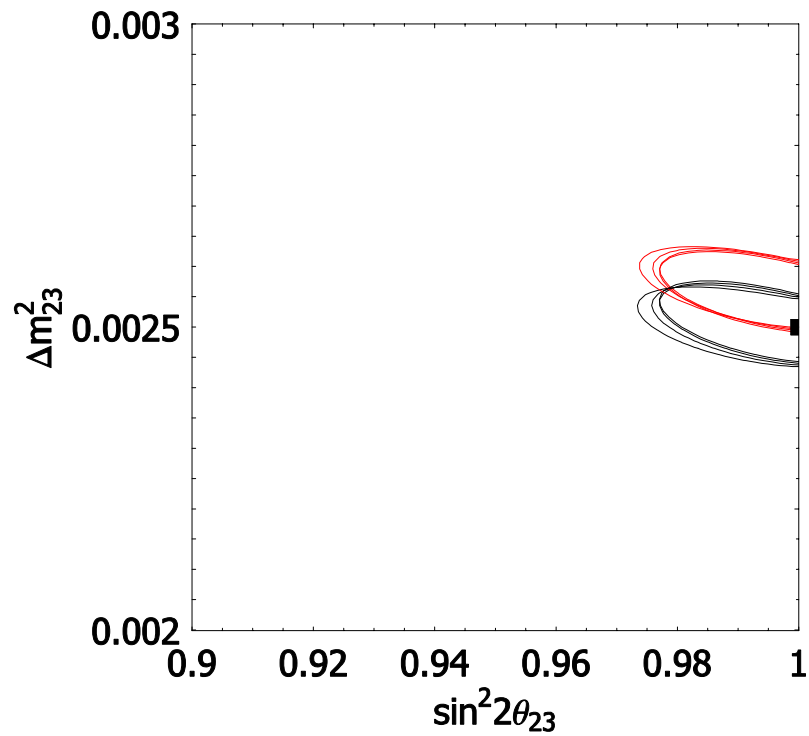
4 bins of 200MeV



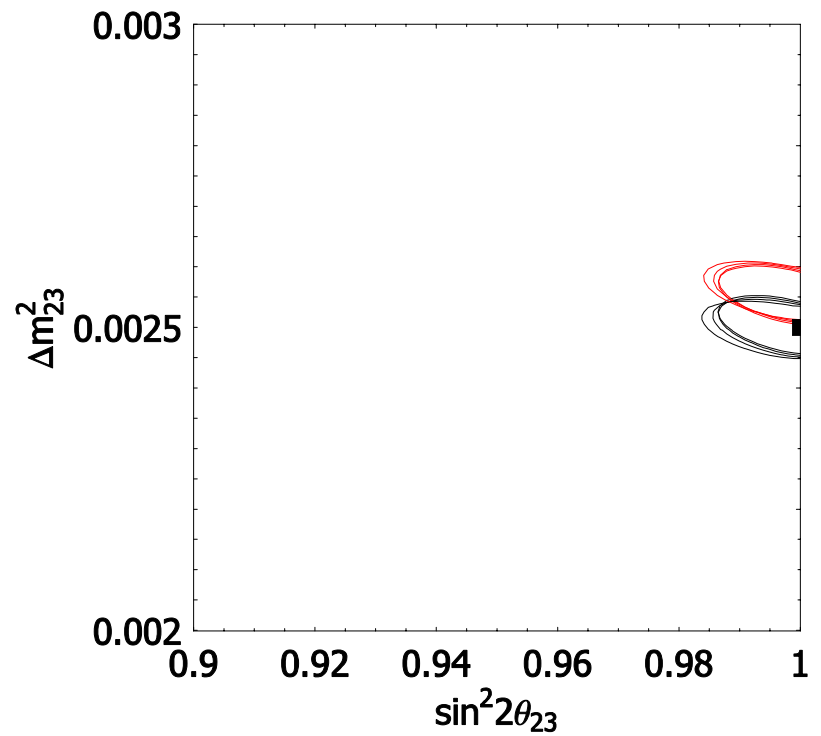
8 bins of 100MeV



# Double energy resolution



4 bins of 200MeV



8 bins of 100MeV