

main experimental predictions

in particle colliders

- Longit. TeV dims $\Rightarrow$ gauge interactions
- Transverse submm dims $\Rightarrow$ strong gravity
- TeV strings $\Rightarrow$ Regge excitations, black holes?

TeV dims: tower of Kaluza-Klein excitations
for SM particles

$$X \equiv X + 2\pi R \quad \Rightarrow \quad p = \frac{n}{R} \quad n=0, \pm 1, \dots$$

$$m_n^2 = m_0^2 + \frac{n^2}{R_{//}^2}$$

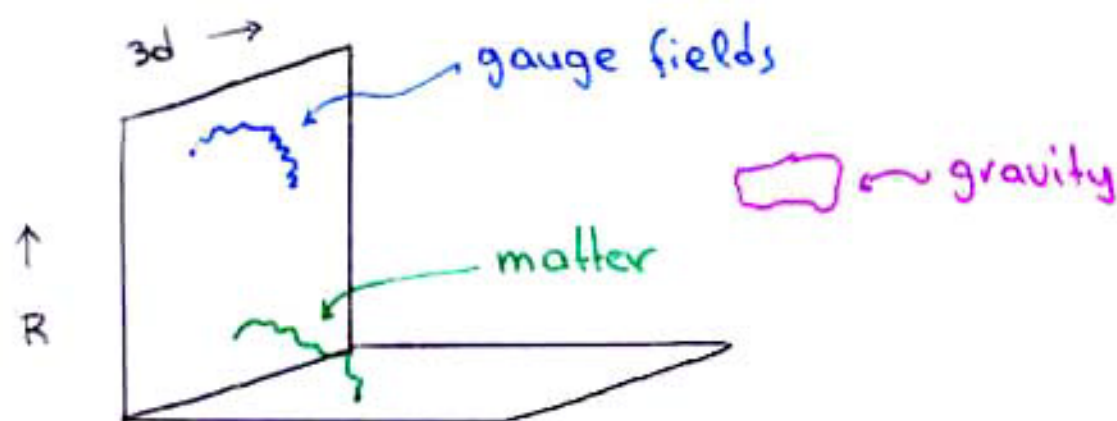
$$R_{//}^{-1} \lesssim M_s$$

$$\Rightarrow \gamma_n, Z_n, W_n^{\pm}, G_n^a$$

- In general SM gauge groups on different branes
- Quarks + leptons : on brane intersections
 - localized in 3d $\Rightarrow$ no KK
 - automatic chirality
 - KK of gauge bosons unstable

otherwise KK momentum conservation

$\Rightarrow$ stable charged excitations



similar to Z_2 -orbifolds of heterotic string I.A. '90

I.A. - Benakli '94

Exp. constraints

* equal couplings

$$= g$$

$$= g \int d^4x \frac{n^2}{R^4 M_s^2} \sim g \quad R M_s \rightarrow \infty$$

$\delta \geq 1$

* bounds from 4-fermion eff ops (compositeness)

$$\sum_{n \neq 0} \approx E \ll R^{-1} \sim R^2 \sum_{\vec{n} \neq 0} \frac{1}{n^2}$$

for more than $d=2$ dims $\Rightarrow$ regulated sum

$$\Rightarrow \sim R^2 (R M_s)^{d-2} \quad \text{modulo logs for } d=2$$

$$\Rightarrow \boxed{R^{-1} \gtrsim \text{TeV}} \quad \text{up to 2 dims if } R M_s \gg 1$$

I.A. - Benakli '94

high precision of z -width + $G_F \Rightarrow R^{-1} \gtrsim 2.5 \text{ TeV} \quad d=4$

Nath-Yamaguchi
Masip-Pomarol
Marciano '99

Experimental constraints

bounds from 4-fermion effective operators (compositeness)

$$\sum_{n \neq 0} \text{diagram with wavy line } n \simeq_{E \ll R^{-1}} \text{diagram with point} \sim R^2 \sum_{n \neq 0} \frac{1}{n^2}$$

more than 2 dims $\Rightarrow$ regulated sum

$$\Rightarrow \sim R^2 (RM_s)^{d-2} \text{ modulo logs for } d=2$$

$$\Rightarrow R^{-1} \gtrsim \text{TeV}$$

I.A. - Benakli '94

high precision of Z -width + $G_F \Rightarrow R^{-1} \gtrsim 3 \text{ TeV}$

Nath-Yamaguchi

Masip-Pomarol

Marciano, Strumia

Delgado-Pomarol-Quiroz

'99

$\Rightarrow$ LHC: production at most one KK resonance $R^{-1} \leq 6 \text{ TeV}$

I.A. - Benakli - Quiroz '94, '99

Nath-Yamada-Yamaguchi

Rizzo - Wells '99

I.A. - Accomando - Benakli

I.A.-Benakli-Quiroz '94, '99

$$q\bar{q} \rightarrow \gamma_n, Z_n \rightarrow \ell^+\ell^-$$

I.A.-Accomando-Benakli '99

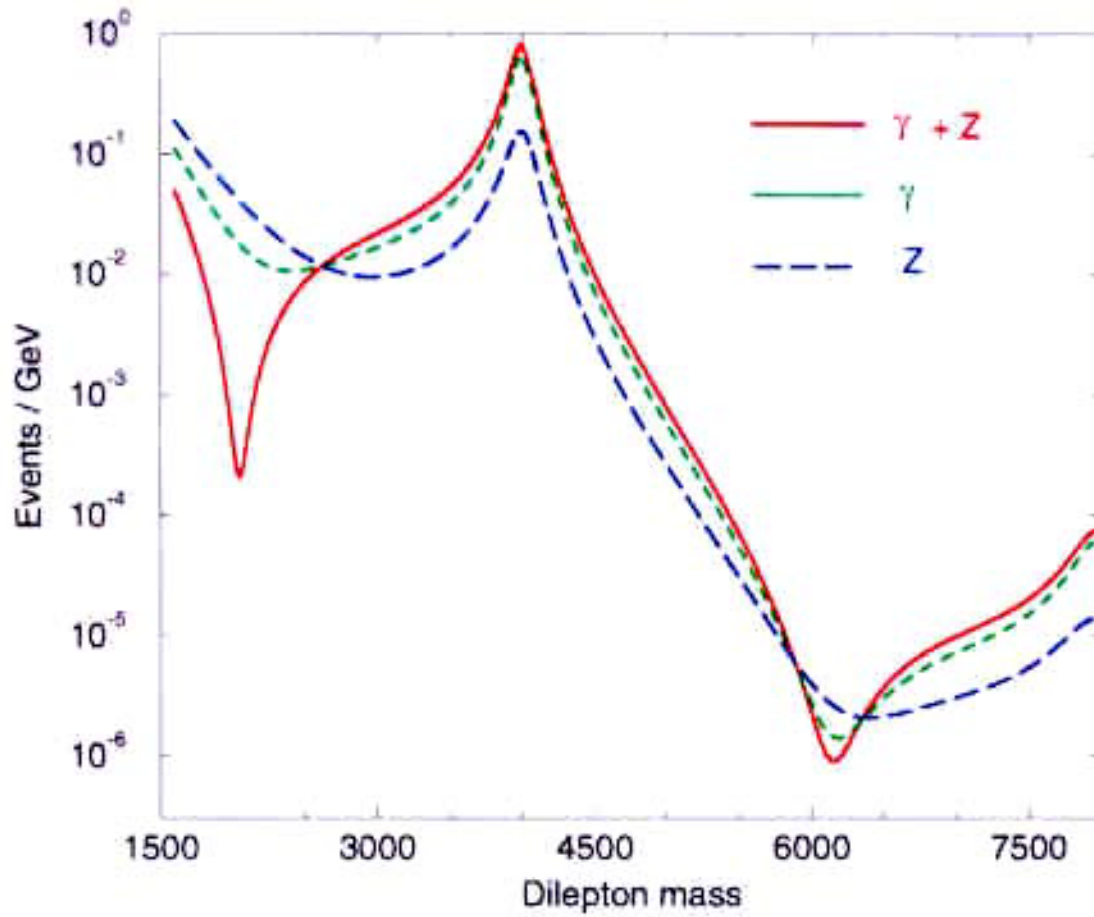


Figure 5: First resonances in the LHC experiment due to a KK excitation of photon and Z for one extra-dimension at 4 TeV. From highest to lowest: excitation of photon+Z, photon and Z boson.

$$Q\bar{Q} \rightarrow W_n^{\pm} \rightarrow \ell^{\pm} \nu$$

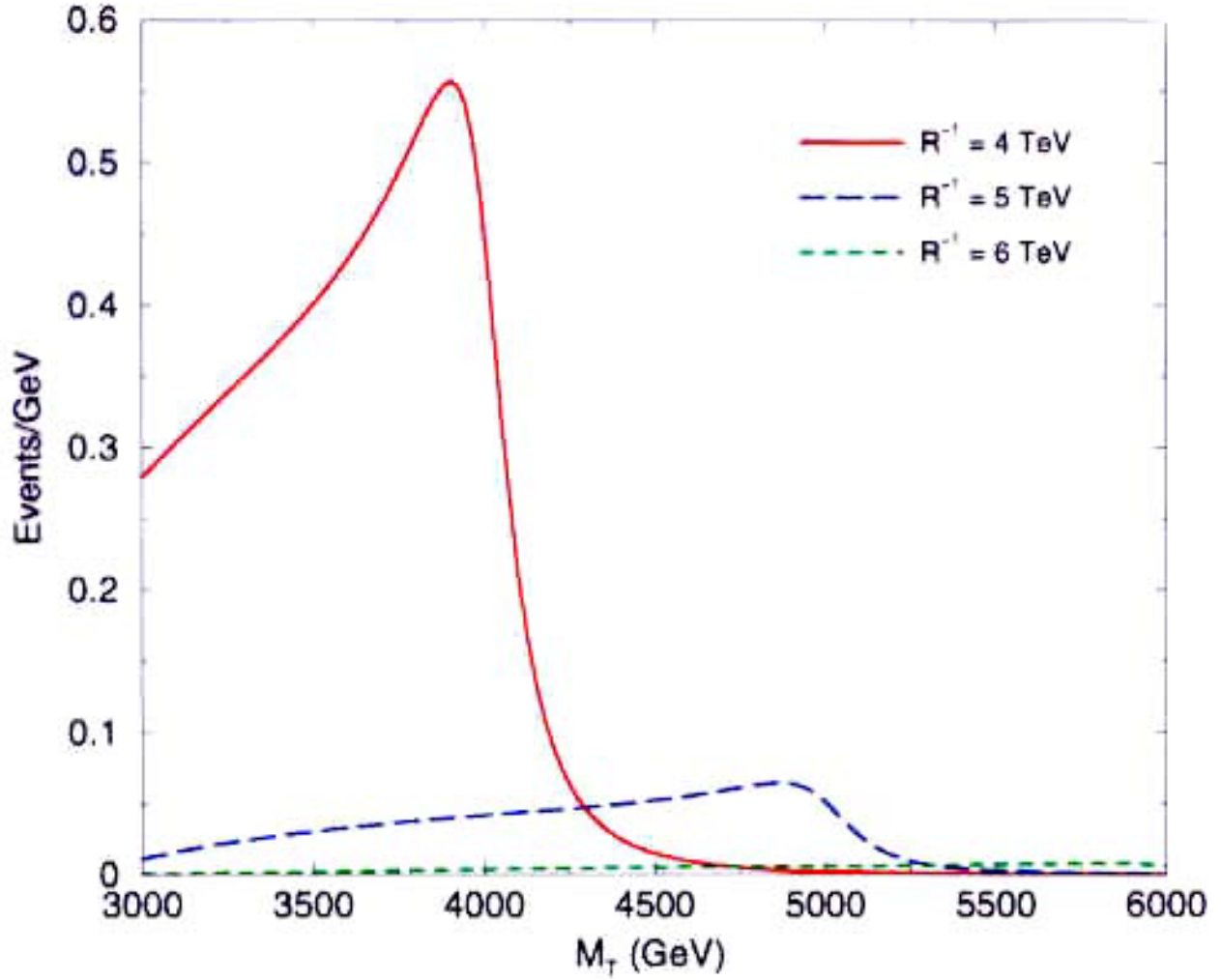


Figure 3: First resonances in the LHC experiment due to a KK excitation of $W_{\pm}^{(n)}$ for one extra-dimension at 4, 5 and 6 TeV. We plot the differential cross section as function of the transverse mass for the W 's.

$$q\bar{q} \rightarrow G_n \rightarrow 2 \text{ jets}$$

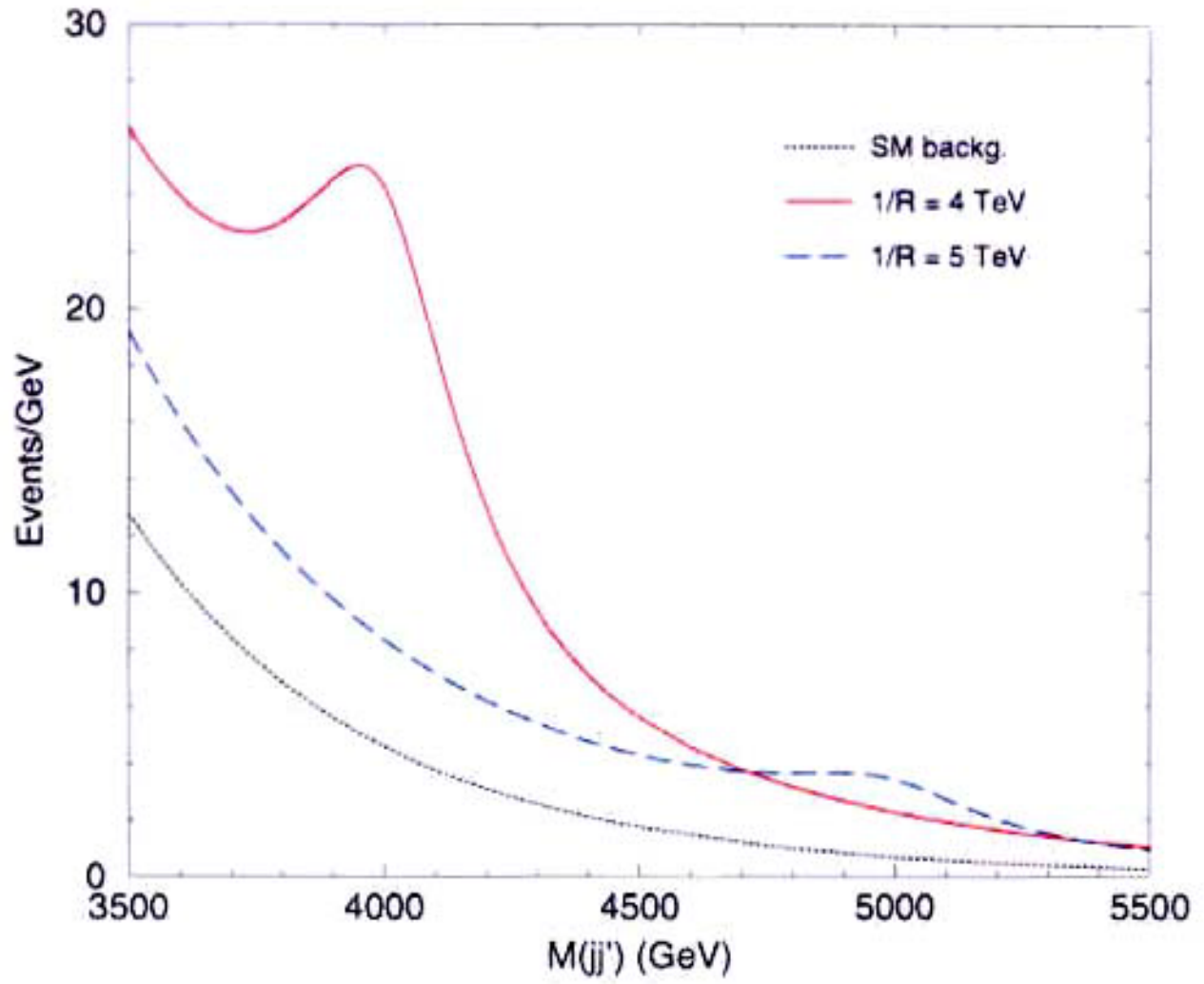


Figure 1: First resonances in the LHC experiment due to a KK excitation of gluon for one extra-dimension at 4 and 5 TeV.

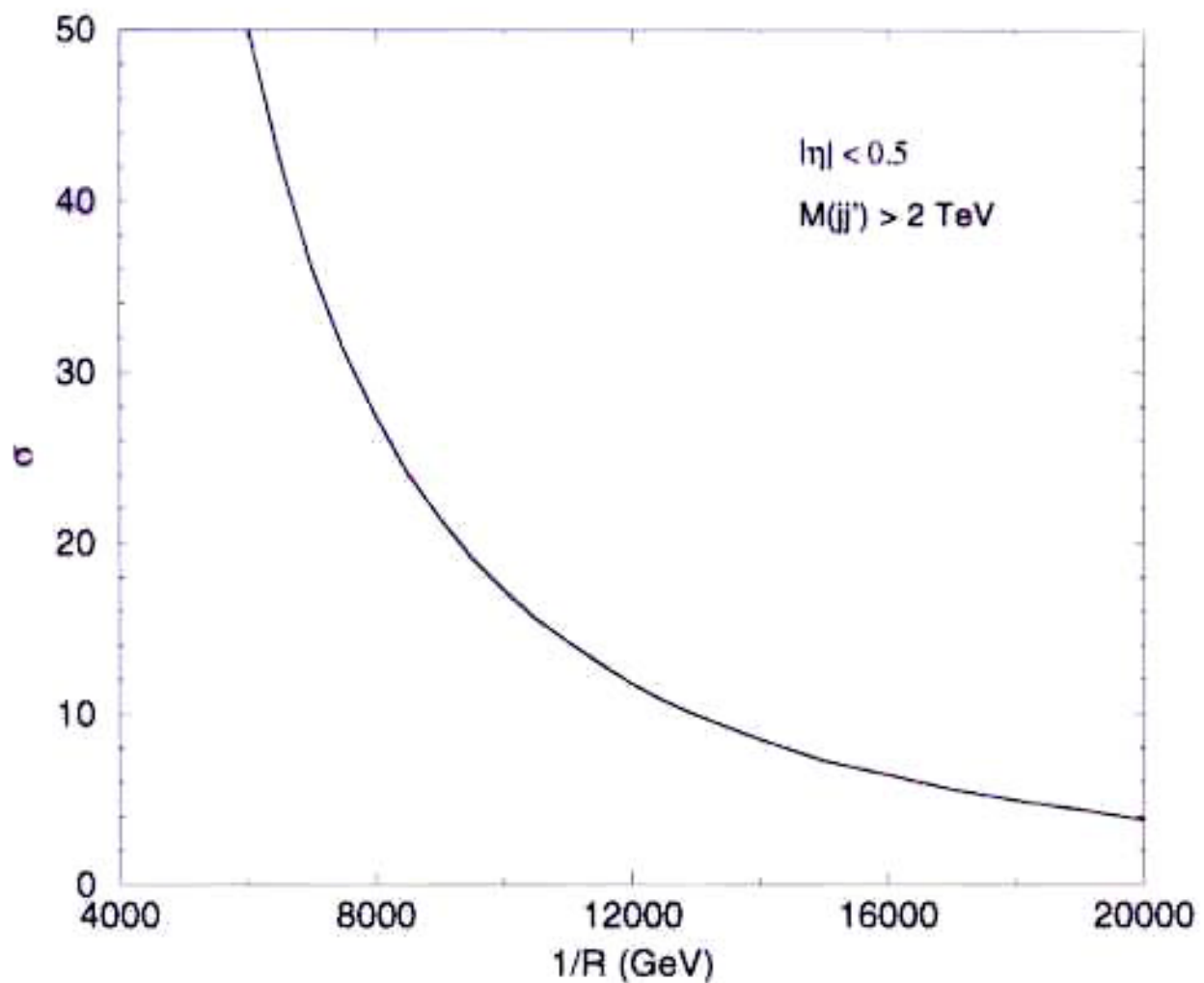


Figure 2: Number of standard deviation in number of observed dijets from the expected standard model value, due to the presence of a TeV-scale extra-dimension of compactification radius R .

$$R^{-1} \gtrsim 20 \text{ TeV} \quad 95\% \text{ CL}$$

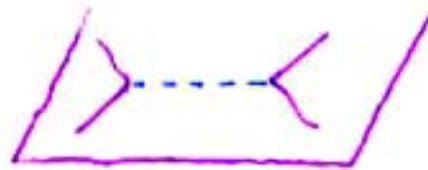
TEVATRON : 4 TeV

If no longitudinal dims with $R_{||}^{-1} \lesssim M_I \Rightarrow$

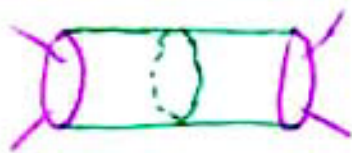
main signal :- graviton emission in the bulk

$\Rightarrow$ low scale quantum gravity

- indirect effects $\Rightarrow$ probe string modes



Type I string theory: **graviton emission** subdominant
compared to Regge exchanges



1-loop $\Rightarrow \lambda^2$

disk $\Rightarrow \lambda$

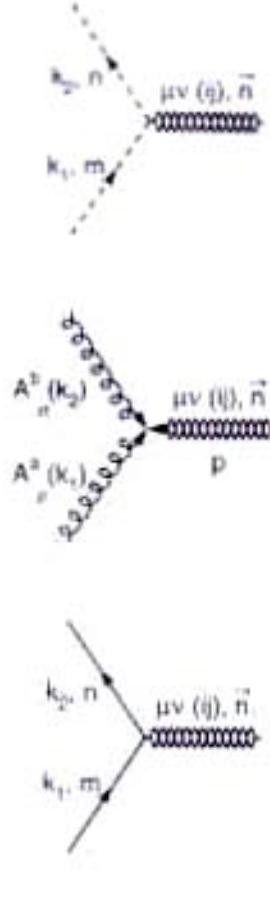
$\lambda \sim g^2 \Rightarrow$ loop factor enhancement $\uparrow$

A-A-B '99

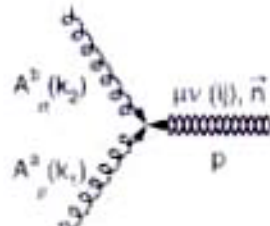
Lallen-Perelstein-Peskin '00

for each KK level.

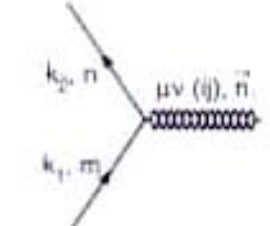
A.2 Vertex Feynman Rules



$$\begin{aligned}
 \vec{h}_{\mu\nu}^{\vec{n}} \Phi\Phi &: -i\kappa/2 \delta_{mn} (m_\Phi^2 \eta_{\mu\nu} + C_{\mu\nu,\rho\sigma} k_1^\rho k_2^\sigma) \\
 \vec{\phi}_i^{\vec{n}} \Phi\Phi &: i m_\Phi \delta_i \delta_{mn} (k_1 \cdot k_2 - 2 m_\Phi^2)
 \end{aligned}$$



$$\begin{aligned}
 \vec{h}_{\mu\nu}^{\vec{n}} AA &: -i\kappa/2 \delta^{ab} ((m_A^2 + k_1 \cdot k_2) C_{\mu\nu,\rho\sigma} + D_{\mu\nu,\rho\sigma}(k_1, k_2) \\
 &\quad + \xi^{-1} E_{\mu\nu,\rho\sigma}(k_1, k_2)) \\
 \vec{\phi}_i^{\vec{n}} AA &: i\omega \kappa \delta_i \delta^{ab} (\eta_{\mu\nu} m_A^2 + \xi^{-1} (k_{1\mu} p_\nu + k_{2\mu} p_\nu))
 \end{aligned}$$



$$\begin{aligned}
 \vec{h}_{\mu\nu}^{\vec{n}} \Psi\Psi &: -i\kappa/8 \delta_{mn} (\gamma_\mu (k_{1\nu} + k_{2\nu}) + \gamma_\nu (k_{1\mu} + k_{2\mu}) \\
 &\quad - 2 \eta_{\mu\nu} (k_1 + k_2 - 2 m_\Psi)) \\
 \vec{\phi}_i^{\vec{n}} \Psi\Psi &: i\omega \kappa \delta_i \delta_{mn} (3/4 k_1 + 3/4 k_2 - 2 m_\Psi)
 \end{aligned}$$

Figure 4: Three-point vertex Feynman rules. The KK states are plot in double-sinusoidal curves. The symbols $C_{\mu\nu,\rho\sigma}$, $D_{\mu\nu,\rho\sigma}(k_1, k_2)$ and $E_{\mu\nu,\rho\sigma}(k_1, k_2)$ are defined in Eqs. (A.10), (A.11) and (A.12) respectively. m_Φ , m_A and m_Ψ are masses of the scalar, vector and fermion. $\omega = \sqrt{\frac{2}{\eta(1+\eta)}}$, $\kappa = \sqrt{16\pi G_N}$ and ξ is the gauge-fixing parameter.

In the following we list the complete leading order Feynman rules in three figures, Figs. 4, 5 and 6. Some of the symbols used are defined as follows:

$$C_{\mu\nu,\rho\sigma} = \eta_{\mu\rho}\eta_{\nu\sigma} + \eta_{\mu\sigma}\eta_{\nu\rho} - \eta_{\mu\nu}\eta_{\rho\sigma} , \quad (\text{A.10})$$

$$D_{\mu\nu,\rho\sigma}(k_1, k_2) = \eta_{\mu\nu} k_{1\rho} k_{2\sigma} - \left[\eta_{\mu\rho} k_{1\nu} k_{2\sigma} + \eta_{\mu\sigma} k_{1\nu} k_{2\rho} - \eta_{\rho\sigma} k_{1\mu} k_{2\nu} + (\mu \leftrightarrow \nu) \right] , \quad (\text{A.11})$$

$$pp \rightarrow \text{jet} + G$$

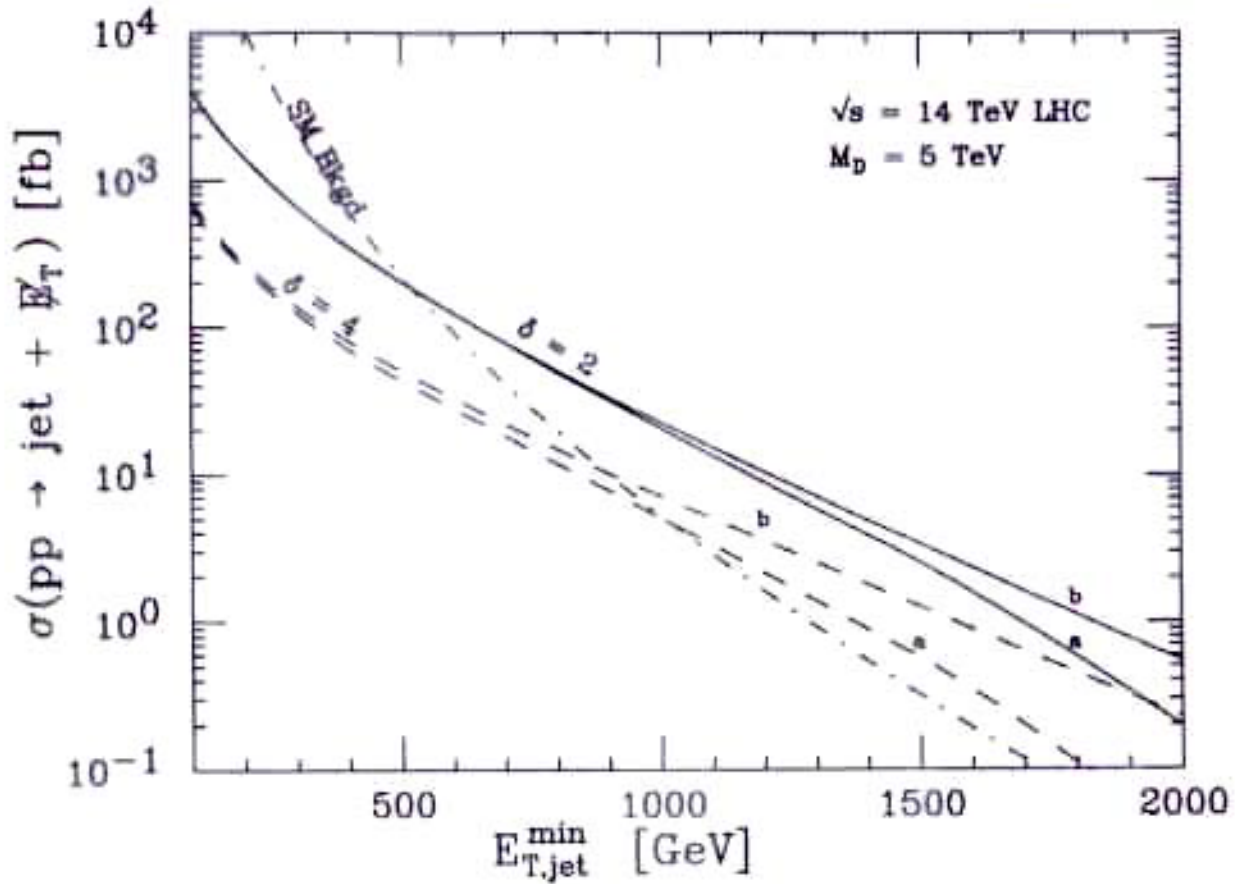
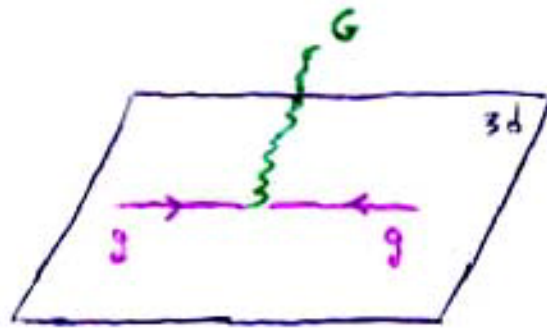


Figure 3: The total jet + nothing cross-section at the LHC integrated for all $E_{T,\text{jet}} > E_{T,\text{jet}}^{\min}$ with the requirement that $|\eta_{\text{jet}}| < 3.0$. The Standard Model background is the dash-dotted line, and the signal is plotted as solid and dashed lines for fixed $M_D = 5$ TeV with $\delta = 2$ and 4 extra dimensions. The a (b) lines are constructed by integrating the cross-section over $\hat{s} < M_D^2$ (all $\hat{s}$).

$$g g \rightarrow G$$



$$\sigma(E) \sim \frac{E^p}{M_I^{p+2}} \frac{\Gamma(1 - 2E^2/M_I^2)^2}{\Gamma(1 - E^2/M_I^2)^4}$$

- $E < M_I \Rightarrow \sim \frac{E^p}{M_I^{p+2}}$ gravity in $4+p$ dims
- $E \sim M_I \Rightarrow$ sequence of poles due to RR resonances
- $E > M_I \Rightarrow$ exp decay due to the UV softness of strings

J.A. - Arkani Hamed - Dimopoulos - Dvali '98

$E < M_I$: reliable computations within eff. field theory
 $\Rightarrow$ model independent predictions

$$p\bar{p} \rightarrow \text{jet} + G$$

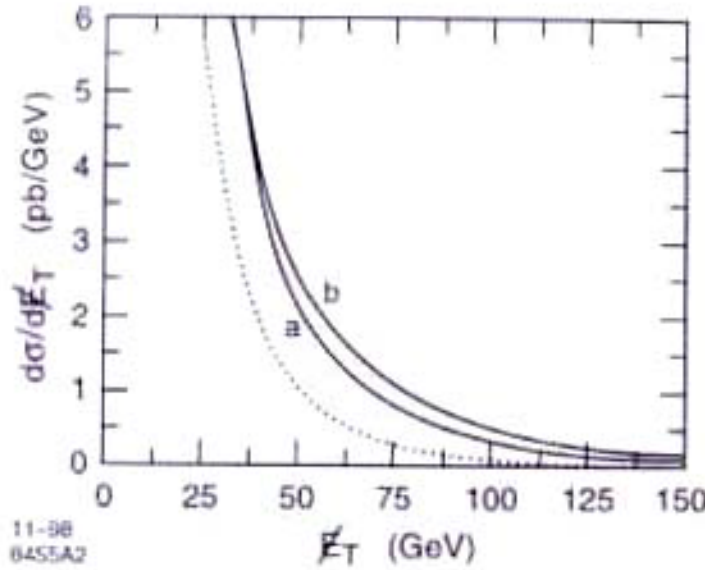


Figure 2: Spectrum of missing energy in events with one jet, computed for $p\bar{p}$ collisions at $\sqrt{s} = 1.8$ TeV, with a rapidity cut $|y| < 2.4$. The dotted curve is the Standard Model expectation. The solid curves show the additional cross section expected in the model of ref. [4] with (a) $n = 2$, $M = 750$ GeV, (b) $n = 6$, $M = 610$ GeV.

Current and future limits on $M_{p(n+4)}$

Collider	R / M ($n = 2$)	R / M ($n = 4$)	R / M ($n = 6$)
Present: LEP 2	$4.8 \times 10^{-2} / 1200$	$1.9 \times 10^{-9} / 730$	$6.9 \times 10^{-12} / 520$
Tevatron	$11.0 \times 10^{-2} / 750$	$2.4 \times 10^{-9} / 610$	$5.8 \times 10^{-12} / 610$
Future: Tevatron	$3.9 \times 10^{-2} / 1300$	$1.4 \times 10^{-9} / 900$	$4.0 \times 10^{-12} / 810$
LC	$1.2 \times 10^{-3} / 7700$	$1.2 \times 10^{-10} / 4500$	$6.5 \times 10^{-13} / 3100$
LHC	$3.4 \times 10^{-3} / 4500$	$1.9 \times 10^{-10} / 3400$	$6.1 \times 10^{-13} / 3300$

Table 1: Current and future sensitivities to large extra dimensions, expressed as 95% confidence limits on the size of extra dimensions R (in cm) and the effective Planck scale M (in GeV). The assumptions of each analysis are explained in the text.

- virtual graviton exchange



$$\lim_{r \rightarrow \infty} \frac{1}{M_P^2} \sum_{\vec{n}} \frac{1}{s - \frac{\vec{n}^2}{r^2}}$$

$$M_P^2 = \frac{1}{g^4} M_I^{2+p} r^p$$

more than 2 dims $\Rightarrow$ regulated sum

$$\simeq \frac{g^4}{M_I^{2+p} r^{p-2}} \underbrace{\sum_{\vec{n}} \frac{\delta^{-\frac{\vec{n}^2}{r^2} M_I^2}}{\vec{n}^2 - s r^2}}_{\frac{2}{p-2} (r M_I)^{p-2}}$$

$\Rightarrow$ dimension 8 effective operator $(\bar{\psi} \psi)^2$:

$$\frac{g^4}{M_I^4} \cdot \begin{cases} \frac{2}{p-2} & p > 2 \\ \ln M_I^2/s & p = 2 \end{cases}$$

- open string Regge exchanges: $\frac{g^2}{M_I^4}$ for every p

$\Rightarrow$ stronger limits for $p > 2$: $M_I \gtrsim 1 \text{ TeV}$

Cullen - Perelstein - Peskin

Gravity modification at submm scales

1) Change of Newton's law

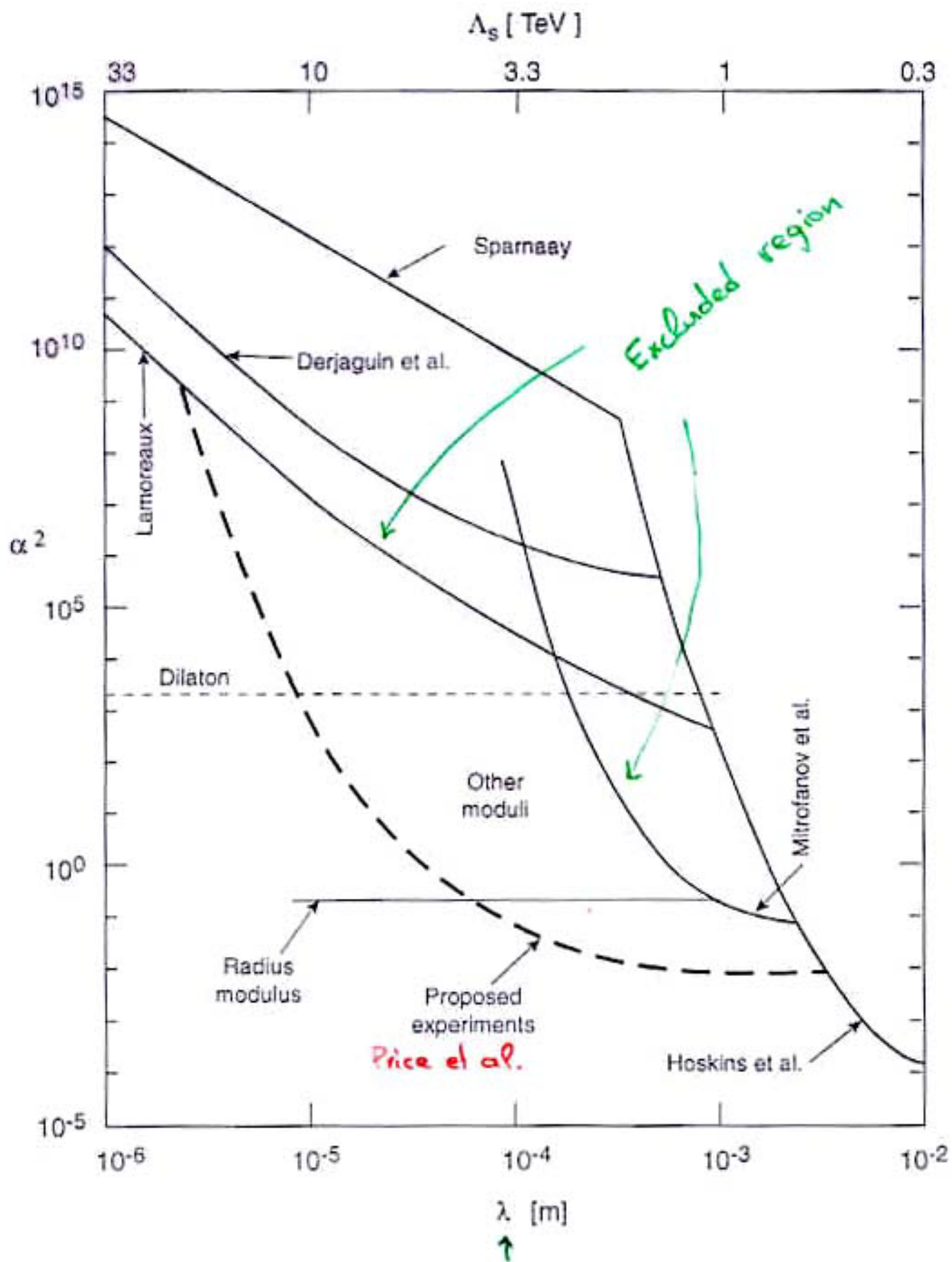
$$\frac{1}{r^2} \longrightarrow \frac{1}{r^{2+p}}$$
$$\begin{array}{l} p=2 \Rightarrow \frac{1}{r^4} \text{ in (sub)mm} \leftarrow \text{testable} \\ \vdots \\ p=6 \Rightarrow \frac{1}{r^8} \text{ in (sub)fm} \end{array}$$

2) susy breaking $\Rightarrow$ light scalars: $\frac{(TeV)^2}{M_P} \sim 10^{-6} eV = 1 mm^{-1}$

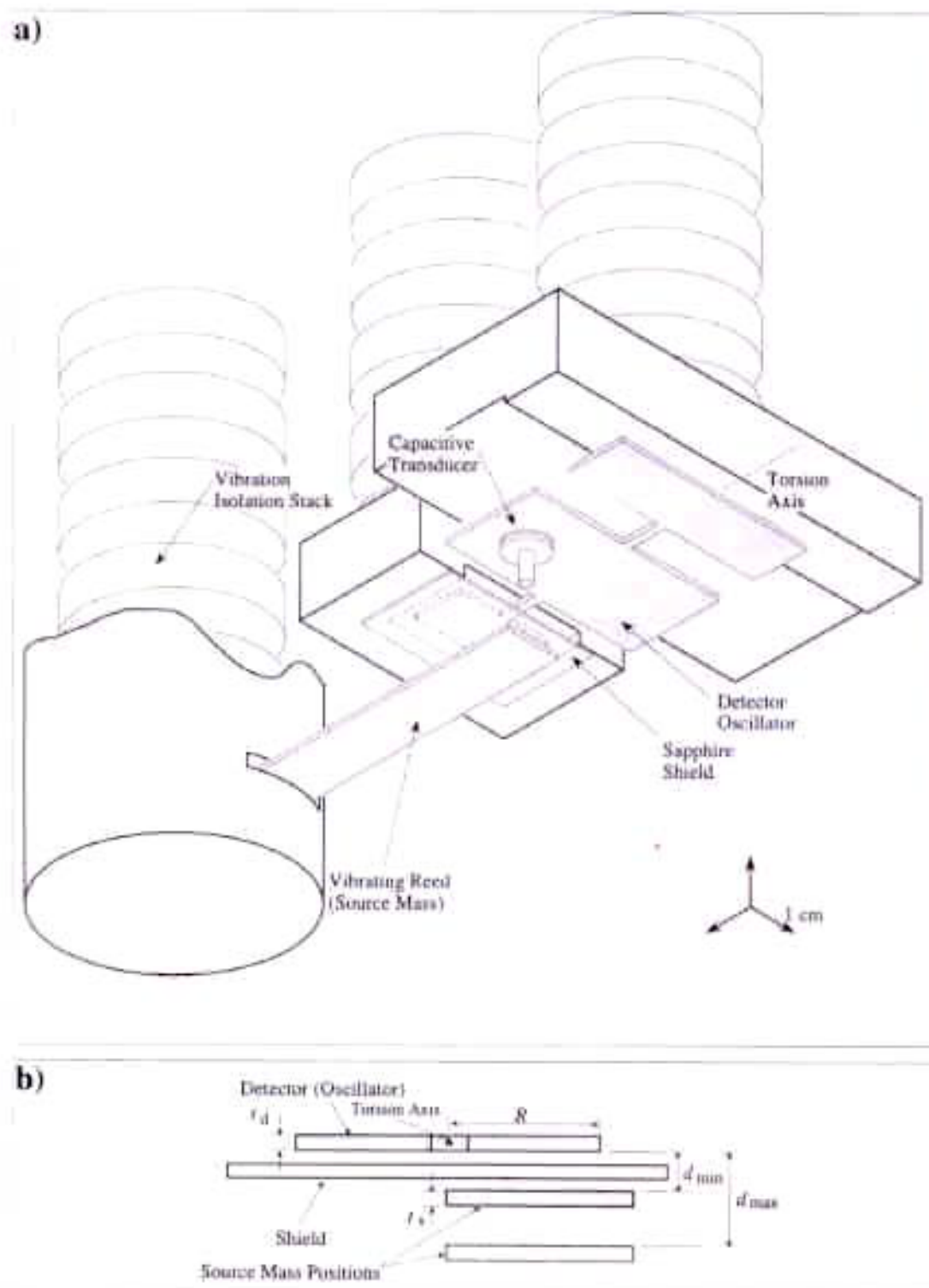
(i) primordially on the brane $\Rightarrow$ suppressed in the bulk

$$P_{\text{brane}} \sim M^4 \Rightarrow P_{\text{bulk}} \sim M^4 / r^p$$

$$\Rightarrow m_{\text{modulus}}^2 \sim \frac{M^4}{M_P^2} \quad \text{modulus} \equiv \ln r$$



Van der Waals $\approx$ gravitational



Gauge hierarchy

$M_P \gg M_Z \Rightarrow$ why large transverse dims?

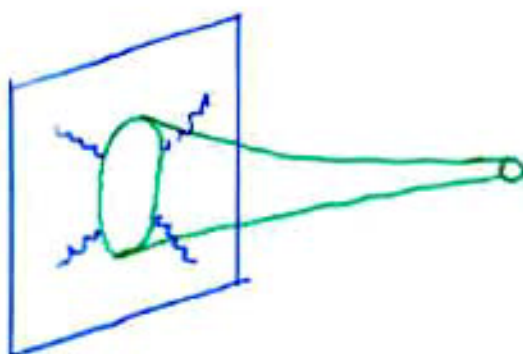
$$r M_I \simeq \left(g^2 \frac{M_P}{M_I} \right)^{2/n} \sim \begin{cases} n=2 & 10^{15} \\ n=6 & 10^5 \end{cases} \quad \text{or } \lambda_{II} \simeq 10^{-14}$$

Technical aspect: stability in a non susy vacuum

no large corrections to SM couplings as $r M_I \rightarrow \infty$

In general no decoupling if massless bulk fields propagate in less than 2 large transv. dims

I.A. - Bachas '98



UV divergence: open string loop

IR divergence: emission of massless closed string

$d_I = 1$: linear IR div $\Rightarrow$ quadratic UV $r \sim M_P^2$

Condition: no bulk propagation in one large dim

or local tadpole cancellation $\Rightarrow$ severe constraints

$d_1=2$: log divergences

can be absorbed into a finite number of parameters:

values of bulk massless fields at the brane position

similar to renormalizable field theory

RGE resum $\Rightarrow$ classical 2d eqs in the transverse space

log dependence $\Rightarrow$ higher orders irrelevant

$\Rightarrow$ hierarchy could be determined by minim SM eff. potential

$\Rightarrow$ No susy TeV strings:

same protection of hierarchy as softly susy at TeV

Do we need susy if $M_{str} \sim \text{TeV}$?

Type I: non susy string models $\Rightarrow$

$$\Lambda_{\text{bulk}} \sim M_I^{4+n} \Rightarrow \Lambda_{\text{brane}} \sim M_I^{4+n} r^n \sim M_I^2 M_P^2$$

analog of quadratic div. to Λ in softly broken susy

absence of quadratic sensitivity:

- $\Lambda = 0$ (special models)

$$- \Lambda_{\text{brane}} \sim M_I^4 \Rightarrow \Lambda_{\text{bulk}} \sim M_I^4 / r^n$$

satisfied if approximate susy in the bulk

e.g. susy is broken primordially only on the brane

explicit realization: Brane susy breaking

I.A. - Dudas - Sagnotti '99

Aldazabal - Uranga '99

No susy in our world (brane)

but it may exist 1 mm away!

to protect the gauge hierarchy against gravit. corrections

Prediction: possible new forces at submm scales

e.g. light scalars: $\frac{(\text{TeV})^4}{M_P} \sim 10^{-4} \text{ eV} = 1 \text{ mm}^{-1}$

modulus $\equiv \ln r$

coupling to nucleons relative to gravity:

$$\frac{1}{m_N} \frac{\partial m_N}{\partial \ln r} = \frac{\partial \ln \Lambda_{\text{QCD}}}{\partial \ln r} \quad m_N \sim \Lambda_{\text{QCD}} \sim e^{-\frac{1}{b_{\text{QCD}}} \frac{2\pi}{\alpha_{\text{QCD}}}}$$
$$\sim \frac{\partial}{\partial \ln r} \alpha_{\text{QCD}}$$

α_s in models with log sensitivity in r e.g. $d_1 = 2$

$\Rightarrow$ can be experimentally tested

Origin of EW symmetry breaking?

mild hierarchy: $\frac{m_W}{M_S} \lesssim O(10^{-1})$

string tree-level: - $m_W = 0$

- $m_W \sim n M_S$ $n \gtrsim O(1)$

- $m_W \sim a \curvearrowright$ flat direction

$\Rightarrow$ only possibility: radiative breaking

$$V = \lambda (h^\dagger h)^2 + \mu^2 (h^\dagger h)$$

(a) $\mu = 0$ at tree

$\mu^2 < 0$ at one loop non susy vacuum

(b) $\langle h \rangle$ flat at tree

lifted radiatively (2 higgses)

I.A. - Benakli - Quiros hep-ph/0004091

simplest case: one SM higgs on our brane-world

open string with both ends on the brane $\Rightarrow$

(1) tree-level potential $\equiv$ same as susy

$$\lambda = \frac{1}{8} (g^2 + g'^2) \quad \text{D-terms}$$

$$(2) \quad \mu^2 = - \frac{1}{N} g^2 \epsilon^2 M_s^2$$

N : order of the orbifold group

ϵ : estimated by an explicit computation in
a toy model

$$\langle h \rangle = (0, v/\sqrt{2}) : \quad v^2 = -\mu^2/\lambda \quad \Rightarrow$$

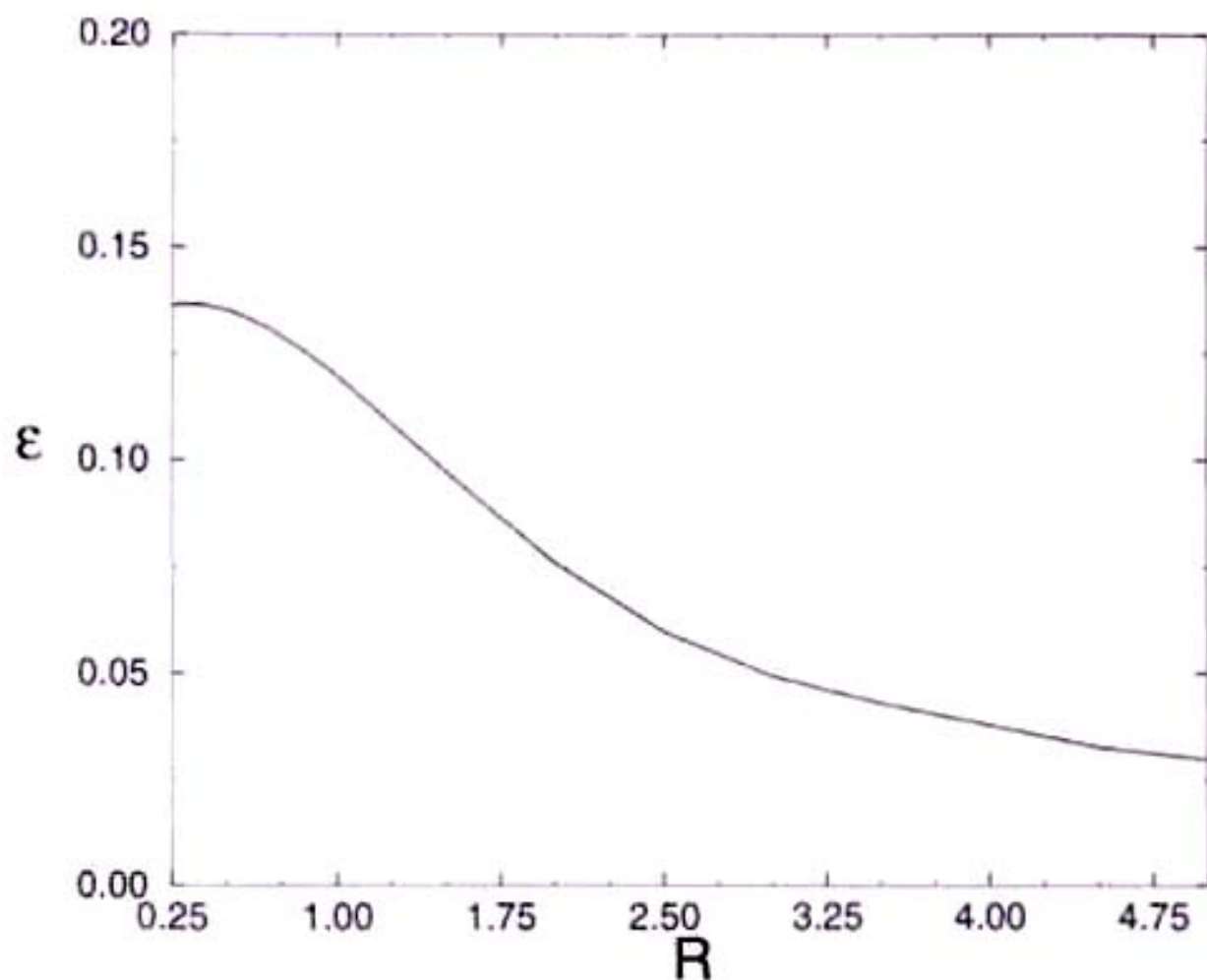
Expansion around the origin

gauge invariance : $a^2 \rightarrow \text{Tr } \Phi^2$

$$V_{\text{eff}} = V_0 + \frac{1}{2} \mu^2 \text{Tr } \Phi^2 + \mathcal{O}(\Phi^4)$$

$$\mu^2 = -\varepsilon^2(R) g^2 M_s^2$$

$$\varepsilon^2(R) = \frac{R^3}{2\pi^2} \int_0^\infty \frac{d\ell}{(2\ell)^{5/2}} \frac{\theta_2^4}{4\eta^{12}} \left(i\ell + \frac{1}{2} \right) \sum_n n^2 e^{-2\pi n^2 R \ell}$$



$$(1) \quad M_h = M_Z$$

same as MSSM for $\tan\beta, m_A \rightarrow \infty$

$$(2) \quad M_s = \frac{M_h \sqrt{N}}{\sqrt{2} g \epsilon}$$

- Low-energy SM radiative corrections top-quark sector

$$\Rightarrow M_h \sim 120 \text{ GeV}$$

$$M_s \sim \text{a few TeV} \quad \mathcal{O}(1-10)$$

- String threshold corrections

- model dependent

- can be very important

e.g. for two large dims in the bulk

$$\ln R_\perp M_s \sim \ln M_P / M_s$$

$\Rightarrow$ both M_s and M_h can be increased

sufficiently low M allows power-law running

in the region $R^{-1} \rightarrow M$

$$\frac{1}{g_i^2(\mu)} = \frac{1}{g_i^2(R^{-1})} - \frac{b_i^{\text{SM}}}{8\pi^2} \ln(\mu R) - \frac{b_i^{\text{KK}}}{8\pi^2} \left(2 \overset{\substack{\text{number of KK} \\ \downarrow}}{\mu R} - \ln \mu R \right)$$

$N=1$ SUSY $\Rightarrow$ KK at least $N=2$

minimal choice: No KK for chiral states

• $SU(3) \times SU(2) \times U(1)$ 1 vector + 2 W-fermions + 2 scalars

• 1 Higgs hypermultiplet 2 fermions + 4 scalars

$\Rightarrow$ approximate unification at weak coupling

Dienes-Dudas-Gherghetta '98

e.g. $R^{-1} = 1 \text{ TeV} \Rightarrow M \simeq 45 \text{ TeV}$ within 2%

however power UV sensitivity $\Rightarrow$

(string) thresholds are important

Alternative possibility:

Bochos '98

M_I is at TeV but gauge couplings unify at a higher scale due to log sensitivity for $d_\perp = 2$

$$\ln r M_I^2 = \ln \frac{r M_I^2}{M_I}$$

winding scale: $r_D^{-1} \equiv r M_I^2 \sim 10^{16}$ GeV for $\begin{cases} d_\perp = 2 \\ M_I = 10 \text{ TeV} \end{cases}$
T-dual to $r^{-1} \sim \text{mm}^{-1}$

effective "running" in the transverse space (bulk)

however in $N=1$ orientifolds logs are also controlled

by $N=2$ β -functions $\Rightarrow$ model dependent

I.A. - Bochos - Dudas '99

$$\frac{1}{g_a^2} = \underbrace{\frac{1}{g^2}}_{\text{tree}} + s_a m + \underbrace{\sum_{i=1}^3 \frac{\lambda_i}{b_a}}_{\text{1-loop}} \left(\ln T_i + f(U_i) \right)$$

- $\frac{1}{g^2}$: same collection of branes

- m : twisted (blow-up) moduli

in all examples: anomalous $U(1)$'s $\Rightarrow m=0$

- $T = R_1 R_2$ $U = \frac{R_1}{R_2}$ for every torus $i=1,2,3$

$$R_1 \sim R_2 \rightarrow \infty \Rightarrow b_a \ln R_1 R_2 \quad d_{\perp} = 2$$

$$R_1 \gg R_2 \Rightarrow b_a \frac{R_1}{R_2} \quad d_{\perp} = 1$$

$\uparrow$
 $f(U) \sim \frac{R_1}{R_2}$

A D-brane embedding of the Standard Model

I.A. - Kiritsis - Tomaroi hep-th/0006114

N coincident branes $\Rightarrow U(N)$

$U(1)$: coupling = $g_N / \sqrt{2N}$

with charge of $\frac{N}{2} = 1$

$\Rightarrow$ gauged "baryon" number

$\Rightarrow$ minimal choice: $U(3) \times U(2) \times U(1)$

color branes (g_3) weak branes (g_2) g_1

$U(1)$ brane with $\begin{cases} U(3) \Rightarrow g_1 = g_3 \\ U(2) \Rightarrow g_1 = g_2 \end{cases}$

fermion generation

$$U(3) \times U(2) \times U(1)$$

$$Q \quad (3, 2; 1, w, 0)_{1/6} \quad w = \pm 1$$

$$u^c \quad (\bar{3}, 1; -1, 0, x)_{-1/3} \quad x = \pm 1 \text{ or } 0$$

$$d^c \quad (\bar{3}, 1; -1, 0, y)_{1/3} \quad y = \pm 1 \text{ or } 0$$

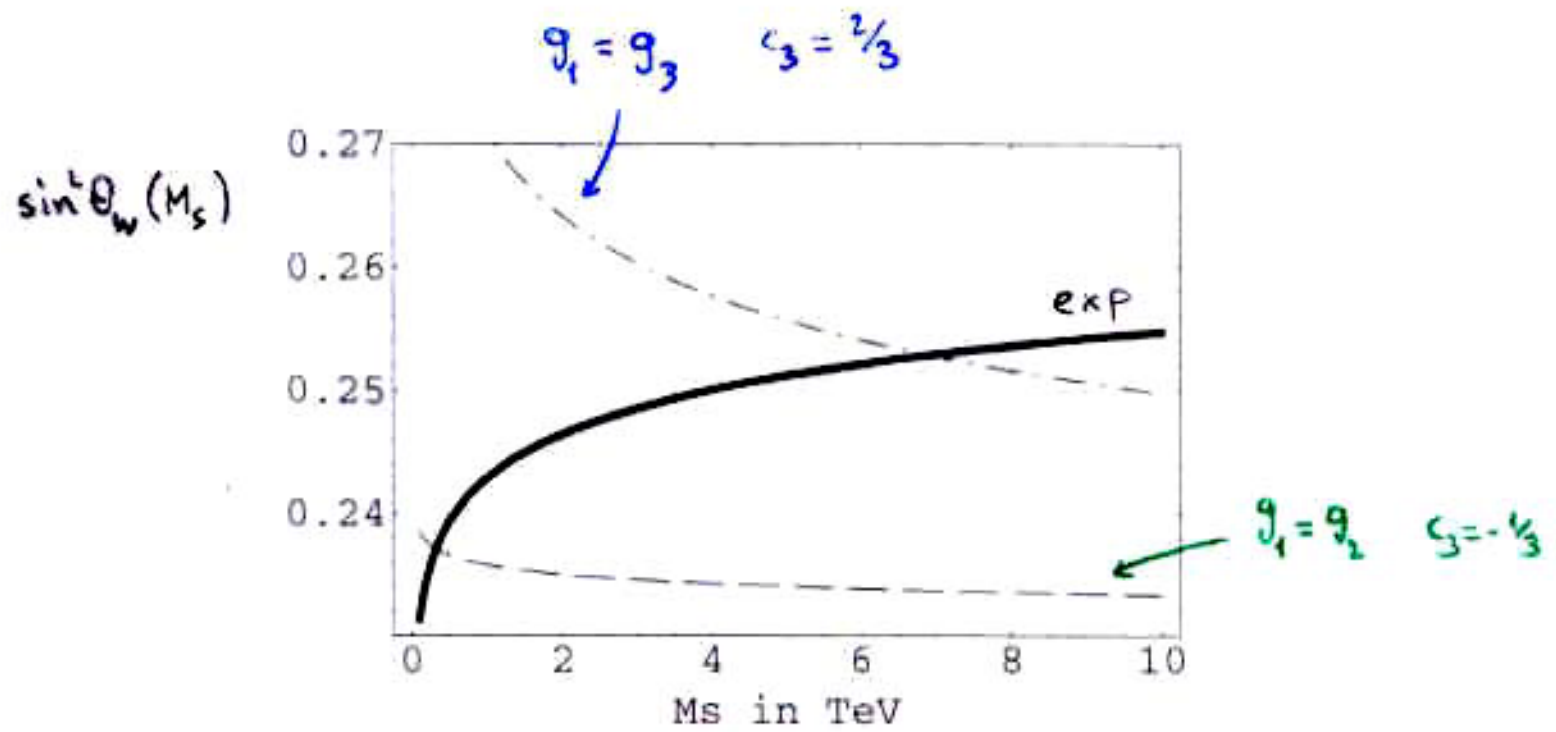
$$L \quad (1, 2; 0, 1, z)_{-1/2} \quad z = \pm 1 \text{ or } 0$$

$$e^c \quad (1, 1; 0, 0, 1)_1$$

hypercharge $Y = c_1 Q_1 + c_2 Q_2 + c_3 Q_3 \Rightarrow 4$ possibilities

$$c_3 = -1/3 \quad c_2 = \pm 1/2 \quad x = -1 \quad y = 0 \quad w = \pm 1 \quad z = -1/0$$

$$c_3 = 2/3 \quad c_2 = \pm 1/2 \quad x = 0 \quad y = 1 \quad w = \mp 1 \quad z = -1/0$$



correct prediction for $\sin^2 \theta_w$ for $M_s \sim \text{few TeV}$

$U(1)$ with color branes

- masses to all quarks + leptons $\Rightarrow$ 2 Higgs doublets
- the remaining two $U(1)$'s : anomalous

Green - Schwarz anomaly cancellation:

shifting of 2 axions $\Rightarrow$ $U(1)$'s become massive

$\Rightarrow$ global (perturbative) symmetries:

- baryon number $\Rightarrow$ proton stability
- PQ-type symmetry $\Rightarrow$ electroweak axion
 $\nearrow$
 can be explicitly broken by moving slightly
 away from the orbifold point
- R-neutrinos : open strings in the bulk

Arkani Hamed - Dimopoulos - Dvali - March Russell

Dienes - Dudas - Gherghetta '98

$$\text{Higgs : } c_2 = -\frac{1}{2} \quad H (1, 2; 0, 1, 1)_{\frac{1}{2}} \quad H' (1, 2; 0, -1, 0)_{\frac{1}{2}}$$

$$H' Q u^c \quad H^+ L e^c \quad H^+ Q d^c$$

$$c_2 = \frac{1}{2} \quad \tilde{H} (1, 2; 0, -1, 1)_{\frac{1}{2}} \quad \tilde{H}' (1, 2; 0, 1, 0)_{\frac{1}{2}}$$

$$\tilde{H}' Q u^c \quad \tilde{H}' L e^c \quad \tilde{H}^+ Q d^c$$

$$H' L \nu_R$$

$$\tilde{H} L \nu_R$$

R-neutrinos in the bulk

$$\int d^4x \bar{\nu} \not{\partial} \nu \quad \nu = (\nu_R, \bar{\nu}_R^c) \Rightarrow$$

$$\int d^4x (r M_s)^p \sum_n \left\{ \bar{\nu}_{Rn} \not{\partial} \nu_{Rn} + \bar{\nu}_{Rn}^c \not{\partial} \nu_{Rn}^c + \frac{n}{r} \nu_{Rn} \nu_{Rn}^c + \text{c.c.} \right\}$$

$$S_{\text{int}} = \lambda \int d^4x H(x) L(x) \nu_R(x, \vec{y}=0)$$

$$\langle H \rangle = v \Rightarrow \frac{\lambda v}{\sqrt{(r M_s)^p}} \sum_n \nu_L \nu_{Rn}$$

$$\frac{\lambda v}{(r M_s)^{p/2}} \ll \frac{1}{r} \Leftrightarrow \lambda \frac{v}{M_s} \ll (r M_s)^{\frac{p}{2}-1} \Rightarrow$$

• $n \neq 0$: masses of KK ν_n unaffected

• $n=0$: Dirac mass for neutrino $m_\nu = \frac{\lambda v}{(r M_s)^{p/2}}$

$$M_P = \frac{1}{g^2} M_s^{4+p/2} r^{1/2} \Rightarrow$$

$$m_\nu = \frac{\lambda}{g^2} v \frac{M_s}{M_P} \sim 10^{-2} \text{ eV} \quad \text{for } M_s \simeq 10 \text{ TeV}$$