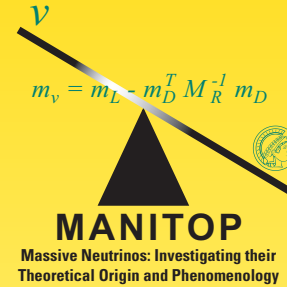


# Neutrinoless Double Beta Decay



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based on

- W.R., *Neutrinoless Double Beta Decay and Particle Physics*, Int. J. Mod. Phys. **E20** (2011) 1833
- W.R., *Neutrinoless Double Beta Decay and Neutrino Physics*, J. Phys. **G39** (2012) 124008
- H. Päs and W.R., *Neutrinoless Double Beta Decay*, New J. Phys. **17** (2015), 115010

## Neutrinoless double beta decay

$$(A, Z) \rightarrow (A, Z + 2) + 2 e^{-} \quad (0\nu\beta\beta)$$

- rare process, not yet observed
- violates lepton number
- best chance to confirm Majorana nature of light oscillating neutrinos
- constraints a lot of BSM models
- observation would have a lot of implications

## Upcoming/running experiments: exciting time!!

best limit was from 2001, improved 2012

Name	Isotope	Source = Detector; calorimetric with			Source $\neq$ Detector
		high $\Delta E$	low $\Delta E$	topology	topology
AMoRE	$^{100}\text{Mo}$	✓	-	-	-
CANDLES	$^{48}\text{Ca}$	-	✓	-	-
COBRA	$^{116}\text{Cd}$ (and $^{130}\text{Te}$ )	-	-	✓	-
CUORE	$^{130}\text{Te}$	✓	-	-	-
DCBA/MTD	$^{82}\text{Se} / ^{150}\text{Nd}$	-	-	-	✓
EXO	$^{136}\text{Xe}$	-	-	✓	-
GERDA	$^{76}\text{Ge}$	✓	-	-	-
CUPID	$^{82}\text{Se} / ^{100}\text{Mo} / ^{116}\text{Cd} / ^{130}\text{Te}$	✓	-	-	-
KamLAND-Zen	$^{136}\text{Xe}$	-	✓	-	-
LUCIFER	$^{82}\text{Se} / ^{100}\text{Mo} / ^{130}\text{Te}$	✓	-	-	-
LUMINEU	$^{100}\text{Mo}$	✓	-	-	-
MAJORANA	$^{76}\text{Ge}$	✓	-	-	-
MOON	$^{82}\text{Se} / ^{100}\text{Mo} / ^{150}\text{Nd}$	-	-	-	✓
NEXT	$^{136}\text{Xe}$	-	-	✓	-
SNO+	$^{130}\text{Te}$	-	✓	-	-
SuperNEMO	$^{82}\text{Se} / ^{150}\text{Nd}$	-	-	-	✓
XMASS	$^{136}\text{Xe}$	-	✓	-	-

## Why should we probe Lepton Number Violation?

- $L$  and  $B$  accidentally conserved in SM
- effective theory:  $\mathcal{L} = \mathcal{L}_{\text{SM}} + \frac{1}{\Lambda} \mathcal{L}_{\text{LNV}} + \frac{1}{\Lambda^2} \mathcal{L}_{\text{LFV, BNV, LNV}} + \dots$
- baryogenesis:  $B$  is violated
- $B, L$  often connected in GUTs
- GUTs have seesaw and Majorana neutrinos
- (chiral anomalies:  $\partial_\mu J_{B,L}^\mu = c G_{\mu\nu} \tilde{G}^{\mu\nu} \neq 0$  with  $J_\mu^B = \sum \bar{q}_i \gamma_\mu q_i$  and  $J_\mu^L = \sum \bar{\ell}_i \gamma_\mu \ell_i$ )

⇒ Lepton Number Violation as important as Baryon Number Violation

( $0\nu\beta\beta$  is much more than a neutrino mass experiment)

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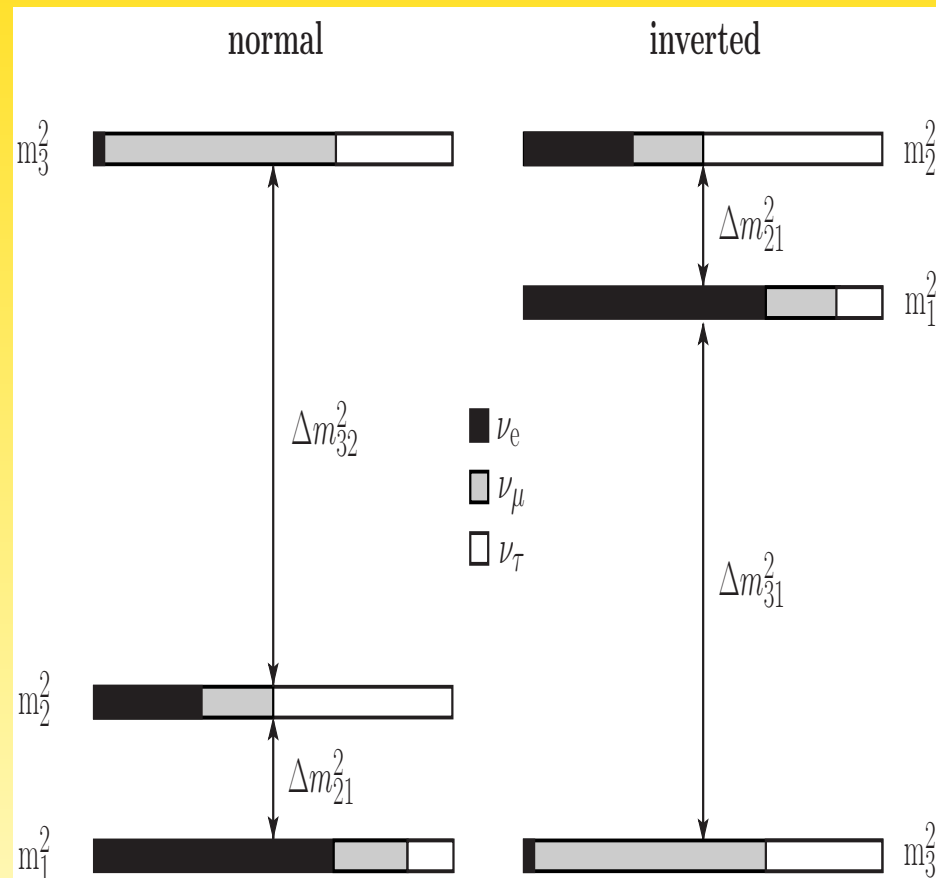
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# I) Dirac vs. Majorana neutrinos

Neutrinos oscillate  $\Rightarrow m_\nu \neq 0$



according to Standard Model, they shouldn't

Limits (later):

- direct (Kurie plot):  $m_\beta = \sqrt{\sum |U_{ei}|^2 m_i^2} \leq 2.3 \text{ eV}$
- neutrinoless double beta decay:  $|m_{ee}| = |\sum U_{ei}^2 m_i| \lesssim 0.3 \text{ eV}$
- cosmology:  $\sum m_i \lesssim 0.7 \text{ eV}$

$\Rightarrow$  “ $m_\nu$ ” less than 1 eV

How can we introduce neutrino mass terms?

## I1) Basics

### a) Dirac masses

with  $L \sim (2, -1)$  and  $\Phi \sim (2, 1)$ , add  $\nu_R \sim (1, 0)$ :

$$\mathcal{L}_D = g_\nu \bar{L} \tilde{\Phi} \nu_R \xrightarrow{\text{SSB}} \frac{v}{\sqrt{2}} g_\nu \bar{\nu}_L \nu_R = m_\nu \bar{\nu}_L \nu_R$$

But  $m_\nu \lesssim \text{eV}$  implies  $g_\nu \lesssim 10^{-12} \lll g_e$

highly unsatisfactory fine-tuning...

actually,  $m_e = 10^{-6} m_t$ , so WTF?

point is that

$$\begin{pmatrix} u \\ d \end{pmatrix} \text{ with } m_u \simeq m_d$$

has to be contrasted with

$$\begin{pmatrix} \nu_e \\ e \end{pmatrix} \text{ with } m_\nu \simeq 10^{-6} m_e$$

what suppresses neutrino mass **for every generation?**

## b) Majorana masses

need charge conjugation:

$$\text{electron } e^-: [\gamma_\mu (i\partial^\mu + e A^\mu) - m] \psi = 0 \quad (1)$$

$$\text{positron } e^+: [\gamma_\mu (i\partial^\mu - e A^\mu) - m] \psi^c = 0 \quad (2)$$

Try  $\psi^c = S \psi^*$ , evaluate  $(S^*)^{-1} (2)^*$  and compare with (1):

$$\Rightarrow S = i\gamma_2$$

and thus

$$\psi^c = i\gamma_2 \psi^* = i\gamma_2 \gamma_0 \bar{\psi}^T \equiv C \bar{\psi}^T$$

flips all charge-like quantum numbers

Properties of  $C$ :

$$C^\dagger = C^T = C^{-1} = -C$$

$$C \gamma_\mu C^{-1} = -\gamma_\mu^T$$

$$C \gamma_5 C^{-1} = \gamma_5^T$$

$$C \gamma_\mu \gamma_5 C^{-1} = (\gamma_\mu \gamma_5)^T$$

properties of charged conjugate spinors:

$$(\psi^c)^c = \psi$$

$$\overline{\psi^c} = \psi^T C$$

$$\overline{\psi_1} \psi_2^c = \overline{\psi_2^c} \psi_1$$

$$(\psi_L)^c = (\psi^c)_R$$

$$(\psi_R)^c = (\psi^c)_L$$

$C$  flips chirality: LH becomes RH

common source of confusion (added here to create or add confusion):

that was actually particle-antiparticle conjugation ( $\nu_L \rightarrow \bar{\nu}_R$ )

charge conjugation leaves chirality untouched ( $\nu_L \rightarrow \bar{\nu}_L$ )

same effect for total state  $\nu = \nu_L + \nu_R$ , different for chiral states

## Fermion mass terms

$$\mathcal{L} = m_\nu \bar{\psi} \psi = m_\nu \overline{\psi_L} \psi_R + h.c. = m_\nu (\overline{\psi_L} \psi_R + \overline{\psi_R} \psi_L)$$

both chiralities a must for mass term!

there are two and only two possibilities:

(i)  $\psi_L$  independent of  $\psi_R$ : **Dirac particle**

(ii)  $\psi_L = (\psi_R)^c$ : **Majorana particle**

$$(ii) \Rightarrow \psi^c = (\psi_L + \psi_R)^c = (\psi_L)^c + (\psi_R)^c = \psi_R + \psi_L = \psi : \psi^c = \psi$$

$\Rightarrow$  Majorana fermion is identical to its antiparticle, a truly neutral particle

$\Rightarrow$  does not carry conserved additive quantum numbers

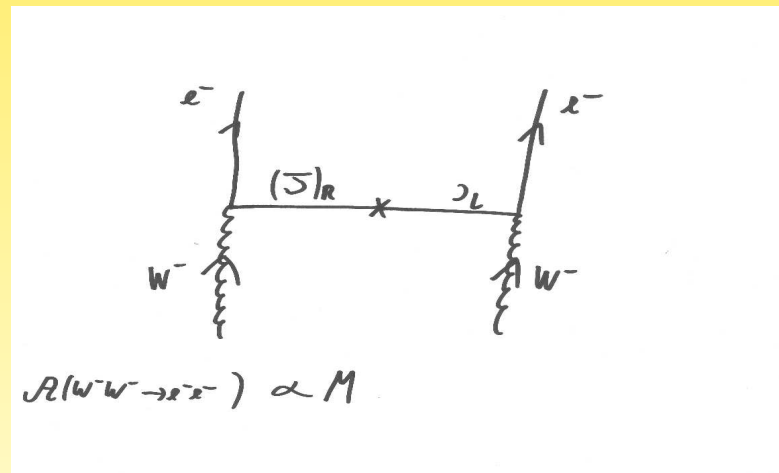


mass term for Majorana particles:

$$\mathcal{L}_M = \frac{1}{2} \bar{\psi} M \psi = \frac{1}{2} \overline{\psi_L + (\psi_L)^c} M (\psi_L + (\psi_L)^c) = \frac{1}{2} \bar{\psi}_L M \psi_L^c + h.c.$$

- Majorana mass term (not allowed by gauge symmetry, never mind for now)
- $\mathcal{L}_M \propto \psi^* \psi^\dagger \Rightarrow$  NOT invariant under  $\psi \rightarrow e^{i\alpha} \psi$   
 $\Rightarrow$  breaks Lepton Number by 2 units!

in standard approach to  $0\nu\beta\beta$ : origin of decay



## Dirac vs. Majorana

in  $V - A$  theories: observable difference always suppressed by  $(m/E)^2$

- In terms of degrees of freedom (helicity and particle/antiparticle):  
 $\nu_D = (\nu_\uparrow, \nu_\downarrow, \bar{\nu}_\uparrow, \bar{\nu}_\downarrow)$  versus  $\nu_M = (\nu_\uparrow, \nu_\downarrow)$
- weak interactions act on chirality (left-/right-handed)
- chirality is not a good quantum number ("spin flip"):  $L = \downarrow + \frac{m}{E} \uparrow$
- Dirac:
  - what we produce from a  $W^-$  is  $\ell^- (\bar{\nu}_\uparrow + \frac{m}{E} \bar{\nu}_\downarrow)$
  - the  $\bar{\nu}_\downarrow$  CANNOT interact with another  $W^-$  to generate another  $\ell^-$ :  
 $\Delta L = 0$
- Majorana:
  - what we produce from a  $W^-$  is  $\ell^- (\nu_\uparrow + \frac{m}{E} \nu_\downarrow)$
  - the  $\nu_\downarrow$  CAN interact with another  $W^-$  to generate another  $\ell^-$ :  $\Delta L = 2$   
 $\Rightarrow$  amplitude  $\propto (m/E) \Rightarrow$  probability  $\propto (m/E)^2$

## Dirac vs. Majorana

- Z-decay:

$$\frac{\Gamma(Z \rightarrow \nu_D \nu_D)}{\Gamma(Z \rightarrow \nu_M \nu_M)} \simeq 1 - 3 \frac{m_\nu^2}{m_Z^2}$$

- Meson decays

$$\text{BR}(K^+ \rightarrow \pi^- e^+ \mu^+) \propto |m_{e\mu}|^2 = \left| \sum U_{ei} U_{\mu i} m_i \right|^2 \sim 10^{-30} \left( \frac{|m_{e\mu}|}{\text{eV}} \right)^2$$

- neutrino-antineutrino oscillations

$$P(\nu_\alpha \rightarrow \bar{\nu}_\beta) = \frac{1}{E^2} \left| \sum_{i,j} U_{\alpha j} U_{\beta j} U_{\alpha i}^* U_{\beta i}^* m_i m_j e^{-i(E_j - E_i)t} \right|$$

- Majorana only works for neutral fermions  $\Rightarrow$  neutrinos only candidate in SM
- $\psi^c = \psi$  makes spinor have two degrees of freedom, not four:

$$\psi_D = \begin{pmatrix} \Phi \\ \xi \end{pmatrix} \longrightarrow \psi_M = \begin{pmatrix} \Phi \\ -i\sigma_2\Phi^* \end{pmatrix}$$

( $\Phi, \xi$  are 2-component Weyl spinors)

- in terms of creators/annihilators,  $\Psi = \Psi^c$  happens for

$$\Psi = \int \frac{d^3p}{(2\pi)^3 \sqrt{2E_p}} \sum_s [b_s(\vec{p})u_s(\vec{p})e^{-ipx} + b_s^\dagger(\vec{p})v_s(\vec{p})e^{+ipx}]$$

where  $v_s = C\bar{u}_s^T$  and  $u_s = C\bar{v}_s^T$

- mass term  $\bar{\nu}_L \nu_L^c$  not allowed as  $\nu_L$  is in  $(2, -1)$
- if  $N_R \sim (1, 0)$  exists:  $\bar{N}_R N_R^c$  allowed!

Majorana nature implies new relation for spinors:

$$\begin{aligned} \sum_s u_s \bar{u}_s &= \not{p} + m & , & & \sum_s v_s \bar{v}_s &= \not{p} - m \\ \sum_s u_s v_s^T &= (\not{p} + m) C^T & , & & \sum_s \bar{u}_s^T \bar{v}_s &= C^{-1} (\not{p} - m) \\ \sum_s \bar{v}_s \bar{u}_s &= C^{-1} (\not{p} + m) & , & & \sum_s v_s u_s^T &= (\not{p} - m) C^T \end{aligned}$$

and new propagators

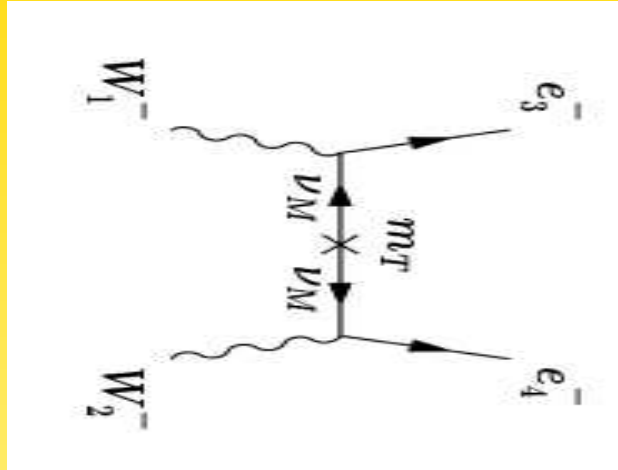
$$\begin{aligned} (1) \quad \langle 0 | T \{ \Psi(x) \bar{\Psi}(y) \} | 0 \rangle &= S(x - y) \\ (2) \quad \langle 0 | T \{ \Psi(x) \Psi(y) \} | 0 \rangle &= -S(x - y) C \\ (3) \quad \langle 0 | T \{ \bar{\Psi}(x) \bar{\Psi}(y) \} | 0 \rangle &= C^{-1} S(x - y) \end{aligned}$$

with standard propagator

$$S(x - y) = i \int \frac{d^4 p}{(2\pi)^4} \left( \frac{\not{p} + m}{p^2 - m^2 + i\epsilon} \right) e^{-ip(x-y)}$$

$\Rightarrow$  can create or annihilate two neutrinos ((2) and (3))

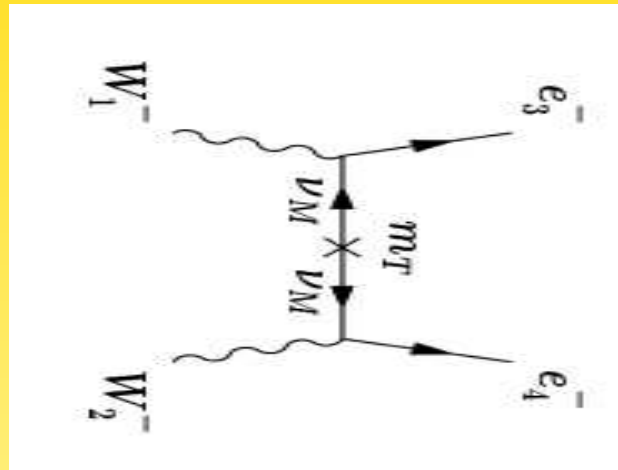
consider  $t$ -channel  $W^-W^- \rightarrow e^-e^-$  from  $\mathcal{L} = -\frac{g}{2\sqrt{2}}W_\mu\bar{e}\gamma^\mu(1-\gamma_5)\nu_M$



$$\begin{aligned}
 -i\mathcal{M} &= \left(\frac{-ig}{2\sqrt{2}}\right)^2 \epsilon_1^\mu \epsilon_2^\nu \left[ \bar{u}_3 \gamma_\mu (1-\gamma_5) i \frac{\not{k} + m}{k^2 - m^2} (-C) \bar{u}_4 \gamma_\nu (1-\gamma_5) \right] \\
 &\propto (\bar{u}_3 \gamma_\mu (1-\gamma_5)) (\not{k} + m) (-C) (\bar{u}_4 \gamma_\nu (1-\gamma_5)) \\
 &= (\bar{u}_3 \gamma_\mu (1-\gamma_5)) (\not{k} + m) (-(1-\gamma_5) \gamma_\nu v_4) \\
 &= m \bar{u}_3 \gamma_\mu (1-\gamma_5) \gamma_\nu v_4
 \end{aligned}$$

proportional to (suppressed by) neutrino mass!!

consider  $t$ -channel  $W^-W^- \rightarrow e^-e^-$  from  $\mathcal{L} = -\frac{g}{2\sqrt{2}}W_\mu\bar{e}\gamma^\mu(1-\gamma_5)\nu_M$

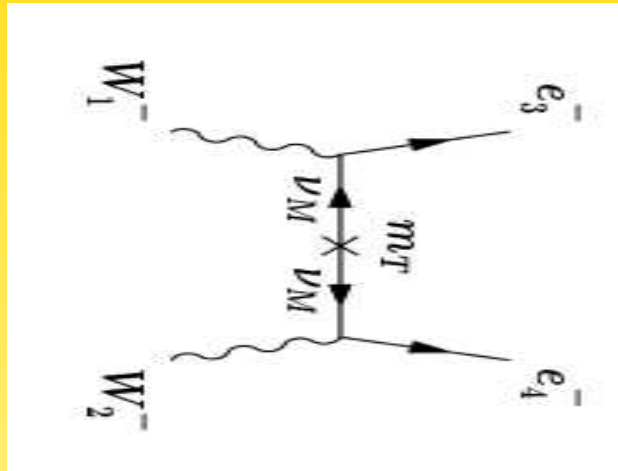


$$-i\mathcal{M} \propto m \bar{u}_3 \gamma_\mu (1 - \gamma_5) \gamma_\nu v_4$$

proportional to (suppressed by) neutrino mass!!

same result by replacing one  $\bar{e}\gamma^\mu(1-\gamma_5)\nu_M$  with  $-\bar{\nu}_M\gamma^\mu(1+\gamma_5)e^c$

consider  $t$ -channel  $W^-W^- \rightarrow e^-e^-$  from  $\mathcal{L} = -\frac{g}{2\sqrt{2}}W_\mu\bar{e}\gamma^\mu(1-\gamma_5)\nu_M$



$$\begin{aligned}\mathcal{M} &\propto \bar{e}_3\gamma_\mu(1-\gamma_5)\nu\bar{e}_4\gamma_\nu(1-\gamma_5)\nu \\ &= \bar{e}_3\gamma_\mu(1-\gamma_5)\nu\bar{\nu}^c\gamma_\nu(1+\gamma_5)e_4^c(-1)\end{aligned}$$

time-ordered product  $\nu\bar{\nu}$  is propagator  $\propto (\not{k} + m)$

use  $e = fu + \bar{f}^\dagger v$ ,  $\bar{e} = f^\dagger \bar{u} + \bar{f} \bar{v}$ ,  $e^c = f^\dagger v + \bar{f} u \Rightarrow e_4^c$  gets a  $v_4$



how to calculate identities:

$\bar{e} \gamma_\mu (1 - \gamma_5) \nu$		start
$= \overline{(e^c)^c} \gamma_\mu (1 - \gamma_5) (\nu^c)^c$		$\psi = (\psi^c)^c$
$= -(e^c)^T C^{-1} \gamma_\mu (1 - \gamma_5) C \overline{\nu^c}^T$		$\overline{\psi^c} = -\psi^T C^{-1}$ and $\psi^c = C \overline{\psi}^T$
$= -(e^c)^T (-\gamma_\mu^T - (\gamma_\mu \gamma_5)^T) \overline{\nu^c}^T$		$C \gamma_\mu C^{-1} = -\gamma_\mu^T$ , $C \gamma_\mu \gamma_5 C^{-1} = (\gamma_\mu \gamma_5)^T$
$= (e^c)^T [\gamma_\mu (1 + \gamma_5)]^T \overline{\nu^c}^T$		collecting signs
$= -\overline{\nu^c} \gamma_\mu (1 + \gamma_5) e^c$		is scalar and fermion exchange

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## I2) Seesaw Mechanisms

how to generate Majorana neutrino masses...

### a) Higher dimensional operators

include  $d \geq 5$  operators, gauge and Lorentz invariant, only SM fields:

$$\mathcal{L} = \mathcal{L}_{\text{SM}} + \mathcal{L}_5 + \mathcal{L}_6 + \dots = \mathcal{L}_{\text{SM}} + \frac{1}{\Lambda} \mathcal{O}_5 + \frac{1}{\Lambda^2} \mathcal{O}_6 + \dots$$

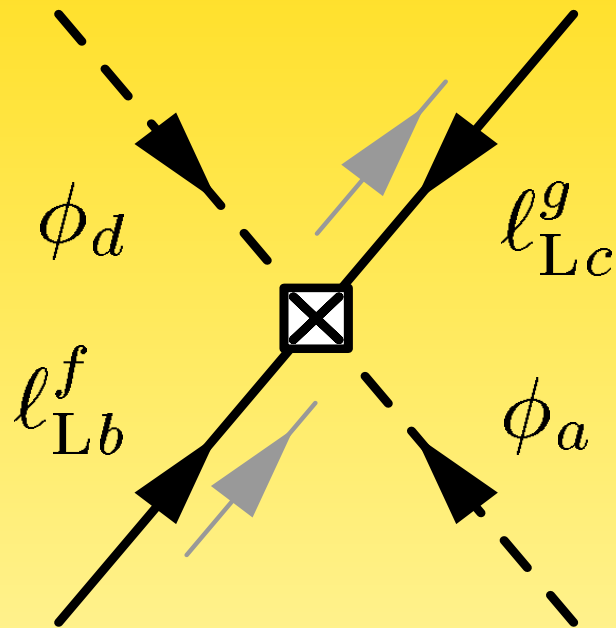
there is only one dimension 5 term! “leading order new physics”

$$\frac{1}{\Lambda} \mathcal{O}_5 = \frac{c}{\Lambda} \bar{L} \tilde{\Phi} \tilde{\Phi}^T L^c \xrightarrow{\text{SSB}} \frac{c v^2}{2\Lambda} \bar{\nu}_L \nu_L^c \equiv m_\nu \bar{\nu}_L \nu_L^c$$

Majorana mass term!  $\Delta L = 2$

it follows  $\Lambda \gtrsim c \left( \frac{0.1 \text{ eV}}{m_\nu} \right) 10^{14} \text{ GeV}$  Weinberg 1979

Weinberg operator is  $LL\Phi\Phi$



seesaw mechanisms are “UV-completions” of this effective operator by integrating out heavy physics

Master formula:  $2 \times 2 = 3 + 1$

$SU(2)_L \times U(1)_Y$  with  $2 \otimes 2 = 3 \oplus 1$ :

$$\bar{L} \tilde{\Phi} \sim (2, +1) \otimes (2, -1) = (3, 0) \oplus (1, 0)$$

to make a singlet, couple to  $(1, 0)$  or  $(3, 0)$ , because  $3 \otimes 3 = 5 \oplus 3 \oplus 1$

Alternatively:

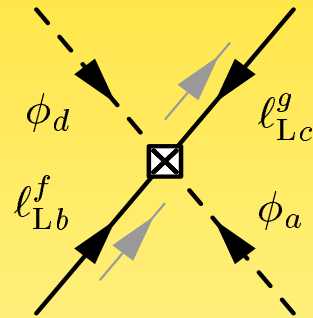
$$\bar{L}^c L \sim (2, -1) \otimes (2, -1) = (3, -2) \oplus (1, -2)$$

to make a singlet, couple to  $(1, +2)$  or  $(3, +2)$ . However, singlet combination  $(1, -2)$  is  $\bar{\nu} \ell^c - \bar{\ell} \nu^c$ , which cannot generate neutrino mass term

$$\begin{array}{ccccc} \implies & (1, 0) & \text{or} & (3, +2) & \text{or} & (3, 0) \\ & \text{type I} & & \text{type II} & & \text{type III} \end{array}$$

## Origin of small masses

$$\mathcal{L} = \frac{c}{\Lambda} \overline{L^c} \tilde{\Phi}^* \tilde{\Phi}^\dagger L$$



has only 3 tree-level realizations

- $N_R \sim (1, 0)$  type I seesaw
- $\Delta \sim (3, 2)$  type II seesaw
- $\Sigma \sim (3, 0)$  type III seesaw

seesaws include **new** representations, new energy scales, new concepts

## Type I Seesaw

introduce  $N_R \sim (1, 0)$  and couple to  $g_\nu \bar{L} \tilde{\Phi} \sim (1, 0)$

becomes  $g_\nu v / \sqrt{2} \bar{\nu}_L N_R \equiv m_D \bar{\nu}_L N_R$

in addition: Majorana mass term for  $N_R$ :  $\frac{1}{2} M_R \bar{N}_R^c N_R$

using  $\bar{\nu}_L m_D N_R = \bar{N}_R^c m_D^T \nu_L^c$ :

$$\begin{aligned} \mathcal{L} &= \frac{1}{2} (\bar{\nu}_L, \bar{N}_R^c) \begin{pmatrix} 0 & m_D \\ m_D & M_R \end{pmatrix} \begin{pmatrix} \nu_L^c \\ N_R \end{pmatrix} + h.c. \\ &\equiv \frac{1}{2} \bar{\Psi} \mathcal{M}_\nu \Psi^c + h.c. \end{aligned}$$

Dirac + Majorana mass term is a Majorana mass term!

Diagonalization:

$$\begin{aligned}
 \mathcal{L} &= \frac{1}{2} (\overline{\nu}_L, \overline{N}_R^c) \begin{pmatrix} 0 & m_D \\ m_D & M_R \end{pmatrix} \begin{pmatrix} \nu_L^c \\ N_R \end{pmatrix} \\
 &= \frac{1}{2} \underbrace{(\overline{\nu}_L, \overline{N}_R^c) \mathcal{U}}_{(\overline{\nu}, \overline{N}^c)} \underbrace{\mathcal{U}^\dagger \begin{pmatrix} 0 & m_D \\ m_D & M_R \end{pmatrix} \mathcal{U}^*}_{\text{diag}(m_\nu, M)} \underbrace{\mathcal{U}^T}_{(\nu^c, N)^T} \begin{pmatrix} \nu_L^c \\ N_R \end{pmatrix}
 \end{aligned}$$

with

general formula:

$$\tan 2\theta = \frac{2m_D}{M_R - 0}$$

$$\begin{aligned}
 m_\nu &= \frac{1}{2} \left[ (0 + M_R) - \sqrt{(0 - M_R)^2 + 4m_D^2} \right] \\
 M &= \frac{1}{2} \left[ (0 + M_R) + \sqrt{(0 - M_R)^2 + 4m_D^2} \right]
 \end{aligned}$$

if  $m_D \ll M_R$ :

$$\ll 1$$

$$\simeq -m_D^2/M_R$$

$$\simeq M_R$$



Note:  $m_D$  associated with EWSB, part of SM, bounded by  $v/\sqrt{2} = 174$  GeV

$M_R$  is SM singlet, does whatever it wants:  $\Rightarrow M_R \gg m_D$

Hence,  $\theta \simeq m_D/M_R \ll 1$

$$\begin{aligned}\nu &= \nu_L \cos \theta - N_R^c \sin \theta \simeq \nu_L && \text{with mass } m_\nu \simeq -m_D^2/M_R \\ N &= N_R \cos \theta + \nu_L^c \sin \theta \simeq N_R && \text{with mass } M \simeq M_R\end{aligned}$$

in effective mass terms

$$\mathcal{L} \simeq \frac{1}{2} m_\nu \bar{\nu}_L \nu_L^c + \frac{1}{2} M_R \bar{N}_R^c N_R$$

compare with Weinberg operator:

$$\Lambda = -\frac{c v^2}{m_D^2} M_R$$

also: integrate  $N_R$  away with Euler-Lagrange equation

matrix case: block diagonalization

$$\begin{aligned}
 \mathcal{L} &= \frac{1}{2} (\overline{\nu}_L, \overline{N}_R^c) \begin{pmatrix} 0 & m_D \\ m_D^T & M_R \end{pmatrix} \begin{pmatrix} \nu_L^c \\ N_R \end{pmatrix} \\
 &= \frac{1}{2} \underbrace{(\overline{\nu}_L, \overline{N}_R^c) \mathcal{U}}_{(\overline{\nu}, \overline{N}^c)} \underbrace{\mathcal{U}^\dagger \begin{pmatrix} 0 & m_D \\ m_D^T & M_R \end{pmatrix} \mathcal{U}^*}_{\text{diag}(m_\nu, M)} \underbrace{\mathcal{U}^T \begin{pmatrix} \nu_L^c \\ N_R \end{pmatrix}}_{(\nu^c, N)^T}
 \end{aligned}$$

with  $6 \times 6$  diagonal matrix

$$\mathcal{U} = \begin{pmatrix} 1 & -\rho \\ \rho^\dagger & 1 \end{pmatrix}, \quad \mathcal{U}^\dagger = \begin{pmatrix} 1 & \rho \\ -\rho^\dagger & 1 \end{pmatrix}, \quad \mathcal{U}^* = \begin{pmatrix} 1 & -\rho^* \\ \rho^T & 1 \end{pmatrix}$$

write down individual components:

write down individual components:

$$m_\nu = \rho m_D^T + m_D \rho + \rho M_R \rho^T$$

$$0 = m_D - \rho m_D^T \rho^* + \rho M_R$$

$$M = -\rho^\dagger m_D - m_D^T \rho^* + M_R$$

now,  $\rho$  (aka  $\theta$  from before) will be of order  $m_D/M_R$ :

$$m_\nu = \rho m_D^T + m_D \rho + \rho M_R \rho^T$$

$$0 \simeq m_D + \rho M_R \Rightarrow \rho = -m_D M_R^{-1}$$

$$M \simeq M_R$$

insert  $\rho$  in  $m_\nu$  to find:

$$m_\nu = -m_D M_R^{-1} m_D^T$$

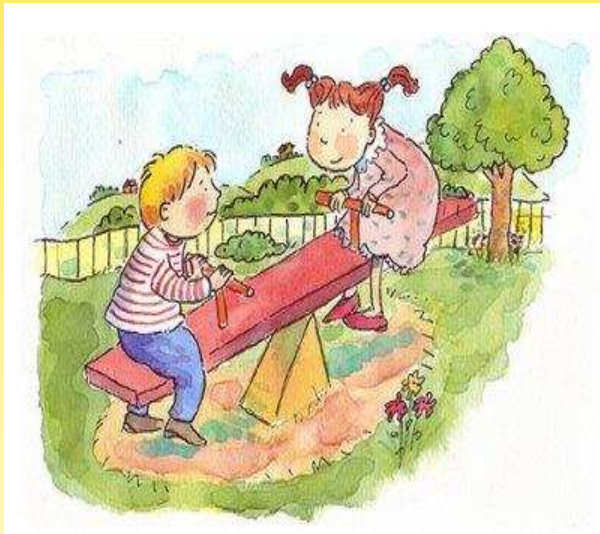
$$m_\nu = \frac{m_D^2}{M_R} = \frac{m_{SM}^2}{M_R} = m_{SM} \epsilon$$

## (type I) Seesaw Mechanism

Minkowski; Yanagida; Glashow; Gell-Mann, Ramond, Slansky; Mohapatra,  
Senjanović (77-80)

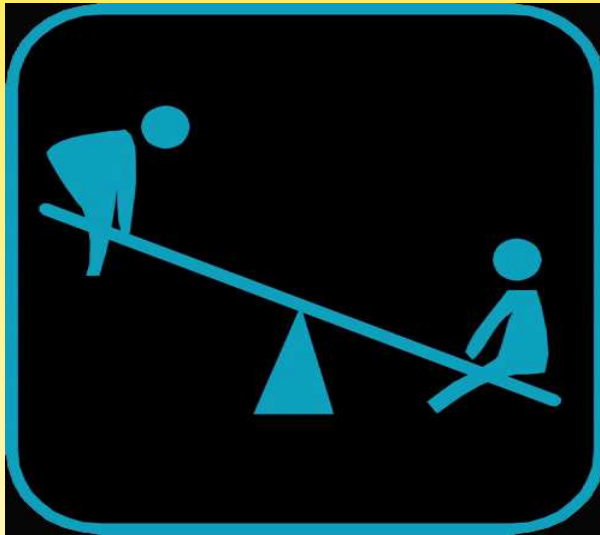
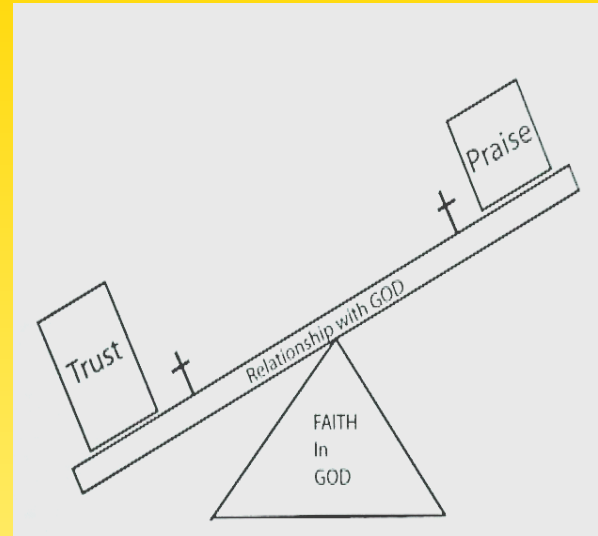
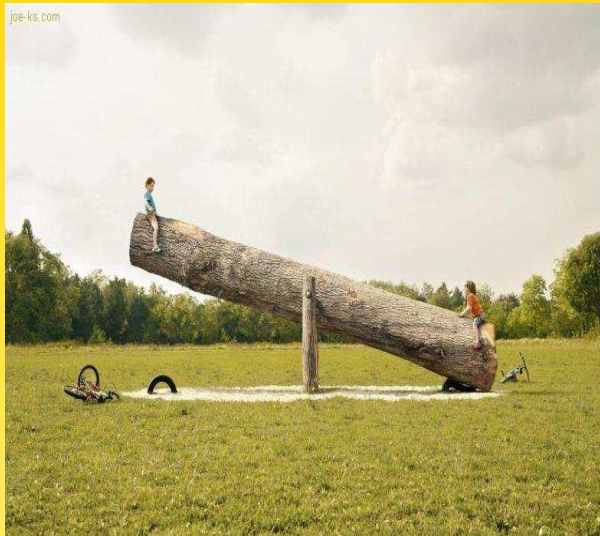


we found the requested suppression mechanism!













## Seesaw Formalism

$$\mathcal{L} = \frac{1}{2} (\bar{\nu}_L, \bar{N}_R^c) \begin{pmatrix} 0 & m_D \\ m_D^T & M_R \end{pmatrix} \begin{pmatrix} \nu_L^c \\ N_R \end{pmatrix} \xrightarrow{M_R \gg m_D} m_\nu = m_D^T M_R^{-1} m_D$$

arbitrary  $6 \times 6$  (?) matrix

with new aspects:

- fermionic singlets  $N_R \sim (1, 0)$
- new energy scale  $M_R (\propto 1/m_\nu)$
- lepton number violation

## See-Saw Phenomenology

Full mass matrix:

$$\mathcal{M} = \begin{pmatrix} 0 & m_D \\ m_D^T & M_R \end{pmatrix} = \mathcal{U} \begin{pmatrix} m_\nu^{\text{diag}} & 0 \\ 0 & M_R^{\text{diag}} \end{pmatrix} \mathcal{U}^T \text{ with } \mathcal{U} = \begin{pmatrix} N & S \\ T & V \end{pmatrix}$$

- $N$  is the PMNS matrix: non-unitary
- $S = m_D^\dagger (M_R^*)^{-1}$  describes mixing of heavy neutrinos with SM leptons

## Type II Seesaw

$$\mathcal{L} \propto \overline{L^c} L \rightarrow \nu^T \nu$$

has isospin  $I_3 = 1$  and transforms as  $\sim (3, -2)$

$\Rightarrow$  introduce Higgs triplet  $\sim (3, +2)$  with  $(I_3 = Q - Y/2)$ :

$$\Delta = \begin{pmatrix} \Delta^+ & -\sqrt{2} \Delta^{++} \\ \sqrt{2} \Delta^0 & -\Delta^+ \end{pmatrix} \rightarrow \begin{pmatrix} 0 & 0 \\ v_T & 0 \end{pmatrix}$$

gives mass matrix:

$$\mathcal{L} = \frac{1}{2} y_\nu \overline{L^c} i\sigma_2 \Delta L \xrightarrow{\text{vev}} \frac{1}{2} y_\nu v_T \overline{\nu_L^c} \nu_L \equiv \frac{1}{2} m_\nu \overline{\nu_L^c} \nu_L$$

$\Leftrightarrow$  neutrino mass without right-handed neutrinos!

$v_T \ll v$  because

$$V = -M_\Delta^2 \text{Tr} (\Delta^\dagger \Delta) + \mu \Phi^\dagger \Delta \tilde{\Phi}$$

with  $\frac{\partial V}{\partial \Delta} = 0$  one has  $v_T = \frac{\mu v^2}{M_\Delta^2}$

$v_T$  can be suppressed by  $M_\Delta$  and/or  $\mu$

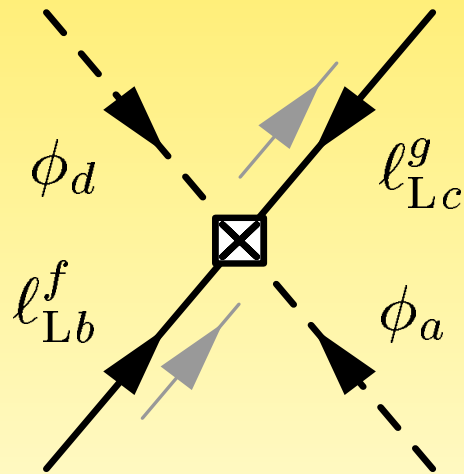
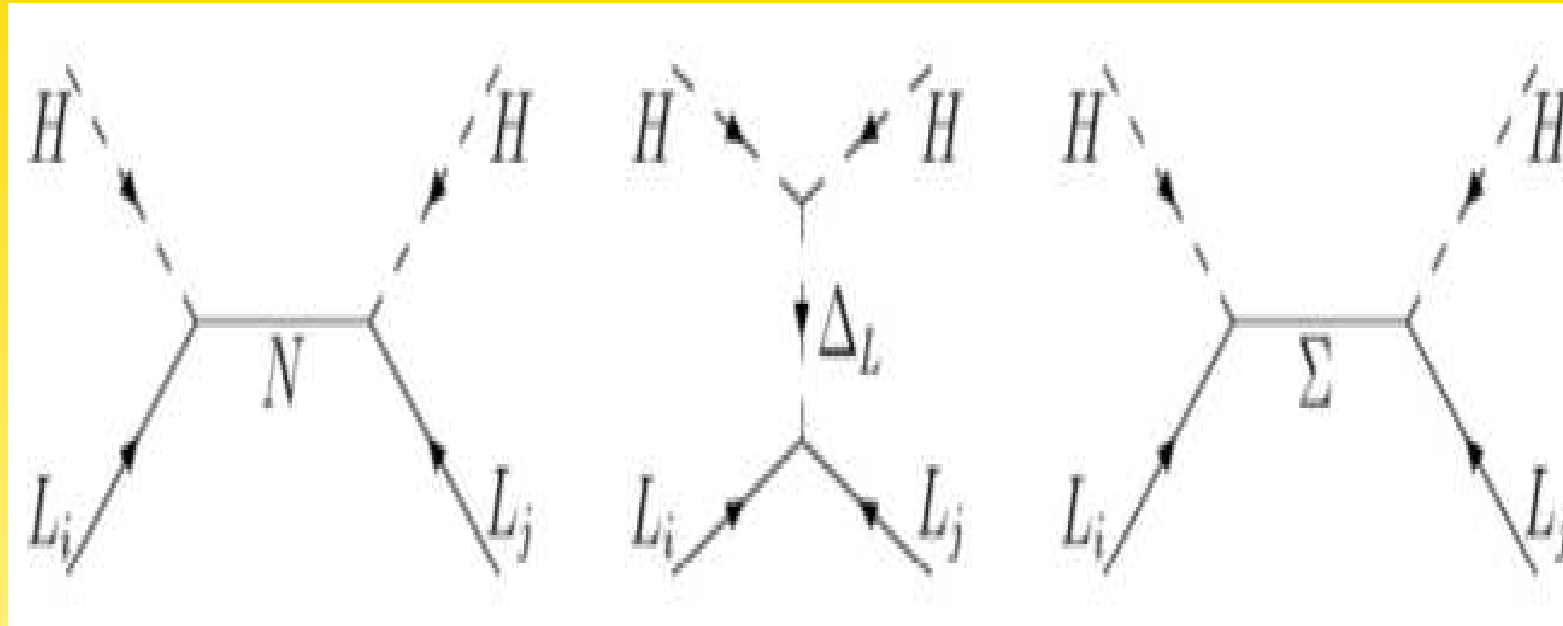
compare with Weinberg operator:

$$\Lambda = \frac{c M_\Delta^2}{g_\nu \mu}$$

## Type II (or Triplet) Seesaw Mechanism

Magg, Wetterich; Mohapatra, Senjanovic; Lazarides, Shafi, Wetterich;  
Schechter, Valle (80-82)

# Seesaw Summary



## Paths to Neutrino Mass

approach	ingredient	quantum number of messenger	$\mathcal{L}$	$m_\nu$	scale
“SM” (Dirac mass)	RH $\nu$	$N_R \sim (1, 0)$	$h \overline{N}_R \Phi L$	$h v$	$h = \mathcal{O}(10^{-12})$
“effective” (dim 5 operator)	new scale + LNV	-	$h \overline{L}^c \Phi \Phi L$	$\frac{h v^2}{\Lambda}$	$\Lambda = 10^{14}$ GeV
“direct” (type II seesaw)	Higgs triplet + LNV	$\Delta \sim (3, 2)$	$h \overline{L}^c \Delta L + \mu \Phi \Phi \Delta$	$h v_T$	$\Lambda = \frac{1}{h \mu} M_\Delta^2$
“indirect 1” (type I seesaw)	RH $\nu$ + LNV	$N_R \sim (1, 0)$	$h \overline{N}_R \Phi L + \overline{N}_R M_R N_R^c$	$\frac{(h v)^2}{M_R}$	$\Lambda = \frac{1}{h} M_R$
“indirect 2” (type III seesaw)	fermion triplets + LNV	$\Sigma \sim (3, 0)$	$h \overline{\Sigma} L \Phi + \text{Tr} \overline{\Sigma} M_\Sigma \Sigma$	$\frac{(h v)^2}{M_\Sigma}$	$\Lambda = \frac{1}{h} M_\Sigma$

plus seesaw variants (linear, double, inverse, . . .)

plus radiative mechanisms

plus extra dimensions

plusplusplus

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II3) Nuclear physics aspects

II4) Exotic physics

### I3) Summary neutrino physics

common prediction of all mechanisms: in basis with diagonal charged leptons

$$\mathcal{L} = \frac{1}{2} \nu^T m_\nu \nu \quad \text{with} \quad m_\nu = U \text{diag}(m_1, m_2, m_3) U^T$$

with PMNS matrix

$$U = \begin{pmatrix} c_{12} c_{13} & s_{12} c_{13} & s_{13} e^{-i\delta} \\ -s_{12} c_{23} - c_{12} s_{23} s_{13} e^{i\delta} & c_{12} c_{23} - s_{12} s_{23} s_{13} e^{i\delta} & s_{23} c_{13} \\ s_{12} s_{23} - c_{12} c_{23} s_{13} e^{i\delta} & -c_{12} s_{23} - s_{12} c_{23} s_{13} e^{i\delta} & c_{23} c_{13} \end{pmatrix} P$$

with  $P = \text{diag}(e^{i\alpha}, e^{i\beta}, 1)$  ( $\leftrightarrow$  Majorana, lepton number violation)

$\Rightarrow$  3 angles, 3 phases, 3 masses

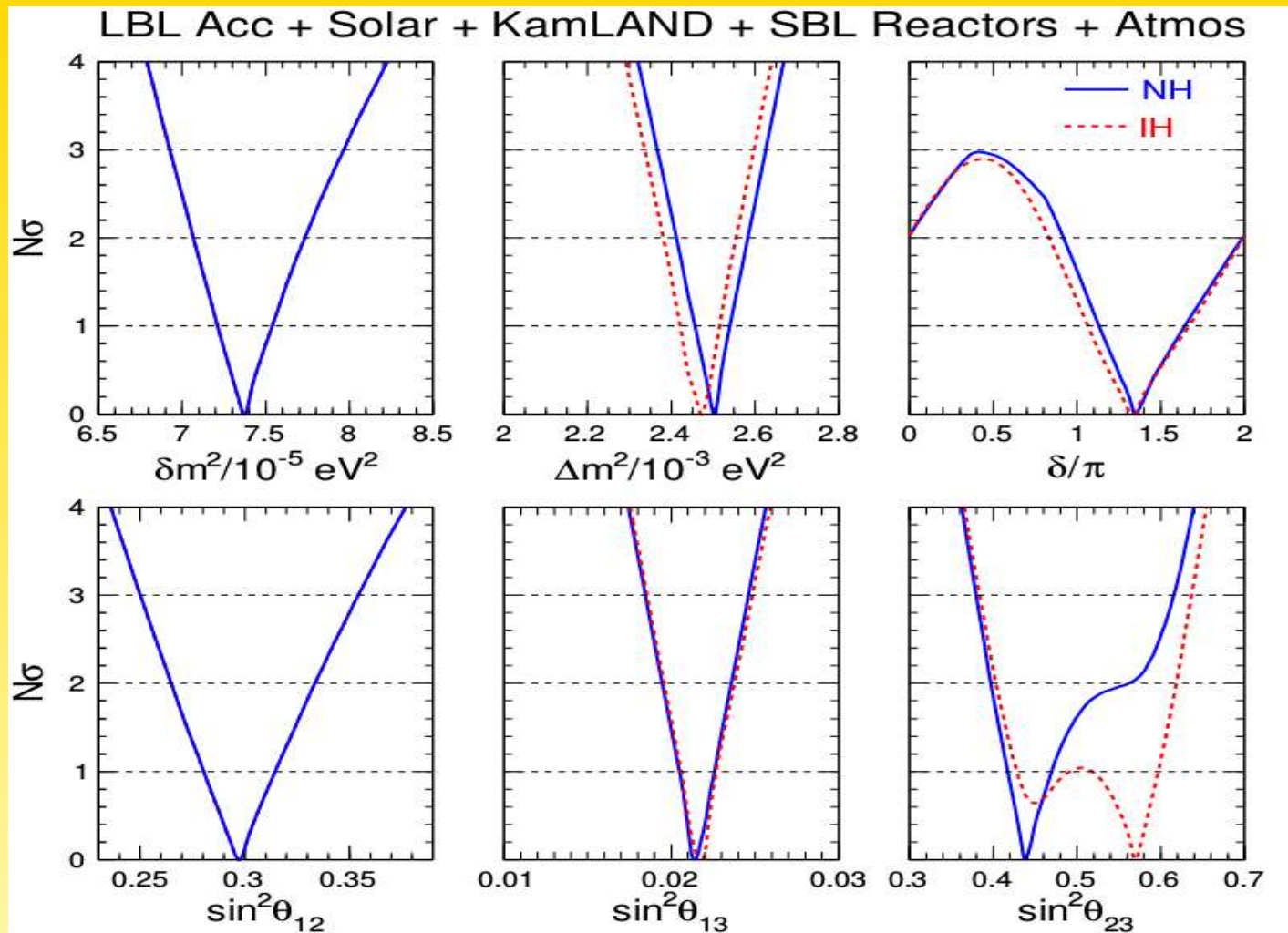
**“three Majorana neutrino paradigm”**



## Status 2016

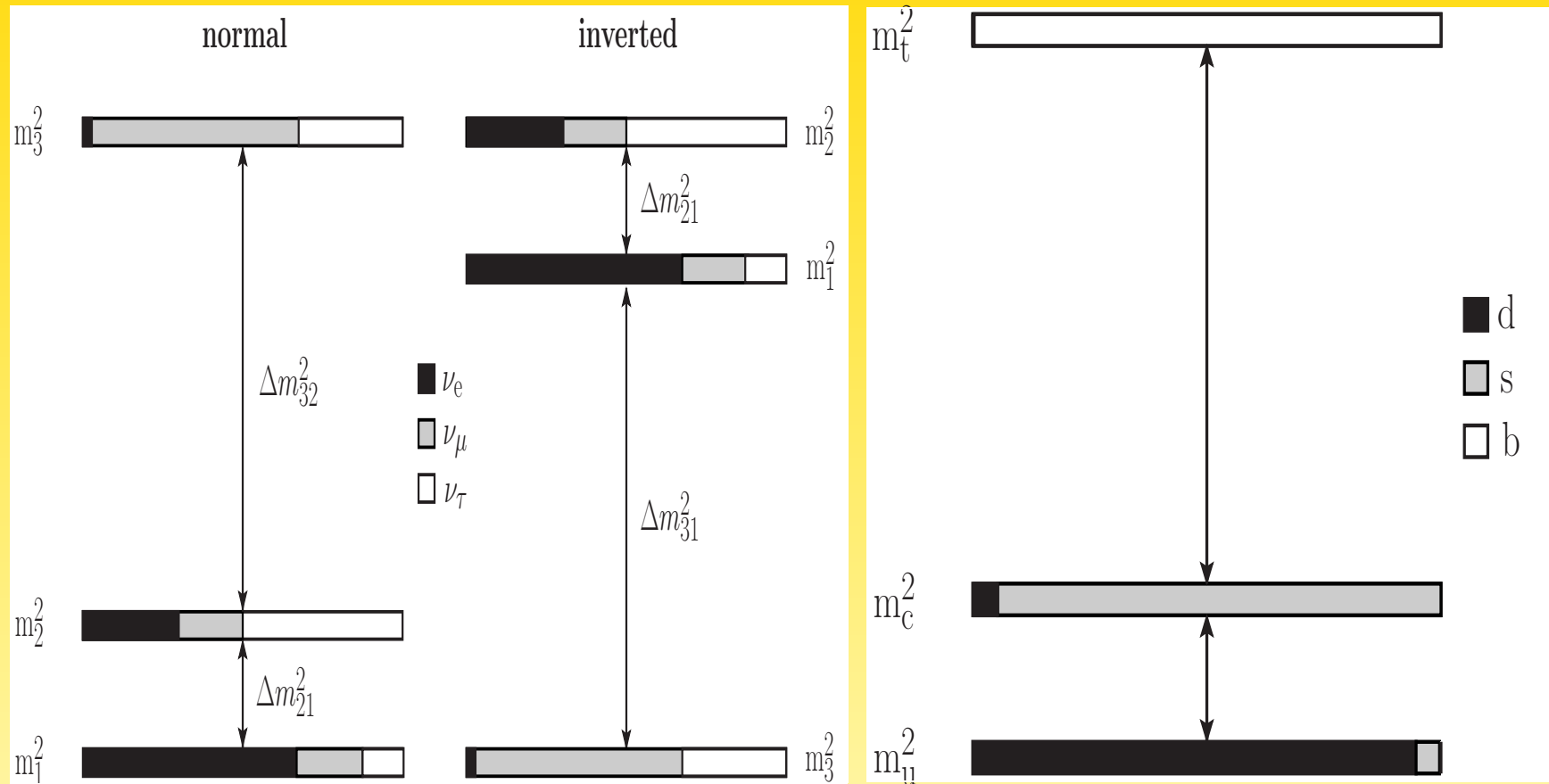
9 physical parameters in  $m_\nu$

- $\theta_{12}$  and  $m_2^2 - m_1^2$
- $\theta_{23}$  and  $|m_3^2 - m_2^2|$
- $\theta_{13}$
- $m_1, m_2, m_3$
- $\text{sgn}(m_3^2 - m_2^2)$
- Dirac phase  $\delta$
- Majorana phases  $\alpha$  and  $\beta$



*Lisi et al., 1601.07777*

(3-flavor, matter effects in solar, atm., LBL expts)



**Why so different?  $\leftrightarrow$  Flavor symmetries!**

PMNS-matrix:

$$|U| = \begin{pmatrix} 0.801 \dots 0.845 & 0.514 \dots 0.580 & 0.137 \dots 0.158 \\ 0.225 \dots 0.517 & 0.441 \dots 0.699 & 0.614 \dots 0.793 \\ 0.246 \dots 0.529 & 0.464 \dots 0.713 & 0.590 \dots 0.776 \end{pmatrix}$$

CKM-matrix:

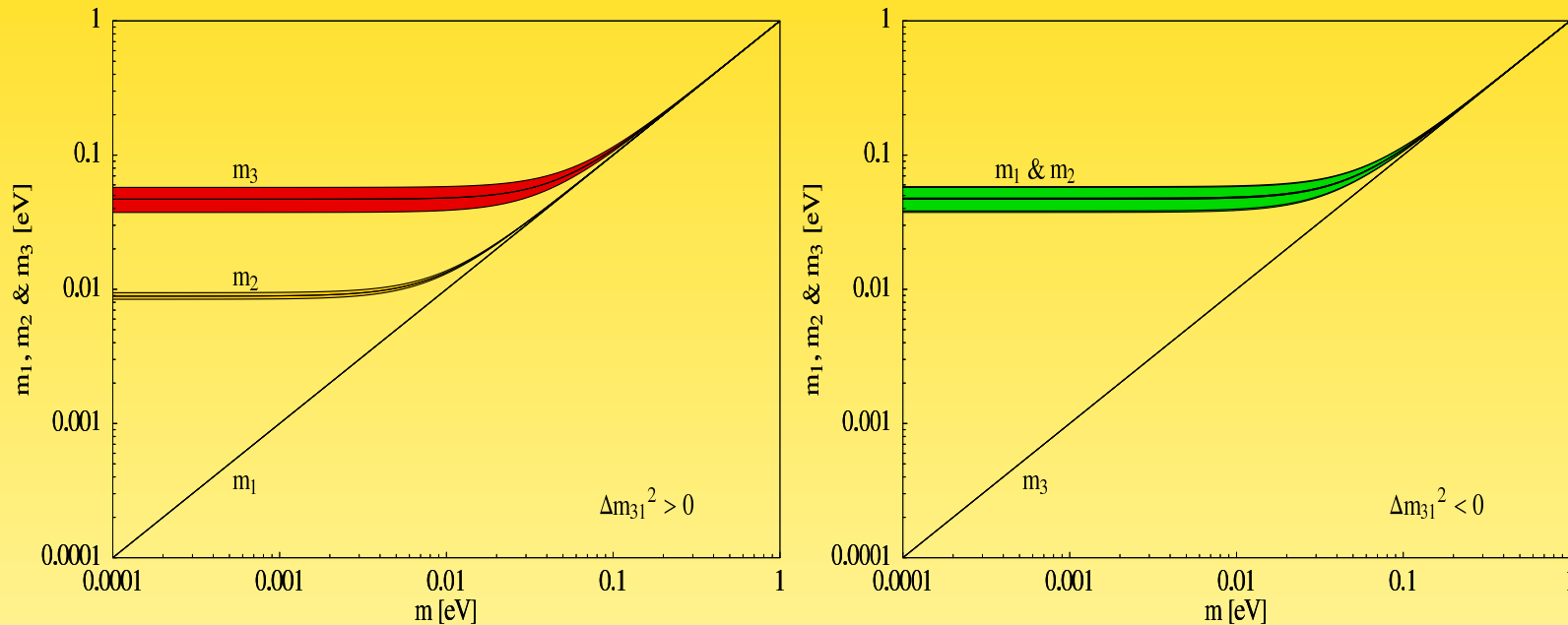
$$|V| = \begin{pmatrix} 0.97427 \pm 0.00014 & 0.22536 \pm 0.00061 & 0.00355^{+0.00015}_{-0.00014} \\ 0.22522 \pm 0.00061 & 0.97341 \pm 0.00015 & 0.0414 \pm 0.0012 \\ 0.00886^{+0.00033}_{-0.00032} & 0.0405^{+0.0011}_{-0.0012} & 0.99914 \pm 0.00005 \end{pmatrix}$$

**large** mixing in PMNS, small mixing in CKM

**Why so different? ↔ Flavor symmetries!**

## Neutrino masses

- neutrino masses  $\leftrightarrow$  scale of their origin
- neutrino mass ordering  $\leftrightarrow$  form of  $m_\nu$



- $m_3^2 \simeq \Delta m_A^2 \gg m_2^2 \simeq \Delta m_\odot^2 \gg m_1^2$ : normal hierarchy (NH)
- $m_2^2 \simeq |\Delta m_A^2| \simeq m_1^2 \gg m_3^2$ : inverted hierarchy (IH)
- $m_3^2 \simeq m_2^2 \simeq m_1^2 \equiv m_0^2 \gg \Delta m_A^2$ : quasi-degeneracy (QD)

## Contents

### I) Dirac vs. Majorana neutrinos

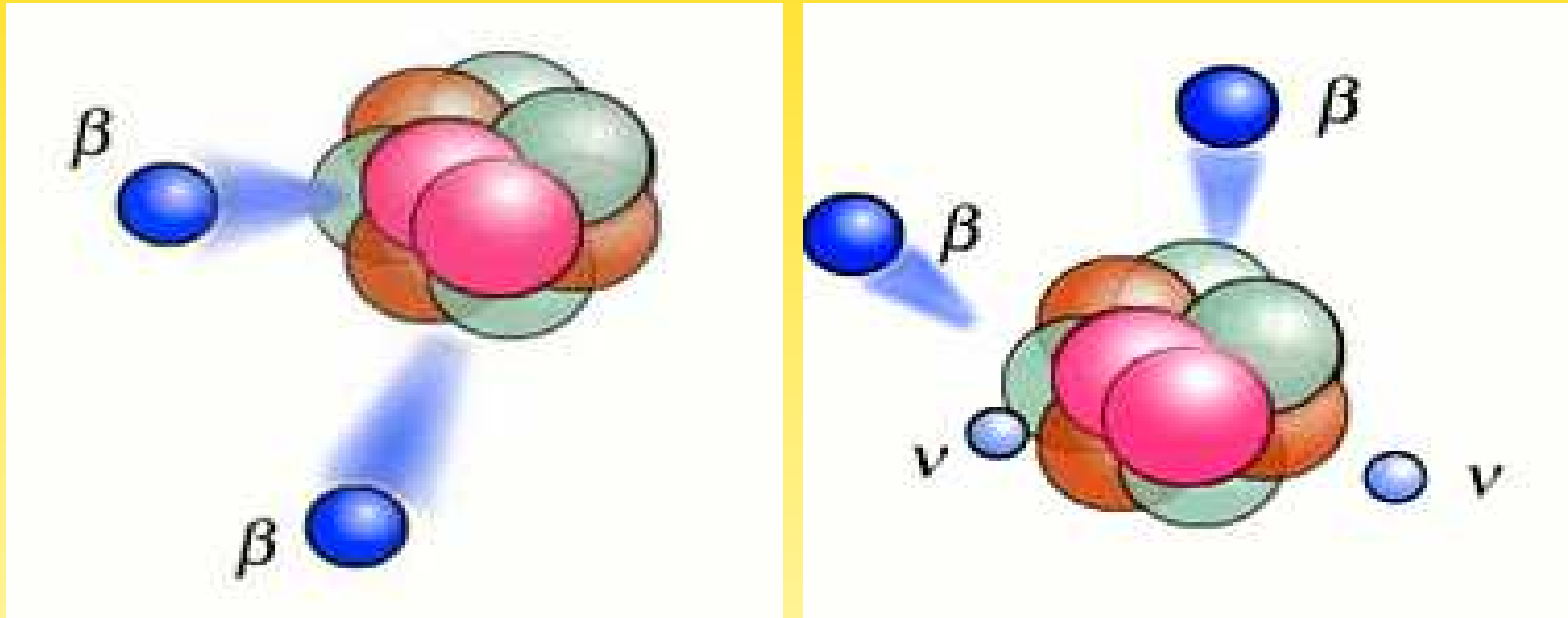
- I1) Basics
- I2) Seesaw mechanisms
- I3) Summary neutrino physics

### II) Neutrinoless double beta decay: standard interpretation

- II1) Basics
- II2) Implications for and from neutrino physics
- II3) Nuclear physics aspects
- II4) Exotic physics

## II) Neutrinoless Double Beta Decay: Standard Interpretation

$$(A, Z) \rightarrow (A, Z + 2) + 2 e^{-} \quad (0\nu\beta\beta)$$



- second order in weak interaction:  $\Gamma \propto G_F^4 \Rightarrow$  rare!
- not to be confused with  $(A, Z) \rightarrow (A, Z + 2) + 2 e^{-} + 2 \bar{\nu}_e \quad (2\nu\beta\beta)$   
(which occurs more often but is still rare)

## Contents

### **I) Dirac vs. Majorana neutrinos**

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### **II) Neutrinoless double beta decay: standard interpretation**

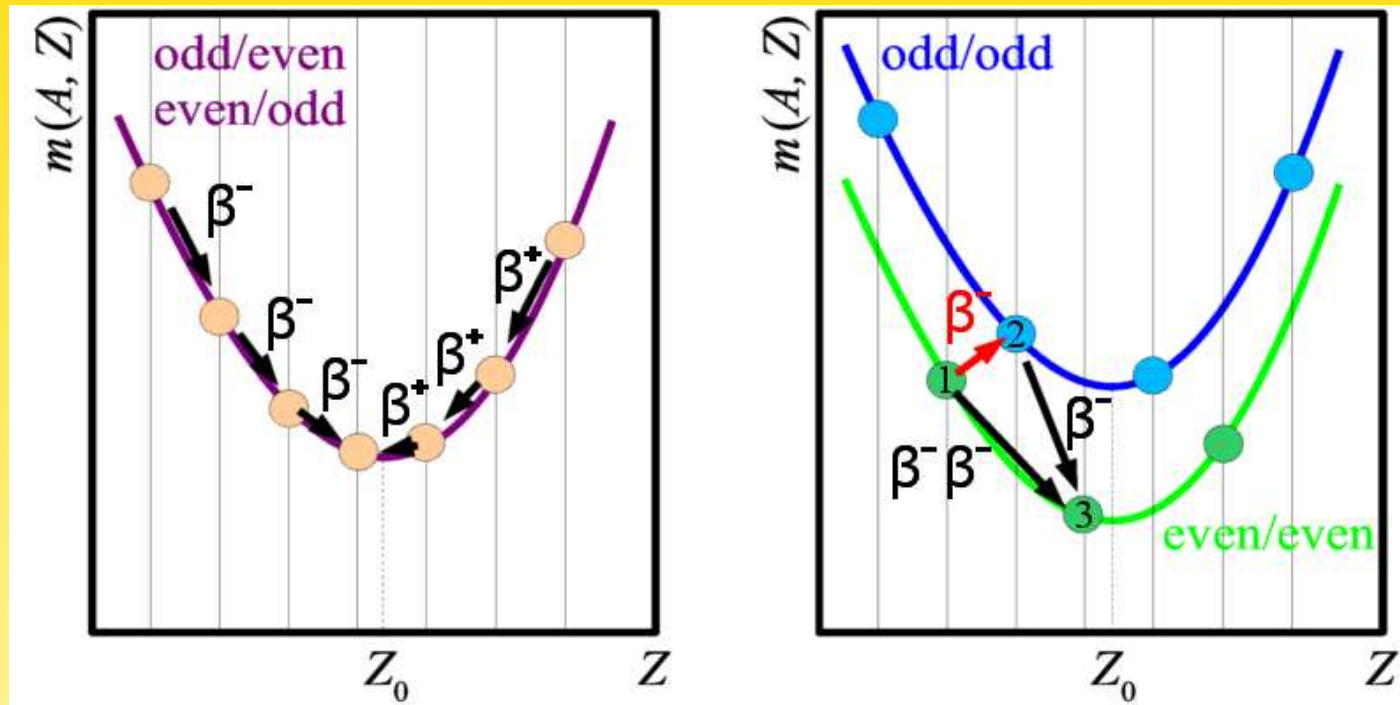
#### **II1) Basics**

- II2) Implications for and from neutrino physics
- II3) Nuclear physics aspects
- II4) Exotic physics



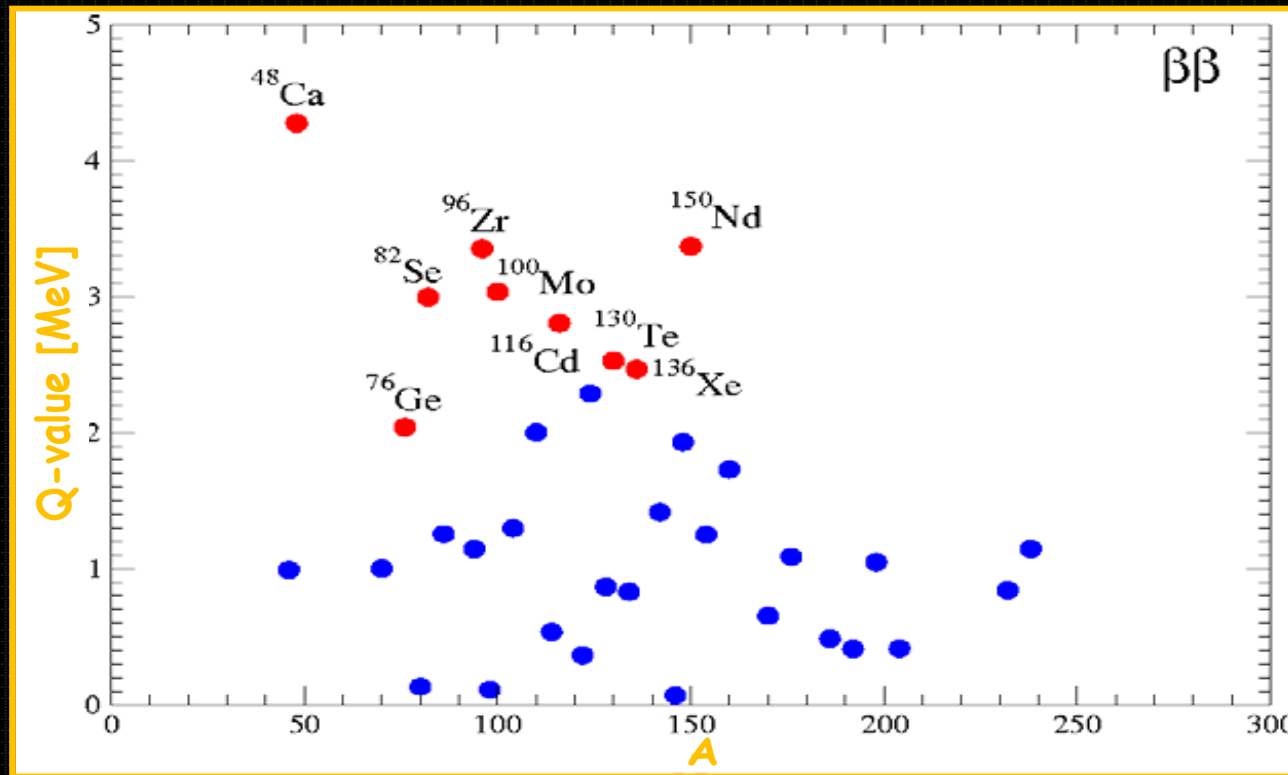
## III1) Basics

Need to forbid single  $\beta$  decay:



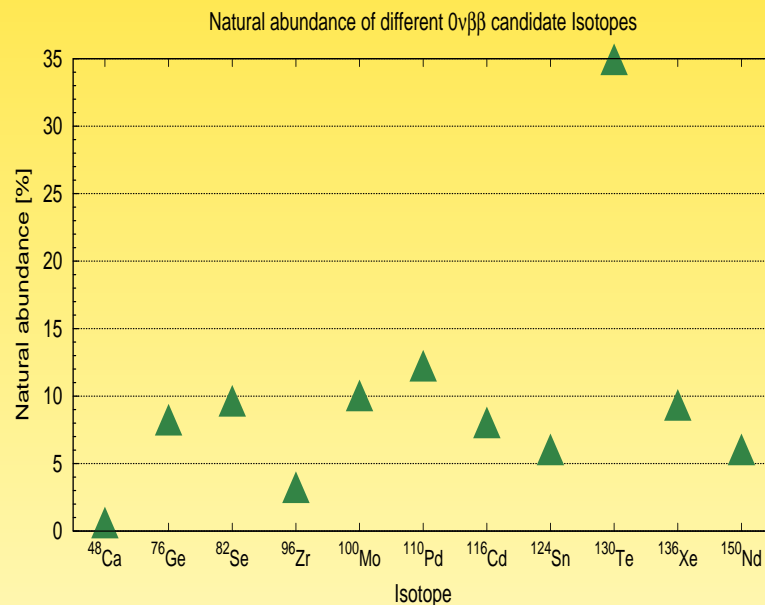
- $E_{\text{Bindung}} = E_{\text{Volumen}} - E_{\text{Oberfläche}} - E_{\text{Coulomb}} - E_{\text{Symmetrie}} \pm E_{\text{Paarbildung}} \Rightarrow \text{even/even} \rightarrow \text{even/even}$
- either direct ( $0\nu\beta\beta$ ) or two simultaneous decays with virtual (energetically forbidden) intermediate state ( $2\nu\beta\beta$ )

How many nuclei in this condition?

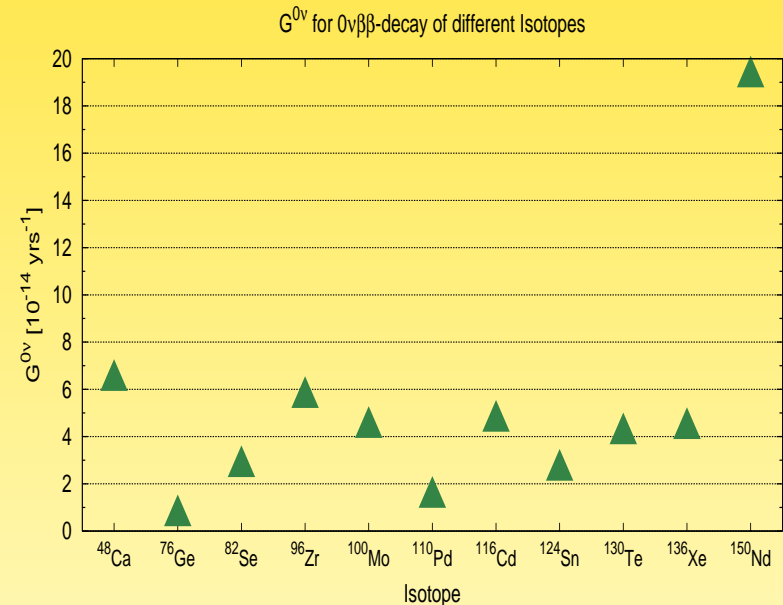


Slide by A. Giuliani

- 35 candidate isotopes
  - 9 are interesting:  $^{48}\text{Ca}$ ,  $^{76}\text{Ge}$ ,  $^{82}\text{Se}$ ,  $^{96}\text{Zr}$ ,  $^{100}\text{Mo}$ ,  $^{116}\text{Cd}$ ,  $^{130}\text{Te}$ ,  $^{136}\text{Xe}$ ,  $^{150}\text{Nd}$
  - $Q$ -value vs. natural abundance vs. reasonably priced enrichment vs. association with a well controlled experimental technique vs. . . .
- ⇒ no superisotope



$$T_{1/2}^{0\nu} \propto 1/a$$



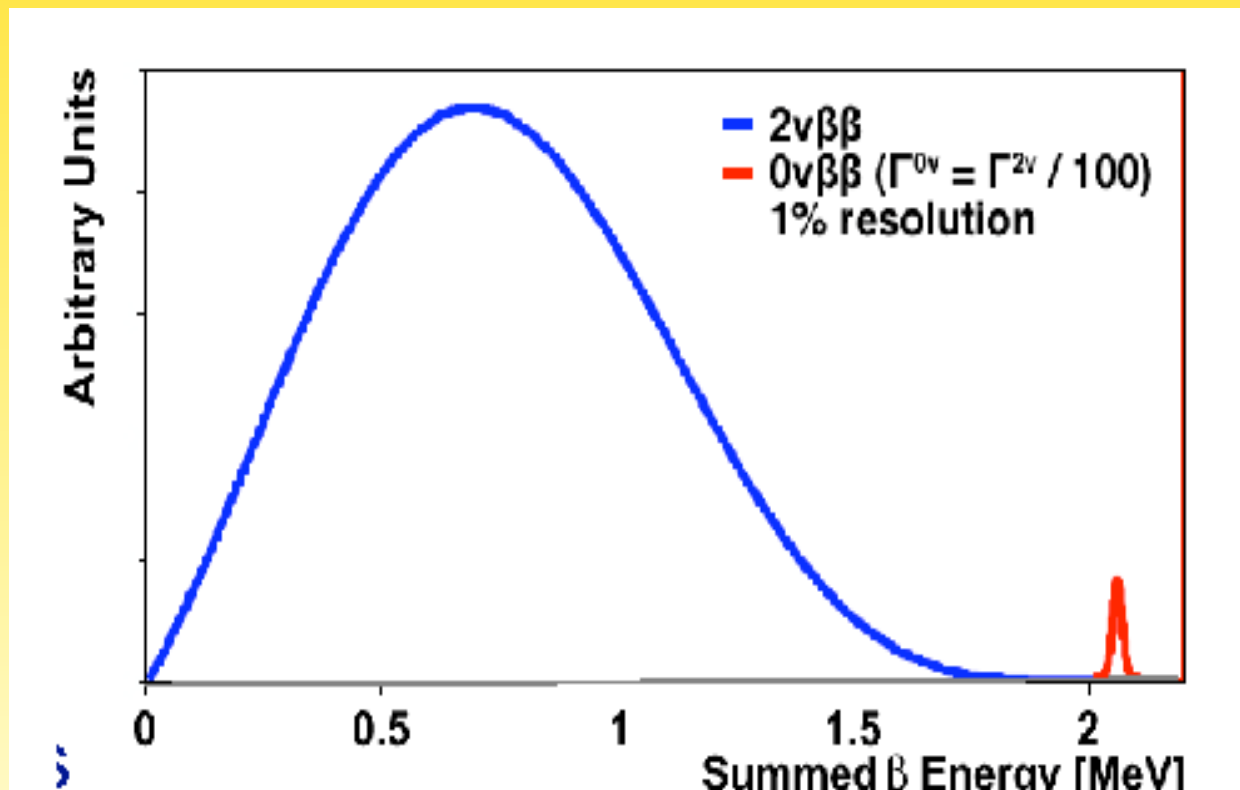
$$T_{1/2}^{0\nu} \propto Q^{-5}$$

Isotope	$G$ [ $10^{-14}$ yrs $^{-1}$ ]	$Q$ [keV]	nat. abund. [%]
$^{48}\text{Ca}$	6.35	4273.7	0.187
$^{76}\text{Ge}$	0.623	2039.1	7.8
$^{82}\text{Se}$	2.70	2995.5	9.2
$^{96}\text{Zr}$	5.63	3347.7	2.8
$^{100}\text{Mo}$	4.36	3035.0	9.6
$^{110}\text{Pd}$	1.40	2004.0	11.8
$^{116}\text{Cd}$	4.62	2809.1	7.6
$^{124}\text{Sn}$	2.55	2287.7	5.6
$^{130}\text{Te}$	4.09	2530.3	34.5
$^{136}\text{Xe}$	4.31	2461.9	8.9
$^{150}\text{Nd}$	19.2	3367.3	5.6

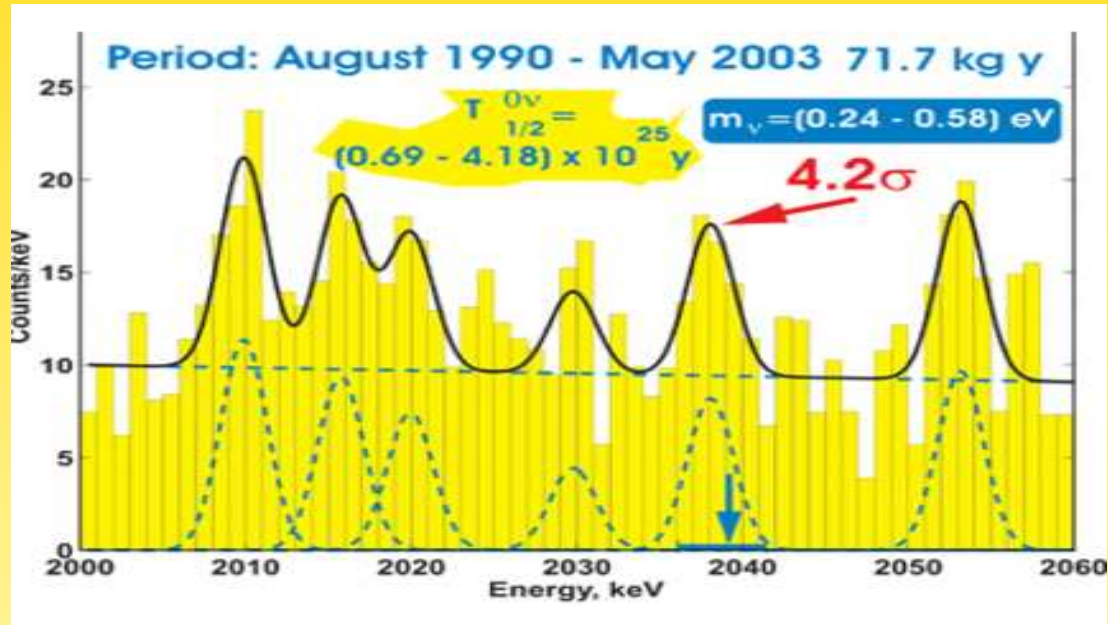
Most mechanisms:  $G_x \propto Q^5$

## Experimental Aspects

- experimental signature: sum of electron energies =  $Q$   
(plus: 2 electrons and daughter isotope)
- background of  $2\nu\beta\beta \leftrightarrow$  resolution



if claimed, typical spectrum will look like:



⇒ first reason for multi-isotope determination

## Experimental Aspects

number of events (measuring time  $t \ll T_{1/2}^{0\nu}$  life-time):

$$N = \ln 2 a M t N_A (T_{1/2}^{0\nu})^{-1}$$

with

- $a$  is abundance of isotope
- $M$  is used mass
- $t$  is time of measurement
- $N_A$  is Avogadro's number

## Experimental Aspects

Number of events (measuring time  $t \ll T_{1/2}^{0\nu}$  life-time):

$$N = \ln 2 a M t N_A (T_{1/2}^{0\nu})^{-1}$$

suppose there is no background:

- if you want  $10^{26}$  yrs you need  $10^{26}$  atoms
- $10^{26}$  atoms are  $10^3$  mols
- $10^3$  mols are 100 kg

From now on you can only loose: efficiency, background, natural abundance,...



## Experimental Aspects

$$(T_{1/2}^{0\nu})^{-1} \propto \begin{cases} a M \epsilon t & \text{without background} \\ a \epsilon \sqrt{\frac{M t}{B \Delta E}} & \text{with background} \end{cases}$$

with

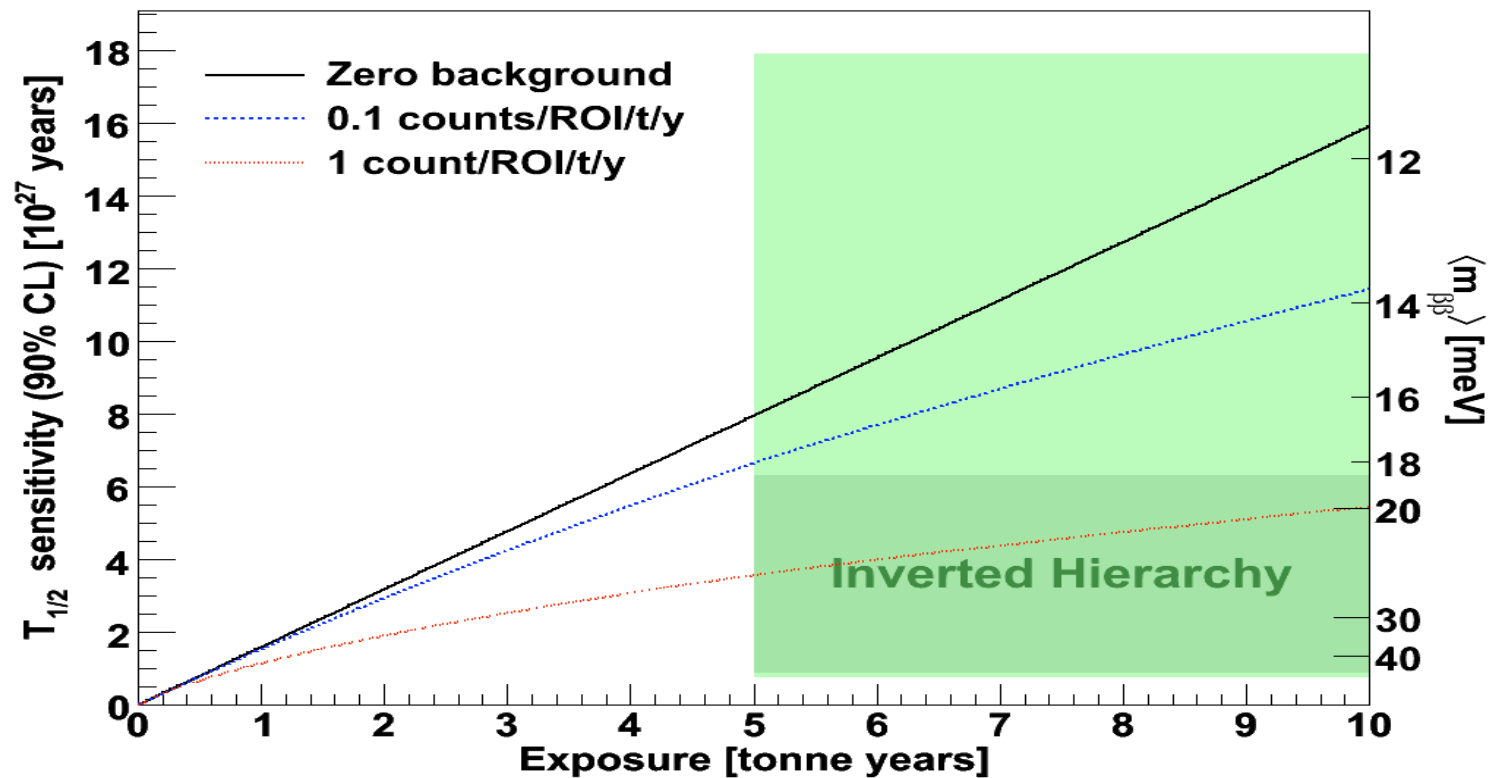
- $B$  is background index in counts/(keV kg yr)
- $\Delta E$  is energy resolution
- $\epsilon$  is efficiency
- $(T_{1/2}^{0\nu})^{-1} \propto (\text{particle physics})^2$

*Note: factor 2 in particle physics is combined factor of 16 in  $M \times t \times B \times \Delta E$*

# Sensitivity and backgrounds

I-tonne  $^{76}\text{Ge}$  Example

$$T_{1/2}^{0\nu} = \ln(2)N\epsilon t/UL(B)$$



Neutrinoless Double Beta Decay

Goldhaber-Grodzins-Sunyar Celebration

slide by J.F. Wilkerson

## Interpretation of Experiments

Master formula:

$$\Gamma^{0\nu} = G_x(Q, Z) |\mathcal{M}_x(A, Z) \eta_x|^2$$

- $G_x(Q, Z)$ : phase space factor
- $\mathcal{M}_x(A, Z)$ : nuclear physics
- $\eta_x$ : particle physics

## Interpretation of Experiments

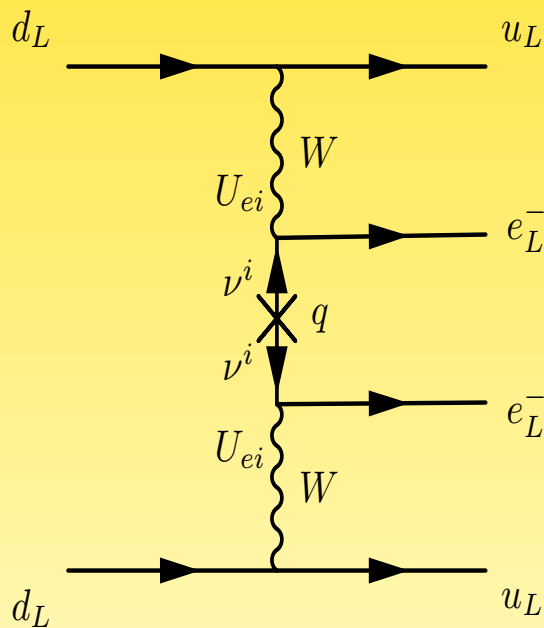
Master formula:

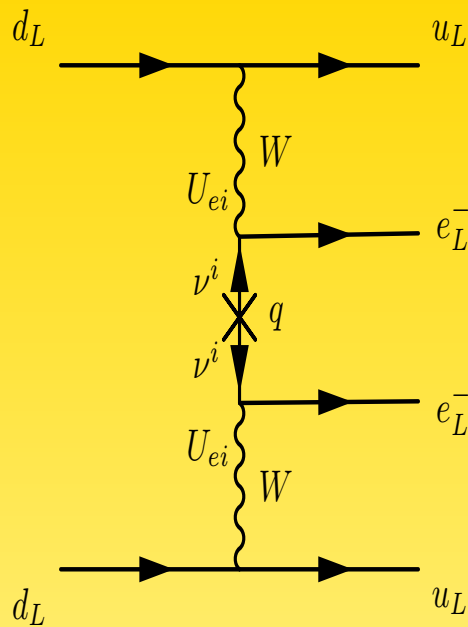
$$\Gamma^{0\nu} = G_x(Q, Z) |\mathcal{M}_x(A, Z) \eta_x|^2$$

- $G_x(Q, Z)$ : phase space factor; **calculable**
- $\mathcal{M}_x(A, Z)$ : nuclear physics; **problematic**
- $\eta_x$ : particle physics; **interesting**

## Standard Interpretation

Neutrinoless Double Beta Decay is mediated by light and massive Majorana neutrinos (the ones which oscillate) and all other mechanisms potentially leading to  $0\nu\beta\beta$  give negligible or no contribution





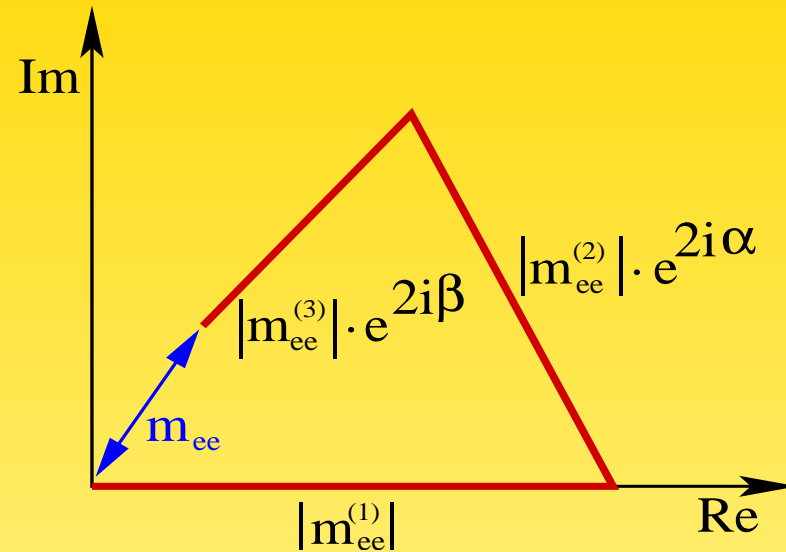
- $U_{ei}^2$  from charged current
- $m_i/E_i$  from spin-flip and *if neutrinos are Majorana particles*

amplitude proportional to coherent sum (“effective mass”)

$$|m_{ee}| = \left| \sum U_{ei}^2 m_i \right|$$

$m/E \simeq \text{eV}/100 \text{ MeV}$  is tiny: only  $N_A$  can save the day!

## The effective mass



amplitude proportional to coherent sum (“effective mass”):

$$|m_{ee}| \equiv \left| \sum U_{ei}^2 m_i \right| = \left| |U_{e1}|^2 m_1 + |U_{e2}|^2 m_2 e^{2i\alpha} + |U_{e3}|^2 m_3 e^{2i\beta} \right|$$

$$= f(\theta_{12}, |U_{e3}|, m_i, \text{sgn}(\Delta m_A^2), \alpha, \beta)$$

7 out of 9 parameters of neutrino physics!

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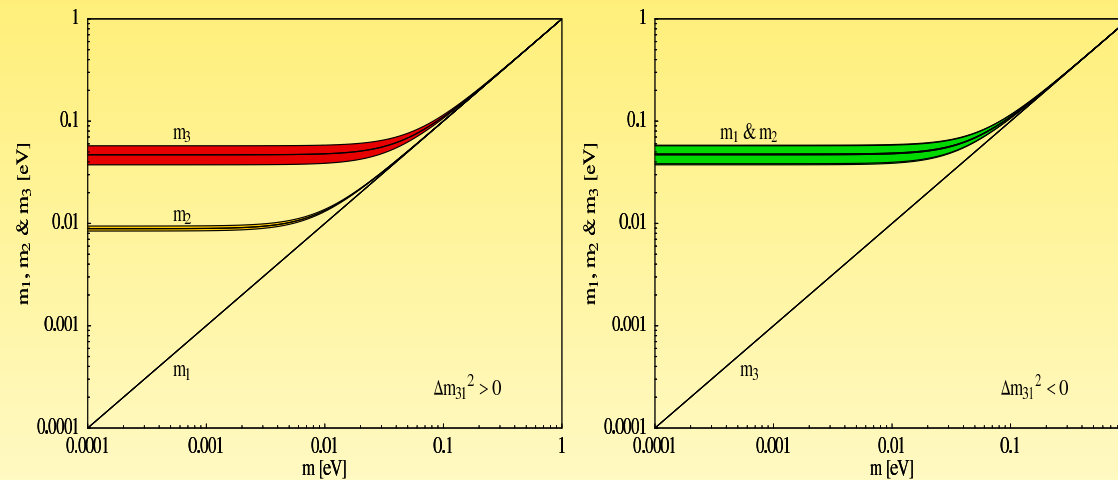
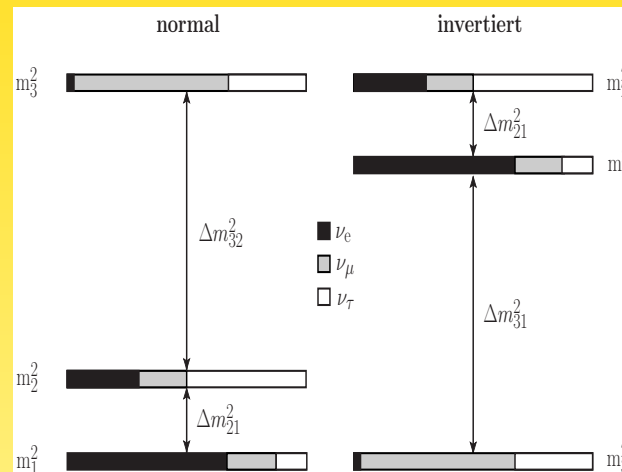
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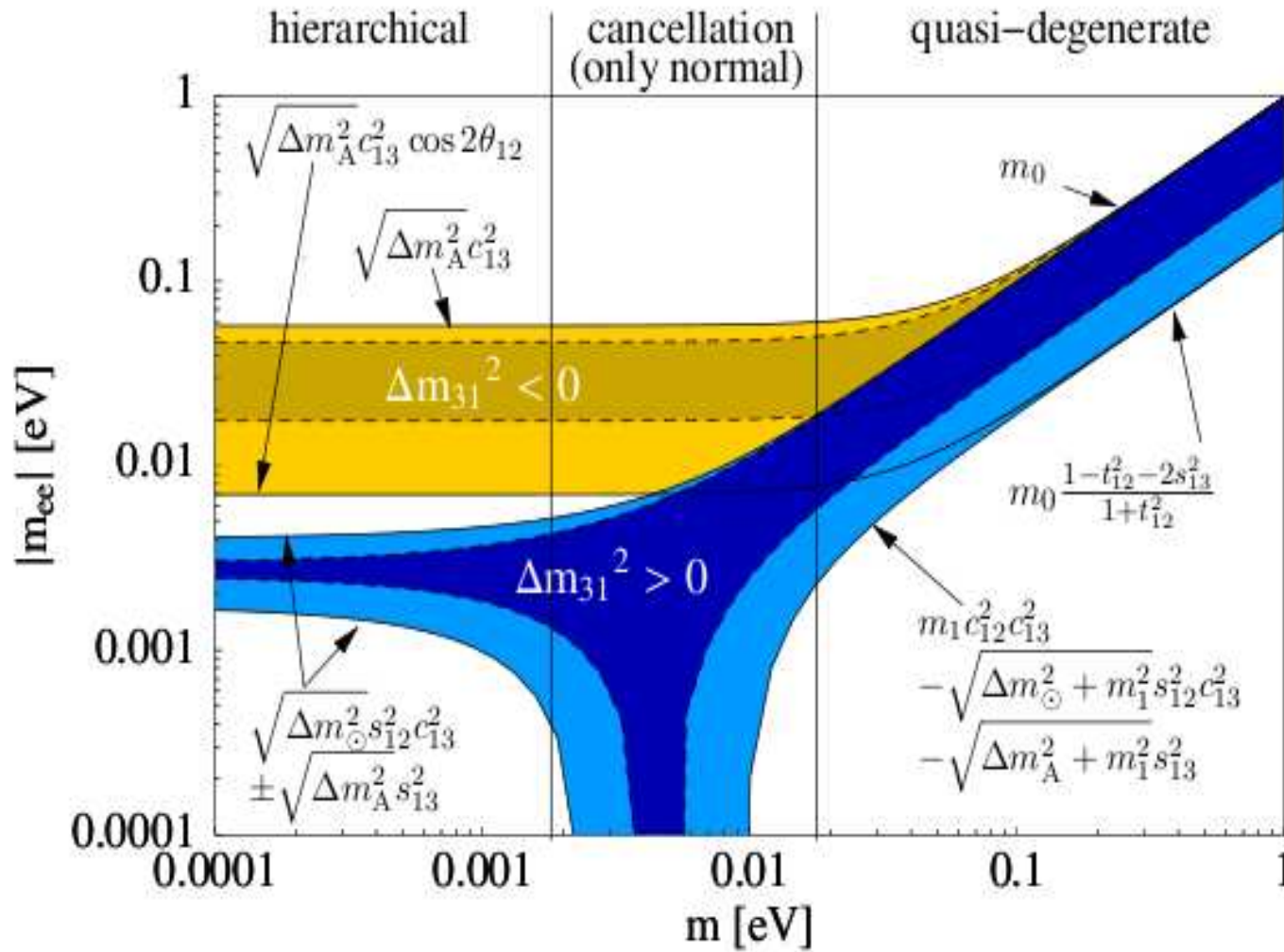


## II2) Implications for and from neutrino physics

$|m_{ee}| = \left| \sum U_{ei}^2 m_i \right|$ : fix known things, vary known unknown things, assume no unknown unknowns are present:

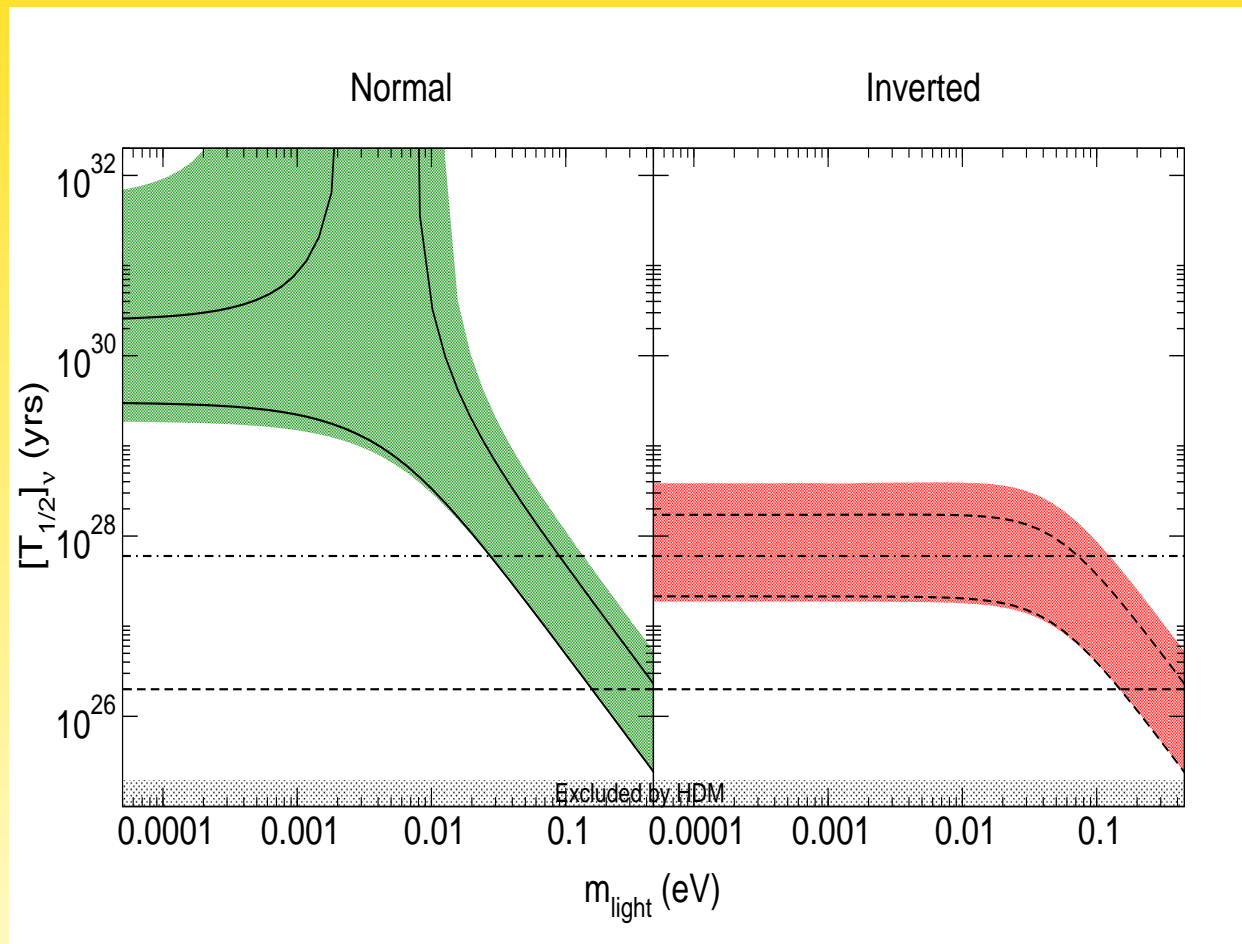


## The usual plot



# The usual plot: the other way around

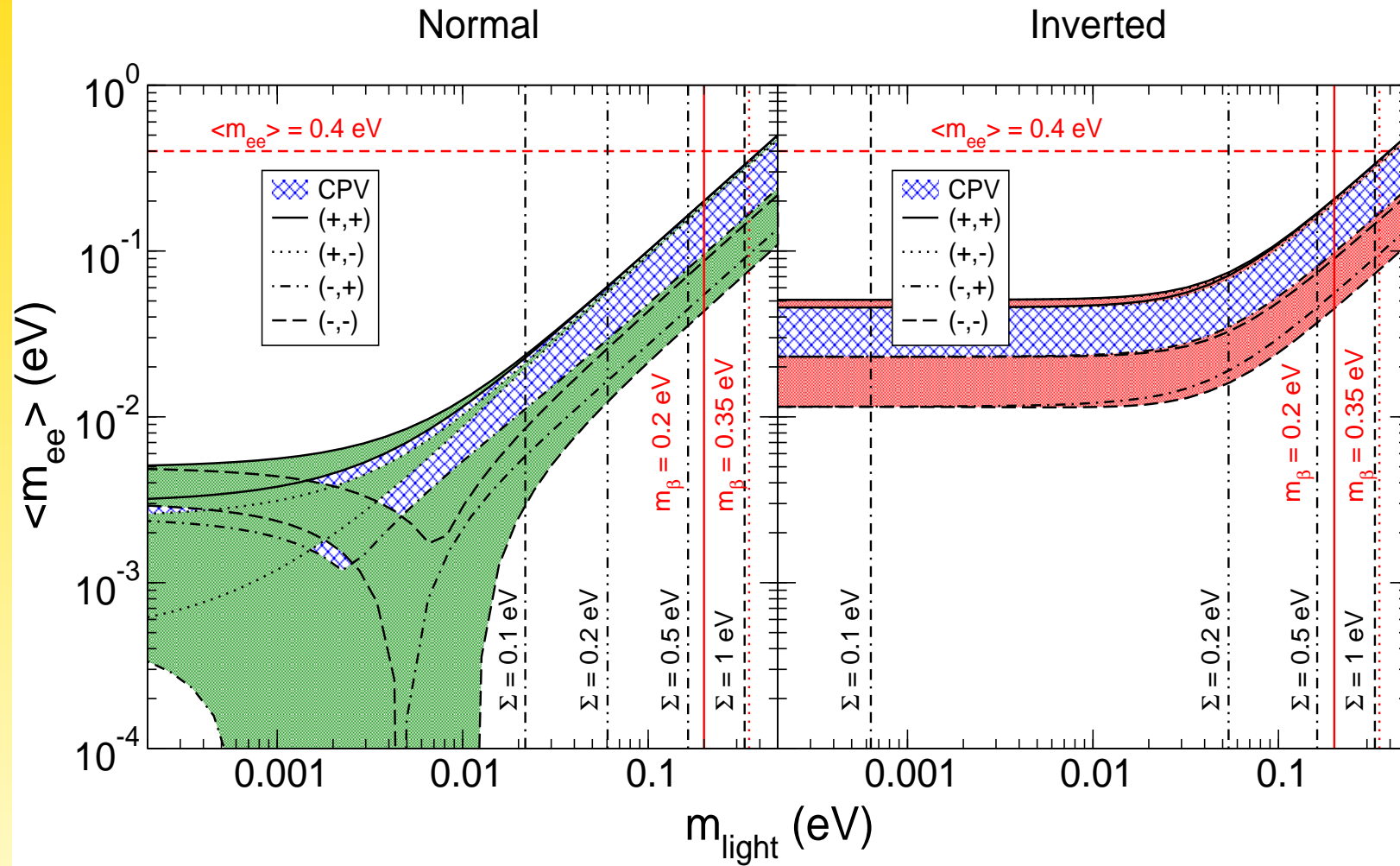
(life-time instead of  $|m_{ee}|$ )



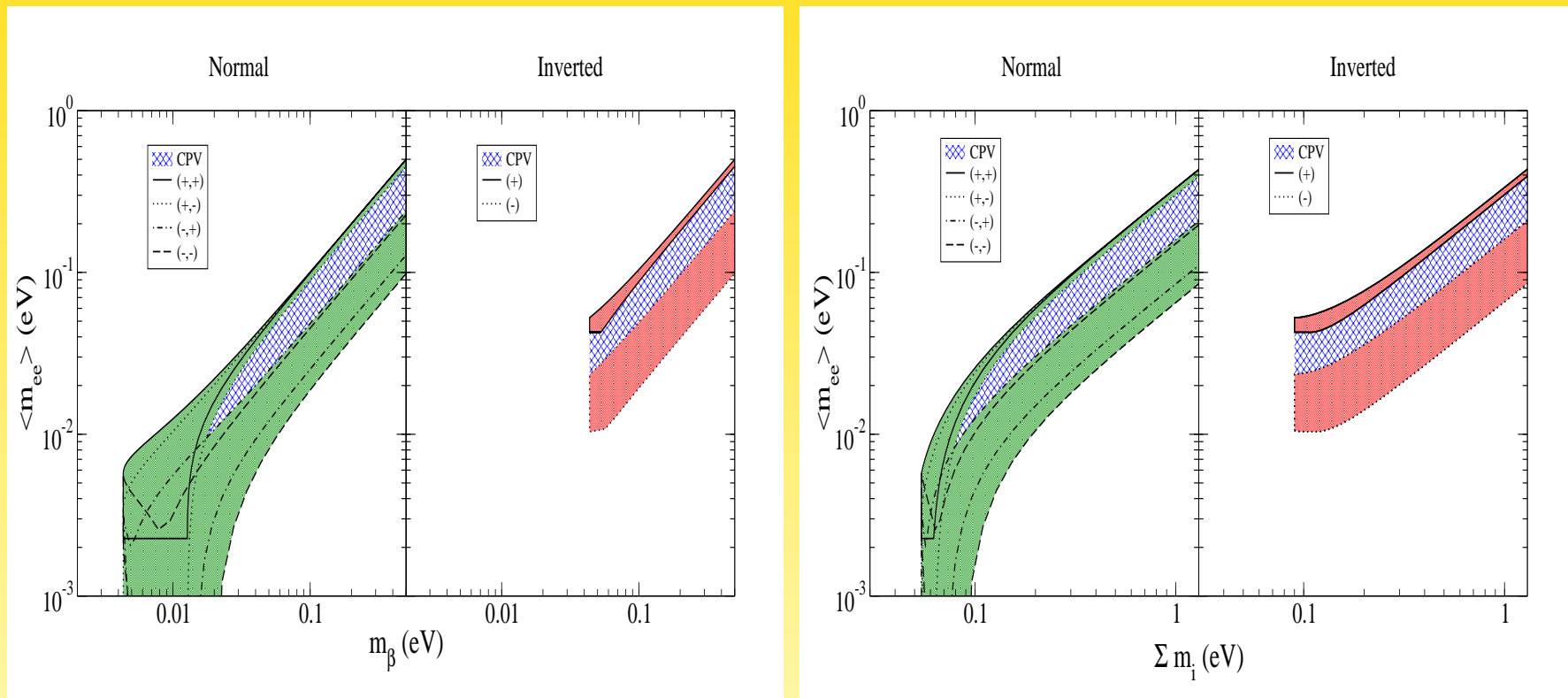
## Which mass ordering with which life-time?

	$\Sigma$	$m_\beta$	$ m_{ee} $
NH	$\sqrt{\Delta m_A^2}$ $\simeq 0.05 \text{ eV}$	$\sqrt{\Delta m_\odot^2 +  U_{e3} ^2 \Delta m_A^2}$ $\simeq 0.01 \text{ eV}$	$\left  \sqrt{\Delta m_\odot^2 +  U_{e3} ^2 \Delta m_A^2} e^{2i(\alpha-\beta)} \right $ $\sim 0.003 \text{ eV} \Rightarrow T_{1/2}^{0\nu} \gtrsim 10^{28-29} \text{ yrs}$
IH	$2\sqrt{\Delta m_A^2}$ $\simeq 0.1 \text{ eV}$	$\sqrt{\Delta m_A^2}$ $\simeq 0.05 \text{ eV}$	$\sqrt{\Delta m_A^2} \sqrt{1 - \sin^2 2\theta_{12} \sin^2 \alpha}$ $\sim 0.03 \text{ eV} \Rightarrow T_{1/2}^{0\nu} \gtrsim 10^{26-27} \text{ yrs}$
QD	$3m_0$	$m_0$	$m_0 \sqrt{1 - \sin^2 2\theta_{12} \sin^2 \alpha}$ $\gtrsim 0.1 \text{ eV} \Rightarrow T_{1/2}^{0\nu} \gtrsim 10^{25-26} \text{ yrs}$

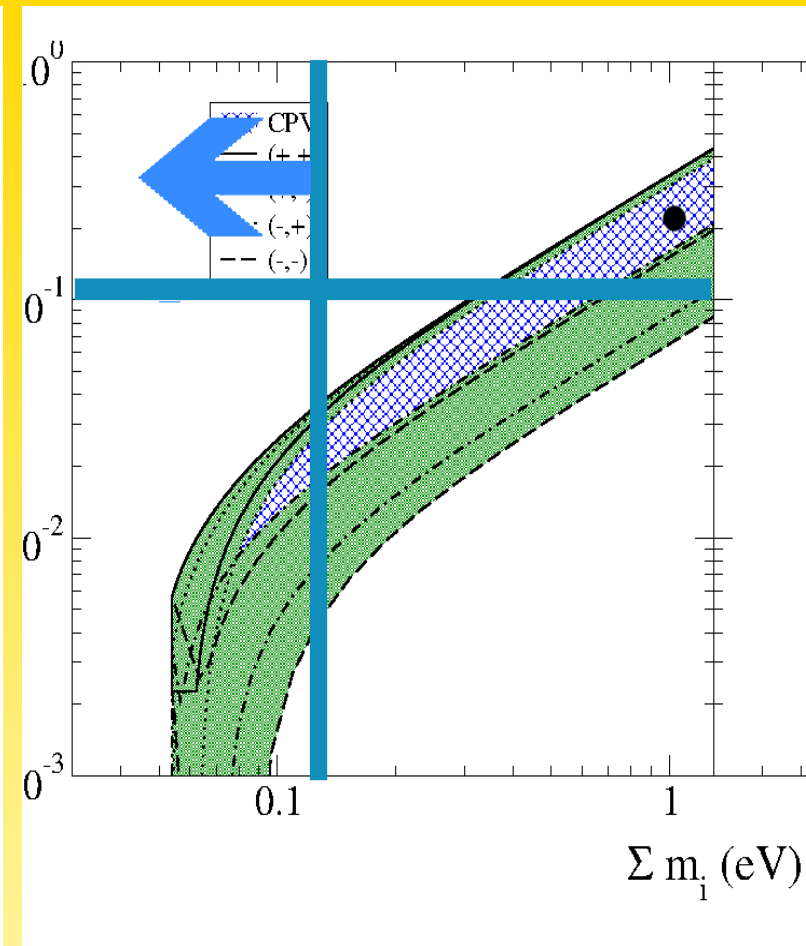
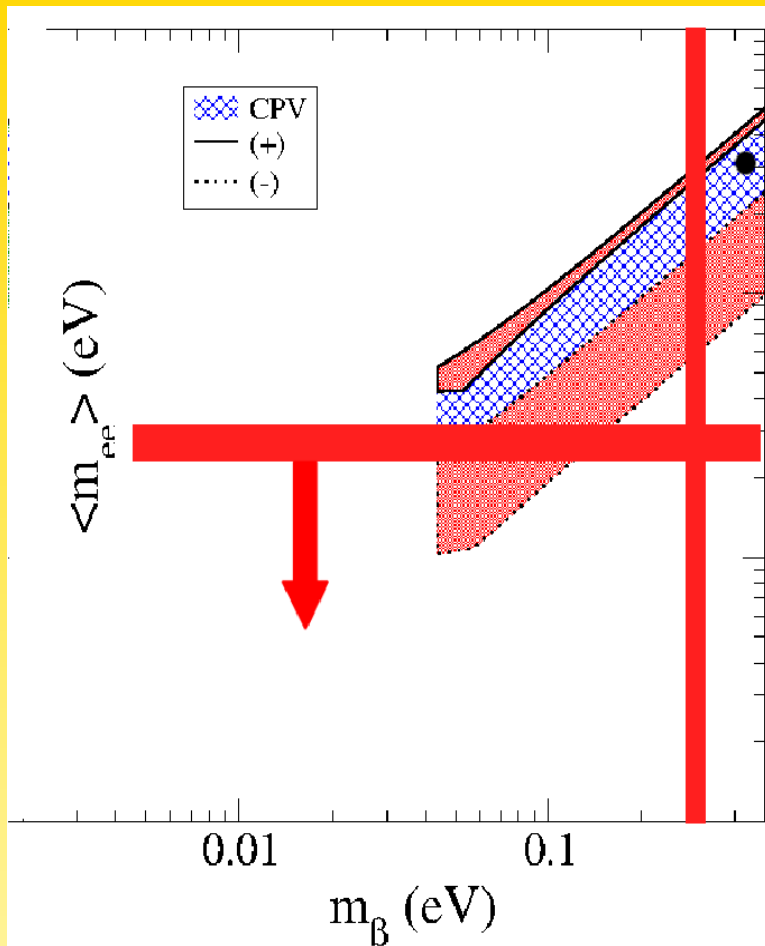
# The usual plot: include other neutrino mass approaches



## Plot against other observables



**Complementarity** of  $|m_{ee}| = U_{ei}^2 m_i$  ,  $m_\beta = \sqrt{|U_{ei}|^2 m_i^2}$  and  $\Sigma = \sum m_i$



CP violation!

Dirac neutrinos!

something else does  $0\nu\beta\beta$ !

## Neutrino Mass

$$m(\text{heaviest}) \gtrsim \sqrt{|m_3^2 - m_1^2|} \simeq 0.05 \text{ eV}$$

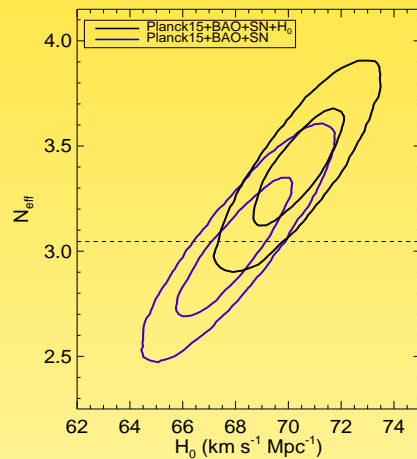
3 **complementary** methods to measure neutrino mass:

Method	observable	now [eV]	near [eV]	far [eV]	pro	con
Kurie	$\sqrt{\sum  U_{ei} ^2 m_i^2}$	2.3	0.2	0.1	model-indep.; theo. clean	final?; worst
Cosmo.	$\sum m_i$	0.5	0.2	0.05	best; NH/IH	systemat.; model-dep.
$0\nu\beta\beta$	$ \sum U_{ei}^2 m_i $	0.3	0.1	0.05	fundament.; NH/IH	model-dep.; theo. dirty

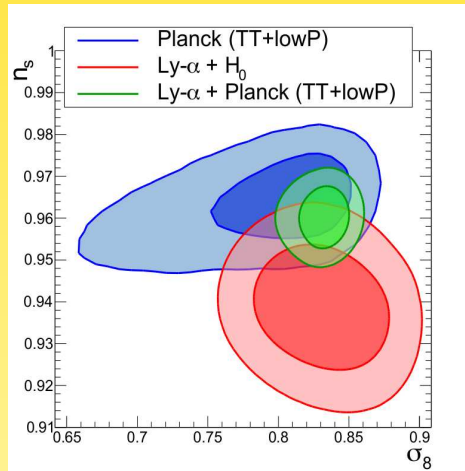


## Cosmological Limits

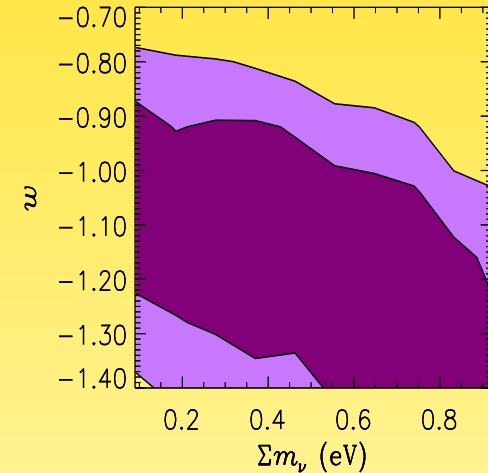
stacking more and more data sets on top of each other, limits become stronger  
including more and more parameters, limits become weaker  
needed to break degeneracies, but induces systematic issues



Riess *et al.*,  
1604.01424



Palanque-Delabrouille *et al.*,  
1506.05976



Hannestad,  
PRL 95

most extreme (1410.7244): at  $1\sigma$ :  $\Sigma m_\nu \leq 0.08$  eV, disfavors inverted ordering...

## 2 kinds of neutrino masses

1)  $ee$ -element of mass matrix

$$\sqrt{\frac{1}{T_{1/2}^{0\nu}}} \propto |(m_\nu)_{ee}| \quad \text{with } (m_\nu)_{ee} = \frac{h_{ee} v^2}{\Lambda} \quad \text{in } \mathcal{L}_{\text{eff}} = \frac{1}{2} \frac{h_{\alpha\beta}}{\Lambda} \overline{L_\alpha} \tilde{\Phi} \tilde{\Phi}^T L_\beta^c$$

direct probe of fundamental object in low energy Lagrangian!

2) neutrino mass scale: QD neutrinos

$$|m_{ee}|^{\text{QD}} = m_0 |c_{12}^2 c_{13}^2 + s_{12}^2 c_{13}^2 e^{2i\alpha} + s_{13}^2 e^{2i\beta}|$$
$$\Rightarrow m_0 \leq |m_{ee}|_{\text{min}}^{\text{exp}} \frac{1 + \tan^2 \theta_{12}}{1 - \tan^2 \theta_{12} - 2|U_{e3}|^2} \simeq 3 |m_{ee}|_{\text{min}}^{\text{exp}} \simeq \text{eV}$$

same order as Mainz/Troitsk!

## Alternative processes

$$(A, Z) \rightarrow (A, Z + 2)^* + 2 e^- \quad (0\nu\beta\beta)^*$$

$$(A, Z) \rightarrow (A, Z - 2) + 2 e^+ \quad (0\nu\beta^+\beta^+)$$

$$e_b^- + (A, Z) \rightarrow (A, Z - 2) + e^+ \quad (0\nu\beta^+EC)$$

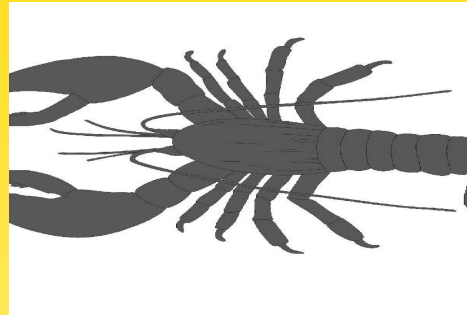
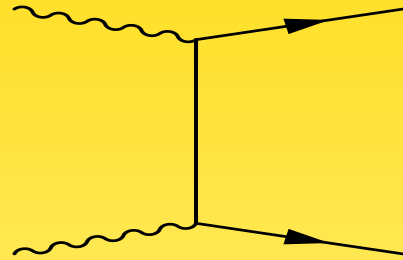
$$2 e_b^- + (A, Z) \rightarrow (A, Z - 2)^* \quad (0\nuECEC)$$

depend on same particle physics parameters, but more difficult to realize/test

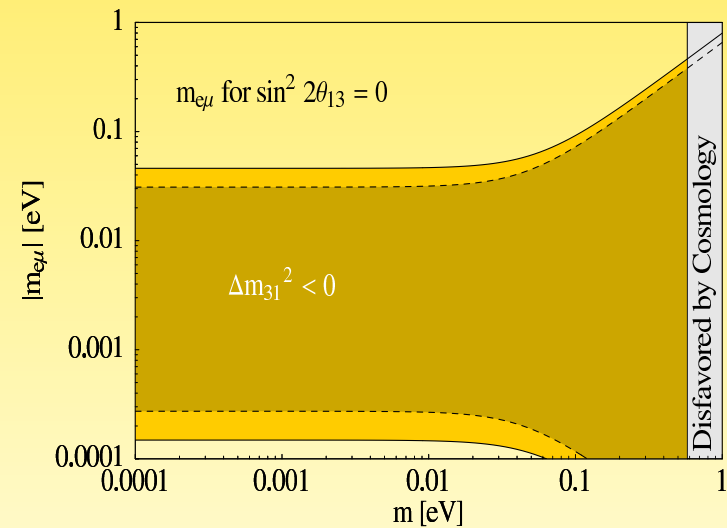
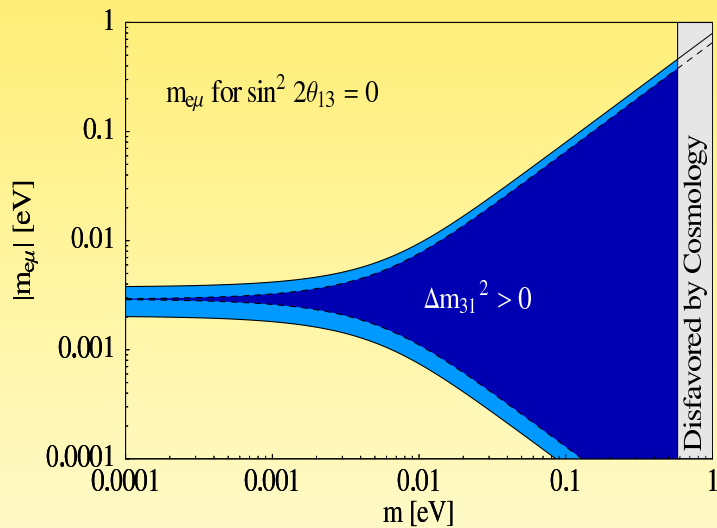
BUT: ratio to  $0\nu\beta\beta$  is test of NME calculation and mechanism

# Alternative processes

the lobster:

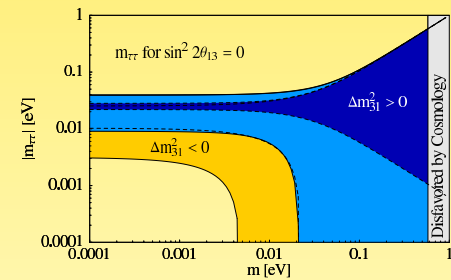
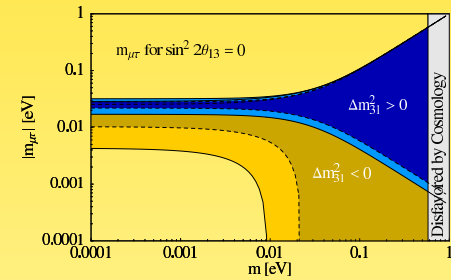
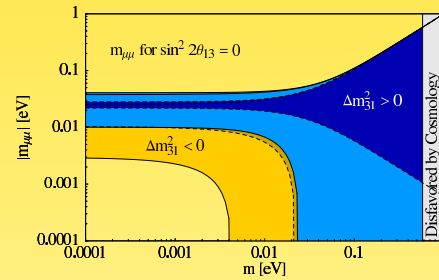
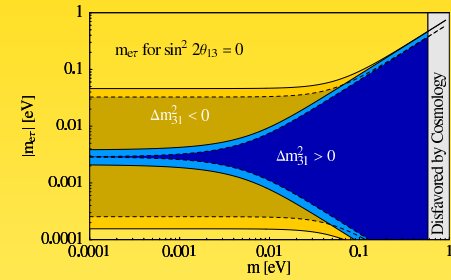
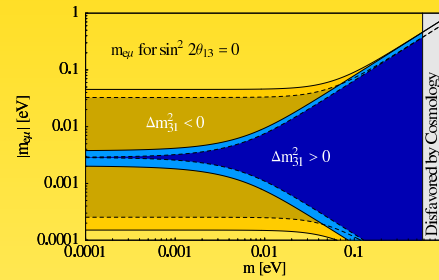
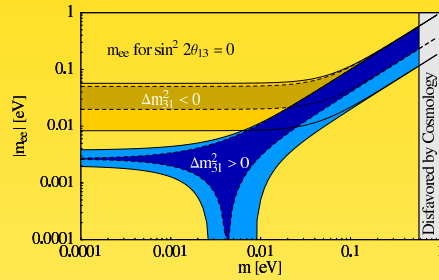


$$\text{BR}(K^+ \rightarrow \pi^- e^+ \mu^+) \propto |m_{e\mu}|^2 = \left| \sum U_{ei} U_{\mu i} m_i \right|^2 \sim 10^{-30} \left( \frac{|m_{e\mu}|}{\text{eV}} \right)^2$$



# Reconstructing $m_\nu$

$m_\nu =$



## Recent Results

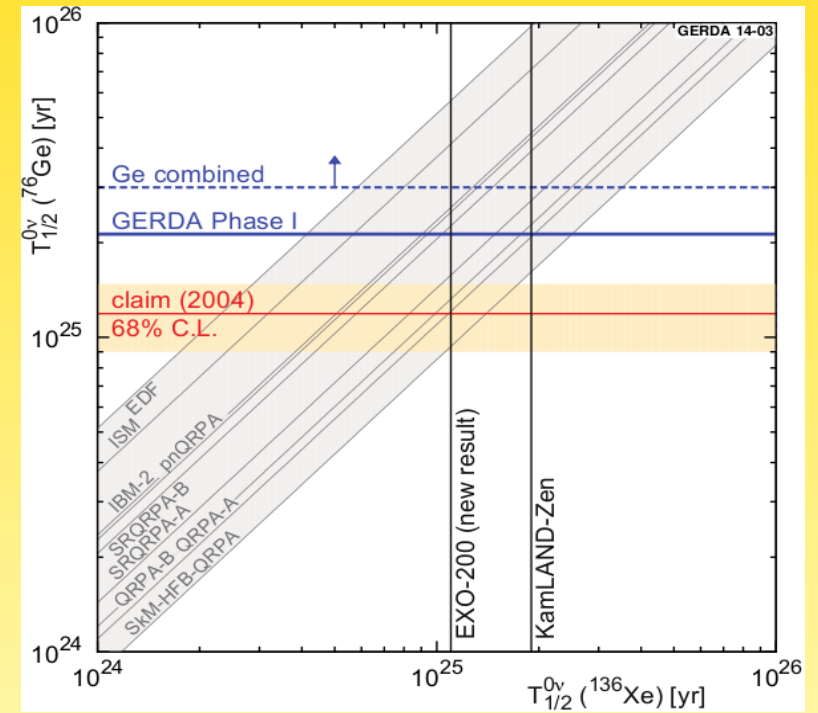
- $^{76}\text{Ge}$ :
  - GERDA:  $T_{1/2} > 2.1 \times 10^{25}$  yrs
  - GERDA + IGEX + HDM:  $T_{1/2} > 3.0 \times 10^{25}$  yrs
- $^{136}\text{Xe}$ :
  - EXO-200:  $T_{1/2} > 1.1 \times 10^{25}$  yrs (first run with less exposure:  $T_{1/2} > 1.6 \times 10^{25}$  yrs. . .)
  - KamLAND-Zen:  $T_{1/2} > 2.6 \times 10^{25}$  yrs

Xe-limit is stronger than Ge-limit when:

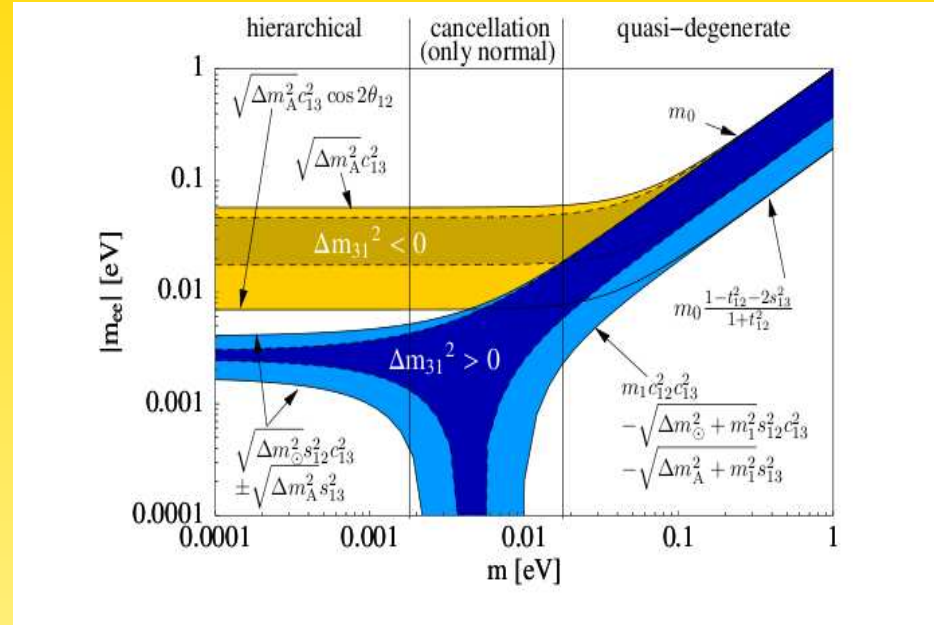
$$T_{\text{Xe}} > T_{\text{Ge}} \frac{G_{\text{Ge}}}{G_{\text{Xe}}} \left| \frac{\mathcal{M}_{\text{Ge}}}{\mathcal{M}_{\text{Xe}}} \right|^2 \text{ yrs}$$

## Current Limits on $|m_{ee}|$

NME	$^{76}\text{Ge}$		$^{136}\text{Xe}$	
	GERDA	comb	KLZ	comb
EDF(U)	0.32	0.27	0.13	—
ISM(U)	0.52	0.44	0.24	—
IBM-2	0.27	0.23	0.16	—
pnQRPA(U)	0.28	0.24	0.17	—
SRQRPA-A	0.31	0.26	0.23	—
QRPA-A	0.28	0.24	0.25	—
SkM-HFB-QRPA	0.29	0.24	0.28	—



## Inverted Ordering



Nature provides 2 scales:

$$|m_{ee}|_{\max}^{\text{IH}} \simeq c_{13}^2 \sqrt{\Delta m_A^2} \quad \text{and} \quad |m_{ee}|_{\min}^{\text{IH}} \simeq c_{13}^2 \sqrt{\Delta m_A^2} \cos 2\theta_{12}$$

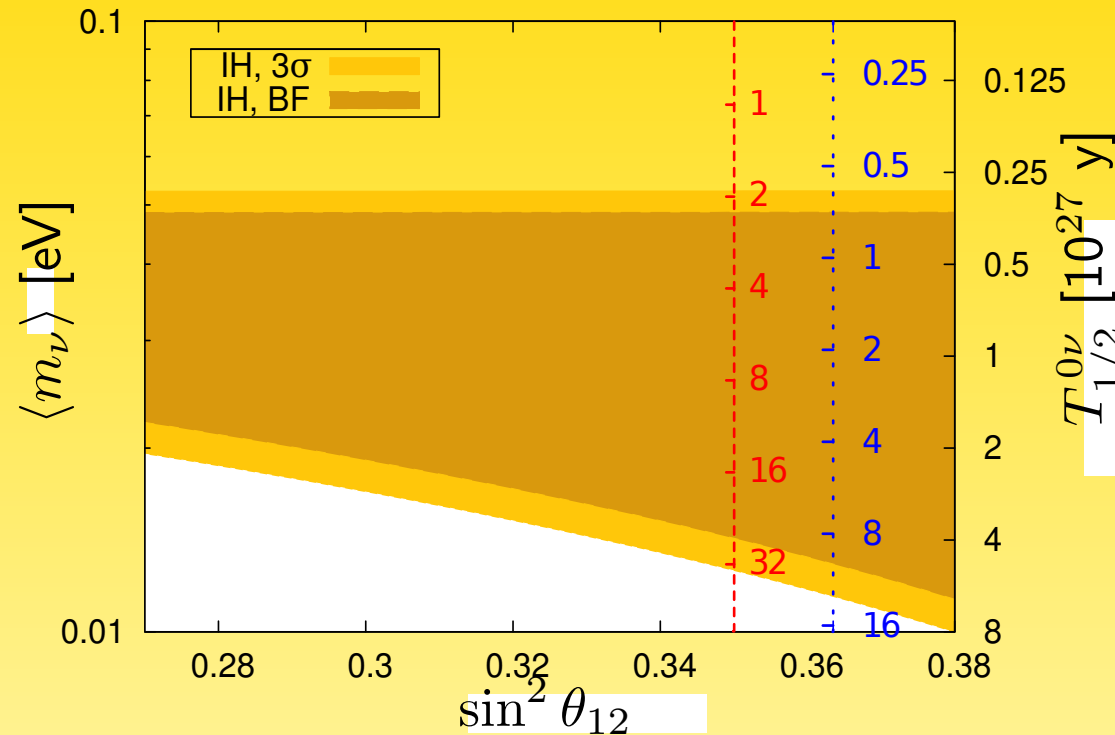
requires  $\mathcal{O}(10^{26} \dots 10^{27})$  yrs

is the lower limit  $|m_{ee}|_{\min}^{\text{IH}}$  fixed?



# Inverted Hierarchy

$m_3 = 0.001 \text{ eV}$

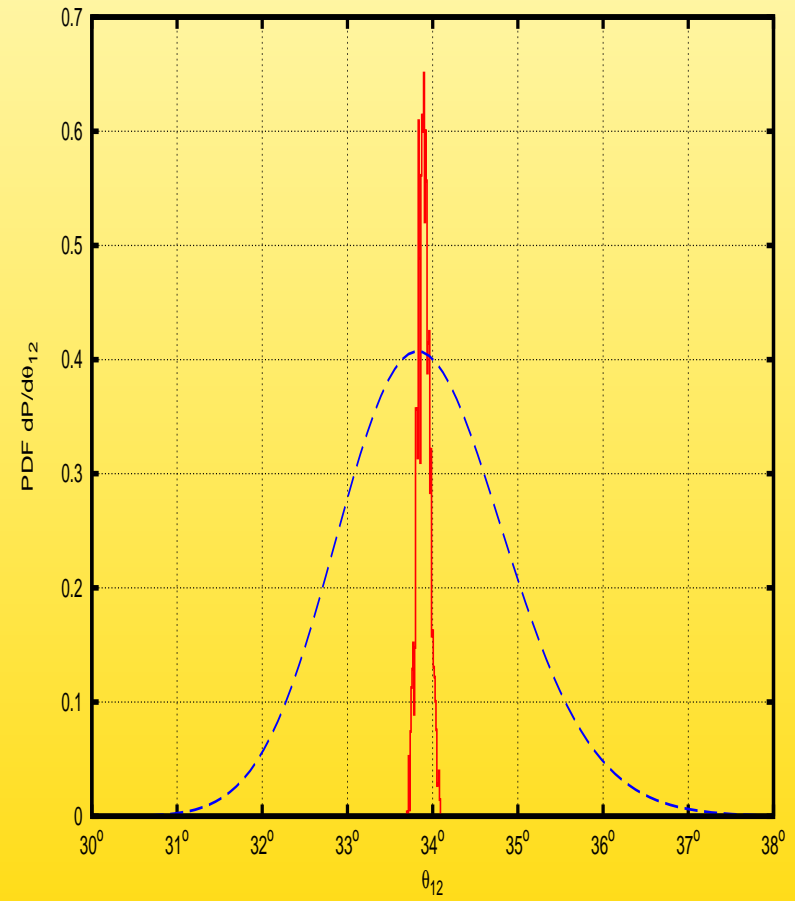
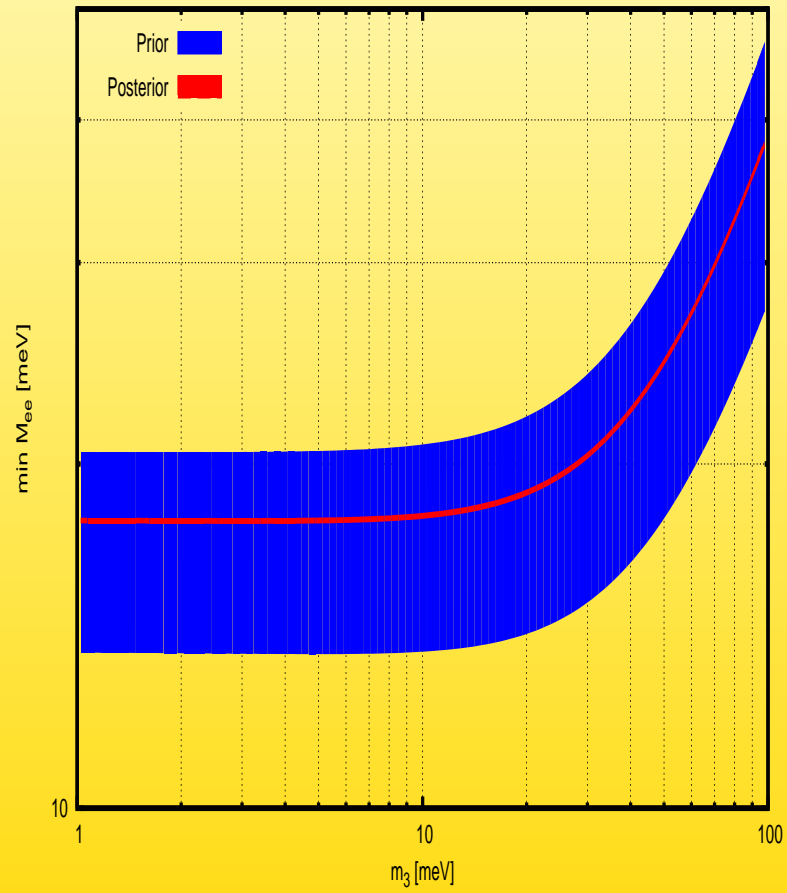


Current  $3\sigma$  range of  $\sin^2 \theta_{12}$  gives factor of  $\sim 2$  uncertainty for  $|m_{ee}|_{\min}^{\text{IH}}$

$\Rightarrow$  combined factor of  $\sim 16$  in  $M \times t \times B \times \Delta E$

$\Rightarrow$  need precision determination of  $\theta_{12}$ !  $\leftrightarrow$  JUNO

Ge, W.R., PRD **92**: <http://npdpro.hepforge.org>



## CP Violation?

Majorana phases: consider IH spectrum

$$|m_{ee}| \propto |\cos^2 \theta_{12} + e^{2i\alpha} \sin^2 \theta_{12}| = \sqrt{1 - \sin^2 2\theta_{12} \sin^2 \alpha}$$

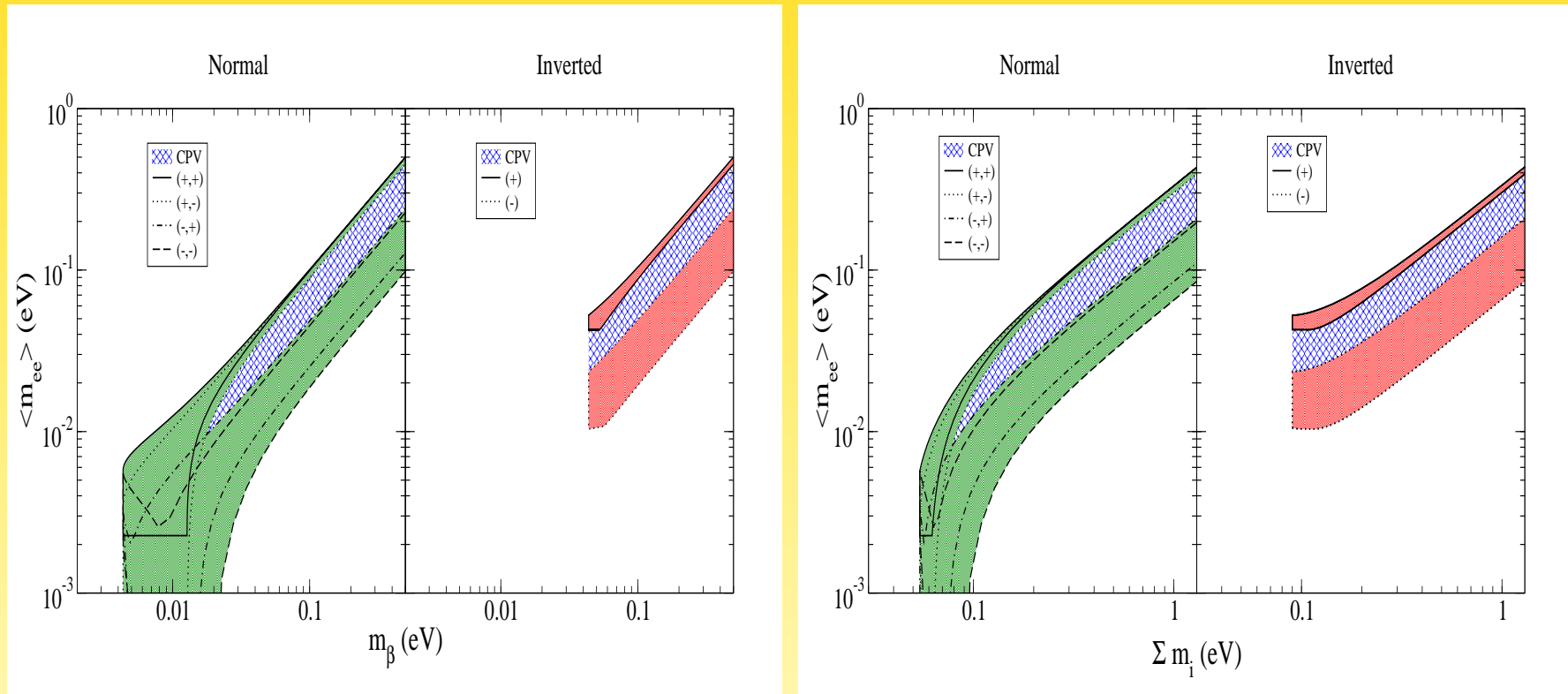
$\alpha$  can be probed if

- uncertainties on  $|m_{ee}|$  from NME smaller than 2
- $\sigma(|m_{ee}|) \lesssim 15\%$
- $\sigma(\Delta m_A^2) \lesssim 10\%$  (IH) or  $\sigma(m_0) \lesssim 10\%$  (QD)
- $\sin^2 \theta_{12} \gtrsim 0.29$
- $2\alpha \in [\pi/4, 3\pi/4]$  or  $[5\pi/4, 7\pi/4]$

Pascoli, Petcov, W.R., PLB **549**

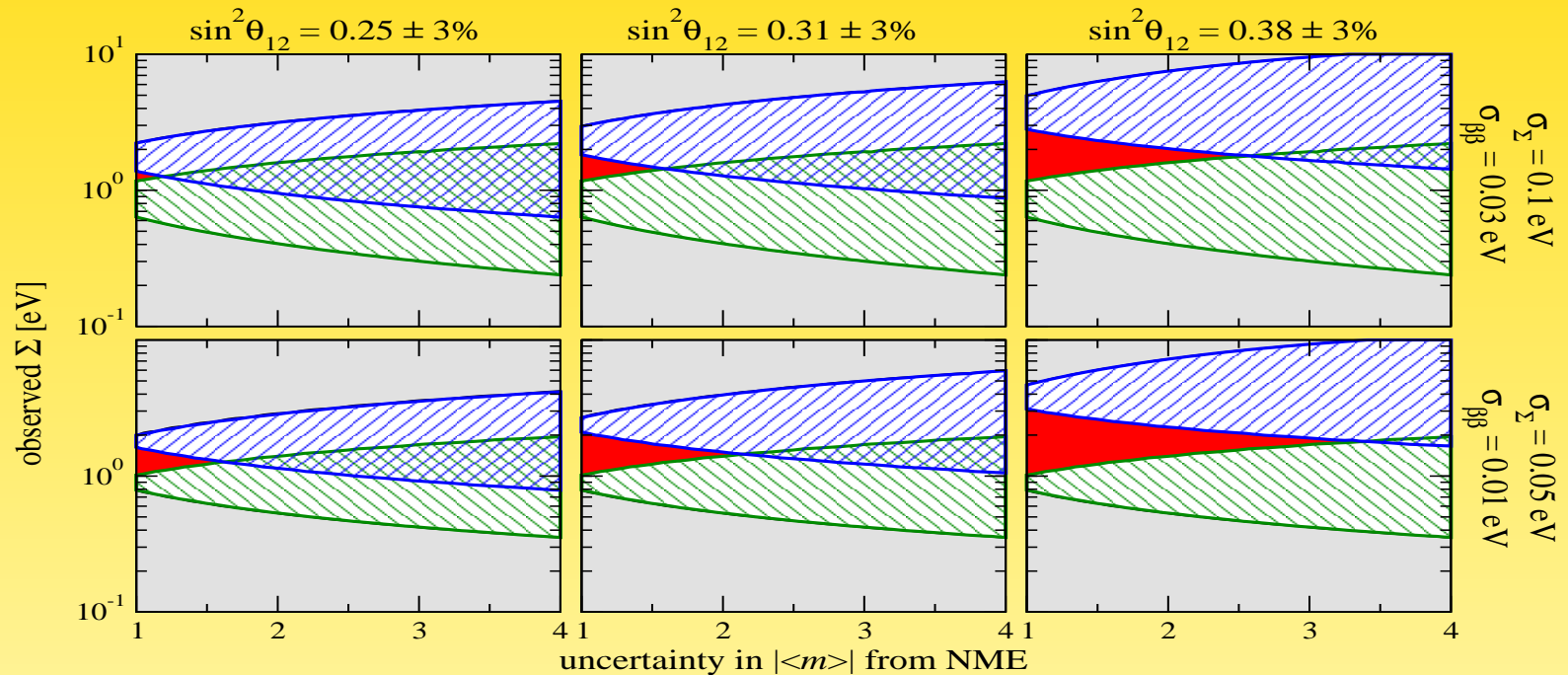
No to “no-go” from Barger *et al.*, PLB **540**

# Majorana phases



## Majorana phases

$$|m_{ee}| \propto |\cos^2 \theta_{12} + e^{2i\alpha} \sin^2 \theta_{12}| = \sqrt{1 - \sin^2 2\theta_{12} \sin^2 \alpha}$$



  data consistent with  $\alpha_{21} = \pi$

  data consistent with  $\alpha_{21} = 0$

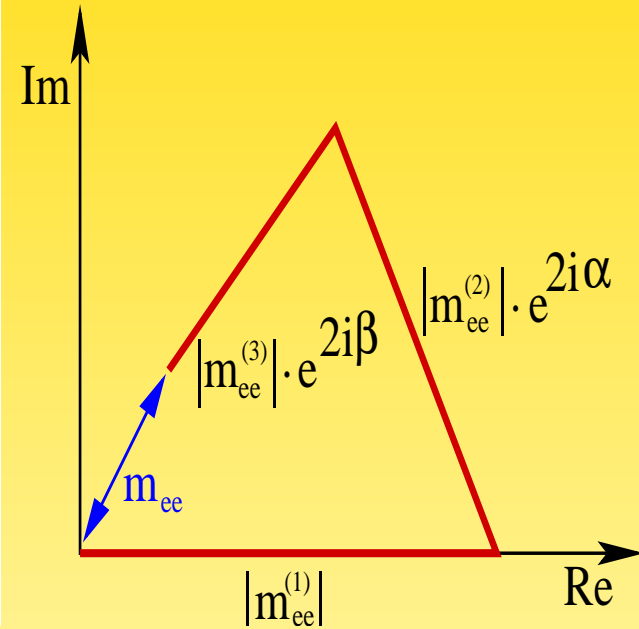
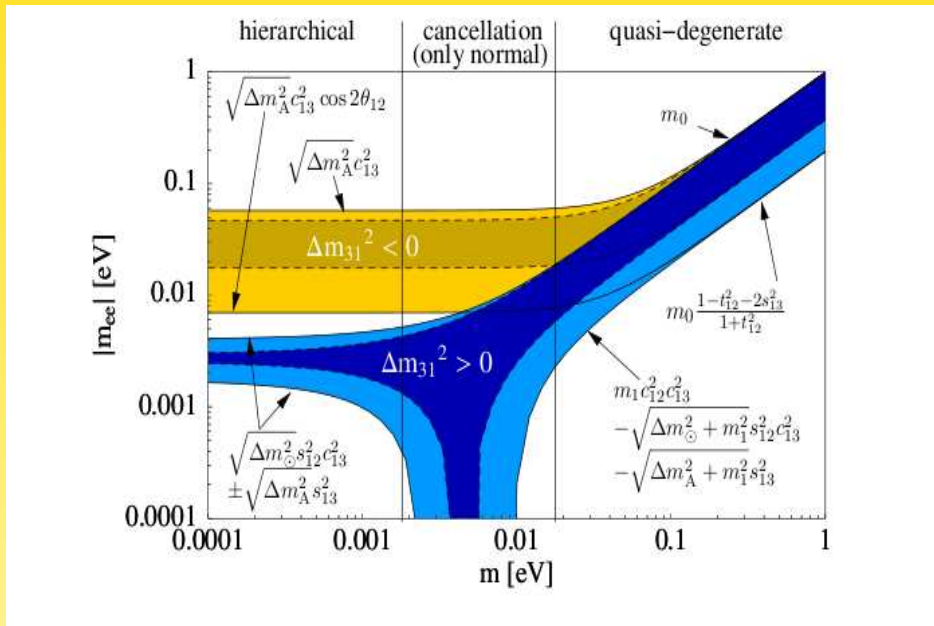
   $\langle m \rangle$  and  $\Sigma$  inconsistent at  $2\sigma$

  CP violation established at  $2\sigma$

$\sin^2 \theta_{13} = 0 \pm 0.002$ ,  $\Delta m_{21}^2 = 8 \times 10^{-5} \pm 2\%$ ,  $\Delta m_{31}^2 = 2.2 \times 10^{-3} \pm 3\%$       observed  $\langle m \rangle = 0.3 \text{ eV}$

Pascoli, Petcov, Schwetz, hep-ph/0505226

## Vanishing $|m_{ee}|$



only for NH  $\Rightarrow$  rule out possibility by ruling out NH

unnatural? texture zero!?!

## Vanishing $|m_{ee}|$

does it stay zero?

- seesaw RG effects
- NLO see-saw terms:

$$m_\nu = m_D^2/M_R + \mathcal{O}(m_D^4/M_R^3)$$

- actually:

$$\mathcal{A} \propto \frac{U_{ei}^2 m_i}{q^2 - m_i^2} \simeq \frac{|m_{ee}|}{q^2} + \mathcal{O}(m_i^3/q^4)$$

- Planck scale (Weinberg operator with Planck mass)

$$|m_{ee}| = \frac{v^2}{M_{\text{Pl}}} \simeq 10^{-6} \text{ eV}$$

## Renormalization

$$m_\nu = I_{\alpha_\nu} \begin{pmatrix} (m_\nu^0)_{ee} I_e^2 & (m_\nu^0)_{e\mu} I_e I_\mu & (m_\nu^0)_{e\tau} I_e I_\tau \\ \cdot & (m_\nu^0)_{\mu\mu} I_\mu^2 & (m_\nu^0)_{\mu\tau} I_\mu I_\tau \\ \cdot & \cdot & (m_\nu^0)_{\tau\tau} I_\tau^2 \end{pmatrix}$$

where

$$I_\alpha \simeq 1 + \frac{C}{16\pi^2} y_\alpha^2 \ln \frac{\lambda}{\Lambda} \quad \text{and} \quad I_{\alpha_\nu} \simeq 1 + \frac{1}{16\pi^2} \alpha_\nu \ln \frac{\lambda}{\Lambda}$$

and

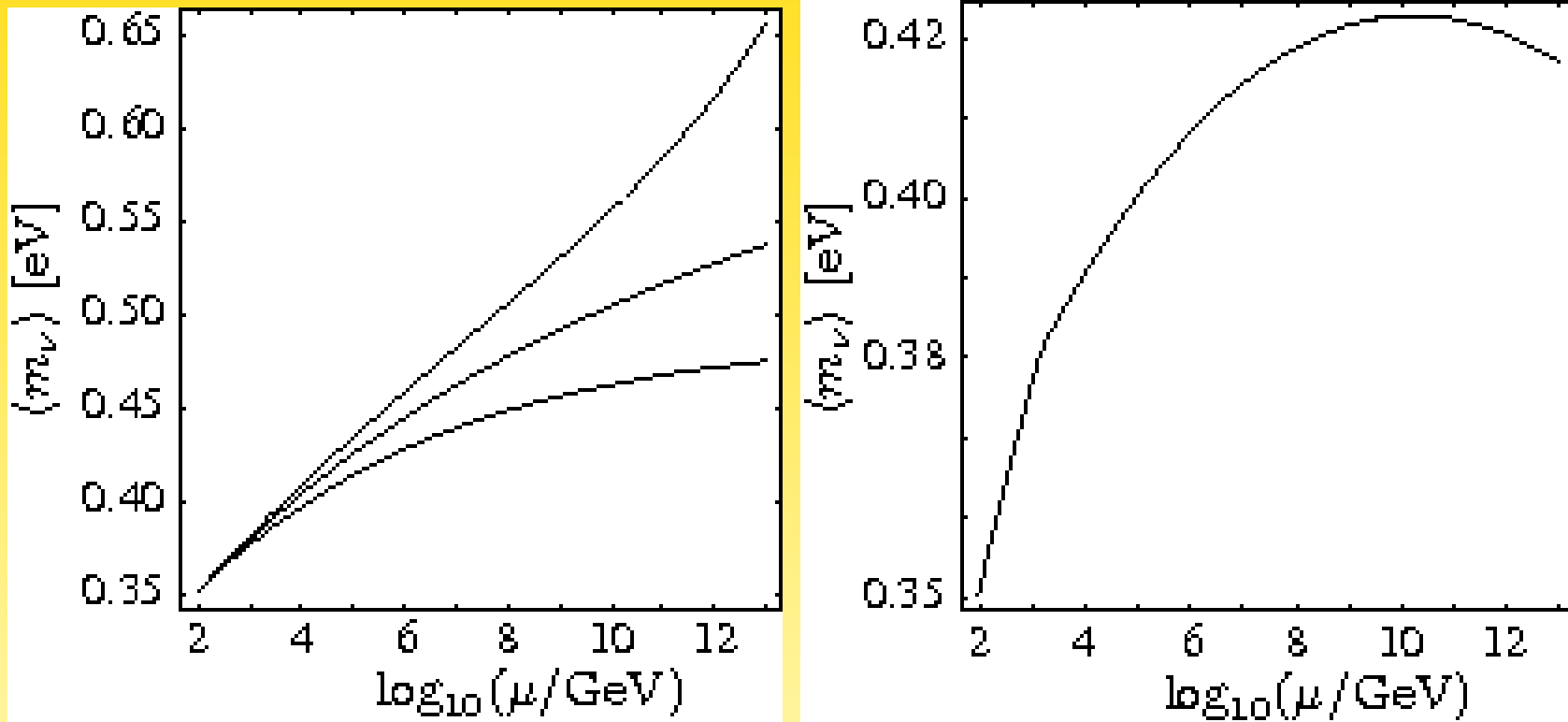
$$\alpha_\nu^{\text{SM}} = -3g_2^2 + 2(y_\tau^2 + y_\mu^2 + y_e^2) + 6(y_t^2 + y_b^2 + y_c^2 + y_s^2 + y_d^2 + y_u^2) + \lambda_H$$

$$\alpha_\nu^{\text{MSSM}} = -\frac{6}{5}g_1^2 - 6g_2^2 + 6(y_t^2 + y_c^2 + y_u^2)$$

$\Rightarrow$  main effect: rescaling of  $|m_{ee}|$ , typically increases from low to high scale

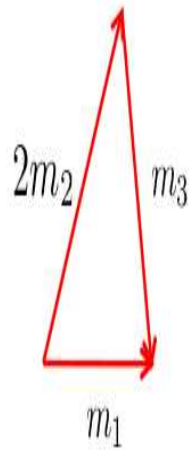


## Renormalization

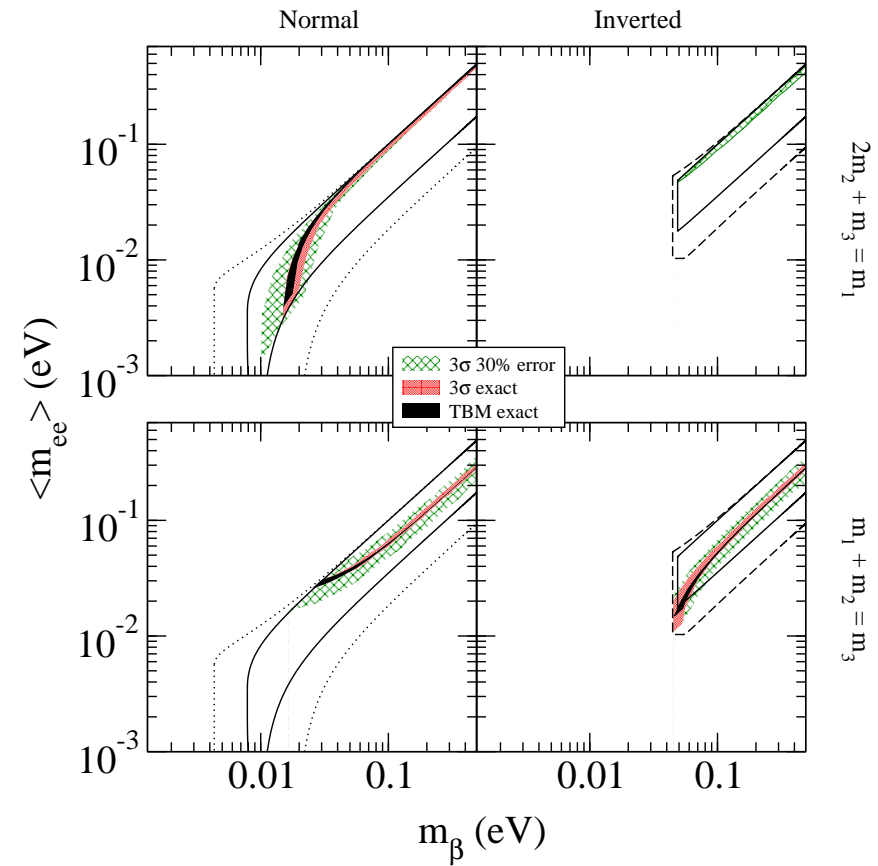


Antusch *et al.*, hep-ph/0305273

# Flavor Symmetry Models: sum-rules



Sum-rule	Flavour symmetry
$2m_2 + m_3 = m_1$	$A_4, T', (S_4)$
$m_1 + m_2 = m_3$	$S_4, (A_4)$
$\frac{2}{m_2} + \frac{1}{m_3} = \frac{1}{m_1}$	$A_4, T'$
$\frac{1}{m_1} + \frac{1}{m_2} = \frac{1}{m_3}$	$S_4$



constraints on masses and Majorana phases

Barry, W.R., Nucl. Phys. **B842**

## Predictions of $SO(10)$ theories

Yukawa structure of  $SO(10)$  models depends on Higgs representations

$$10_H (\leftrightarrow H), \overline{126}_H (\leftrightarrow F), 120_H (\leftrightarrow G)$$

Gives relation for mass matrices:

$$m_{\text{up}} \propto r(H + sF + it_u G)$$

$$m_{\text{down}} \propto H + F + iG$$

$$m_D \propto r(H - 3sF + it_D G)$$

$$m_\ell \propto H - 3F + it_l G$$

$$M_R \propto r_R^{-1} F$$

Numerical fit including RG, Higgs,  $\theta_{13}$

$$10_H + \overline{126}_H: 19 \text{ free parameters}$$

$$10_H + \overline{126}_H + 120_H: 18 \text{ free parameters}$$

$$20 \text{ (19) observables to be fitted}$$

## Predictions of $SO(10)$ theories

Model	Fit	$ m_{ee} $ [meV]	$m_0$ [meV]	$M_3$ [GeV]	$\chi^2$
$10_H + \overline{126}_H$	NH	0.49	2.40	$3.6 \times 10^{12}$	23.0
$10_H + \overline{126}_H + SS$	NH	0.44	6.83	$1.1 \times 10^{12}$	3.29
$10_H + \overline{126}_H + 120_H$	NH	2.87	1.54	$9.9 \times 10^{14}$	11.2
$10_H + \overline{126}_H + 120_H + SS$	NH	0.78	3.17	$4.2 \times 10^{13}$	$6.9 \times 10^{-6}$
$10_H + \overline{126}_H + 120_H$	IH	35.52	30.2	$1.1 \times 10^{13}$	13.3
$10_H + \overline{126}_H + 120_H + SS$	IH	24.22	12.0	$1.2 \times 10^{13}$	0.6

Dueck, W.R., JHEP **1309**

## Contents

### **I) Dirac vs. Majorana neutrinos**

- I1) Basics
- I2) Seesaw mechanisms
- I3) Summary neutrino physics

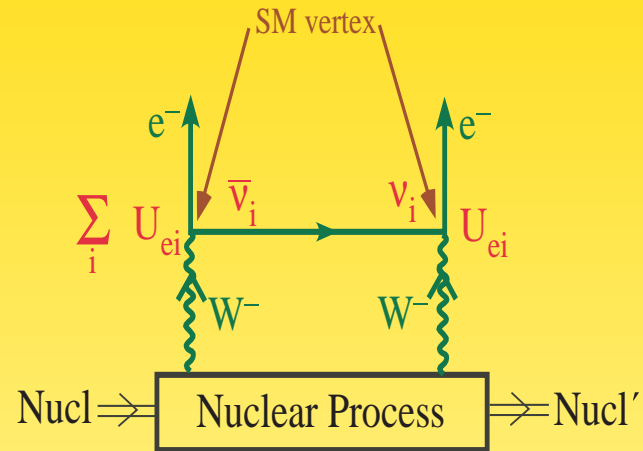
### **II) Neutrinoless double beta decay: standard interpretation**

- II1) Basics
- II2) Implications for and from neutrino physics
- II3) Nuclear physics aspects
- II4) Exotic physics

## II3) Nuclear physics aspects



## Nuclear physics



- 2 point-like Fermi vertices
- “long-range” neutrino exchange
- momentum exchange  $q \simeq 1/r \simeq 0.1$  GeV
- wave functions of nucleons not exactly known, depend on nuclear model

typical model for NME: set of single particle states with a number of possible wave function configurations; obtained by solving Dirac equation in a mean background field; interaction known, treatment of fields differs

- Quasi-particle Random Phase Approximation (QRPA) (many single particle states, few configurations)
- Nuclear Shell Model (NSM) (many configurations, few single particle states)
- Interacting Boson Model (IBM) (many single particle states, few configurations)
- Generating Coordinate Method (GCM) (many single particle states, few configurations)
- projected Hartree-Fock-Bogoliubov model (pHFB)

tends to overestimate NMEs; tends to underestimate NMEs

gives uncertainty due to different approach



plus uncertainty due to model details:

- short range correlations ( $\leftrightarrow$  repulsive  $NN$  force): multiply two-body wave function with
  - Jastrow function
  - Unitary Correlation Operator method
  - Coupled Cluster method
  - (if heavy particles exchanged: NMEs suppressed)
- model parameters
  - correlated: e.g.,  $g_A$ ,  $g_{pp}$ , SRC
  - uncorrelated: e.g., model space of single particle states
  - some calculations include errors/ranges, some don't. . .

$$\mathcal{M}^{0\nu} = \left( \frac{g_A}{1.25} \right)^2 \left( \mathcal{M}_{\text{GT}}^{0\nu} - \frac{g_V^2}{g_A^2} \mathcal{M}_{\text{F}}^{0\nu} \right)$$

with Gamov-Teller (axial) and Fermi (vector) parts

$$\mathcal{M}_{\text{GT}}^{0\nu} = \langle f | \sum_{lk} \sigma_l \sigma_k \tau_l^- \tau_k^- H(r_{lk}, E_a) | i \rangle$$

$$\mathcal{M}_{\text{F}}^{0\nu} = \langle f | \sum_{lk} \tau_l^- \tau_k^- H(r_{lk}, E_a) | i \rangle$$

- $r_{lk} \simeq 1/p \simeq 1/(0.1 \text{ GeV})$  distance between the two decaying neutrons
- $E_a$  average energy (closure approx.);  $|i\rangle$  and  $|f\rangle$  nucleon wave functions
- $J_\alpha = \sum \delta(\vec{x} - \vec{r}_n) \tau_+^n (g_{\alpha 0} J_n^0 + g_{\alpha k} J_n^k)$  with  $J_n^0 = g_V$  and  $J_n^k = g_A J_n^k \vec{\sigma}_n$
- 'neutrino potential' integrates over the virtual neutrino momenta

$$H(r, E_a) = \frac{2R}{\pi r} \int_0^\infty dq \frac{\sin qx}{q + E_a - (M_i + M_f)/2}$$

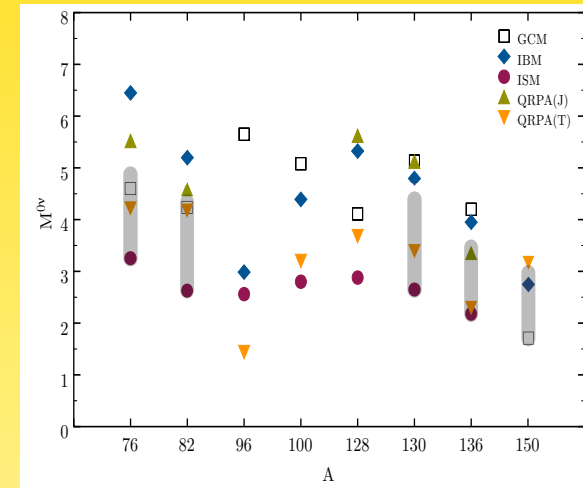
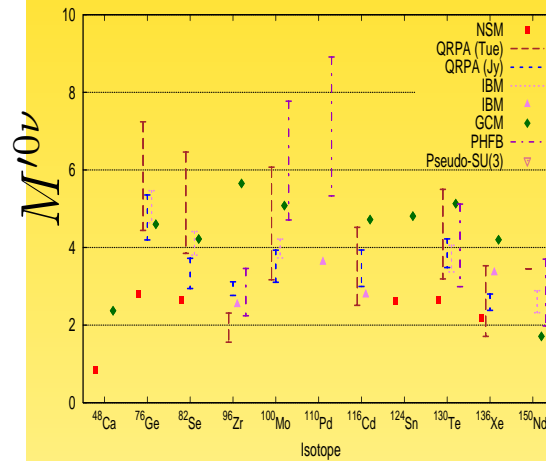
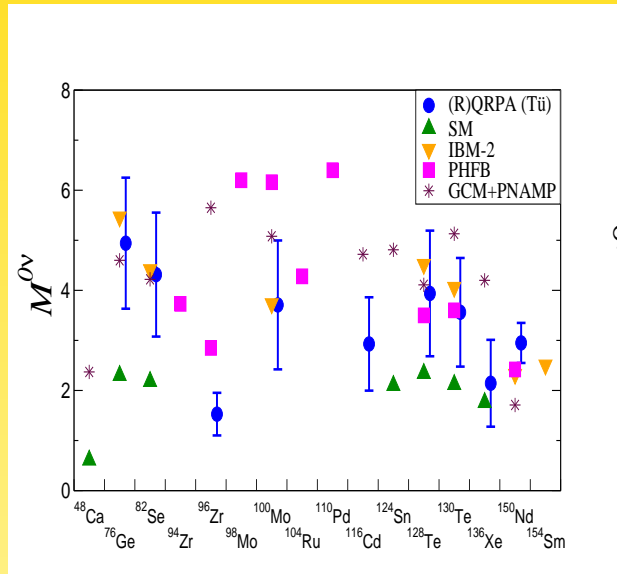
The  $2\nu\beta\beta$  matrix elements can be written as

$$\mathcal{M}_{\text{GT}}^{2\nu} = \sum_n \frac{\langle f | \sum_a \sigma_a \tau_a^- | n \rangle \langle n | \sum_b \sigma_b \tau_b^- | i \rangle}{E_n - (M_i - M_f)/2}$$

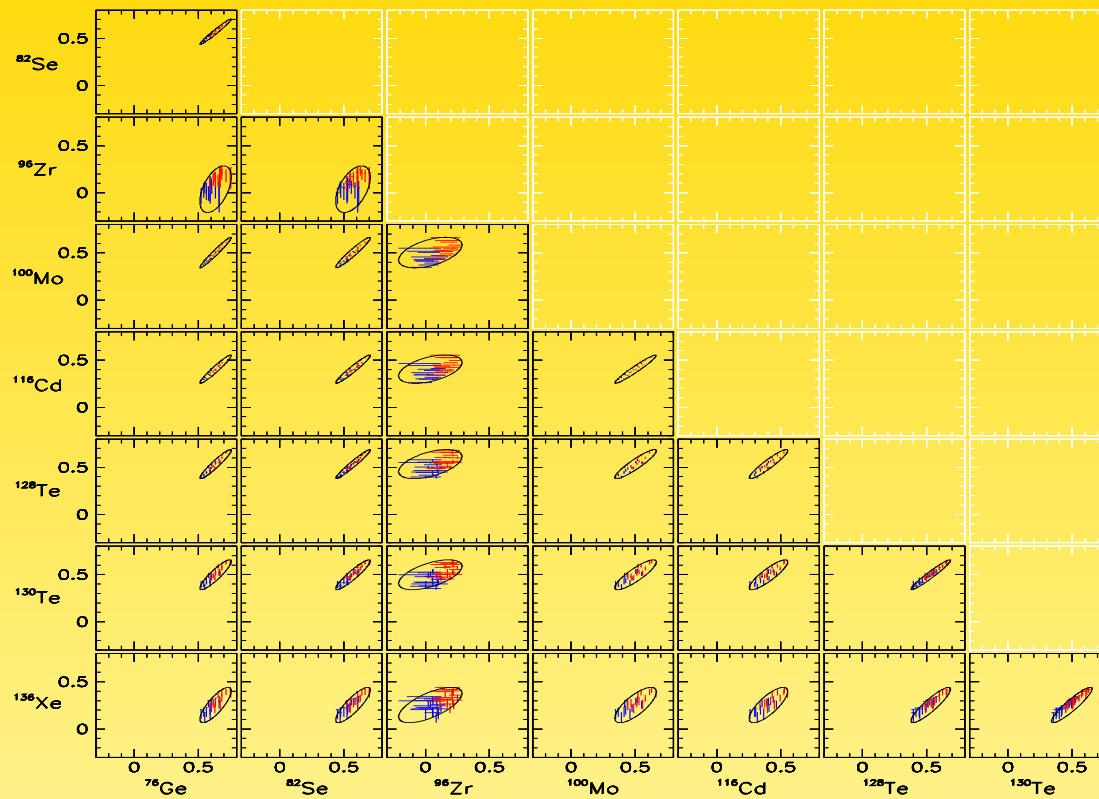
$$\mathcal{M}_{\text{F}}^{2\nu} = \sum_n \frac{\langle f | \sum_a \tau_a^- | n \rangle \langle n | \sum_b \tau_b^- | i \rangle}{E_n - (M_i - M_f)/2}$$

- no direct connection to  $0\nu\beta\beta$ ...
- sum over  $1^+$  states (low momentum transfer); in  $0\nu\beta\beta$  all multipolarities contribute
- adjust some parameters to reproduce  $2\nu\beta\beta$ -rates
- $2\nu\beta\beta$  observed in  $^{48}\text{Ca}$ ,  $^{76}\text{Ge}$ ,  $^{82}\text{Se}$ ,  $^{96}\text{Zr}$ ,  $^{100}\text{Mo}$  (plus exc. state),  $^{116}\text{Cd}$ ,  $^{128}\text{Te}$ ,  $^{130}\text{Te}$ ,  $^{150}\text{Nd}$  (plus exc. state) with half-lives from  $10^{18}$  to  $10^{24}$  yrs

## From life-time to particle physics: Nuclear Matrix Elements



to better estimate error range: correlations need to be understood



Faessler, Fogli *et al.*, PRD 79

ellipse major axis: SRC (blue, red) and  $g_A$

ellipse minor axis:  $g_{pp}$

## Quenching

$T_{1/2}^{0\nu} \propto g_A^{-4}$ , where in  $2\nu\beta\beta$ -decay (Iachello)

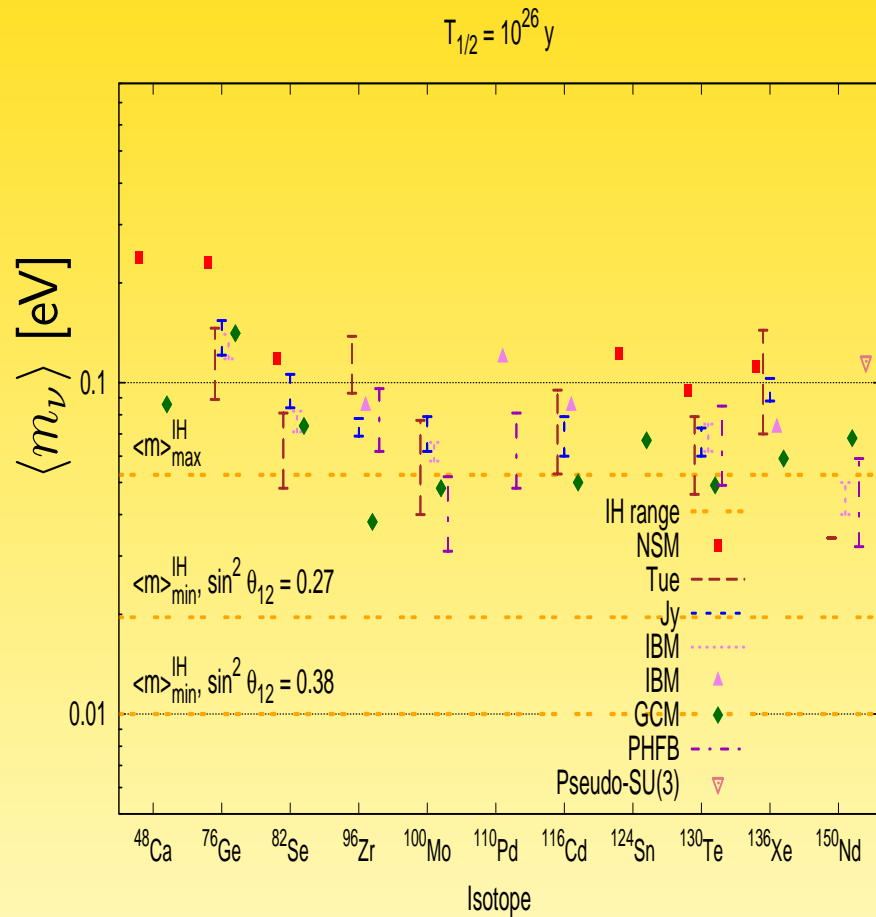
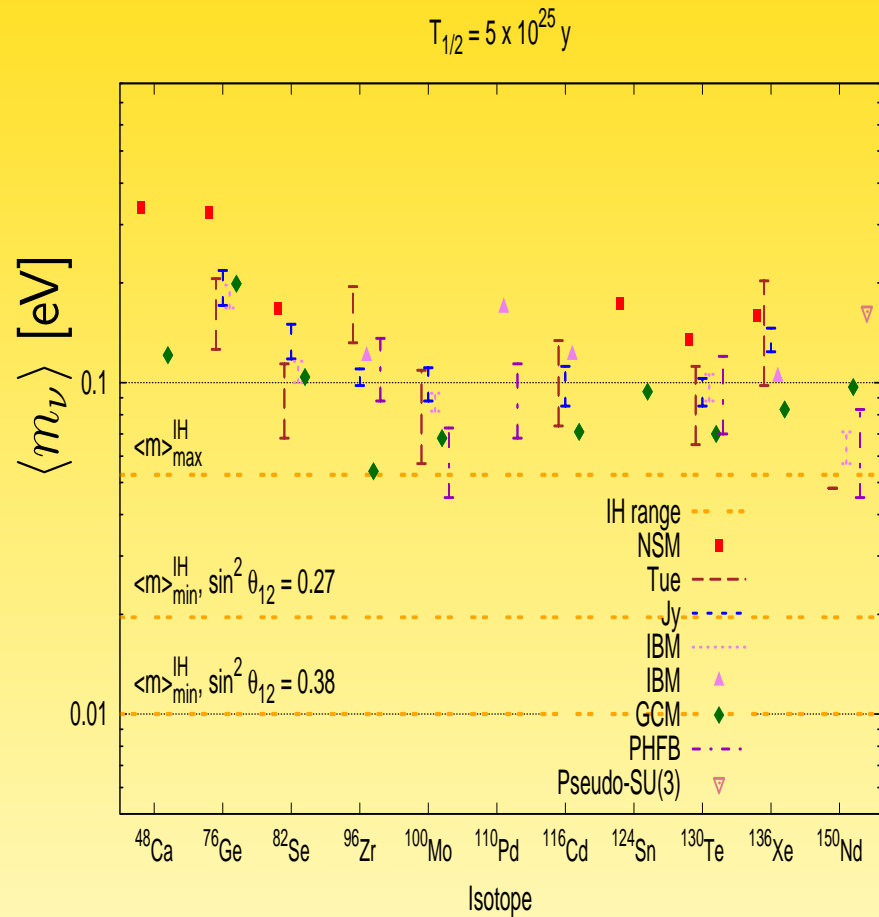
$$g_A^{\text{IBM}} \simeq 1.27 A^{-0.18} = \begin{cases} 0.58 & \text{Ge} \\ 0.53 & \text{Te} \\ 0.52 & \text{Xe} \end{cases}$$

would shift lifetimes an order of magnitude (in wrong direction...)

- model-space truncation ( $A$ -dep.)/omission of  $N^*$ ,  $\Delta$  (not  $A$ -dep.)
- energy scale (no quenching for muon capture)/many multipolarities in  $0\nu$
- include 2 body currents (creation of 2p2h), weaker quenching in  $0\nu$  (Engel, Simkovic, Vogel, PRC89)

lots of nuclear theory work on the way, possibly no showstopper (20%?)

# $0\nu\beta\beta$ and Neutrino physics



## Testing NMEs

$$\frac{1}{T_{1/2}^{0\nu}(A, Z)} = G(Q, Z) |\mathcal{M}(A, Z) \eta|^2$$

if you measured two isotopes:

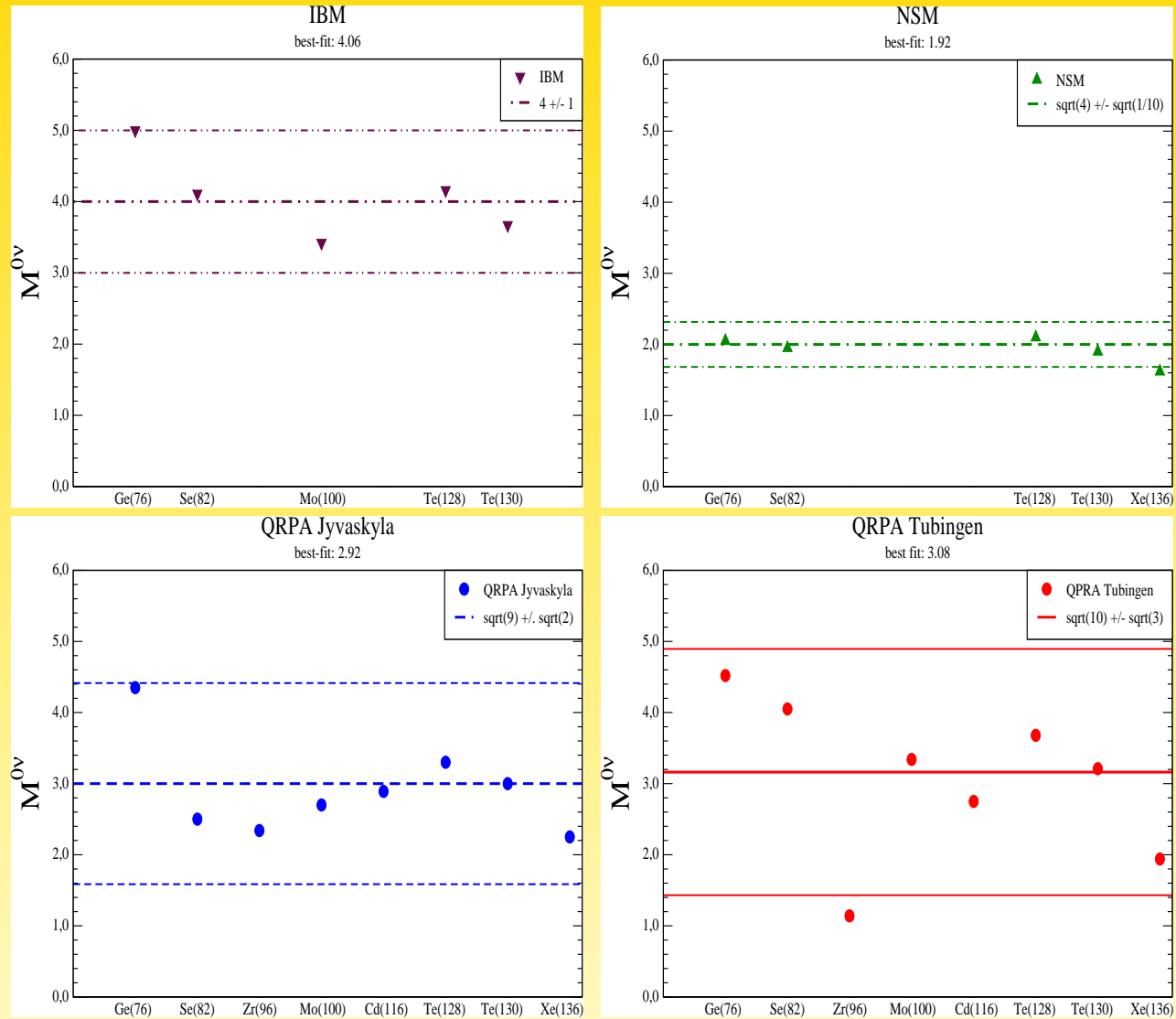
$$\frac{T_{1/2}^{0\nu}(A_1, Z_1)}{T_{1/2}^{0\nu}(A_2, Z_2)} = \frac{G(Q_2, Z_2) |\mathcal{M}(A_2, Z_2)|^2}{G(Q_1, Z_1) |\mathcal{M}(A_1, Z_1)|^2}$$

systematic errors drop out, ratio sensitive to NME model

⇒ **second reason for multi-isotope determination**



# NMEs are order one numbers :-)



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## II4) Exotic physics

- Dirac neutrinos
- Pseudo-Dirac neutrinos: for each mass state

$$m_i \begin{pmatrix} \epsilon & 1 \\ 1 & 0 \end{pmatrix} \rightarrow U = \sqrt{\frac{1}{2}} \begin{pmatrix} 1 + \frac{\epsilon}{4} & -1 + \frac{\epsilon}{4} \\ 1 - \frac{\epsilon}{4} & 1 + \frac{\epsilon}{4} \end{pmatrix} \text{ and } m_i^\pm = m_i \left( \pm 1 + \frac{\epsilon}{2} \right)$$

and  $|m_{ee}|^{(i)} = \epsilon m_i = \frac{1}{2} \delta m^2 / m_i$ , with  $\delta m^2 = (m_i^+)^2 - (m_i^-)^2$

- one neutrino could be (Pseudo-)Dirac, the other Majorana!

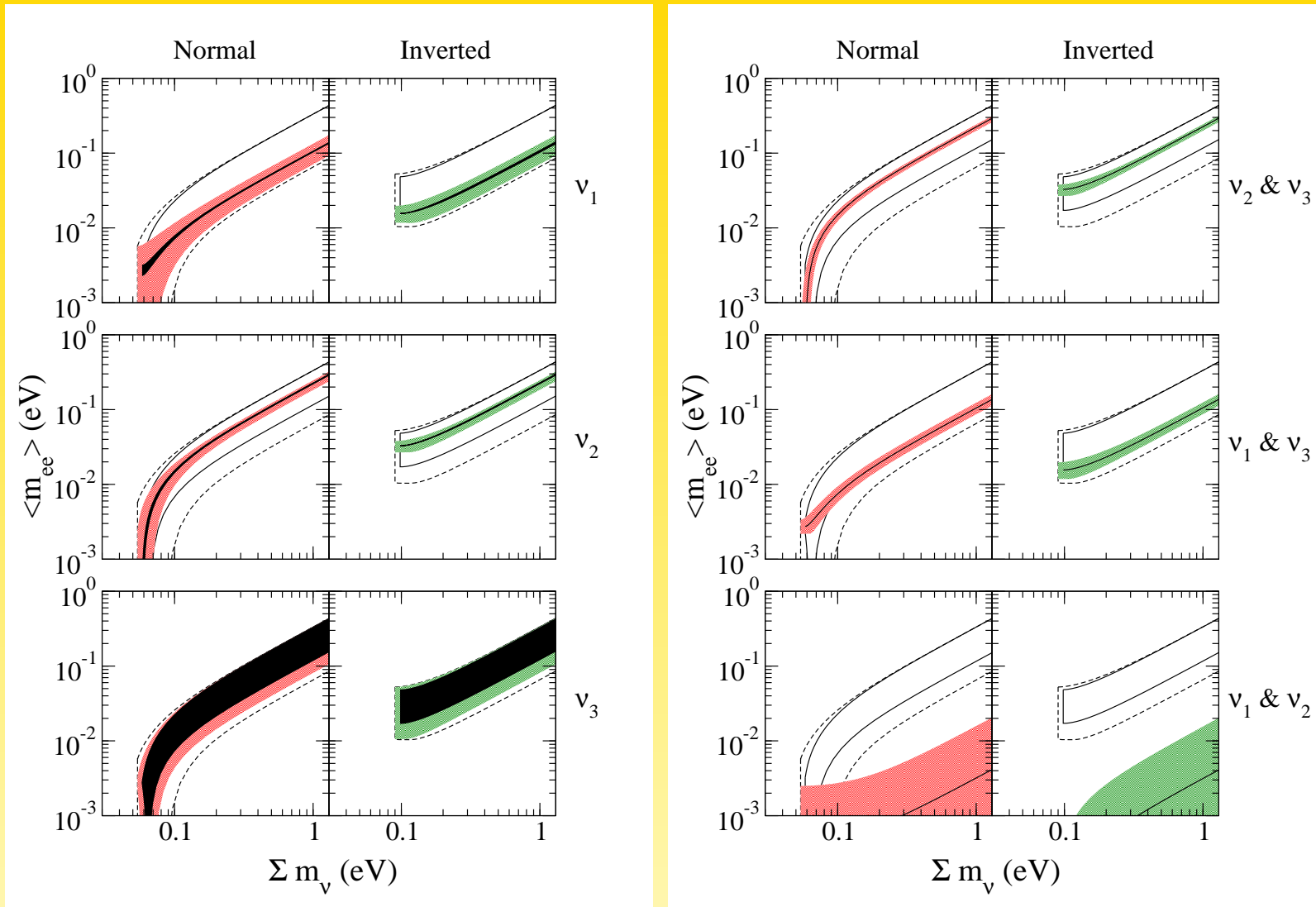
“Schizophrenic neutrinos”

For instance: inverted hierarchy,  $m_2$  Dirac:

$$|m_{ee}| = c_{12}^2 c_{13}^2 \sqrt{\Delta m_A^2}$$

factor 2 larger than lower limit

Mohapatra *et al.*, 1008.1232



Barry, Mohapatra, W.R., 1012.1761

## Lorentz invariance violation

Lorentz and CPT violating term for Majorana neutrinos (Kostelecky *et al.*)

$$\mathcal{L} \propto \nu_e^T g_{ee}^{\lambda\mu\nu} \sigma_{\lambda\mu} q_\nu \nu_e$$

gives contribution to propagator

$$S(q) = \frac{\not{q} + g_{ee}^{\lambda\mu\nu} \sigma_{\lambda\mu} q_\nu}{q^2}$$

and leads to  $0\nu\beta\beta$  even for massless neutrinos

can identify

$$|m_{ee}| \leftrightarrow \frac{\frac{1}{3} (|\langle g_{ee}^{\lambda\mu 1} \rangle|^2 + |\langle g_{ee}^{\lambda\mu 2} \rangle|^2 + |\langle g_{ee}^{\lambda\mu 3} \rangle|^2)}{R}$$

and set limits of order  $10^{-8}$  (Diaz, 1311.0930)

## CPT Violation

MINOS, MiniBooNE, solar/KamLAND...

what about Dirac/Majorana??

Barenboim, Beacom, Borisso, Kayser, hep-ph/0203261

Define  $\xi = \text{CPT}$ . A Majorana particle is defined as

$$\xi|\nu_i\rangle = e^{i\zeta_i} |\nu_i\rangle$$

Now we introduce CPT-violation

$$M_\nu = \begin{pmatrix} \mu + \Delta & y^* \\ y & \mu - \Delta \end{pmatrix} \text{ for basis } \nu, \bar{\nu}$$

$y$  mixes  $\nu$  and  $\bar{\nu}$  and hence violates  $L$

diagonalize  $M_\nu$  to get eigenstates

$$|\nu_+\rangle = \cos\theta |\nu\rangle + e^{i\phi} \sin\theta |\bar{\nu}\rangle, \quad m_+ = \mu + \sqrt{|y^2| + \Delta^2}$$

$$|\nu_-\rangle = -\sin\theta |\nu\rangle + e^{i\phi} \cos\theta |\bar{\nu}\rangle, \quad m_- = \mu - \sqrt{|y^2| + \Delta^2}$$

$$\text{mixing angle } \tan 2\theta = \frac{|y|}{\Delta}$$

CPT properties are  $\xi|\nu\rangle = e^{i\zeta} |\bar{\nu}\rangle$  and  $\xi|\bar{\nu}\rangle = e^{i\zeta} |\nu\rangle$ , hence

$$\xi|\nu_+\rangle = e^{i(\zeta-\phi)} (\sin\theta |\nu\rangle + e^{i\phi} \cos\theta |\bar{\nu}\rangle),$$

$$\xi|\nu_-\rangle = -e^{i(\zeta-\phi)} (-\cos\theta |\nu\rangle + e^{i\phi} \sin\theta |\bar{\nu}\rangle)$$

$\Rightarrow \nu_{\pm}$  are Majorana particles if and only if  $\theta = \pi/4$

but  $\theta = \pi/4$  means  $\Delta = 0$  and thus CPT conservation:-)

if  $\Delta \neq 0$ : CPT is violated: neutrinos are no longer Majorana fermions; if  $y \neq 0$   
then  $L$  is violated and  $0\nu\beta\beta$  can occur

observation of  $0\nu\beta\beta$  implies non-zero  $y$  but not that neutrinos are CPT  
self-conjugate

Neutrinoless double beta decay?

mass eigenstates are for 3 generations

$$|\nu_i\rangle = U_{\alpha i}|\nu_\alpha\rangle + \bar{U}_{\alpha i}^*|\bar{\nu}_\alpha\rangle$$

amplitude for  $0\nu\beta\beta$  is

$$\mathcal{A} \propto \sum m_i U_{ei} \bar{U}_{ei}$$

for 1-flavor case:

$$\begin{aligned} \mathcal{A} &\propto m_+ U_{+\nu} U_{+\bar{\nu}} + m_- U_{-\nu} U_{-\bar{\nu}} \\ &= m_+ \cos\theta e^{i\phi} \sin\theta + m_- (-\sin\theta) e^{i\phi} \cos\theta \\ &\propto (m_+ - m_-) \cos\theta \sin\theta = y \end{aligned}$$

as expected

CPT counter part of  $0\nu\beta\beta$  gives same result



## Light Sterile Neutrinos

↔ is there a **sterile** neutral lepton with  $\Delta m^2 \sim \text{eV}^2$  and mixing  $\mathcal{O}(0.1)$ ?

- LSND ( $\bar{\nu}_\mu \rightarrow \bar{\nu}_e$ )  $3.8\sigma$
- MiniBooNE ( $\bar{\nu}_\mu \rightarrow \bar{\nu}_e$  and  $\nu_\mu \rightarrow \nu_e$ )  $3.8\sigma$
- Gallium anomaly ( $\nu_e \rightarrow \nu_e$ )  $2.9\sigma$
- Reactor anomaly ( $\bar{\nu}_e \rightarrow \bar{\nu}_e$ )  $2.8\sigma$
- Cosmology and Astroparticle Physics ( $N_{\text{eff}}$ , “Dark Radiation”)

	$\Delta m_{41}^2 [\text{eV}^2]$	$ U_{e4} $	$ U_{\mu 4} $	$\Delta m_{51}^2 [\text{eV}^2]$	$ U_{e5} $	$ U_{\mu 5} $
3+2/2+3	0.47	0.128	0.165	0.87	0.138	0.148
1+3+1	0.47	0.129	0.154	0.87	0.142	0.163

or  $\Delta m_{41}^2 = 1.78 \text{ eV}^2$  and  $|U_{e4}|^2 = 0.151$

Kopp, Maltoni, Schwetz, 1103.4570

## Sterile Neutrinos and $0\nu\beta\beta$

- recall:  $|m_{ee}|_{\text{NH}}^{\text{act}}$  can vanish and  $|m_{ee}|_{\text{IH}}^{\text{act}} \sim 0.03$  eV cannot vanish
- $|m_{ee}| = \left| \underbrace{|U_{e1}|^2 m_1 + |U_{e2}|^2 m_2 e^{2i\alpha} + |U_{e3}|^2 m_3 e^{2i\beta}}_{m_{ee}^{\text{act}}} + \underbrace{|U_{e4}|^2 m_4 e^{2i\Phi_1}}_{m_{ee}^{\text{st}}} \right|$
- sterile contribution to  $0\nu\beta\beta$  (assuming 1+3):

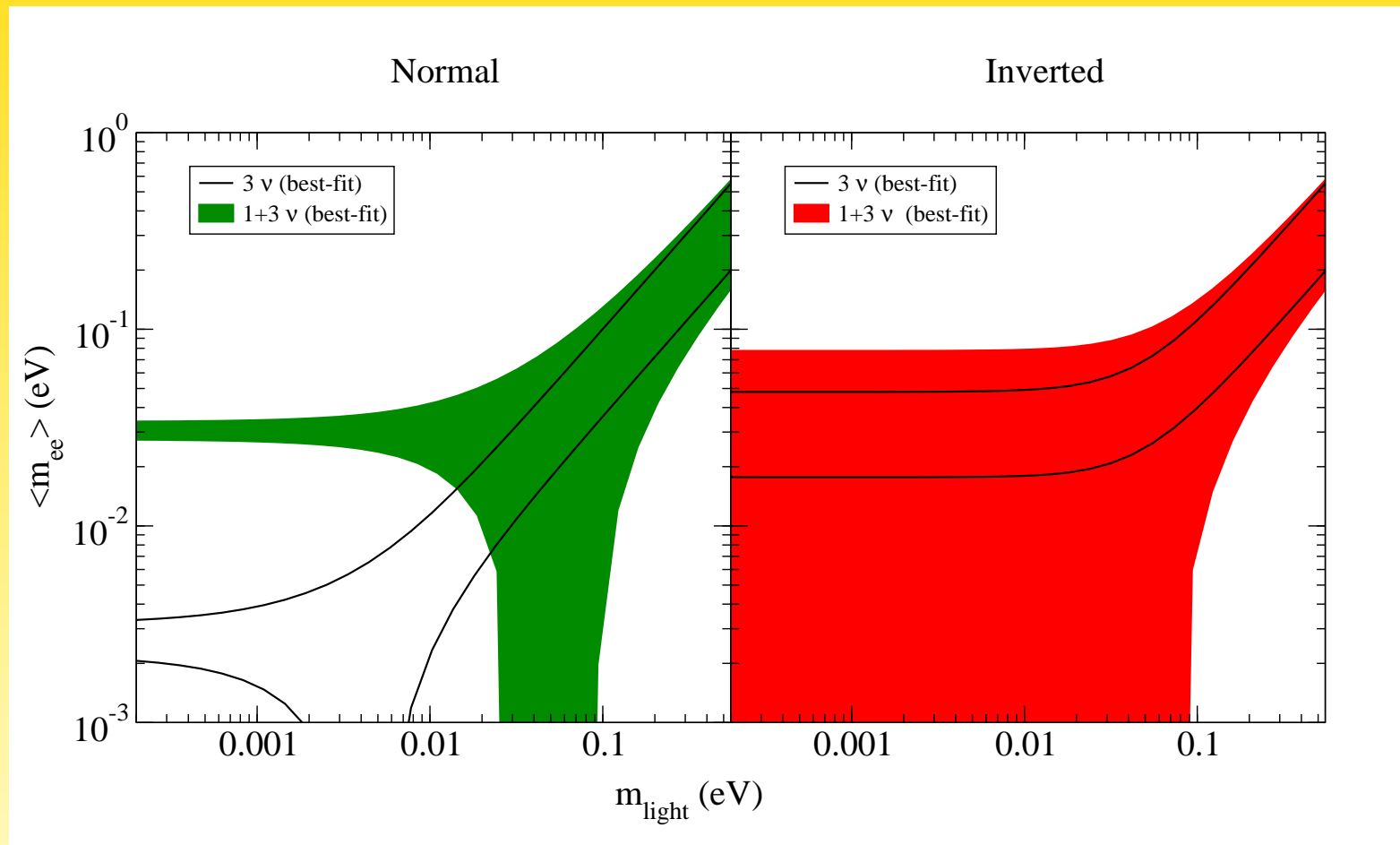
$$|m_{ee}|^{\text{st}} \simeq \sqrt{\Delta m_{\text{st}}^2} |U_{e4}|^2 \begin{cases} \gg |m_{ee}|_{\text{NH}}^{\text{act}} \\ \simeq |m_{ee}|_{\text{IH}}^{\text{act}} \end{cases}$$

$\Rightarrow |m_{ee}|_{\text{NH}}$  cannot vanish and  $|m_{ee}|_{\text{IH}}$  can vanish!

$\Rightarrow$  usual phenomenology gets completely turned around!

Barry, W.R., Zhang, JHEP **1107**; Giunti *et al.*, PRD **87**; Girardi, Meroni, Petcov, JHEP **1311**; Giunti, Zavanin, 1505.00978

Usual plot gets completely turned around!



## Do Dirac neutrinos mean there is no Lepton Number Violation?

model based on gauged  $B - L$ , broken by 4 units

$\Rightarrow$  Neutrinos are Dirac particles,  $\Delta L = 2$  forbidden, but  $\Delta L = 4$  allowed...

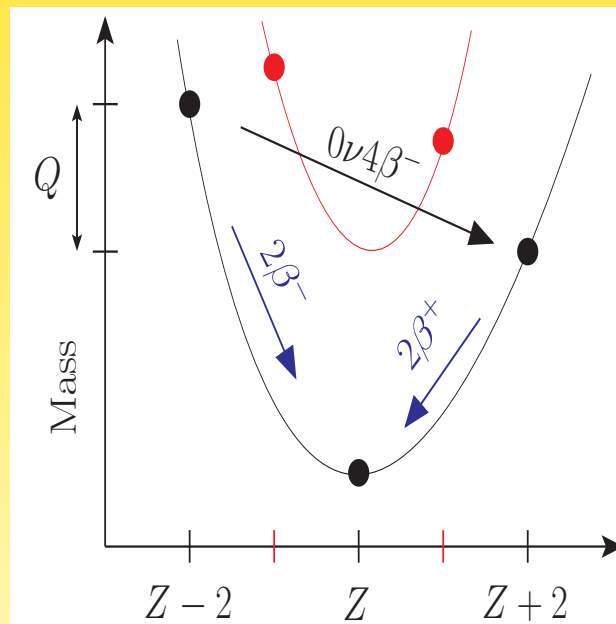
Heeck, W.R., EPL **103**

## Do Dirac neutrinos mean there is no Lepton Number Violation?

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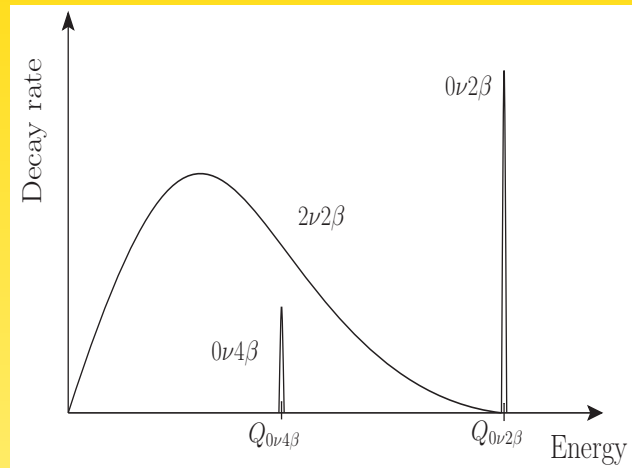
⇒ Neutrinos are Dirac particles,  $\Delta L = 2$  forbidden, but  $\Delta L = 4$  allowed...

⇒ observable: neutrinoless quadruple beta decay  $(A, Z) \rightarrow (A, Z + 4) + 4e^-$



Heeck, W.R., EPL 103

## Candidates for neutrinoless quadruple beta decay



	$Q_{0\nu 4\beta}$	Other decays	NA
${}_{40}^{96}\text{Zr} \rightarrow {}_{44}^{96}\text{Ru}$	0.629	$\tau_{1/2}^{2\nu 2\beta} \simeq 2 \times 10^{19}$	2.8
${}_{54}^{136}\text{Xe} \rightarrow {}_{58}^{136}\text{Ce}$	0.044	$\tau_{1/2}^{2\nu 2\beta} \simeq 2 \times 10^{21}$	8.9
${}_{60}^{150}\text{Nd} \rightarrow {}_{64}^{150}\text{Gd}$	2.079	$\tau_{1/2}^{2\nu 2\beta} \simeq 7 \times 10^{18}$	5.6

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## With $0\nu\beta\beta$ one can

- test lepton number violation
- test Majorana nature of neutrinos
- probe neutrino mass scale
- extract Majorana phase
- constrain inverted ordering

conceptually, it would increase our believe in

- GUTs
- seesaw mechanism
- leptogenesis