

Revamping LLR @ MLRO

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Finder cameras

Cassegrain port

entrales

Nasmyth platform

Contraves

Coudé path

MLRO facts

152 cm diameter, diffraction limited, altaz Coudé Cassegrain telescope
Shared aperture
10 Hz, 100 mJ, 40 ps @ 532 nm laser
MCP/PMT (Photek) + CFD detection

Issues

- MLRO was designed as an LLR capable system, but the production has been very sporadic due to different reasons:
 - LLR is inherently difficult
 - LLR needs a dedicated and competent crew
 - Untimely spare parts procurement

Why the revamping

MLRO IS a LLR station, what the heck... And...



Critical points

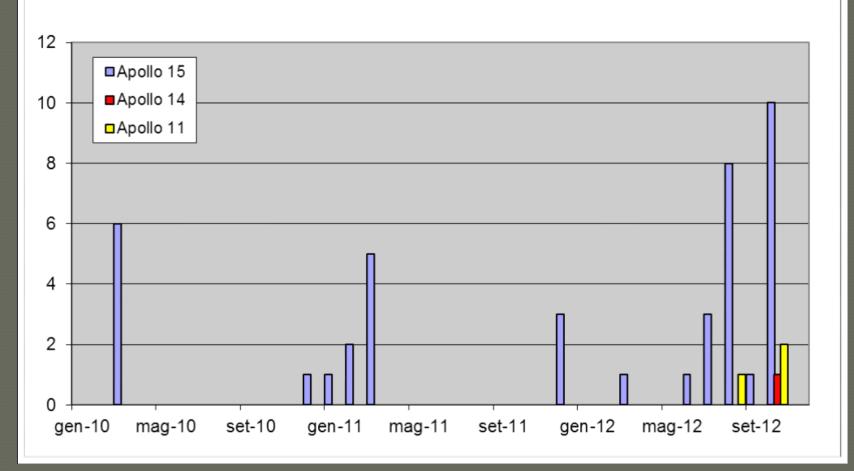
- Divergence: this is difficult to measure and characterize, at least on MLRO. Thanks to Mark Davis and colleagues at NRL for their efforts!
- Optical Efficiency: the MLRO telescope has 7 reflections: a 10% loss on each means a signal loss of 75% (two-way)
- Seeing
- About everything else (PMT, interference filter, alignments, etc.)

System improvements

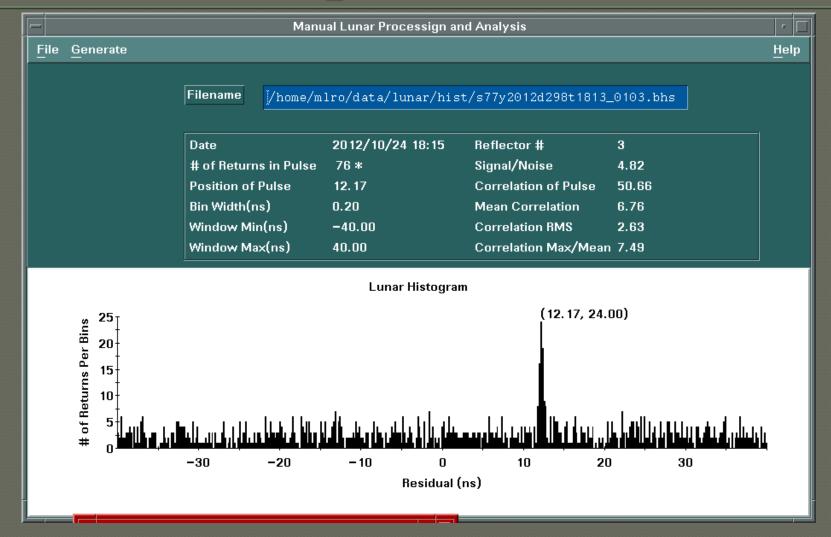
- New MPC PMT from Photek
 New 3 Å filter
- Optical path efficiency monitoring & (more) systematic cleaning
 (More) careful alignments



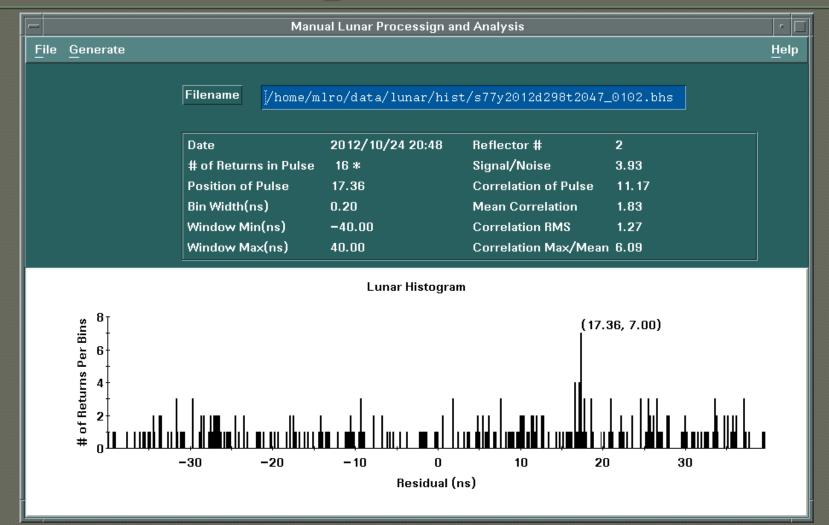
MLRO LLR Normal points



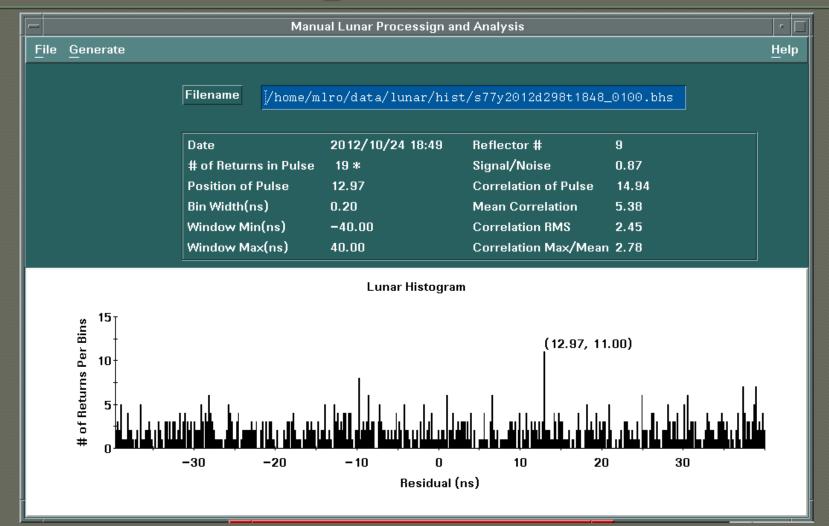
Apollo 15



Apollo 14

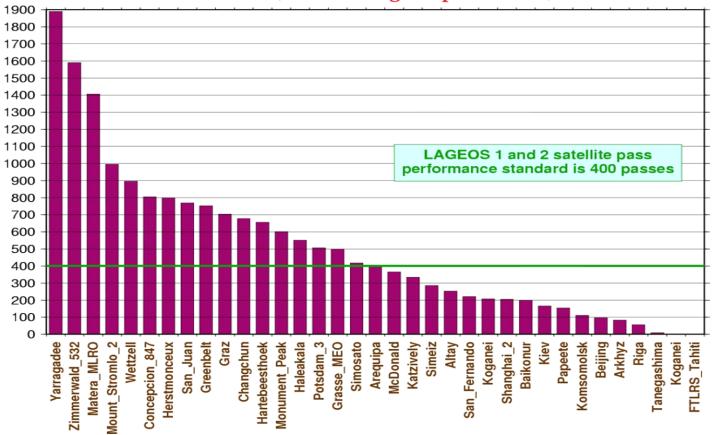


Apollo 11



SLR @ MLRO

LAGEOS 1 and 2 passes from October 1, 2011 through September 30, 2012



20121001

Plans

- Continue operations and improve LLR capability
- Telescope control system being replaced
- Main mirror recoating to be done asap
- MLRO general upgrade under study

Thank you!

